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1 **Crustal Structure of the Southern Margin of the African**
2 **Continent: Results from Geophysical Experiments.**

3

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5

6 **Abstract**

7 A number of geophysical onshore and offshore experiments were carried out along a profile across the
8 southern margin of the African Plate in the framework of the Inkaba yeAfrica project. Refraction
9 seismic experiments show that Moho depth decreases rapidly from over 40 km inland to around 30 km
10 at the present coast, before gently thinning out towards the Agulhas-Falkland Fracture Zone, which
11 marks the transition zone between continental and oceanic crust. In the region of the abruptly
12 decreasing Moho depth, in the vicinity of the boundary between the Namaqua-Natal Mobile Belt and
13 the Cape Fold Belt, lower crustal P-wave velocities up to 7.4 km/s are observed. This is interpreted as
14 metabasic lithologies of Precambrian age in the Namaqua-Natal Mobile Belt, or mafic intrusions added
15 to the base of the crust by younger magmatism. The velocity model for the upper crust has excellent
16 resolution, and is consistent with the known geological record. A joint interpretation of the velocity
17 model with an electrical conductivity model, obtained from magnetotelluric studies, makes it possible
18 to correlate a high velocity anomaly north of the centre of the Beattie Magnetic Anomaly with a highly
19 resistive body.

20

21 **1. Introduction**

22

23 A number of features make southern Africa a key site for integrated geophysical and geological
24 research on continental accretion and break-up. Bounded to the south by the Agulhas-Falkland Fracture
25 Zone (AFFZ, Fig. 1), this region represents one of the world's best examples of a sheared continental
26 margin which formed during break-up of Western Gondwana from about 130 Ma (Ben-Avraham et al.,
27 1997). At the same time, a classic volcanic rifted margin was developing at the western margin of
28 South Africa. The deep crustal structures and physical properties on these two contrasting margin
29 types, and the nature of the continent-ocean transition across them, are key to understanding the
30 processes of continental break-up and passive margin development. The southern margin of Africa is
31 also of great interest because progressively older crustal provinces are crossed as one moves inboard
32 from the coast (Cape Fold Belt, Namaqua-Natal Mobile Belt, Kaapvaal Craton). The boundaries
33 between these crustal provinces are first-order features relating to continental assembly and information
34 on their 3-D geometry is largely lacking. Particularly important in this respect are the Beattie Magnetic
35 Anomaly and the Southern Cape Conductivity Belt (Fig. 1), which run roughly parallel to the boundary
36 of the Cape Fold Belt and Namaqua-Natal Belt, and are among the largest and strongest geophysical
37 anomalies on the African continent. Despite the nearly 100 years since their discovery, the geologic
38 nature of these anomalies is still a mystery. Finally, and superimposed on the crustal mosaic, is the
39 huge and economically important Karoo sedimentary basin of Triassic to Jurassic age and the
40 associated lavas and sills of the Karoo Large Igneous Province (LIP) of about 175-185 Ma age.

41 All of these aspects are addressed in a coordinated program of marine and terrestrial geophysical
42 experiments that has been carried out along the so-called Agulhas-Karoo Geoscience Transect
43 extending from the Agulhas Plateau across the sheared margin and inland as far as the Kaapvaal Craton
44 border (Fig. 1). Beginning in 2004, this work has been one of the main research targets of the Inkaba ye
45 Africa project (de Wit & Horsfield, 2006). Seismic experiments included wide angle refraction
46 (Stankiewicz et al., 2007), and near-vertical reflection (Lindeque et al., 2007) surveys across the Cape
47 Fold Belt and into the Karoo Basin, magnetotelluric profiles (partially coincident with the seismic
48 survey) across the Beattie Anomaly and Cape Conductivity Belt (Branch et al., 2007; Weckmann et al.,
49 2007a,b), and offshore refraction and reflection seismics (Parsieglä et al., 2007, 2008). This paper
50 combines and jointly interprets these different data sets along the Western Karoo-Agulhas Profile.

51

52 **2. Profile Setting**

53

54 The northern part of the profile runs through the Karoo Basin (KB, Fig. 1). This retroarc foreland basin
55 is believed to have formed during the accretion of the palaeo-Pacific plate to the Gondwana plate
56 during the Late Carboniferous (Cole, 1992). The basin contains up to 12 km of Late Carboniferous to
57 Mid Jurassic sedimentary strata (Karoo Supergroup) dominated by shale, siltstone and sandstone
58 (Broquet, 1992; Cole, 1992; Cloetingh et al., 1992; Catuneanu et al., 1998). Cole (1992) and Cloetingh
59 et al. (1992) report a thickness of about 4600 m near the northern end of our profile. The Karoo
60 Supergroup is subdivided from the top downwards into the Beaufort, Ecca and Dwyka Groups. An
61 important marker horizon in the Karoo Basin is a series of deep water carbonaceous shales within the
62 Ecca Group (~40 m thick Whitehill and ~180 m thick Prince Albert Formations). In the southern part
63 of the basin along the geophysical profile, the Karoo strata are tightly folded with north-vergence
64 whereas deformation intensity decreases northward and strata have gentler dips.

65 To the south the Karoo Basin is terminated by the Cape Fold Belt (CFB, Fig. 1). This belt comprises a
66 thick sequence of Neoproterozoic to early Palaeozoic metasedimentary rocks (Cape Supergroup),
67 which were deposited in a marginal basin created by inversion of a Pan-African mobile belt (de Wit,
68 1992). The stratigraphy of the Cape Supergroup consists of a lower series up to ~3 km of
69 Neoproterozoic to Cambrian metasediments in local basal rift sequences (e.g., Kango Inlier, Barnett et
70 al., 1997), which are overlain by ~8 km Ordovician to Carboniferous clastic sediments and
71 orthoquartzites (Hälbich, 1983, 1993; Tankard et al., 1982; Broquet, 1992; Catuneanu et al., 1998).
72 Together with the lower units of the Karoo Basin (see above), the Cape Supergroup rocks were
73 deformed at ~250 Ma, with the formation of north-vergent asymmetric or overturned folds and thrust
74 faults (Hälbich, 1993; Hälbich & Swart, 1983). The dominant CFB thrusts are south dipping and may
75 coalesce into a common décollement (Hälbich, 1983, 1993; Newton 1992; Paton et al. 2006). Drill
76 cores discussed by Eglington & Armstrong (2003) show that some of the Cape Supergroup underlies
77 the Karoo Basin, but it is not known how far north the Cape Supergroup extends.

78 The Cape Supergroup, and farther north the KB, unconformably overlie the Namaqua-Natal Mobile
79 Belt (NNMB, Fig. 1). This Mesoproterozoic complex has accreted to the lithospheric core of the
80 Kaapvaal Craton, and experienced successive periods of extension and compression between 2.0 and
81 1.0 Ga (de Wit, 1992). During the Late Proterozoic / Early Cambrian, eroded material from uplifts in
82 the NNMB filled basins to the south, forming the Kango and Kaaimans Inliers (Hälbich, 1993). The
83 NNMB is a highly complex polymetamorphic province that constitutes three sub-provinces or terranes,

84 which were amalgamated and attached to the Archaean Kaapvaal Craton during the late
85 Mesoproterozoic (Robb et al., 1999; Eglington and Armstrong, 2003; Raith et al., 2003; Dewey et al.,
86 2006; Eglington, 2006 and references therein). The age of peak metamorphism across all terranes was
87 1.02-1.04 Ga (Eglington, 2006), and assemblages indicate upper amphibolite to granulite-facies
88 conditions with locally extensive generation of crustal melts now represented by megacrystic
89 orthogneisses. The structural history of the NNMB includes regional ductile deformation with refolded
90 recumbent folds, thrusting and listric shear zones that produced strong shallow to steeply- dipping
91 fabrics and regionally extensional shear zones. Interpretation of a seismic refraction profile across the
92 NNMB in the western Cape Province indicated crustal thickness of about 42 km and intermediate P-
93 wave velocity of 6.2-6.9 km/s in the lower crust (Green and Durrheim, 1990). Similar results were
94 reported by Hirsch et al. (2008) from a combined interpretation and modelling of gravity and seismic
95 data along an offshore-onshore profile which crosses the Orange Basin off the western coast of
96 Southern Africa, and extends 100 km inland. The crustal thickness and lower-crustal P-wave velocity
97 values at the inland end of this profile are 36-38 km and 6.7 to 7.0 km/s, respectively. The eastern
98 exposures of the NNMB in KwaZulu-Natal comprise a number of litho-tectonic terranes dominated by
99 upper amphibolite to granulite- facies gneisses with occasional charnockites, separated by subvertical
100 ductile shear zones (Jacobs et al.1993; Jacobs and Thomas, 1994; Eglington, 2006; McCourt et al.,
101 2006).

102 The profile crosses two major continental geophysical anomalies: the Beattie Magnetic Anomaly
103 (BMA), and the Southern Cape Conductive Belt (SCCB). Both features stretch for more than 1000 km
104 in a roughly east-west orientation (Fig. 1). The BMA, a positive anomaly, was first observed by Beattie
105 (1909). Following a magnetometer array study, Gough et al. (1973) concluded that a large body with
106 high electrical conductivity underlies the NNMB and parts of the KB. Gough (1973) interpreted this
107 Southern Cape Conductive Belt as the signature of a linear plume in the upper mantle. However, de
108 Beer et al. (1974) suggested that, due to their spatial coincidence, the two geophysical anomalies are
109 likely to have a common source, and that such a plume could not account for the BMA. de Beer &
110 Gough (1980) then used Curie isotherms to show that if a common source exists, it must be a crustal
111 feature. Pitts et al. (1992) suggested a sliver of serpentinised oceanic crust, 30 km broad and dipping
112 south from a depth of 7 to 30 km as the source. However, a recent magnetotelluric study by
113 Weckmann et al. (2007a) found no evidence for the existence of such an extensive high conductivity
114 body, which could be the source of the magnetic anomaly at the same time.

115 The offshore section of this western profile starts about 20 km off the coast and stretches over 400 km

116 south, past the Agulhas-Falkland Fracture Zone (AFFZ, Fig. 1). Between the AFFZ and the present
117 coast a network of basins, collectively referred to as the Outeniqua Basin (Fig. 1), is located (e.g.
118 Fouche et al., 1992, McMillan et al., 1997, Broad et al., 2006). This complex basin developed due to
119 extensional episodes in Oxfordian / Kimmeridgian and Valangian times exploiting the structural gain
120 of the underlying Cape Fold Belt (e.g. McMillan et al., 1997). The Outeniqua Basin is bounded to the
121 south by a marginal ridge, which was first described by Scrutton and du Plessis (1973) and renamed
122 into Diaz Marginal Ridge (DMR) by Ben-Avraham et al. (1997). While earlier studies (e.g. Scrutton &
123 du Plessis, 1973) describe it as a basement ridge, more recent studies by Parsieglia et al. (2007) suggest
124 a metasedimentary composition of this ridge. South of the DMR, the AFFZ, a large scale fracture zone,
125 is located. The magnetic signature of this fracture zone was first identified by Oguti (1964) south of
126 Africa. The AFFZ marks the southern boundary of the African continent (Talwani & Eldholm, 1973).
127 This sheared continental margin developed as a consequence of dextral strike-slip motion between
128 today's African and South American continents during the Cretaceous break-up of Gondwana (e.g.,
129 Barker, 1979; Rabinowitz & LaBrecque, 1979). During this shear process the originally juxtaposed
130 Southern Outeniqua and Falkland Plateau Basins were separated (e.g. Martin et al., 1981; Ben-
131 Avraham et al., 1997) and the major faults of the Outeniqua Basin were bent (e.g. Ben-Avraham et al.,
132 1993; Thomson, 1999; Parsieglia et al., 2007).

133

134 **3. Seismic data acquisition**

135

136 The wide angle seismic experiment was carried out in April-May 2005. The 240 km onshore part
137 (Stankiewicz et al., 2007) consisted of 13 shots (each 75-125 kg of explosives) fired from boreholes 20-
138 30 metres deep (Fig. 1). Forty-eight stations (average spacing of ~5 km) each consisting of a GPS
139 synchronised electronic data logger (EDL) and a three-component seismic sensor were used to record
140 the data. The offshore part (AWI-20050100, Parsieglia et al., 2007) consisted of 20 four-component
141 (three component seismometer and a hydrophone component) ocean bottom seismometers (OBS)
142 deployed over 400 km profile length. Eight G-guns and one Bolt-airgun (volume of 96 l) were fired
143 every 60 seconds during cruise SO-182 of the RV Sonne. This gave a shot spacing of approximately
144 150 metres. As the onshore and offshore parts were carried out simultaneously, the airgun shots were
145 recorded by the land receivers, improving the profile's ray coverage, especially beneath the coast. Land
146 shots were not detected by OBSs.

147

148 4. Wide-Angle Seismic Data Processing

149

150 Figure 2 shows the high quality of the traces resulting from land shots, recorded by the land receivers.
151 The direct (refracted) P-wave arrival (P_g – red), as well the Moho reflection (P_mP – blue) are clearly
152 seen. Examples of airgun shots picked up by land receivers are illustrated in Fig. 3, where, due to the
153 longer offsets, the P-waves refracted in the upper mantle (P_n) are also seen. The airguns produce much
154 less energy than the land shots, but because of their small spacing prominent arrivals can be easily
155 identified. Examples of airgun signals registered by the OBS can be seen in Fig. 4. All together, 24045
156 arrivals were manually picked (Table 1). Pick uncertainty was in the range of 0.05 – 0.25 seconds,
157 depending on the signal to noise ratio.

158 Two travel-time inversion techniques have been used in this study. Standard 2-dimensional
159 tomography involves dividing the cross section beneath the profile into rectangular cells. An initial
160 velocity model needs to be provided, and its values temporarily assigned to the cells. A simple one-
161 dimensional model was used here. Synthetic travel times are then computed, and compared to real
162 (picked) ones. An iterative algorithm adjusts the individual values until an optimal model is reached.
163 This study uses the software package FAST (First Arrival Seismic Tomography), which was released
164 by Zelt & Barton (1998). This package is a modification of the algorithm developed by Vidale (1988)
165 to ensure a better detectability of high velocity contrasts. To minimize the influence of the starting
166 model, we have used an iterative approach developed by Ryberg et al. (2007), which repeats the
167 inversion 5 times, using increasingly smaller cell sizes. We have performed the inversion using vastly
168 different starting models, and found no significant differences in the resulting models. A ratio of 1:8
169 was used for the vertical-to-horizontal smoothing constraints.

170 The above technique was used to compute the velocity model for the upper crust beneath the land
171 section of the profile, as well as the northernmost 100 km of the offshore section. The final cell size
172 used was 2 km horizontal by 1 km vertical. For first-arrival tomography calculations it is important to
173 have individual rays crossing each other in every individual cell, so having airgun shots spaced in only
174 one direction from the receivers is not an optimal situation. For this reason only a small part of the
175 offshore section was used, with the purpose of improving the onshore ray coverage rather than trying to
176 compute a detailed offshore model. The resolving capabilities of the algorithm can be tested with
177 checkerboard tests, where checkerboard of alternate positive and negative velocity anomalies is added
178 to the originally computed velocity model. Synthetic travel times are then generated, and an inversion
179 is performed using these times. The inversion result is then subtracted from the initial model. If the

180 blocks can be observed in the final model, we can assume that a real feature corresponding to the
181 block's size, position and velocity perturbation would be resolved by our inversion algorithm, and
182 therefore an observed feature like that is likely to be real, and not an inversion artifact.

183 A disadvantage of FAST is that reflected phases cannot be incorporated in the model. For this reason,
184 the travel-time routine RAYINVR (Zelt & Smith, 1992) was used. In this program the 2-D velocity
185 model is defined in terms of layers. Velocity is specified at a number of nodes along the layer
186 boundaries, and a linear velocity gradient is assumed between nodes (both along boundaries, and with
187 depth). Rays may refract in a particular layer, reflect off any boundary, or travel as head waves along it.
188 As with FAST, model quality is determined by comparing synthetic travel times to the manually picked
189 real ones. Modelling is done as an iterative combination of forward modelling and inversion (Zelt &
190 Forsyth, 1994). The resolution of the velocity model depends (among other things) on the number of
191 layers defined, and the node density along them. The marine model of Parsieglia et al. (2007) was
192 computed using this package. Their model consists of 6 layers: a water layer, two sedimentary layers,
193 two crustal layers and the mantle.

194 To produce a starting velocity model for the joint onshore-offshore analysis of the entire profile, the
195 two separate models had to be merged. As a number of reflected phases are available, RAYINVR was
196 considered the more appropriate package to use. For this reason, the onshore FAST model needed to be
197 converted into a layered model, with layers matching the ones in the offshore model. The uppermost
198 sedimentary layer (marine sediments), and obviously the water layer, exist only in the offshore section,
199 so the onshore part consisted of 4 layers. The uppermost of these was a shallow (not exceeding 3 km)
200 zone with velocities typical for sedimentary rocks. The next 2 layers represented the crystalline crust.
201 In the starting model they were separated at a depth of 15 km, using velocities obtained from the FAST
202 at nodes spaced 20 km apart. The boundary between the 3rd and 4th layer was the Moho, set in the
203 starting model at 42 km (after Lindeque et al., 2007). The combined starting velocity model is then
204 iteratively updated, using forward and inverse modelling, until the calculated synthetic travel times
205 match the data as well as possible. While converting from FAST to RAYINVR decreased the
206 resolution of small-scale features in the onshore part, the two modelling techniques complement each
207 other, and our final results are better than what could be achieved using either of the techniques alone.

208

209 **5. Results**

210

211 *5.1. Crustal features beneath the continent*

212 The upper crust model computed with FAST using first arrivals of land shots (445 P-wave arrivals), as
213 well as airgun shots less than 100 km from coast (3351 P-wave arrivals), is shown in Fig. 5. The RMS
214 travel time residual was 0.045 seconds, and the chi-squared misfit parameter 2.60. Fig. 5 needs to be
215 viewed in conjunction with Fig. 6, which shows the checkerboard tests conducted on the model. These
216 tests confirm that the model is better resolved beneath the land, where features barely twice as wide as
217 the receiver spacing can be resolved in the upper 10 or 15 km. Large scale features, 40 km wide, can be
218 confidently resolved almost to the maximum ray penetration, even beneath the ocean.

219 The model in Fig. 5 is consistent with the model computed using only onshore shots (Stankiewicz et
220 al., 2007). However, the improved ray coverage increased the maximum depth of the model from less
221 than 30 km to almost 40 km, and many of the upper crustal features are better resolved. In the northern
222 half of the profile, the slow velocities (4.6 – 5.3 km/s, red and orange in Fig. 5) characterising the
223 Karoo Basin are clearly visible to depths not exceeding 5 km. Slightly faster (5.4 – 5.8 km/s, yellow
224 and pale green) velocities beneath indicate the relatively thin Cape Supergroup. The resolution of the
225 model does not clearly indicate the contact of the two Supergroups. The checkerboard tests indicate
226 that the thinning of the basin at profile length of 100 km is a real feature – this is consistent with our
227 earlier results (Stankiewicz et al., 2007), which suggested a blind Paleozoic Thrust Fault, which could
228 mark the northernmost deformation of the Cape Supergroup. The thickening of the basin at profile
229 length of approximately 130 km (also well within the resolving capability of the model) is consistent
230 with the large asymmetric syncline inferred from field observations (Cole, 1992).

231 Farther south, the geometry of the listric Kango Fault (KF) is resolved much more clearly than in the
232 earlier results (Stankiewicz et al., 2007). Near the surface (at ~170 km profile length) this fault marks
233 the northern edge of the Jurassic Uitenhage basin, characterised by very low velocities (~4.5 km/s). The
234 offshore section of the model is not well resolved, which was expected from the shot-receiver geometry
235 explained in the previous section. However, the very low velocities representing the sediment cover (<
236 5 km/s) are clearly seen to be significantly deeper than onshore.

237

238 *5.2. Combined onshore-offshore model*

239 The final RAYINVR model is shown in Fig. 7. The programme used 21,868 travel times (over 90 % of
240 the picks). The RMS travel time residual was 0.134 seconds, which is within the uncertainty bounds of
241 individual picks. The chi-squared misfit parameter was 1.74. Table 2 shows how these uncertainties
242 vary for different phases. A value greater than 1 for the chi-squared parameter means the small-scale
243 features of the model have not been resolved (Zelt & Forsyth, 1994), so we will concentrate our

244 discussion on the large-scale features. A likely explanation for a large value of chi-squared is the
245 presence of 3-D effects, in particular the fact that the onshore and offshore parts of the profile are not
246 perfectly aligned (Fig. 1). Examples of ray paths in the model are shown in Fig. 8. The relative ray
247 coverage available is shown in Fig. 9, with red areas indicating poor ray coverage.

248 In addition to the standard phases P_g and P_mP , an unusual phase has been observed on the traces
249 recorded from the two southernmost shots (P_x – Fig. 2). The amplitude of this phase is of similar order
250 of magnitude to the P_mP . Furthermore, the phase is most prominent in the vertical component, so we
251 interpret it as P-waves reflected inside the crust. As a reflected phase, these travel times could not be
252 included in the FAST model, and the corresponding reflector location was derived with RAYINVR
253 using the floating reflector technique. The position of the reflector that best fits the observed travel
254 times was found to be between profile km 190 and 220, rising steeply southwards from a depth of 35
255 km to 23 km (Fig. 7). The steep landward dip of this reflector makes it impossible to detect reflections
256 of airgun shots off it, the same way the steeply rising Moho was invisible.

257 The most interesting feature of the joint model shown in Fig. 7 is the Moho discontinuity, clearly
258 visible as a high-velocity contrast. The Moho depth beneath the Karoo Basin is ~40 km, and slightly
259 deeper (~42 km) beneath the CFB. The crustal thickening beneath the CFB is consistent with receiver
260 function analyses (Harvey et al., 2001; Nguuri et al., 2001). However, the study by Nguuri et al.
261 (2001), while detecting the thickening, consistently locates the Moho approximately 5 km deeper than
262 reported here. This could be due to the fact that these authors use an average crustal velocity of 6.5
263 km/s. This value is typical for the Kaapvaal Craton at which their study was aimed, but is too high for
264 off craton analysis. Our results are much closer to those of Harvey et al. (2001), who subdivided the
265 crust into regions of different velocity at different depths.

266 South of the CFB the Moho depth becomes more shallow very abruptly, reaching 30 km at the present
267 coast. The depth is consistent with values obtained from an east-west reflection profile of Durrheim
268 (1987), which runs approximately 10 km south of the coast. Farther south the crust continues to thin,
269 albeit much more gradually, for another 250 km, underneath the Agulhas Bank, Outeniqua Basin and
270 the Diaz Marginal Ridge, until it reached the Agulhas-Falkland Fracture Zone (AFFZ). This fracture
271 zone marks the continent-ocean transition (COT). This transition to oceanic crust begins around profile
272 distance of 480 km, where the Moho depth of 20 km is reduced to 12 km over a little over 50 km
273 horizontal difference (Parsieglia et al., 2007). This is a typical length, as well as depth change, for the
274 transition at sheared margins (e.g., White et al., 1992; Bird, 2001). The depth of 11-12 km (i.e. 6-7 km
275 of crust under 5 km of ocean) is observed in the Agulhas Passage, where the southernmost 130 km of

276 the profile stretches.

277

278 **6. Discussion**

279

280 *6.1. Crustal Features beneath the continent*

281 Most of our shallow results are consistent with the results of the magnetotelluric profile (Weckmann et
282 al., 2007a) coinciding with the northern 140 km of the seismic line. These authors traced the highly
283 conductive Whitehill Formation (pyrite-rich black shales at the bottom of the Eccu Group) for virtually
284 the entire length of their profile, which correlates well with the geometry of the basin obtained with our
285 tomography. The accuracy of the Whitehill's location in the MT depth section is confirmed by
286 excellent correlation with drill cores (Eglington & Armstrong, 2003; Branch et al., 2007). Weckmann
287 et al. (2007a) also observe an offset in the depth to the shales at the same location as our basement
288 thinning, and also suggest a south dipping thrust as a likely explanation.

289 There is also some correlation between the velocity model and the electrical conductivity image of
290 Weckmann et al. (2007a) for deeper features (Fig. 10). Our model shows a zone of anomalously high
291 velocity (~7 km/s) at a depth of ~15 km between profile km 60 and 90. The size of the body certainly
292 falls within the model's resolving capability. The southern edge of this anomaly is in the vicinity of the
293 centre of the BMA. In the 2D image of the electrical conductivity distribution, a zone of high electrical
294 resistivity is found at the same location. To the north this zone is flanked by a large mid-crustal region
295 of stacked layers of high electrical conductivity, possibly imaging mineralisations in synforms, and, to
296 the south, by a narrow, southward dipping conductor at 7-15 km depth under the maximum of the
297 BMA. The top of this anomaly is exactly coincident with what a seismic reflection study (Lindeque et
298 al., 2007) called a "complex seismic reflectivity patch". However, a comparison of magnetic models
299 explaining the magnetic response of the BMA with the electrical conductivity model clearly shows that
300 the electrical conductivity anomaly located beneath the centre of the BMA is too narrow to be a
301 possible source of the BMA (Weckmann et al, 2007b). Fig. 10 shows the two magnetic bodies outlined
302 by black lines, each with a magnetic susceptibility of 0.05 SI. Weckmann et al. (2007b) show that these
303 simple magnetic bodies would produce a response similar to the signature of the BMA. The bodies are
304 separated by a fault, which cuts through at the same inclination as the conductivity anomaly. The gap
305 (some 100 metres wide) representing the fault has an induced magnetic susceptibility of 0.0 SI, the
306 same value as the background susceptibility. The location of the fault correlates well (at least in the
307 upper 20 km) with the velocity contrast marking the northern edge of the synclinal low velocity zone

308 between 100 and 140 km along the profile, while the top of the northern body correlates very well with
309 both the high velocity anomaly and the zone of high resistivity north of the surface maximum of the
310 BMA (Weckmann et al., 2007a, Quesnel et al., 2008).

311

312 *6.2. Combined onshore-offshore model*

313 The geometry and seismic velocity structure derived from a joint interpretation of onshore and
314 offshore Vp tomography (Fig. 7) provides one of the best available geophysical images of a sheared
315 continental margin in cross-section. It also allows a detailed comparison with features from profiles
316 across the classic volcanic rifted margins on the western coast of South Africa. The most direct
317 comparison is with the seismic profile across the Orange Basin (Hirsch et al., 2008), because this
318 profile also extends oceanward from the Namaqua-Natal Mobile Belt. The sheared and volcanic margin
319 profiles are compared at the same scale on Figure 11. Both profiles show similar wedge-shaped
320 geometry and furthermore, the crustal thickness (Moho depth) does not reduce uniformly but stepwise,
321 with inflection points separating segments of rapid and gradual change. Comparing these segments in
322 turn, from the continental crust (A in Fig. 11) outward, we consider the following observations most
323 important:

- 324 • Both profiles begin at the landward end in the Namaqua-Natal Mobile Belt, in the case of the
325 Karoo-Agulhas profile this is covered by the Karoo sequence. The crustal thickness at this end
326 of both profiles is about 40 km, but there is a major difference in the seismic velocity structures
327 of the lower crust. The Karoo-Agulhas profile shows a zone up to 7 km thick with seismic
328 velocity above 7.0 km/s (maximum Vp = 7.4 km/s), whereas the basal velocity at the Springbok
329 profile does not exceed 6.8 km/s. The significance of the high velocity lower crust is discussed
330 in more detail below. The high-velocity zone on the Karoo-Agulhas profile is interrupted at
331 profile distance 120 km by a “keel” of material with intermediate velocity (green on Figs. 7,
332 11). The higher resolution of the upper crustal seismic image (Fig. 5) shows that this feature has
333 a complex shape that appears consistent with the north-vergent folds and thrust faults known in
334 the upper crust.
- 335 • Over much of the segment of stretched continental crust (B in Fig. 11) between continental
336 crust (segment A) and the steep decrease in Moho depth (segment C), the crustal thickness at
337 the Karoo-Agulhas profile is significantly greater at the Springbok profile. This is likely due to
338 the presence of the Cape Fold Belt, which is not intersected by the Springbok profile on the
339 west coast.

- 340 • The stretched crust shows a gradual and uniform thinning on the sheared Agulhas margin. The
341 lower crustal seismic velocity in this segment is also uniform, with V_p values of 6.5 to 7.0 km/s
342 that are in the same range as the lower crust observed in other parts of the NNMB (Green and
343 Durrheim, 1990; Durrheim and Moony, 1994; Hirsch et al., 2008). This is in marked contrast to
344 the corresponding segment on the Springbok VRM, which shows a down-warping of the lower
345 crust in the middle part of the segment. This "keel" has a higher seismic velocity ($V_p > 7.0$
346 km/s) than the lower crust elsewhere along the profile, and even the crust above the keel has a
347 higher velocity than at the same depth outside the keel region. High-velocity lower crust are
348 very common features at volcanic rifted margins and interpreted as underplated basaltic magma
349 that intruded and ponded at the crust-mantle boundary (Menzies et al., 2002). Hirsch et al.
350 (2008) showed that this interpretation is consistent with gravity modelling of the Springbok
351 profile and also noted that the high-velocity keel underlies a zone of seaward-dipping reflector
352 wedges in the upper crust, also characteristic features of volcanic rifted margins and interpreted
353 as submerged basalt flows on the foundering margin (Menzies et al., 2002). Note that if the
354 high-velocity keel on the Springbok profile is considered to be accretion of new crust by
355 magmatism, the geometry of the Springbok margin in this segment is very similar to that of the
356 Agulhas margin, with a uniform, moderate thinning (dashed line, Fig. 11).
- 357 • Both margins exhibit an abrupt rise of the Moho over about 50 km (segment C). In the case of
358 the Agulhas margin, the lower crust at this rise shows a higher seismic velocity than in the
359 "stretched" segment, with values of about 7 km/s. This represents the continent-ocean transition
360 across the sheared margin. For the Springbok profile this increase in seismic velocity at the
361 steep rise is not observed, because it is part of a much longer (~200 km) continent-ocean
362 transition zone.
- 363 • The oceanic crust (segment D) on both sections has the global average thickness of about 7 km
364 and a velocity structure that is also typical of oceanic crust worldwide. The post-rift
365 sedimentary cover on the oceanic crust at the Springbok profile is considerably thicker than at
366 the Agulhas profile because it crosses the Orange River Basin.

367 A number of theoretical and empirical models exist that link observed geometry and seismic properties
368 with the processes of rifting and magmatism at volcanic rifted margins (e.g., Menzies et al., 2002). It is
369 clear from their work that seaward-dipping reflector sequences in the upper crustal section land-ward of
370 the ocean-continent transition zone, and thick, high seismic-velocity lower crust (>7.0 km/s)
371 underneath the reflector sequences result from massive intrusion of breakup-related magma. Classic

372 examples of these features were documented from the Walvis Basin in Namibia by Bauer et al. (2000),
373 and Trumbull et al. (2002) showed that the thickness (20 km) and high V_p-values (7.2-7.4 km/s) of
374 underplated lower crust in NW Namibia require high mantle temperatures and active mantle upwelling,
375 consistent with the proximity to the Walvis Ridge and Paraná-Etendeka Large Igneous Province. The
376 volcanic rifted margin at the Orange River Basin (Springbok profile) is located some 1500 km south of
377 the Walvis Ridge and the underplated crustal body is considerably thinner and has a lower average V_p
378 velocity (Hirsch et al., 2008). Trumbull et al. (2007) found that dolerite dykes exposed along the coast
379 from NW Namibia to the Cape Province showed a systematic decrease in their maximum MgO
380 contents, both in whole-rock and in olivine, which is consistent with a waning plume influence from
381 north to south expressed in lower mantle potential temperature and less active upwelling.

382 On the sheared Agulhas margin, neither seaward-dipping reflectors nor high-velocity lower crust are
383 revealed by the geophysical surveys, nor is there geologic evidence for syn-rift magmatism onshore.
384 This margin can thus be considered as non-magmatic with respect to the time of breakup. There is,
385 however, seismic evidence of post-breakup magmatism on the profile. Seismic reflection sections in
386 the Agulhas Passage found signs of volcanic seamounts and intra-basement reflectors interpreted as
387 basaltic flows (Parsieglä et al., 2007). A late Cretaceous age of extensive volcanism (ca. 100 Ma) is
388 also suggested from seismic data from the Agulhas Plateau (Uenzelmann-Neben et al., 1999; Parsieglä
389 et al., 2008).

390 One feature of the Karoo-Agulhas profile that deserves further comment is the zone of high seismic
391 velocity (V_p > 7.0 km/s) in the lower crust inland of the coast (from profile km 150 to the coast).
392 Whereas the velocity of this zone is like that encountered in underplated lower crust at volcanic rifted
393 margins, the position of this zone in relation to the continental margin makes it unlikely to represent
394 breakup-related magmatism. Also, as mentioned above, there are no matching expressions of
395 magmatism in the upper crustal seismic data, or on land. Crustal rocks with P-wave velocities greater
396 than 7 km/s are most likely to represent mafic igneous rocks rich in olivine, or their metamorphosed
397 equivalents (amphibolites, mafic granulites). Seismic velocities of metabasic rocks are particularly high
398 if they are rich in garnet (e.g., Durrheim & Mooney, 1994; Christensen & Mooney, 1995). Gerignon et
399 al. (2004) discussed the likelihood of older high-pressure rocks (garnet granulite, eclogite) as an
400 explanation for high V_p lower crust on parts of the rifted Vøring Basin in the Norwegian Sea.
401 Considering the geologic setting of the Karoo-Agulhas profile, there are principally two possible
402 explanations for the high-velocity crust. Either it represents metabasic lithologies of Precambrian age in
403 the NNMB, or mafic intrusions added to the base of the crust by younger magmatism. The dominant

404 exposed lithologies of the high-grade unit in the NNMB (Namaqualand terrane) are intermediate to
405 felsic gneisses and meta-granitoids (Dewey et al., 2006). These authors suggested that the heat source
406 for granulite-grade metamorphism and crustal melting was “massive underplating” of mafic magmas at
407 around 1050 Ma. The evidence cited for this underplating was a seismic refraction study by Green &
408 Durrheim (1990), who reported lower crustal velocities of 6.6 – 6.9 km/s for the western Namaqualand
409 crust. Similar values of 6.8 km/s were derived for the base of the crust in the NNMB at the landward
410 end of the Springbok profile (Hirsch et al., 2008). Even under the thick Kaapvaal Craton, seismic data
411 indicate a generally sharp transition from the Moho to crustal rocks with intermediate to felsic
412 properties (Durrheim & Mooney, 1994; Nair et al., 2006) and no evidence for widespread high-Vp
413 lower crust. Crustal xenoliths of garnet granulite are not uncommon in kimberlites from the NNMB
414 (Schmitz & Bowring, 2004), and thus some high-Vp material must be present in the crust, however the
415 seismic surveys so far undertaken indicate that such material cannot be a major component of the lower
416 crust. Therefore, although we do not rule out the possibility that the high-Vp zones at the base of the
417 crust in the Karoo-Agulhas profile are part of the Proterozoic basement, all other studies of the
418 NNMB and cratonic crust failed to detect velocities above 6.9 km/s.

419 The alternative explanation for the high Vp values observed at the landward end of the Karoo-Agulhas
420 profile is that Phanerozoic magmatism added thick intrusions of gabbroic material in the lower crust.
421 This is of course the same scenario offered to explain the very common thick underplated crust at
422 volcanic rifted margins, and it is clear from many studies that appropriate volumes of material and bulk
423 seismic properties would fit the observations (7 km thickness, $V_p = 7.0\text{-}7.4$ km/s – see Fig. 7). From
424 the regional geology and location of the profile, the most likely candidate for producing extensive
425 mafic underplating is the mid-Jurassic Karoo-Ferrar-Chon Aike Large Igneous Province. This was one
426 of the major episodes of mafic magmatism on a global scale, and it is unlikely that the huge volumes of
427 magma represented by Karoo lavas, dykes and sill complexes could be emplaced at upper crustal levels
428 without an intrusive equivalent at depth. Unfortunately, the seismic experiments used in this study did
429 not produce enough S-wave information to allow calculation of V_p/V_s ratios which would provide
430 better lithologic discrimination, but from the available shape, location and Vp values of the lower
431 crustal zone we believe an origin from underplated magmas related to the Karoo LIP provides the best
432 explanation.

433 de Wit (2007) already noted the surprising lack of seismic evidence for mafic underplating of the
434 Karoo magmas, whereas such evidence is clear and abundant for the Paran-Etendeka LIP on the
435 western margin. We argue from the results of this study that significant Karoo underplating did occur.

436 Its paucity in seismic studies under the craton probably relates to the deflecting action of the thick
437 Kaapvaal lithosphere. Why there is no seismic expression for underplating in the other surveys of the
438 NNMB surveys is less clear but it may be related to the fact those surveys were located farther west
439 than the Karoo-Agulhas profile whereas the locus of strongest Karoo magmatism is to the east, in the
440 Lebombo-Natal and Lesotho regions.

441

442 **7. Summary**

443

444 In this study we have used tomographic inversion of reflected and refracted travel times to construct a
445 P-wave velocity model along an off-onshore profile across the southern margin of Africa. The findings
446 are summarized in Fig. 12. A number of the model's features agree with known geological record, and
447 are consistent with the results of the separate onshore (Stankiewicz et al., 2007) and offshore (Parsiegla
448 et al., 2007) analyses. These include the rough geometry of the Karoo and Cape Supergroups, the
449 presence of a blind Paleozoic Thrust Fault ~ 40 km north of the farthest Witteberg Group outcrop
450 which possibly marks the northern edge of deformation in the Cape Supergroup, and the geometry of
451 the Agulhas-Falkland Fracture Zone which marks the continent – ocean transition.

452 New findings of this study are:

- 453 • A high velocity anomaly north of the centre of the BMA is much better resolved than
454 previously. This feature is coincident with an extensive zone of high resistivity and a magnetic
455 body required to reproduce the magnetic signature of the BMA.
- 456 • A synclinal low velocity feature was identified in the Mesoproterozoic basement beneath the
457 front of the Cape Fold Belt, south of the above mentioned feature. The northern edge of this
458 feature correlates with the second magnetic body necessary to account for the BMA's signature.
- 459 • A steep decrease in crustal thickness (from 40 to 30 km over lateral distance of 40 km) occurs
460 under the present coast.
- 461 • A zone of high velocity material is observed in the lower-most crust beneath the present coast.
462 This either represents metabasic lithologies of the Mesoproterozoic Namaqua-Natal
463 Metamorphic Complex, or intrusions of gabbroic material added to the base of the crust by
464 younger magmatism.

465

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730

731 Table 1: Number of picked arrival times

732

phase	Air Gun – OBS	Air Gun – EDL	Land Shot – EDL	Σ
Crustal Refraction	2329	5030	445	7804
Mantle Refraction	225	13861	0	14086
Moho Reflection	642	1028	132	1802
Intracrustal Reflection	332	0	21	353
Σ	3528	19919	598	24045

733

734

735

736 Table 2: Model quality for different phases

737

phase	rms	X²
Crustal Refraction	0.132	1.70
Mantle Refraction	0.132	1.70
Moho Reflection	0.161	2.40
Intracrustal Reflection	0.173	0.88
Σ	0.134	1.74

738

739

739 **Figure Captions**

740

741 1. Study area: Map of southern Africa shown the relevant large-scale features, like the Kaapvaal
742 Craton, Karoo Basin (KB), Cape Fold Belt (CFB), the location of the Namaqua Natal Mobile Belt
743 (NNMB), the Agulhas-Falkland Fracture Zone (AFFZ), as well the Southern Cape Conductivity Belt
744 and the Beattie Magnetic Anomaly. The detailed map shows the geological setting of the study area.
745 The onshore section of the profile is indicated by the red triangles (13 shots, numbered X01-X13) and
746 smaller black triangles (48 receivers, numbered 01-48). Four cities are shown on the map to aid
747 location: Fraserburg (FR), Prince Albert (PA), Oudtshoorn (OU) and Mossel Bay (MB). The 20 Ocean
748 Bottom Seismometers used in the offshore part are marked as blue triangles, and are numbered 101-120
749 following the notation of Parsieglä et al. (2007).

750

751 2. Examples of data recorded by land receivers from land shots. Top: shot X10. The refracted P-waves
752 (P_g) marked in red, and Moho reflections (P_mP) in blue. Bottom: shot X13. Phases reflected off the
753 “floating” crustal reflector (P_x) are marked in green. Reduced travel times are shown with a reduction
754 velocity of 6 km/s.

755

756 3. Example of the data from air gun shots picked up by land receiver 46. Direct P-waves from the crust
757 (P_g), from the mantle (P_n) and Moho reflections (P_mP) are clearly visible. Reduced travel times are
758 shown with a reduction velocity of 6 km/s.

759

760 4. Air gun shots recorded by OBS 105. Phases P_g , P_n and P_mP as in Fig. 3. P_2 corresponds to P-waves
761 refracted in the upper-crust low-velocity sediments, and P_{c2} P-waves reflected inside the crust (notation
762 from Parsieglä et al., 2007). Reduced travel times are shown with a reduction velocity of 8 km/s.

763

764 5. P-wave velocity model computed for the onshore part of the profile, as well as the first 100 km of the
765 offshore part. Upper panel shows just the model, the lower panel includes interpretations. Starting in
766 the north, 140 km of the inferred distribution of the Karoo Supergroup, and the Cape Supergroup
767 underlying it, has been drawn. The location of the blind thrust fault which we believe marks the
768 northernmost deformation of the Cape Supergroup has been marked by an arrow (PTF). We see the
769 thickening of the basement at profile km 130, which is consistent with the syncline at which it tapers
770 out. The actual edge of it, however, is not resolved, and has been drawn in dotted lines. Farther south,
771 the geometry of the normal listric Kango Fault (KF, surfacing at 170 km) has been drawn. A high
772 velocity anomaly most likely related to the BMA has been indicated with an ellipse at a depth of ~15
773 km, at profile km 80. The centre of the surface trace of the BMA has also been indicated.

774

775 6. Resolution tests for the model in Fig. 5. Original checkerboard sizes are 40 km by 10 km (top), 25km
776 by 10 km (middle) and 12 km by 5 km (bottom); velocity perturbation of 5% used in each case.

777

778 7. P-wave velocity model for the entire profile. To help location with Fig. 1, land shots (red triangles)
779 and OBSs (blue triangles) have been indicated. Velocity contours corresponding to 6.0, 6.5 and 7.0
780 km/s have been drawn. The thick white line marks the “floating” reflector. Black rectangle indicates
781 the northern 340 km that are shown in Figs 5 and 6.

782

783 8. Airgun shots recorded by 4 land receivers (a: receiver 12; b: 17; c: 24; d: 44). In each figure the top
784 panel shows the raypaths; the bottom panel the observed travel time (steeples corresponding to
785 uncertainty bars), and the travel time expected from the model for the given source-receiver distance
786 (solid line).

787

788 9. Ray coverage for the model presented in Fig. 7. A logarithmic scale of number of rays crossing each
789 2 km by 1 km cell highlights regions of relatively high or low coverage.

790

791 10. Top panel: Surface magnetic field (red line) along the profile, compared to the modeled magnetic
792 signature of two magnetic bodies intersected by a few 100 m wide non-magnetic region (Weckmann et
793 al., 2007b). Middle panel: the two magnetic bodies drawn over the MT depth section of Weckmann et
794 al. (2007a). Lower panel: the magnetic bodies drawn over the northern 180 km of the P-wave velocity
795 model shown in Fig. 5. Vertical dashed lines across all three panels project the locations of the maxima
796 of the BMA onto the depth sections.

797

798 11. P-wave velocity model comparison between the Springbok profile on the western coast of South
799 Africa (top panel, after Hirsch et al., 2008), and this study (bottom panel). Based on Moho geometry,
800 four segments of the crust can be distinguished in both profiles: thick continental (A), stretched
801 gradually thinning (B), steeply thinning (C), and thin oceanic crust (D).

802

803 12. The combined onshore/offshore profile, showing the summary of the findings of this study. To the
804 north the Karoo and Cape Supergroups taper out in an asymmetric syncline. The division between
805 Upper and Middle/Lower Crust roughly follows the P-wave velocity of 6.5 km/s contour. Near its most
806 shallow point, the high velocity anomaly most likely is related to the BMA is found in the Middle
807 Crust. Immediately south a large synclinal feature was identified in the Upper Crust. In the Lower
808 Crust high velocity material has been observed on top of the steeply rising Moho, as was a reflector
809 roughly parallel to the Moho. Offshore the layer of sediments can be seen, as is the structure of the
810 AFFZ, which separates continental from oceanic crust.























