



Originally published as:

Smith, B., Jordan, M., Steinberger, B., Puskas, C., Farrell, J., Waite, G., Husen, S., O'Connell, R. J., Klingele, E. (2009): Geodynamics of the Yellowstone hotspot and mantle plume: Seismic and GPS imaging, kinematics and mantle flow. - Journal of Volcanology and Geothermal Research, 188, 1-3, 26-56

DOI: [10.1016/j.jvolgeores.2009.08.020](https://doi.org/10.1016/j.jvolgeores.2009.08.020)

Manuscript Number: VOLGEO2707

Title: Geodynamics of the Yellowstone Hotspot and Mantle Plume: Seismic and GPS  
Imaging, Kinematics, Mantle Flow

Article Type: Review Article

Keywords: mantle plume; tomography; volcanism; earthquakes; dynamics; intraplate;  
kinematics; convection; hotspot; plume; Yellowstone; Snake River Plain

Corresponding Author: Professor Robert B. Smith, PhD

Corresponding Author's Institution: University of Utah

First Author: Robert B. Smith, PhD

Order of Authors: Robert B. Smith, PhD; Michael Jordan; Bernhard Steinberger;  
Christine M Puskas; Jamie Farrell; Gregory P Waite; Stephan Husen; Wu-Lung Chang;  
Richard O'Connell

Abstract: The Yellowstone hotspot resulted from interaction of a mantle plume with the overriding North America plate highly modifying the lithosphere by magmatic-tectonic processes and producing the 17 Ma Yellowstone-Snake River Plain (YSRP) volcanic system. The accessibility of the YSRP has allowed large-scale geophysical experiments to seismically image the hotspot and to evaluate its kinematic and dynamic properties using geodetic measurements. Tomography reveals a Yellowstone crustal magma body with 8-15% melt that is fed by an upper-mantle plume extending from 80 km to 660 km deep and tilting 60° west. Contemporary deformation of the Yellowstone caldera is dominated by SW-extension at up to ~3 mm/yr, a fourth of the total Basin-Range opening rate, but with superimposed volcanic uplift and subsidence at decade scales, averaging ~2 cm/yr and unprecedented caldera uplift from 2004-2008 at up to 7 cm/yr. Convection models reveal eastward upper-mantle flow beneath Yellowstone at relatively high rates of 5 cm/yr and opposite in direction to the overriding N. American Plate. This strong flow deflects the ascending plume melt into a tilted configuration, i.e., the plume is caught in a mantle "wind". Dynamic models of the Yellowstone plume revealed relatively low excess temperatures, up to 120°K, with up to 1.5% melt, properties consistent with a weak buoyancy flux of ~0.25 Mg/s. The flux is several times smaller than for oceanic plumes, but it produced a ~600-km wide topographic ~300-m high swell. Employing the plume-geometry we extrapolated the location of the Yellowstone mantle-source southwestward to its initial position at 17 million years beneath eastern Oregon and the southern edge of the LIP Columbia Plateau basalt field suggesting a common origin. Our model suggests that the original plume head rose vertically behind the subducting Juan de Fuca plate, but at ~12 Ma it lost the protection of the subducting plate and encountered cooler, thicker continental lithosphere and became affected by the eastward upper-mantle flow. Regionally, excess gravitation potential energy of the swell drives the SW motion of the YSRP lithosphere that becomes part of a general clockwise rotation pattern of intraplate western U.S. tectonism. Our models thus demonstrate that plume-plate processes of the YSRP have

"continentalized" oceanic lithosphere enhancing intraplate extension and highly modifying topography, deep into the continental interior. Our results demonstrate that the dynamic properties of the Yellowstone hotspot deserved its recognition as a "window into the Earth's interior".

1 **Geodynamics of the Yellowstone Hotspot and Mantle Plume: Seismic and GPS Imaging,**  
2 **Kinematics, Mantle Flow**

3

4 Smith<sup>a</sup>, Robert B., Michael Jordan<sup>a,f</sup>, Bernhard Steinberger<sup>b</sup>, Christine M. Puskas<sup>a</sup>, Jamie  
5 Farrell<sup>a</sup>, Gregory P. Waite<sup>a,c</sup>, Stephan Husen<sup>a,d</sup>, WuLung Chang<sup>a</sup>, and Richard O'Connell<sup>e</sup>

6

7 <sup>a</sup>Department of Geology and Geophysics, University of Utah, Salt Lake City, Utah.

8 <sup>b</sup>Center for Geodynamics, Norwegian Geological Survey, Trondheim, Norway.

9 <sup>c</sup>Department of Geological and Mining Engineering and Sciences, Michigan Technological  
10 University, Houghton, Michigan.

11 <sup>d</sup>Swiss Seismological Service, Swiss Federal Institute of Technology, Zurich, Switzerland.

12 <sup>e</sup>Department of Earth and Planetary Sciences, Harvard University, Boston, Ma.

13 <sup>f</sup>SINTEF Petroleum Research, Seismic and Reservoir Technology, Trondheim, Norway

14

15 \*Corresponding author. Department of Geology and Geophysics, 135 S. 1460 E., rm. 702,  
16 University of Utah, Salt Lake City, Utah, 84112; E-mail address: r.smith@earth.utah.edu.

17

18 **Abstract**

19 The Yellowstone hotspot resulted from interaction of a mantle plume with the overriding North  
20 America plate, highly modifying the lithosphere by magmatic-tectonic processes and producing  
21 the 17 Ma Yellowstone-Snake River Plain (YSRP) volcanic system. The accessibility of the  
22 YSRP has allowed large-scale geophysical experiments to seismically image the hotspot and to  
23 evaluate its kinematic and dynamic properties using geodetic measurements. Tomography



24 reveals a Yellowstone crustal magma body with 8-15% melt that is fed by an upper-mantle  
25 plume extending from 80 km to 660 km deep and tilting 60° west. Contemporary deformation of  
26 the Yellowstone caldera is dominated by SW-extension at up to ~3 mm/yr, a fourth of the total  
27 Basin-Range opening rate, but with superimposed volcanic uplift and subsidence at decadal  
28 scales averaging ~2 cm/yr and unprecedented caldera uplift from 2004-2008 at up to 7 cm/yr.  
29 Convection models reveal eastward upper-mantle flow beneath Yellowstone at relatively high  
30 rates of 5 cm/yr and opposite in direction to the overriding North American Plate. This strong  
31 flow deflects the ascending plume melt into a tilted configuration, i.e., the plume is caught in a  
32 mantle “wind”. Dynamic models of the Yellowstone plume revealed relatively low excess  
33 temperatures, up to 120°K, with up to 1.5% melt, properties consistent with a weak buoyancy  
34 flux of ~0.25 Mg/s. The flux is several times smaller than for oceanic plumes, but it produced a  
35 ~600-km wide topographic ~300-m high swell. Employing the plume-geometry we extrapolated  
36 the location of the Yellowstone mantle-source southwestward to its initial position at 17 million  
37 years beneath eastern Oregon and the southern edge of the LIP Columbia Plateau basalt field  
38 suggesting a common origin. Our model suggests that the original plume head rose vertically  
39 behind the subducting Juan de Fuca plate, but at ~12 Ma it lost the protection of the subducting  
40 plate and encountered cooler, thicker continental lithosphere and became affected by the  
41 eastward upper-mantle flow. Regionally, excess gravitation potential energy of the swell drives  
42 the SW motion of the YSRP lithosphere that becomes part of a general clockwise rotation pattern  
43 of intraplate western U.S. tectonism. Our models thus demonstrate that plume-plate processes of  
44 the YSRP have “continentalized” oceanic lithosphere, enhancing intraplate extension and highly  
45 modifying topography of the continental interior. Our results demonstrate that the dynamic  
46 properties of the Yellowstone hotspot deserve recognition as a “window into the Earth’s

47 interior”.

48

49 **Keywords:** mantle plume, tomography, volcanism, earthquakes, dynamics, intraplate,

50 kinematics, convection, hotspot, plume, Yellowstone, Snake River Plain

51

## 52 **1. Introduction**

53 In its isolation from a plate boundary, the Yellowstone hotspot is the classic example of a  
54 continental hotspot. Age-transgressive volcanism has systematically modified the composition  
55 of the overriding North America plate, creating the 700 km long Yellowstone-Snake River Plain  
56 (YSRP) volcanic province over the last 17 Ma (Fig. 1, 2). The Yellowstone hotspot is the  
57 youthful part of this dynamic system centered on the Yellowstone volcanic field and its  
58 magmatic processes of have played an important role in the intraplate Cenozoic evolution of the  
59 western U.S. However, lithospheric structure of the YSRP and its hotspot-related magmatic  
60 processes, effects on surface deformation and seismicity have been poorly understood and  
61 widely debated, largely because of a lack of definitive data.

62 The focus of this paper is on the youngest component of the 1.8 Ma Yellowstone volcanic  
63 field, centered at Yellowstone National Park (Fig. 3), because of the importance of characterizing  
64 its active volcano-tectonic properties that bear on the evolution of the entire YSRP. With the  
65 availability of new seismic and GPS data summarized here, we model and interpret Yellowstone  
66 upper mantle and lithospheric structure, magmatic sources, and swell deformation. We also  
67 address kinematic and dynamic interactions within this system and how it has affected the rest of  
68 western U.S.

69           We first review our current knowledge of the Yellowstone hotspot and its volcano-  
70 tectonic field (hereafter called Yellowstone). We then present the seismic imaging of  
71 lithospheric structure focused on the origin and properties the Yellowstone hotspot, its mantle  
72 melt properties and evidence for a plume, and the resultant kinematics and dynamics from  
73 geodetic measurements. We conclude with a discussion of the large-scale properties of whole  
74 mantle convective flow and how mantle flow has influenced the geometry and location of the  
75 Yellowstone plume as well as on the dynamics of an upper mantle plume origin for Yellowstone.

76

## 77 **1.1 Hotspots and Plumes**

78           Largely because of their presumed association with the Earth's deep interior, plumes of  
79 ascending magma have been commonly thought of as the sources of volcanic hotspots, i.e. areas  
80 of long-lived concentrated volcanism (Wilson, 1963; Dietz and Holden; 1970, Crough, 1978).  
81 While most of Earth's volcanism is associated with plate boundaries including mid-ocean ridges  
82 or subduction zones, some hotspots occur within plates (Fig. 1). Yellowstone is an example of a  
83 continental hotspot and is located on the western side of the North American plate, about 1000  
84 km from its nearest plate boundary. Because of its accessibility, large-scale geophysical  
85 experiments were readily implemented and provide the key data used in this paper.

86           Hotspots have distinct physical properties. The most notable feature is that the passage  
87 of a plate over a plume results in a linear, time-transgressive volcanic chain and broad  
88 topographic swells (Ito and van Keken, 2007). Geochemically, the compositions of intruded and  
89 erupted hotspot magmas generally contain a component of mantle-derived melt, while the plume  
90 properties differ from the surrounding mantle in terms of composition, density, and temperature.  
91 Because they are hotter and less dense than typical mantle rock, plumes rise buoyantly and create

92 a corresponding topographic swell at the surface. Combined with the low-density plume, the  
93 swell creates a mass deficit at the hotspot that can be produce a notable gravity and geoid  
94 signature. Moreover, modeling the dynamics of both the shallow and deep mantle requires  
95 information on such factors as thermal buoyancy, lithologic heterogeneity, and laterally varying  
96 rheology.

97         The significance of hotspots can be seen in the Earth's gravity field (Tapley et al., 2005),  
98 where strong, long-wavelength positive anomalies are associated with the Hawaii, Iceland, and  
99 Yellowstone hotspots (Fig. 1). The magnitude of the Yellowstone anomaly of ~35 MGal is  
100 thought to reflect reduced density of the lithosphere and asthenosphere across the 1000 km width  
101 of the hotspot. The Yellowstone hotspot is associated with a strong geoid anomaly of -7 m with  
102 respect to the NAD83 ellipsoid and +15 m compared to the surrounding region over an area with  
103 a 1000 km diameter, approximately the same dimensions as topographic swells of oceanic  
104 hotspots (Fig. 1). An interpretation of the geoid anomaly is that it represents an amalgam of  
105 isostatically uncompensated high topography and a broad lithospheric-asthenospheric low-  
106 density zone.

107         The traditional hotspot hypothesis argues that plumes result from conduits of magma, 100  
108 to 200 km wide, that ascend through the mantle and produce long-lived volcanism at the surface.  
109 As a lithospheric plate moves across the plume, a volcanic center is displaced from the plume  
110 source and dies, while a new center forms above the plume in a cycle that propagates a  
111 characteristic chain of volcanic centers that grow older with horizontal distance from the plume.  
112 This process formed the age-transgressive 6000-km long Hawaiian-Emperor seamount chain in  
113 the Pacific Ocean over the past 80 million years. Continental volcanic features associated with  
114 hotspots include, beside Yellowstone, the European Eifel and Massif Central volcanic fields, and

115 extinct African hotspots expressed by volcanic mountains (Burke, 1996). For the details of  
116 hotspot geology, we recommend an objective summary of hotspots, melting anomalies, and  
117 plume arguments by Ito and van Keken (2008).

118         The plume-plate process presumably also produced the time transgressive 16-17 Ma  
119 Yellowstone-Snake River Plain-Newberry silicic volcanic field of the western U.S., the focus of  
120 this paper (Fig. 2). The northwest time-progression of the Newberry volcanic system and the  
121 High Lava Plains of Oregon (Jordan et al., 2004; Camp and Ross, 2004) is the west or mirror-  
122 imaged branch of the YSRP and we will not include a discussion of it in this paper as it is  
123 beyond the scope of this paper and breadth of the data acquired in the Yellowstone Geodynamics  
124 project.

125         In traditional thinking, plumes of buoyant mantle partial melt ascend from near the core-  
126 mantle boundary in near-vertical conduits and become entrained in convective flow (Morgan,  
127 1972). But new mantle flow models (Steinberger et al., 2004) reveal that plume tracks rise  
128 upward through convecting mantle flow and follow circuitous paths, i.e., material rises buoyantly  
129 along curved paths that follow the directions of mantle flow. Thus hotspots are not necessarily  
130 fixed and non-vertical mantle flow can tilt a plume by as much as 60°, as we show in our later  
131 discussions in Section 4-6.

132         However the plume hypothesis is not without conjecture. One reason is that narrow  
133 plume conduits have not been imaged by seismic tomography. Practical limitations are from the  
134 lateral extent of seismic networks and the low density of seismographs that prevent the resolution  
135 needed to resolve anomalous low wave speed bodies deeper than ~1000 km and less than ~100  
136 km in diameter. This has been a practical restriction in imaging low-velocity plume bodies from  
137 the core-mantle boundary to the lithosphere. In the last decade, many articles have been

138 published on the application of seismic methods to seismic studies and imaging of plumes. We  
139 will not go into details of the methodology and limitations, but refer the reader to a recent review  
140 of tomographic methodology by Nolet et al (2007).

141 There are multiple ideas on the origin of plumes. A widely held hypothesis argues that  
142 plumes form as upwellings associated with upper mantle convection, i.e. from the bottom of the  
143 transition zone at the 660-km discontinuity that separates the upper and lower mantle. Another  
144 theory is that plumes originate in the lower mantle from whole or partial-mantle convection. We  
145 will discuss aspects of these models in Section 6 of our paper.

146

## 147 **1.2 The Yellowstone Geodynamics Project**

148

149 The Yellowstone Geodynamics Project included a large-scale field experiment that was  
150 planned to give the seismic and geodetic data used for studies presented here. The project  
151 focused on a much larger scale than had done been before and included the entire Yellowstone  
152 hotspot, the eastern Snake River Plain, and surrounding 600 km by 600 km area (Fig. 4). Data  
153 from the field projects were then used to model the dynamic properties of the Yellowstone plume  
154 and the regional kinematic and dynamic properties of the western U.S. influenced by the  
155 Yellowstone hotspot.

156 Additional data was derived from long-term volcano and earthquake monitoring by  
157 regional seismic networks and detailed seismic and geodetic studies of Yellowstone National  
158 Park. The 25-station Yellowstone seismic network began operation in 1973 and covers  
159 Yellowstone National Park and the surrounding area to 50 km. This network is primarily  
160 supported by the USGS- and NPS-funded Yellowstone Volcano Observatory. Field GPS

161 measurements were made biannually from 1987 to 1995, and a network of permanent GPS  
162 stations was implemented in Yellowstone and the eastern Snake River Plain starting in 1996.

163 The Yellowstone Geodynamics Project operated from 1999 to 2005 and included  
164 extensive seismic data acquisition, processing, analysis, modeling, and interpretation. The data  
165 acquisition phase of the project consisted of operation of a temporary 80-station broadband and  
166 short-period array (50 temporary IRIS-PASSCAL stations and a special 30-station PASSCAL  
167 telemetered array) over an area ~800 km in diameter centered on Yellowstone (Fig. 4) (see Waite  
168 et al., 2005 for a detailed description of this project). The installation of 15 permanent GPS  
169 stations was also implemented along ~160 temporarily occupied GPS sites (Fig. 4) (see Puskas et  
170 al., 2007a for a detailed description of this project). In addition we also employed seismic data  
171 from regional seismic networks and data from the University of Utah and EarthScope PBO GPS  
172 networks to supplement our analyses.

173

### 174 **1.3 Other Seismic Plume Studies**

175

176 Past studies have not definitively seismically imaged a mantle plume. Most focused  
177 large-scale tomographic studies have relied on permanent and temporary deployments of seismic  
178 stations, i.e. on the islands of Iceland and Hawaii and for the Eifel volcanic field, Europe (Allen  
179 et al., 2002, Wolfe et al., 2002, and Ritter et al., 2001, respectively). Images from these studies  
180 reveal low P-wave velocity bodies beneath these areas of active volcanism but could not resolve  
181 plume-like bodies that could be reliably traced to depths greater than ~400 km. The main  
182 limitation of these studies was that their seismic arrays were not wide enough to resolve seismic

183 arrivals over the range of incidence angles required to sample deep velocity anomaly, a condition  
184 that can smear velocity anomaly over a large depth range without resolving a distinct image.

185 Early tomographic studies revealed a complex velocity structure in the crust and upper  
186 mantle beneath the Snake River Plain, southwest of the Yellowstone hotspot. Structural  
187 heterogeneity of the mid-crust was interpreted to represent compositional variability associated  
188 with the bimodal rhyolite-basaltic volcanism leaving a mid-crustal high density, high velocity  
189 mafic sill along the entire SRP composed of hundreds of individual sill intrusions (Sparlin et al.,  
190 1982; Annen et al., 2006; Shervais et al., 2006).

191 Variability in upper mantle structure including mafic crustal underplating of the SRP  
192 associated with melting (Saltzer and Humphreys, 1997). More recent work suggests that a  
193 narrow, low-velocity feature extends from the upper mantle into the top of the transition zone  
194 (Waite et al., 2006, Yuan and Dueker, 2005). The shallow upper mantle anomaly is present over  
195 a distance of more than 400 km, from the eastern extent of the Snake River Plain the northeast  
196 Yellowstone caldera. The anomaly is strongest at depths of 50 to 200 km with peak anomalies of  
197 -2.3% for  $V_p$  and -5.5% for  $V_s$  (Waite et al., 2006). The velocity reductions are interpreted to  
198 represent 1-2% partial melt at excess temperatures of  $\sim 170$  K (Jordan, et al. 2004, 2005; Schutt  
199 and Humphreys, 2004).

200 Receiver function studies indicate that the mantle transition zone, which separates the  
201 upper from the lower mantle, tends to be thinner when the hot rock of a plume intersects it,  
202 raising the 660 km discontinuity to shallower depths and depressing the 410 km discontinuity  
203 (Bina and Hellfrich, 1996). Initial transition-zone studies showed significant topography of the  
204 410 discontinuity beneath the Snake River Plain (Dueker and Sheehan, 1997). More recent  
205 studies show that the 410 discontinuity deepens by 10 km near the intersection of the low



206 velocity anomaly identified by Waite et al. (2006). Fee and Dueker's (2004) analysis showed  
207 that the 660 km discontinuity shallows by ~20 km beneath Yellowstone, a property that we used  
208 as a constraint for the tomography and dynamic modeling of the plume properties,

209       Seismic anisotropy by shear wave splitting measurements of the YSRP show that fast S-  
210 wave speeds of the upper mantle are aligned primarily with the direction of apparent plate  
211 motion, except for stations in the Yellowstone caldera, which reveal rotation of splitting  
212 directions around the crustal magma chamber (Waite et al., 2005). This method separates  
213 polarity components of shear waves from distant seismic sources and assumes that olivine  
214 crystals have their fast axes aligned with mantle flow, producing a difference in the velocity of  
215 seismic waves with azimuth through birefringence. Moreover the fast-axes were not  
216 significantly perturbed as expected in circular pattern for a strong mantle plume. This  
217 observation lead to Waite et al., (2005) to concludes that a strong plume, i.e., with a high  
218 buoyancy flux, was not responsible for the Yellowstone hotspot.

219       To evaluate the contemporary deformation field of the Yellowstone hotspot, GPS  
220 measurements were made to determine the directions of crustal motion. These data are used to  
221 distinguish volcanically induced deformation from tectonic deformation and with additional new  
222 data from other projects to evaluate the velocity field of the western U.S. particularly on the  
223 overall kinematics.

224       The kinematic data are then used to model the dynamic properties of the hotspot such as  
225 composition, temperature, hydrous content, etc. and to evaluate the effect on the regional  
226 deviatoric stress field from mass variations related to plume-plate interactions.

227       These properties were then compared to new models of plumes rising in a convecting  
228 mantle and to models of plume deflection following the methodology of (Steinberger and

229 O'Connell, 1998) and (Steinberger, 2000). This methodology employs numerical models of  
230 whole mantle convection assuming that mantle flow is driven kinematically by the motion of the  
231 overlying plates and dynamically by internal density variations estimated from seismic  
232 tomography models.

233 We conclude by speculating on the history of the Yellowstone plume source at the  
234 transition zone source depth and its effect on the 16-17-million year history of the YSRP system.  
235 This was done by extrapolating the mantle source backward in time and space to evaluate its  
236 correlation with the surface volcanic features of the YSRP, the High Lava Plains, and the  
237 Columbia Plateau basalt field.

238

## 239 **2. Volcano-Tectonic Setting of Yellowstone**

240 The Yellowstone volcanic field (Fig. 2) is located at the northeastern end of a 700 km-  
241 long, age-progressive, bimodal basalt-rhyolite system that defines the track of the Yellowstone  
242 hotspot (Armstrong et al., 1975). A mirror image of the YSRP extends northwest across the  
243 High Lava Plains to the Newberry caldera, Oregon (Geist and Richards, 1993). The latter zone,  
244 while not as volcanically voluminous as the SRP, reflects a similar silicic volcanic history  
245 associated with sub-lithosphere source of magmatism.

246 An arcuate shaped area of high topography, active faulting, and earthquakes that we call  
247 the tectonic parabola surrounds the subsided and seismically quiet Snake River Plain. The SRP  
248 is seismically quiescent at the M3+ level and lacks faulting but has notable Late Quaternary  
249 basaltic dikes of similar orientation to regional faults (Fig. 2). The dikes are hypothesized to  
250 structurally extend the crust at the same rate as the surrounding area of the northern Basin-Range

251 by non-explosive eruption of Holocene basalts covering older rhyolite flows (Rodgers et al.,  
252 1990; Parsons et al., 1998).

253         The Yellowstone hotspot has been the source of voluminous rhyolite tuffs and lavas, with  
254 eruptions often having volumes of hundreds to thousands of cubic kilometers and representing  
255 some of the most voluminous Quaternary volcanism on Earth over the history of the  
256 Yellowstone-Snake River Plain volcanic field. More than 140 giant silicic eruptions have been  
257 identified by the tephrochronology of ash flow tuffs associated with Yellowstone hotspot  
258 volcanism (Perkins and Nash, 2002). YSRP volcanism is similar to other Late Cenozoic silicic  
259 volcanic centers of the Basin and Range province but have rhyolites with higher magmatic  
260 emplacement temperatures and include tholeiitic rather than alkali olivine basalts, a  
261 characteristic shared with basalts of the Columbia Plateau and the High Lava Plains (Perkins and  
262 Nash, 2002).

263         The active component of the Yellowstone volcanic field is characterized by youthful 2  
264 Ma caldera-forming silicic magmatism, a high topographic swell, and a geoid anomaly, the latter  
265 of which extends outward hundreds of kilometers and has the same scale as oceanic hotspots  
266 (Smith and Braile, 1994; Smith and Siegel, 2000) (Fig. 1). Moreover, Yellowstone is recognized  
267 as a globally significant volcano-tectono system (Christiansen, 2001) that originates from a  
268 vigorous sub-crustal source of heat. Three caldera-forming explosive eruptions at 2, 1.6, and 0.6  
269 Ma formed the currently active Yellowstone volcanic field (Fig. 5). More than 50 post-caldera  
270 rhyolite flows have since covered the Yellowstone Plateau, the youngest 70 000 years ago  
271 (Christiansen, 2001). In addition, the Yellowstone caldera exhibits extraordinary high heat flow  
272 (Fig. 6), high rates of seismicity and earthquake swarms, including a M7.5 earthquake in 1959,  
273 unprecedented decade-long periods of uplift and subsidence at rates of up to 2.5 cm/yr, and the

274 world's greatest display of geysers and hot springs at Yellowstone National Park (e.g. Pierce and  
275 Morgan, 1992; Smith and Braile, 1993, 1994; Morgan et al., 1995).

276         The YSRP is an area of pronounced high heat flow that stands out regionally amongst the  
277 thermal provinces of North America (Blackwell et al., 2006). The averaged regional heat flux of  
278 the Snake River Plain is  $\sim 150 \text{ mWm}^{-2}$ , 50% higher than the Basin-Range to the south and three  
279 times higher than the background flux of the Rocky Mountains to the east (Fig. 6). But the  
280 extraordinarily high heat flow of  $\sim 2000 \text{ mWm}^{-2}$  of the Yellowstone volcanic field is more than  
281 30-40 times the heat flow of continents. About 25% of the total flux is due to conductive heat  
282 transfer fed by thermal conduction from crustal magma sources (to be discussed later) that in  
283 turn are fed by magmas from the Yellowstone mantle plume.

284         Throughout Yellowstone heat flow is highly variable because it is dominated by shallow  
285 convective fluid flow associated with movements of hydrothermal fluids and shallow magma  
286 migration (Fournier, 1989). While the averaged convective heat flux is about  $2000 \text{ mWm}^{-2}$ , it  
287 can attain extremely high values in thermal basins. Because of the highly lateral variability in  
288 convective flow, standard heat flow measurements can only be made in a few restricted sites  
289 (Fig. 6). The best example is from the heat flow studies of the northern Yellowstone Lake that is  
290 underlain by an active hydrothermal area (Morgan et. al, 1977; Smith and Blackwell, 2000). In  
291 this area hundreds of meters of lake sediments provide a thermal blanket capturing the total  
292 conductive and convective flux in localized area of hot springs, some as high as  $20\,000 \text{ mWm}^{-2}$ ,  
293 more than 300 hundred times the average flux of continents.

294         This extraordinarily high heat flow is thus the result of active volcanism from crustal and  
295 mantle magma sources that in turn can cause large variations in density and thus stress potential  
296 and deviatoric stresses, which can lead to earthquakes, faulting, and crustal deformation (see

297 DeNosaquo et al., 2008, this volume, for a discussion of densities and the influence of heat flow  
298 of the YSRP).

299         Visitors to Yellowstone expecting rugged Rocky Mountains are often surprised by the  
300 undulating, relatively flat, forested topography of the Yellowstone Plateau, much of which is  
301 ~500 m above the surrounding area. This flat terrain is the product of youthful Yellowstone  
302 volcanism that has filled and smoothed out the pre-existing topography with rhyolite and basalt  
303 flows and by glaciation, which has scoured off its sharp topographic features and filled the lakes  
304 with sediments.

305         Yellowstone is famous for the world's greatest concentration of hydrothermal features,  
306 with over 10 000 geysers, hot springs, and fumaroles (Fig. 7). These features are the  
307 manifestations of the enormous heat flow resulting from hot water circulating along fracture  
308 systems in the upper crust and heated by crystallizing magma at a depth of 10-15 km (Fournier,  
309 1989; Husen et al., 2004). The majority of Yellowstone's hydrothermal systems are located  
310 within the Yellowstone caldera; however significant hydrothermal systems are located outside  
311 the caldera including the Norris-Mammoth Corridor and Hot Springs Basin.

312         Although Yellowstone volcanic activity has been attributed to a conduit-like mantle  
313 plume (e.g., Morgan, 1972; Smith and Sbar, 1974; Anders and Sleep, 1992; Armstrong et al.,  
314 1975), others have suggested an upper mantle origin. Smith (1977) ascribed it to intraplate  
315 lithospheric extension that allowed hot, buoyant magma to rise from the upper mantle.  
316 Christiansen and McKee (1978) suggested that YSRP magmatism is related to a transition from  
317 extended crust to the south to less extended crust to the north. King and Anderson (1995)  
318 modeled Yellowstone volcanism as a result of asthenospheric flow perturbed by the cratonic

319 root. And Humphreys et al. (2000) suggest an upper mantle convective system of propagating  
320 melt production associated with North America plate motion.

321       Earliest suggestions of a mantle source for Yellowstone magmatism were based on  
322 earthquake imaging from a sparse seismic array (Iyer et al., 1981). Most recently Yuan and  
323 Dueker (2005) and Waite et al., (2006) have used the initial data from the 1999-2005  
324 geodynamics experiment, described in Section 1.2, to infer a mantle plume structure to depths of  
325 ~450 km.

326       However, relative to the Columbia River and contemporaneous Steens Basalts of Oregon,  
327 the subsequent Yellowstone-Snake River olivine tholeiitic basalts and high-alumina olivine  
328 tholeiitic basalts of the High Lava Plains have: 1) a major contribution from depleted (MORB-  
329 like) mantle (Carlson and Hart, 1988; Hart and Carlson, 1987), and 2) a sub-continental  
330 lithospheric mantle component (Leeman, 1982; Hart and Carlson, 1987). In addition, a primitive  
331 mantle component commonly associated with the lower mantle is evidenced by elevated  $^3\text{He}/^4\text{He}$   
332 values of hydrothermal waters and radiogenic isotope signatures of Yellowstone basalts of (Craig  
333 et al., 1978; Doe et al., 1982; Leeman, 1982; Kennedy et al., 1985 and Leeman et al., 2008, this  
334 volume). Thus, petrologic and geochemical data suggest a sub-crustal magmatic origin with  
335 contributions from the mantle, a depleted asthenosphere, and old continental lithosphere. These  
336 properties require unusually high temperatures.

337       Volcanism of the Yellowstone hotspot trend differs from typical mantle plume-related  
338 volcanism in several ways: 1) the associated Newberry hotspot has migrated northwestward  
339 across the High Lava Plains at rates comparable to that of Yellowstone but with much smaller  
340 eruptive volumes, nonetheless producing a petrologically similar mirror-image age progression  
341 of bimodal volcanism (Fig. 2) , 2) parallel upper-mantle thermal anomalies of comparable

342 seismic velocity reduction further south in the Basin and Range Province, 3) a lack of clear age  
343 progression among the related YSRP basalts from Quaternary eruptions extending the length of  
344 the YSRP from Yellowstone to the Cascades, and 4) Yellowstone magmatism and the northern  
345 margin of Basin-Range extension are interrelated.

346         The relationship of the Yellowstone hotspot volcanics to the voluminous LIP (Large  
347 Igneous Province) Columbia River 17 Ma basalt volcanism is problematic. While flood basalts  
348 are commonly associated with the onset of hotspot activity, the Columbia River basalts are less  
349 voluminous than most flood basalt provinces, they lie to the north of the commonly projected  
350 trace of the Yellowstone hotspot, and they are only the most active part of a long, narrow zone of  
351 nearly simultaneously active basaltic fissures that extend south into central Nevada (Fig. 2).

352         In addition to the origin of the Yellowstone hotspot, we are also interested in the relations  
353 among the young and ongoing tectonic and magmatic processes that are actively reconstructing  
354 and modifying the lithosphere. Yellowstone hotspot volcanism is superimposed on the tectonics  
355 of the 30 Ma Basin and Range province, an 800-km wide province of active faulting and  
356 extension driven by the gravitational collapse of a thickened lithosphere in response to cessation  
357 of subduction at the southwest plate margin and the development of the Pacific-North America  
358 transform boundary. Basin-Range extension is in turn superimposed on relic structures of crustal  
359 compression and thickening associated with 80 to 45-Ma Laramide and Sevier orogenies.

360         In short, the origin of the Yellowstone hotspot has been enigmatic. Is it a plume  
361 originating in the shallow or deep mantle? How much of its character is a result of upper mantle  
362 processes in a highly modified Archean lithosphere? How does mantle flow affect the geometry  
363 and structure of the plume? These are important problems that this paper will address.

364           Moreover Yellowstone is a window into the processes of lithospheric construction,  
365   destruction, and magmatism. And Yellowstone is leaving a record that can be compared with  
366   hypothesized hotspot-continent interactions that have been commonly invoked for similar  
367   features around the world.

368

## 369   **2.1 Yellowstone Seismicity**

370           The Yellowstone region is one of the most seismically active areas of the Intermountain  
371   west and Rocky Mountains and occupies much of the central and northern Intermountain  
372   Seismic Belt (ISB) (Smith and Sbar, 1974, Smith and Arabasz, 1991). Seismicity in this belt,  
373   which separates the active tectonism of the interior western U.S. from the stable part of the North  
374   American plate, is the result of a combination of extensional tectonism and local spatial and  
375   temporal variations in stress related to the magmatic and hydrothermal systems.

376           The Yellowstone seismic network spans the entire caldera and surrounding faults of the  
377   Yellowstone Plateau (Fig. 8). More than 30 000 earthquakes for the period 1973 to 2007 were  
378   recorded (Fig. 9), most of which had magnitudes less than 4 (Waite and Smith, 2004). We  
379   specifically note that depths of earthquakes and crustal tomographic images are with respect to a  
380   datum of 2 km, the average elevation of the Yellowstone Plateau.

381           Notably, Yellowstone has experienced large earthquakes including the August 1959  
382   M<sub>s</sub>7.5 Hebgen Lake, MT, earthquake, which was the largest historic earthquake of the western  
383   U.S. interior (Fig. 9 and 10a). It was located 25 km northwest of the north rim of the  
384   Yellowstone caldera (Doser, 1985) and its relationship to regional stresses associated with the  
385   Yellowstone swell are problematic.



386           The Hebgen Lake earthquake broke along a pair of west-trending normal faults totaling  
387 40-km long with up to 5.7 m of slip and resulted in the deaths of 28 people. In contrast, the  
388 Yellowstone caldera is characterized by frequent and smaller earthquakes, the largest the  $M_L 6.1$   
389 Norris Junction earthquake (Pitt et al., 1979), which was also the largest recorded event to occur  
390 within the Yellowstone caldera. Caldera rim faults (Fig. 10b) extend for tens of kilometers  
391 around the caldera and while they are not seismically active, they may have the potential for  
392 large earthquakes related to loading from stress changes caused by magmatic processes.

393           Seismicity in Yellowstone is characterized by extensive earthquake swarms punctuated  
394 with large scarp-forming  $M 7$  earthquakes, described above. Yellowstone historic seismicity is  
395 evaluated from precisely relocated hypocenters of the Yellowstone earthquake catalog,  
396 November 1972 to December 2007 (Husen and Smith, 2004). Hypocenters have been relocated  
397 using new three-dimensional P-wave velocity models and probabilistic earthquake location of  
398 Husen et al. (2004). In addition, new coda magnitudes for events between 1984 and 2007 were  
399 computed with an improved coda magnitude equation. Earthquakes were relocated using a  
400 nonlinear, probabilistic solution to the earthquake location problem.

401           The most intense seismicity in the Yellowstone region occurs northwest of the  
402 Yellowstone caldera between the Hebgen Lake fault and the northern rim of the caldera (Fig. 9).  
403 Cumulative seismic moment release in this region is an order of magnitude higher than inside the  
404 Yellowstone caldera (Puskas et al., 2007a). Epicenters in the northwestern part of Yellowstone  
405 form two distinct bands: one extending in an E-W direction from Hebgen Lake to Norris Geyser  
406 Basin, the other extending in a general NW-SE direction from Hebgen Lake to the northern rim  
407 of the Yellowstone caldera. The majority of earthquakes in this area occur between 5 and 10 km  
408 depth but focal depths  $> 12$  km are observed close to Hebgen Lake (Fig.11).

409           The largest historic earthquake swarm in Yellowstone, which began in October 1985,  
410 also occurred in the area northwest of the Yellowstone caldera. The activity during this swarm  
411 migrated away from the caldera at about 150 m/d, at a time when the deformation of the caldera  
412 was changing from uplift to subsidence. This evidence suggested the swarm was triggered by  
413 the pore pressure diffusion away from the caldera, possibly following an episodic release of  
414 trapped fluids (Waite and Smith, 2002). Many of the earthquakes in the E-W band north of the  
415 caldera also occurred during swarms, but they did not typically exhibit a migration of activity  
416 (Waite, 1999; Farrell, 2007). Most of these earthquake swarms occur as individual clusters  
417 within the band of general seismicity (Fig. 9).

418           Seismicity in the Yellowstone region correlates with the Late Quaternary Hebgen,  
419 Madison, and Gallatin faults (Smith and Arabasz, 1991; Miller and Smith, 1999). Some of the  
420 faults have been mapped and dated (Fig. 5 and 9), but active seismogenic structures are thought  
421 to be buried under rhyolitic flows that postdate the latest caldera-forming eruption. These faults  
422 are theorized to have experienced a significant increase in Coulomb failure stress due to the  
423 rupture of 1959  $M_S7.5$  Hebgen Lake earthquake (Chang and Smith, 2002), explaining the intense  
424 seismicity of northwestern Yellowstone National Park. The areas of largest increase in Coulomb  
425 failure stress coincide very closely with the two seismicity bands stretching east from Hebgen  
426 Lake to Norris Geyser Basin and southeast from Hebgen Lake to the northern rim of the  
427 Yellowstone caldera. The modeled increase in failure stress decreases beyond Norris Geyser  
428 Basin and beyond the northern rim of the caldera (Chang and Smith, 2002).

429           Another explanation for the high seismicity in northwest Yellowstone is thought to be  
430 from stresses arising from postseismic relaxation of the lower crust and upper mantle as a  
431 response to the Hebgen Lake Fault earthquake (Chang and Smith, 2008). Relaxation can

432 increase normal stress along faults in the Hebgen Lake area, hence encouraging the occurrence  
433 of local earthquakes, as evidenced by current GPS N-S extensions across the fault zone of 2-3  
434 mm/yr (Puskas et al., 2007a) and consistent with the local stress field estimated from local focal  
435 mechanisms (Waite and Smith, 2004).

436         The Yellowstone caldera itself is characterized by shallow and diffuse hypocenters with  
437 some individual clusters of earthquakes. The central part of the caldera has relatively low  
438 seismicity, and no distinct seismic patterns are associated with the Mallard Lake resurgent dome  
439 in the western part of the caldera. The majority of earthquakes in the caldera are less than 4 km  
440 deep, with notable northwest and southeast trending zones of earthquakes parallel to the post  
441 caldera volcanic vents (Fig. 5) (Husen and Smith, 2004). Some of the earthquake clusters inside  
442 the Yellowstone caldera are associated with major hydrothermal areas such as Upper and Lower  
443 Geyser Basins, West Thumb Geyser Basin, the central part of the Yellowstone Lake, and the  
444 Mud Volcano area (Fig. 9). The pronounced shallowing of focal depths beneath the Yellowstone  
445 caldera has been explained by a decrease in depth of the brittle-ductile transition zone, as  
446 discussed in Section 2.2.

447         Intense seismicity is also associated with the Norris Geyser Basin close to the northern  
448 rim of the Yellowstone caldera. Seismicity associated with hydrothermal areas commonly  
449 occurs as swarm-like behavior clustering in time and space (Upper Geyser Basin, central part of  
450 Yellowstone Lake), and sometimes seismicity is persistent in time (West Thumb Geyser Basin,  
451 Lower Geyser Basin) (Farrell, 2007).

452         Cumulative seismic moment release is significant for some of the hydrothermal areas,  
453 e.g. West Thumb Geyser Basin, Mud Volcano area, and central part of the Yellowstone Lake, as  
454 compared to the background seismicity. Focal depths of earthquakes close to hydrothermal areas

455 are generally shallow, < 4 km depth, and are in part associated with pressurization of  
456 hydrothermal fluids within the upper shallow crust.

457 In addition to the background seismicity, Yellowstone has experienced triggered  
458 earthquakes due to the passage of large-amplitude surface waves from distant earthquakes that  
459 cause rapid changes in the stress. Earthquake triggering in Yellowstone from distant earthquakes  
460 is known to have occurred due to the 1992 M7.4 Landers earthquake (Hill et al. 1993) and the  
461 2002 M7.9 Denali fault earthquake (DFE) (Husen et. al., 2004a, Husen et. al., 2004b). For  
462 example, 250 earthquakes were triggered within the first 24 hours following the passage of the  
463 DFE surface waves. Within this first 24-hour period 11 earthquakes with  $M_c > 2.5$  were  
464 recorded compared to nine earthquakes with  $M_c > 2.5$  in all of 2002 prior to the DFE (Husen et.  
465 al., 2004b). The triggered seismicity slowly decayed over the next few weeks with the  
466 magnitudes of triggered events ranging from <0.0 to 3.2.

467

468 The DFE also produced pronounced significant changes in geyser activity in  
469 Yellowstone. Of the 22 geysers that were being monitored during the winter of 2002-2003, 8  
470 displayed notable changes in their eruption intervals due to the passage of the DFE surface  
471 waves (Husen et. al. 2004a).

472

## 473 **2.2 Effects of High Temperatures On Earthquake Focal Depths**

474 Lateral variations in focal depths of earthquakes of the Yellowstone caldera are thought  
475 to reflect variations in the depth to the brittle-ductile transition. In Fig. 11, we show the 80<sup>th</sup>  
476 percentile depth of earthquakes as the brittle-ductile isosurface of constant temperature. Taking

477 the brittle-ductile transition temperature of 450°C, as hypothesized by Smith and Bruhn (1984)  
478 for extensional tectonic regimes, allows estimates of the conductive temperature gradient.

479 This distinctive shallowing of the seismogenic layer beneath the caldera is attributed to  
480 high temperatures that reduce the strength of the rock, transforming it from brittle to ductile  
481 behavior above a shallow high-temperature source, namely magma. With the caldera, the crust  
482 appears to behave in a quasi-plastic state at depths exceeding 4 km at temperatures greater than  
483 350°C to 450°C as determined from petrological constraints. Such hot rocks are incapable of  
484 sustaining shear stresses on faults (Smith and Bruhn, 1984). The maximum focal depths of ~15  
485 km occur about 10 km from the west side of the caldera and correspond to a thermal gradient of  
486 ~26°C/km. Inside the caldera, the average 80<sup>th</sup> percentile depth is 4 to 6 km and corresponds to a  
487 gradient of 110°C/km to 65°C/km. These values are considered the conductive component of  
488 heat flow and would correspond to heat flow values of ~250 mWm<sup>-2</sup> which is about 6 times  
489 smaller than the total conductive and convective heat flow of ~2000 mWm<sup>-2</sup> and thus requires a  
490 convective heat transfer of ~1750 mWm<sup>-2</sup>.

491 The relative magnitude of thermal convective heat transport is specified by the Nusselt  
492 number, defined as the ratio of the convective to conductive heat flow. For Yellowstone, the  
493 Nusselt number of 8 compares with the Nusselt number calculated for the Long Valley caldera,  
494 CA, of 6 to 8 (Hill, 1992).

495

### 496 **2.3 Earthquake Hazards**

497 While this paper is not a discussion of earthquake hazard of the region, the historic  
498 seismicity plus Late Quaternary fault data form the basis of a companion paper in this volume on  
499 the probabilistic seismic hazard analysis by White et al. (2008). Their results reveal that the

500 dominant earthquake hazard of the Yellowstone-Teton region are associated with the large  
501 normal faults of the area, the Teton, Red Mountain, Hebgen Lake, Madison, Gallatin and South  
502 Arm faults. None of the magmatic-related faults are considered seismogenic although they can  
503 produce seismic activity during periods of volcanic activity.

504         We note a reduction in earthquake hazard within the Yellowstone caldera that is due to  
505 the very narrow depth of the seismogenic zone caused the high temperature gradients, that  
506 restricts the fault to ~4 km deep and short, i.e., not more than the length of the resurgent dome  
507 graben faults of a few km. Comparing these data with the fault-magnitude parameterization of  
508 Wells and Coppersmith (1994), the maximum magnitude earthquake for the caldera would not be  
509 expected to exceed M6.5 to 6.8. Volcanically induced earthquakes are of course a possibility,  
510 but seldom exceed M6.5 for other volcanic areas including the Snake River Plain (Smith et al.,  
511 1996). For earthquakes directly associated with a dike intrusion, the maximum magnitudes are  
512 estimated to be ~M5.

513

## 514 **2.4 Crustal Structure**

515         The three-dimensional (3D) P-wave velocity and P- to S-wave velocity ratio structure of  
516 Yellowstone has been determined using local-earthquake body-wave tomography from data of  
517 the Yellowstone seismic network. Husen and Smith (2004) selected 3374 highest quality, local  
518 earthquakes between 1995 and 2001 to invert for the three-dimensional (3D) P-wave velocity,  
519  $V_p$ , and P- to S-wave velocity ratio ( $V_p/V_s$ ) structure. The results confirm the existence of a low-  
520  $V_p$  body at depths from ~8 km to the maximum depth of resolution of ~16 km (Fig. 12). The  
521 body is interpreted to represent hot, crystallizing magma beneath the Yellowstone caldera, i.e., a  
522 magma chamber.

523           The low-velocity body, directly beneath the Yellowstone caldera, suggests that it is the  
524 source of Yellowstone’s volcanism. Its apparent “U upward” shape, with the shallowest part of  
525 the velocity beneath the two resurgent domes, may be in part due to the lack of resolution  
526 between the shallowest bodies, or it may reflect shallow magma conduits that feed the resurgent  
527 domes and hence are the main source of Yellowstone caldera magmatism. For a maximum  $V_p$   
528 reduction of  $\sim 6\%$ , the corresponding melt fraction may be as large as 5-10%. Together with the  
529 geochemistry, this magma body is thus likely composed of a partial melt of basaltic and granitic  
530 composition that has fed Yellowstone’s Late Quaternary volcanism and in turn was fed from  
531 basaltic magma from a deeper mantle plume.

532           We note that these data are consistent with model for a silicic magma reservoir that is  
533 responsible for Yellowstone’s youthful volcanism and overlies a middle and lower crust intruded  
534 by plume-derived basalt (Lowenstern and Hurwitz, 2008). Their model, derived from  
535 geochemical arguments of  $\text{CO}_2$  flux and thermal models, suggest that silicic magma is a hybrid  
536 of crustal melts and residual liquid formed as mafic magma cools and crystallizes, and that  
537 magma rises closest to the surface (5–7 km depth) beneath the resurgent domes.

538           Another result of this analysis was the imaging of a shallow volume of anomalously low  
539  $V_p$  and  $V_p/V_s$  in the northwestern part of Yellowstone volcanic field at depths of  $< 2.0$  km.  
540 Models indicate that this anomaly can be interpreted as porous, gas-filled rock. The close spatial  
541 correlation of the observed anomalies and the occurrence of the largest earthquake swarm in  
542 historic time in Yellowstone, 1985, suggest that the gas may have originated as magmatic fluids  
543 released by crystallization of magma beneath the Yellowstone caldera.

544

## 545 **2.5 Yellowstone-Snake River Plain Stress Field**

546           The combination of regional extension and magmatic processes at Yellowstone produces  
547 a complex deformation pattern (Fig. 13) resulting from tectonic and volcanic processes.  
548 Yellowstone has undergone caldera-wide and localized episodes of historic uplift and subsidence  
549 (e.g. Wicks et al. 1998; Chang et al., 2007; Puskas et al., 2007a) as well as late Quaternary  
550 deformation as evidenced by the youthful volcanism and Holocene paleoearthquakes  
551 (Christiansen, 2001; Haller et al., 2004). Fig. 14 shows a summary of horizontal extensional  
552 strain and minimum principal stress directions determined from geological and geophysical  
553 studies of the region (Zoback and Zoback, 1991; Nabelek and Xia, 1995; Smith et al., 1996;  
554 Waite and Smith, 2004; Puskas et al., 2007a). For simplicity, we will approximate the minimum  
555 principal stress directions and focal mechanism T-axes as extension directions.

556           Crustal extension dominates the regional deformation pattern. Extension directions are  
557 generally NNE-SSW immediately south of Yellowstone (see White et al, 2008, this volume) and  
558 E-W from the Teton Range south into southeastern Idaho, but changes north and west of  
559 Yellowstone to NE-SW except in the vicinity of the 1959 Hebgen Lake earthquake where it is  
560 closer to N-S or NNE-SSW.

561           The regional change in extension direction is clearly seen within the Yellowstone  
562 Plateau. Focal mechanism T-axes and minimum principal stress directions from focal  
563 mechanism inversion are roughly N-S west of Yellowstone but rotate to NNE-SSW north of the  
564 Yellowstone caldera and then abruptly rotate to NE-SW at Norris.

565           The earthquake stress indicators are sparse within or south of the caldera so we include  
566 data from local earthquake shear-wave splitting (Waite and Chang, 2007). Fluid-filled cracks are  
567 preferentially aligned with crustal stresses, resulting in an anisotropic seismic medium (e.g.,  
568 Crampin and Chastin, 2003). Shear-wave fast anisotropy directions ( $\hat{\nu}$ ) were estimated by



569 determining shear-wave splitting from local earthquakes and the maximum and minimum  
570 horizontal stress directions from the anisotropy were inferred. Averaged directions of fast  
571 anisotropy were measured at each station. These arrows are approximately perpendicular to the  
572 directions of maximum horizontal extension interpolated from GPS-measured deformation  
573 (Puskas et al., 2007a), focal mechanism T-axes, and minimum principal stress directions.

574 The influence of post-seismic deformation following the 1959 Hebgen Lake earthquake  
575 was suggested as a factor in the abrupt change in the stress direction at Norris (Waite and Smith,  
576 2004). We also note that NW-SE alignments of post-caldera volcanic vents within the caldera  
577 likely represent zones of weakness related to structures that predate the 640 000-year-old  
578 Yellowstone caldera and its post-caldera rhyolite flows (Ruppel, 1972; Christiansen, 2001; Waite  
579 and Smith, 2004). The north to northwest vent alignments appear to link N-S oriented normal  
580 faults south of the caldera with N-S oriented normal faults to the north. For example, the  
581 Gallatin fault and the Red Mountain Fault zone may be linked by the NW-SE trending alignment  
582 of vents through the central caldera.

583 GPS and InSAR studies (Wicks et al., 1998; Wicks et al., 2006; Chang et al., 2007;  
584 Puskas et al., 2007a) have revealed temporal changes in the strain field at Yellowstone, but  
585 similar changes have not been identified in the seismically derived measures of stress or strain  
586 except in the case of the 1985 swarm. The direction of maximum principal stress rotated from  
587 vertical during that swarm to become horizontal and sub-parallel to the direction of earthquake  
588 activity migration.

589 In summary, the stress field of the YSRP is dominated by NE-SW extension that is  
590 hypothesized to have begun with inception of Basin-Range tectonism in this area at ~17 Ma.  
591 Extension implies minimum horizontal compressional stress that would amplify vertical

592 buoyancy of mantle melt that can enhance YSRP volcanism because of the reduced horizontal  
593 compressive stress.

594

### 595 **3. Crustal Deformation of the Yellowstone Plateau**

#### 596 **3.1 Geodetic Measurements and Data Processing**

597       Earliest geodetic measurements in Yellowstone were from precise leveling of  
598 benchmarks in 1923 in conjunction with road construction. The benchmarks were re-surveyed in  
599 1975-76-77 when the highways were upgraded (Pelton and Smith, 1982). Re-observations of  
600 these points discovered the unprecedented uplift of the entire Yellowstone caldera by up to 75  
601 mm.

602

603       These observations lead to the establishment of 15 permanent GPS stations that were  
604 installed in Yellowstone and the central Snake River Plain and surrounding areas for the  
605 Geodynamics project beginning in 1996. In addition ~160 campaign GPS stations were  
606 observed at various times in eight surveys (1985, 1987, 1989, 1991, 1993, 1997, 2000, and 2003)  
607 over the study area (Fig. 13), (see Puskas et al., 2007a for details of these field projects).

608       The purpose of the geodetic observations was to measure deformation rates of the  
609 Yellowstone volcanic field needed to model fault slip, magmatic fluid migration, and regional  
610 extension. The GPS measurements focused primarily on the deformation of the Hebgen Lake,  
611 MT, and Teton, WY, faults, the Yellowstone caldera, the eastern Snake River Plain, and the  
612 Yellowstone tectonic parabola.

613       Up to six GPS stations were occupied continuously and served as local reference points.  
614 Stations from the International GPS Service (IGS) (Mueller, 1991) were included as additional

615 base reference stations in 1995, 2000, and 2003. Coordinates of the IGS and CGPS stations were  
616 available in the global ITRF96 reference frame (Sillard et al., 1998) and allowed the Yellowstone  
617 stations to be tied to the global reference frame. These GPS observations generally achieved a  
618 precision of <3 mm in the horizontal component and < 10 mm in the vertical component (Puskas  
619 et al., 2007a).

620

### 621 **3.2 Kinematics of Yellowstone from GPS Observations**

622 GPS velocities were determined by calculating changes in station coordinates over time  
623 to obtain the station velocities. Velocities were assumed to be linear and calculated for all  
624 stations observed for two or more campaigns. The campaigns were sorted into three periods to  
625 document temporal changes in Yellowstone: 1987-1995, 1995-2000, and 2000-2003. These  
626 observations had horizontal velocity errors of typically 1-2 mm/yr, while vertical errors were ~10  
627 mm/yr for most stations.

628 At least three stations of the 1987-1995 campaigns were selected as fixed stations, and  
629 their coordinates and velocities were tightly constrained to positions and velocities from the  
630 1995-2000 positions and velocities. This was done because the earlier campaigns predated the  
631 establishment of the IGS network in 1993. IGS stations were used as reference points and their  
632 positions and velocities were constrained in the North America fixed reference frame (Bennett et  
633 al. 2001) for the 1995-2000 and 2000-2003 time windows. In this reference frame, it is assumed  
634 that there is no deformation in the stable continental interior, and all velocities are with respect to  
635 the continental interior, also called stable North America.

636 The velocities for the three time periods revealed notable deformation changes in  
637 horizontal and vertical velocities (Fig. 13). These data show an unexpected pattern of alternating

638 subsidence and uplift of the Yellowstone caldera at up to 2 cm/yr, uplift northwest of the caldera,  
639 and regional extension at 2-4 mm/yr across the Hebgen Lake fault zone. During the observing  
640 periods, stations southwest of the caldera and outside the aforementioned regions moved  
641 southwest, indicating regional extension with respect to stable North America (Fig. 13).

642         The caldera floor sank from 1987-1995, rose slightly from 1995-2000, and resumed  
643 subsidence in 2000-2003 – *evidence of a living breathing caldera*. Subsidence occurred on the  
644 long axis of the caldera, reversing a trend of uplift that had been measured by leveling between  
645 1923-1985 (Pelton and Smith, 1982; Dzurisin and Yamashita, 1987; Holdahl and Dzurisin,  
646 1991). Notably, 112 mm maximum subsidence occurred on and around the northern caldera  
647 resurgent dome. In 1995, this subsidence stopped, and uplift shifted to the northwest caldera  
648 boundary area and continued through 2003, for a total of 117 mm. When subsidence resumed in  
649 2000, the caldera sank an additional 27 mm in a three-year period. This subsidence was  
650 concurrent with ongoing uplift to the northwest.

651  
652         The episodic caldera motions were corroborated by the permanent, but less dense, GPS  
653 stations. Five of the permanent sites operated from 2000 and 2003 and showed a net subsidence,  
654 while the single station northwest of the caldera had a net uplift.

655         Vertical displacement fields (Fig. 13) from the individual station velocities clearly  
656 correlate with the variations in horizontal spatial deformation. During periods of subsidence,  
657 caldera vectors pointed inward toward the long axis, while during periods of uplift vectors  
658 diverged radially from the center of uplift. Both these trends were discernible in 2003, when  
659 uplift continued at the northwest caldera and subsidence resumed within the caldera.

660 Notable extension occurred across the Hebgen Lake fault during the study period at rates  
661 from 3 to 5 mm/yr, with the highest rate in 1995-2000. This implies that the deformation here  
662 may have been influenced by uplift in the northwest caldera. Significant uplift was also  
663 observed at most of the stations near the fault (Fig. 13). At least part of this uplift episode that  
664 has been explained by post-seismic viscoelastic relaxation of the lower crust and upper mantle  
665 following the 1959 M7.5 Hebgen Lake earthquake that produced deformation of up to 2 mm/yr  
666 north of the fault (Chang and Smith, 2005).

667 The Snake River Plain area adjacent to Yellowstone moved southwest at  $2.4 \pm 0.4$  mm/yr  
668 from 1995 to 2000. This motion was less than southwest extension rate across the caldera for the  
669 same time period, implying compression between the Yellowstone Plateau and the Snake River  
670 Plain. However, southwest extension rates across the caldera varied for the different time  
671 windows, and the compression may have been a transient phenomenon.

672

### 673 **3.3 Magmatic Source Modeling of Deformation Data**

674 Volumetric strain modeling was employed to determine the configuration and depths of  
675 source bodies responsible for uplift and subsidence of the caldera. The crust of the caldera was  
676 divided into a 3D grid with three depth layers. The modeling algorithm related the observed  
677 surface motions to volumetric changes of grid blocks by a Green's displacement functions  
678 (Vasco et al., 1988). Surface motions, measured by leveling surveys between 1987 and 1993 and  
679 InSAR (Interferometric Synthetic Aperture Radar) between 1992 and 2002, were combined with  
680 the GPS data for 1987-2003. Total surface motions were obtained from the weighted sum of  
681 volume changes in individual grid blocks (Vasco et al., 1990). A linear least squares inversion  
682 was used to minimize the sum of squares of the residuals.

683 For 1992-1995, the modeled volumetric decrease was located on the caldera axis,  
684 between the resurgent domes and extended from 6-10 km depth (Vasco et al., 2007). The  
685 deflation rate during this interval was  $8.7 \times 10^{-3} \text{ km}^3/\text{yr}$  for a total volume loss  $2.6 \times 10^{-2} \text{ km}^3$ .  
686 For 1996-2000, caldera uplift was modeled by volume increase below the northwest caldera  
687 boundary at 6-10 km depth. The inflation rate was  $4.6 \times 10^{-3} \text{ km}^3/\text{yr}$  and the total volume loss  
688 was  $1.86 \times 10^{-2} \text{ km}^3$ . Additional models were for 2000-2001 and 2001-2002 using InSAR data  
689 only. These models imaged volume decrease along the caldera axis at depths of 6-8 km with  
690 uplift along the north caldera boundary at 4-6 km deep for 2000-2001 and 2-4 km deep for 2001-  
691 2002.

692 Interestingly, the deeper parts of the volumetric source models overlap with the top of the  
693 seismically imaged magma chamber (Husen et al., 2004). This argues that the subsidence source  
694 can originate within the upper portions of the magma chamber. The modeling assumes that the  
695 volumetric surfaces are continuous with depth which implies that, at least for caldera subsidence,  
696 source bodies extend from the near-surface through the brittle-ductile transition at  $\sim 6$  km to the  
697 top of the magma chamber.

698 Moreover, the shallow inflating volume models of the northwest caldera may be related  
699 to a low-velocity zone,  $< 2$  km depth, in the same region (Husen et al., 2004). The low velocities  
700 were interpreted as due to porous, gas-filled rock, where the gases were released from the  
701 magma chamber.

702 For fluids to produce crustal uplift, they must be trapped by an impermeable barrier that  
703 allows pore pressure to increase. The brittle-ductile transition has been proposed as a barrier to  
704 fluid flow (e.g., Fournier and Pitt, 1985; Waite and Smith, 2002, see also Section 3.4). At the  
705 near-surface, stratigraphic barriers could impede fluid flow as well.

706

### 707 **3.4 Temporal Correlation of Earthquakes and Deformation**

708         The correlation of earthquakes and deformation of Yellowstone from 1973 to 2006 is  
709 shown in Fig. 15. These data reveal long-term episodes of uplift and subsidence at rates of up to  
710 2.2 cm per year. The discovery of historic Yellowstone uplift from leveling re-observations in  
711 1975-77-76 by Pelton and Smith (1982) was considered an important signal of magmatic  
712 activity. However, the rapid change to subsidence at the time of Yellowstone's largest  
713 earthquake swarm in fall 1985 (Waite and Smith, 2004) was totally unexpected and prompted us  
714 to begin thinking of the causative relationships between earthquakes and magmatism.

715         For example, the seismicity of Yellowstone is characterized by intense swarms of  
716 generally small magnitude ( $M_C < 3$ ) earthquakes (Waite and Smith, 2004). The largest swarm  
717 occurred in late autumn 1985 adjacent to the northwest rim of the caldera and 10 km southeast of  
718 the eastern extent of the 1959 M7.5 Hebgen Lake earthquake rupture. More than 3000  
719 earthquakes were recorded in the swarm during a three-month period. The swarm was  
720 concomitant with a major reversal in caldera deformation, two crater-forming explosions, and  
721 several observed changes in the behavior of hot springs and geysers (Fig. 15).

722         Many of the earthquakes of the 1985 swarm were of normal, dip-slip faulting style, which  
723 is the expected type of faulting in the eastern Basin-Range extensional regime including  
724 Yellowstone. 75% of the focal mechanisms determined for the first month of the swarm were  
725 oblique-normal strike-slip, which is unusual for this region of crustal extension. Principal stress  
726 axis directions computed from the focal mechanisms, as well as the rate of migration of  
727 earthquake activity, were interpreted by Waite and Smith (2004) as consistent with seismicity  
728 induced by magmatic dike propagation. It was thus plausible that magmatic or hydrothermal

729 fluid was ejected from beneath the southern resurgent dome to the northwest and induced  
730 earthquakes once it reached the brittle crust. The volume loss as fluids escaped could be  
731 partially responsible for the change in caldera surface deformation from uplift to subsidence.

732 Following the 1985 caldera reversal, subsidence continued for a decade until 1995, when  
733 the caldera motion began a period of 5 years of minor uplift followed by renewed subsidence  
734 until a sudden change to accelerated uplift of the caldera at unprecedentedly high rates. At the  
735 inception of the 1995 uplift, seismicity began to increase until the onset of accelerated uplift in  
736 late 2004 (Fig. 15).

737

### 738 **3.5 Accelerated Uplift of the Yellowstone Caldera, 2006-2008**

739 GPS and InSAR measurements revealed that the Yellowstone caldera began an episode  
740 of ground uplift in mid-2004 at unexpectedly high rates of up to 7 cm/yr, three times greater than  
741 previously observed deformation episodes (Chang et al. 2007). Source modeling of the  
742 deformation suggests a near-horizontal expanding magma body over an area  $40 \times 60 \text{ km}^2$ , at 9  
743 km beneath the caldera, notably located near the top of the seismically imaged crustal magma  
744 chamber.

745 Importantly, the estimated rate of volumetric expansion of  $\sim 0.1 \text{ km}^3/\text{yr}$  for this uplift  
746 episode is similar to the magma intrusion rate required to supply the extraordinarily high heat  
747 flow of  $\sim 2000 \text{ mW/m}^2$  of Yellowstone (Fournier, 1989). This evidence suggests that these are  
748 the first observations of caldera-wide magma recharge of the Yellowstone volcanic system. In  
749 addition tens to hundreds of small earthquakes ( $M < 3$ ) occurred during the deformation period  
750 and were concentrated within the modeled dilatation zone while the rest of the caldera  
751 experienced low seismicity.



752

#### 753 **4. Tomographic Imaging of the Upper Mantle**

754 In this part of our discussion we evaluate the teleseismic delay-time data that is used to  
755 provide new images of the upper mantle structure of the Yellowstone hotspot. New data are  
756 modeled and interpreted from a special Yellowstone hotspot seismic array deployment from  
757 1999-2002 of 80 broadband/short period stations that were operated in two PASSCAL-supported  
758 deployments (50 temporary stations and the 30-station telemetered array with additional data  
759 from 45 stations of the Yellowstone, Montana, and Utah regional seismic networks, with a total  
760 of nearly 200 seismic stations). The arrays were deployed in a 500 km by 600 km area centered  
761 on Yellowstone for ~11 months each (Fig. 12).

762 Earlier interpretations from these data by Yuan and Dueker (2005) and Waite et al.  
763 (2005, 2006) revealed a 60° west-dipping low P-wave velocity body to depths of ~450km that  
764 was interpreted as a mantle plume. We also note the analysis of anomalous low shear wave  
765 velocities of the YSRP system evaluated from surface wave analysis of these to depths of 200  
766 km (Shutt et al., 2008).

767 In this paper we employ the use of an optimized tomographic inversion (Jordan, 2003;  
768 Barth et al., 2007; Wullner et al., 2006) to develop an updated image of the upper mantle seismic  
769 structure related to the Yellowstone hotspot. The methodology uses seismic inversion of  
770 teleseismic P-wave and S-wave delays constrained by other data such as the seismic structure of  
771 the mantle discontinuity structure (Fee and Dueker, 2004), the crustal structure of the  
772 Intermountain region (Lynch et al, 1997), the detailed crustal structure of Yellowstone (Husen et  
773 al., 2004), and the 2003 GEOID data (Fig. 12).

774 The delay time data were taken from the earlier study of Waite et al. (2006) consisting of

775 recordings of 103 teleseismic events from the special Yellowstone hotspot network of 85 seismic  
776 stations. Based on the estimated accuracy of the timing, the data were sorted into three quality  
777 classes ( $\pm 0.01$  s,  $\pm 0.03$  s and  $\pm 0.05$  s) for weighting in the inversion scheme. Moreover, the  
778 data consistency was extensively tested and problematic data were removed from the data set,  
779 resulting in a total of 4,319 travel time residuals from P and PKiKP phases. P-wave travel time  
780 residuals were calculated by subtracting the theoretical travel time for the IASP91 model  
781 (Kennett and Engdahl, 1991) from the observed travel time. S and SKS data were used in our  
782 earlier models but we did not employ the S-wave data in our inversions, but used the  $V_s$  results  
783 of Waite et al. (2006).

784         To remove the effect of source mis-location and source structure, we removed the mean  
785 residual travel time from all stations for each event. This results in relative travel time residual  
786 data, mainly generated by seismic velocity perturbations beneath the station network, which  
787 show late arrivals of PKiKP phases associated with underlying low seismic velocities in three  
788 main regions, namely the Yellowstone National Park, the Snake River Plain, and the area to the  
789 NW of Yellowstone.

790         For the modeling, we employed the JI-3D inversion method (Jordan, 2003), designed to  
791 provide stable and highly resolved models, both in the mathematical and spatial sense. The  
792 inversion algorithm is based on a Bayes approach (Tarantola and Valette, 1982; Zeyen and  
793 Achauer, 1997), which allows including a priori information, e.g., crustal structure and 1<sup>st</sup>-order  
794 discontinuities, in the inversion via an a priori covariance matrix. This is especially important  
795 since these features can have a significant effect on the observed delay-time data, but usually  
796 cannot be resolved by teleseismic tomography.

797           The forward calculations of ray paths, travel times, and Frechet derivatives are based on  
798 standard ray theory. The full 3D ray tracing is performed iteratively, using a simplex algorithm,  
799 and has been adapted from Steck and Prothero (1991). The chosen step length is 100 m and the  
800 cut-off for the simplex search  $2 \times 10^{-7}$  s. We use 15 harmonics with initial amplitudes of 1 km.  
801 These settings are clearly optimized to yield maximum accuracy at the expense of computation  
802 time. Though there is an ongoing debate about the validity of standard ray theory in  
803 tomographic problems, our reconstruction test modeling as well as other studies that employed  
804 the JI-3D method (e.g., Keyser et al., 2002; Jordan, 2003; Barth et al., 2007) show that the use  
805 of standard ray tracing provides adequate and realistic results. This is also backed by van der  
806 Hilst and de Hoop (2005). To reduce non-uniqueness and to improve stability of the inversion  
807 problem we implemented a variable optimized parameterization scheme that is explained in  
808 detail below.

809

#### 810 **4.1 Tomographic Model Parameterization**

811           We employed a 1-D IASP91 (Kennett and Engdahl, 1991) homogenous starting model.  
812 The model extends to 790 km depth and consists of 22 layers. Layers corresponding to the  
813 crustal structure at 14 km depth, the Moho, and the 410 km and 660 km discontinuities,  
814 respectively, were included as a priori information with their parameters were fixed during the  
815 inversion process by attributing small values in the a priori covariance matrix of the velocity  
816 model parameters. The JI-3D program relies on an optimized parameterization (Jordan, 2003)  
817 that is variable and performed strictly according to the information distribution in the model  
818 space, i.e., the angular distribution of rays.

819           Blocks and nodes define the 3D velocity model. While a particular block defines the  
820 region that contributes to the calculation of the each model parameter, the actual model  
821 parameter is defined at the node, located in the center of the block. Because the velocity is  
822 interpolated linearly between the nodes, this also avoids artificial velocity jumps at block  
823 boundaries. The parameterization itself is an iterative process that uses the diagonal elements of  
824 the resolution matrix as a measure to decide if the model parameters are equally well resolved  
825 and the inversion is stable. The resulting model is defined by small blocks and dense nodal  
826 spacing where the ray distribution has the highest density. The block sizes in our model range  
827 from 10 km in the best-sampled areas of the model space to several hundred kilometers at the  
828 lateral edges of the model.

829           Our tomographic inversion incorporated a priori information from known crustal  
830 structure and the topography of the Moho and the geometry of the 410 km and 660 km  
831 discontinuities. The crustal a priori information comprises P-wave models from high-resolution  
832 local earthquake tomography in the Yellowstone National Park area (Husen et al., 2004) and the  
833 intermountain region (Lynch et al., 1997). For the remaining model area, the a priori  
834 information was provided by the CRUST2.0 model (the model is online at  
835 <http://mahi.ucsd.edu/Gabi/rem.html>) (Basin et al. 2000). After careful assessment of the  
836 resolution properties of those models, a common reliable depth range was determined as -3 km to  
837 14 km.

838           Important contributions to the a priori crustal model are two highly resolved low-velocity  
839 bodies (Husen et al., 2004) in the mid to upper crust beneath Yellowstone. An  $\sim 8\%$   $V_p$  low-  
840 velocity body has been interpreted as a hot body with up to a few percent partial melt and a  
841 second shallower body with up to  $10\%$   $V_p$  reduction is believed to be a gas body. The Moho

842 topography and velocities also are taken from CRUST2.0, while the topography data at 410 and  
843 660 km were obtained from a model by Fee and Dueker (2004). These data are transformed into  
844 velocity contrasts of layers at the respective depths and also treated as a priori information.

845

## 846 **4.2 Tomographic Inversion and Resolution**

847 To derive the model parameters, we invert the travel-time P-wave delay data, starting  
848 with a homogeneous initial model for the non-constrained parameters. Due to the nonlinear  
849 nature of the problem, we perform four iterations. The result depends on the a priori variances of  
850 the model parameters, or the damping parameters. These are determined by a trial and error  
851 approach. Several iterative inversions are calculated for different variances. We define the  
852 optimum solution as a simple model that provides a large reduction of the data misfit. Due to the  
853 stability of the method a moderate change of the optimum damping merely changes the  
854 amplitudes but not the features of the model. Since the resolution depends on the ray  
855 distribution, which can change during the iterative inversion, we recheck the resolution matrix to  
856 ensure that the model still provides optimum stability conditions, which implies that the diagonal  
857 elements of the resolution matrix still are “constant”.

858 The corresponding variance reduction was 30.5%. Although the achieved variance  
859 reduction is relatively low, the resolution power of the inversion is not compromised. This low  
860 value is due to using large blocks, covering heterogeneous structure, so that the data sampling  
861 these blocks often sample different velocity contrasts and thus contradict each other. To reduce  
862 the effect of arbitrarily positioned blocks and nodes (with respect to the true structure) we apply  
863 an offset and average procedure (Evans and Achauer, 1993) during the last iteration step.

864 Checkerboard tests are often applied to examine one or more layers at the same time.  
865 However, due to the irregular parameterization in our approach, the ability to recover the  
866 synthetic input structure depends on the block size and relative position of the cells of the  
867 checkerboard with respect to the blocks of the inversion model. Therefore we do not consider  
868 checkerboard tests as suitable in this case.

869 The resolution of an inversion solution can also be assessed using either the full  
870 resolution matrix or its diagonal elements to evaluate the spatial distribution of well and less  
871 well-resolved model parameters. The inversion method applied here relies on the optimized  
872 parameterization, to keep the diagonal elements of the resolution matrix constant. As a  
873 consequence, the resolution matrix cannot be used to assess the resolution properties of the  
874 inversion since it is designed to possess near identical values, yielding no further insight. Hence  
875 we employ reconstruction tests to evaluate the resolution of our inversion results (See figures in  
876 the supplementary data).

877 Reconstruction tests consist of the forward calculation of a data set according to a  
878 synthetic structure in the model space utilizing the same source and receiver distribution as in the  
879 real inversion. Gaussian errors are added to the resulting data with realistic standard deviations  
880 and the data are inverted in the same way as the real data. The inversion result shows how an  
881 anomaly at the position of the synthetic input structure can be recovered in terms of location and  
882 amplitude and also shows possible smearing along main ray directions.

883

### 884 **4.3 Seismic Images of the Yellowstone Hotspot Mantle**

885 Our tomographic images (Fig. 17 and 18) reveal prominent upper-mantle low-velocity  
886 anomaly bodies beneath Yellowstone and surrounding regions with several distinct features. The

887  $V_P$  (this study) and  $V_S$  (Waite et al., 2006) tomography reveal a strong low-velocity anomaly  
888 from ~30 to 250 km beneath the Yellowstone caldera and from 30 to 100 km beneath the eastern  
889 Snake River Plain. Peak anomalies are  $-2.3\%$  for  $V_P$  and  $-5.5\%$  for  $V_S$  (Waite et al., 2006). A  
890 weaker, smaller-volume anomaly with a peak velocity perturbation of  $\sim -0.75\%$   $V_P$  (and  $-2.5\%$   
891  $V_S$ ) (Waite et al., 2006) is imaged from about 250 km depth beneath the caldera to 650 km depth  
892 at a position  $\sim 100$  km WNW of the caldera and dipping  $60^\circ$  (Fig. 19).

893 A variety of synthetic tests confirm these anomalies (see figures in the supplementary  
894 material) and indicate that the resolved amplitudes are too low. According to the findings of the  
895 reconstruction tests, we determine the real  $V_P$  amplitudes of the low velocity anomalies as  $-3\%$   
896 between 30 and 200 km depth and  $-1\%$  between 200 and 650 km depth. These values are  
897 included in the schematic representation of the geodynamic plume model in Fig. 20. The  
898 combined low-velocity anomaly is interpreted as a plume that extends from the transition zone  
899 and that promotes small-scale convection in the uppermost 200 km of the mantle.

900 In addition to the low-velocity plume structure described above, there are two additional  
901 blob-like low velocity bodies in the transition zone in the layers labeled 428 km 528 km and 572  
902 km in Fig. 17. These depths correspond to the respective nodal layers. These structures are also  
903 somewhat recognizable in the last two model layers, below the transition zone. We again utilize  
904 synthetic tests to determine where these blobs are located and if the image is affected by  
905 smearing.

906 Within the mantle transition zone, small-scale anomalies as small as  $50 \times 100$  km and  $-$   
907  $0.5\%$  velocity contrasts can be reliably resolved, where permitted by the parameterization, with  
908 only little upward and some downward smearing. In case of the plume, the latter shows an offset  
909 of 100 km in a westward direction and is outside of the region underlying the station network

910 that defined as best resolved. Our tests show that the true velocity reduction within the transition  
911 zone again is on the order of -1%. The resolution in the region beneath the transition zone is still  
912 sufficient to recover two additional low velocity structures in the southwestern part and  
913 southeastern part of the model. However, the corresponding amplitudes are only 30% of the  
914 input anomalies. Below the transition zone neither of the weak low velocity anomalies is located  
915 in the densely parameterized part of the model any more. The reconstruction tests indicate that  
916 they might be caused by downward smearing of the structure within the transition zone.

917         Considering the absence of local uplift of the 660 km discontinuity boundary, the position  
918 of the deepest anomaly outside the network region, and no obvious continuation of the plume-  
919 like structure into the lower mantle from global tomography (Montelli et al., 2004, van der Hilst  
920 et al., 2005), the anomaly below 660 km may be due to smearing effects from transition zone  
921 structure into the lower mantle along the main ray direction and is not considered as reliable a  
922 result from our modeling as the shallower structure. Hence, we consider the isosurface in Figs.  
923 19 and 20 as well resolved with a corrected amplitude of -1% instead of -0.5%.

924         The deeper structure to ~800 km, the maximum depth of our resolution, does not have a  
925 strong velocity contrast, but a horizontally wide zone extending beyond the area of our image of  
926 increased temperature would raise the 660 km discontinuity in the whole area around  
927 Yellowstone. Thus there are no horizontal differences that our data can resolve and that there is  
928 no uplift in the local 660 km discontinuity associated with the plume where we see it.

929         We also note a prominent high-velocity anomaly, +1.2%  $V_P$  and +1.9%  $V_S$ , that is  
930 located at ~50 to 200 km depth SE of Yellowstone and coincident with thicker lithosphere of the  
931 stable interior. This area is also above a region of prominent Laramide contractional folds and  
932 thrusts and part of the Precambrian core of North America. Yuan and Dueker (2005) and Waite



933 et al. (2006) described this structure as potentially the downwelling arm of mantle convection,  
934 but the large-scale mantle-convection models described in Section 6 argue that the return flow is  
935 nearly horizontal in this region and flowing in an easterly direction. Thus this anomaly may  
936 reflect a remnant structure of the tectosphere.

937 Mantle structure on a broad scale is revealed by whole-mantle tomographic images. A  
938 NW-SE cross-section through P-wave model by Montelli et al. (2004) and by van der Hilst et al.  
939 (2005) passing through Yellowstone illustrates clearly the location of Yellowstone with a west-  
940 dipping low-velocity body of -1% extending to depths of <1000 km. New data emerging from  
941 the USArray data (Xue and Allen, 2007; Burdick, 2008; Sigloch et al., 2008;) reveal a 0.5% to  
942 1.2% low velocity body to depths of up to 1000 km underlying most of the western U.S. This  
943 pillow of low velocity material may reflect the deep underpinning of the lithospheric thermal  
944 upwelling and extension of the Basin and Range province, with a small area of leakage forming  
945 the Yellowstone plume source.

946 The Yellowstone upper-mantle low-velocity layer is also near the boundary of marked  
947 lithospheric change where a midmantle low-velocity pillow beneath the Basin-Range to the  
948 southwest abuts the North American high-velocity craton to the east. A major high-velocity  
949 body apparently continuous with the craton occupies most of the thickness of the upper mantle  
950 beneath Yellowstone.

951 This leads to another possible explanation of the Yellowstone plume origin: permanent  
952 weak, but large scale, heating below the 660 km discontinuity, could lead to localized thermal  
953 instabilities in the transition zone that appear as blobs of lower velocity (see supplemental  
954 material). One or more of those blobs could have turned into the Yellowstone plume. The  
955 recent western U.S tomographic image of Burdick et al. (2008) reveals a -1% low velocity body

956 from 660 to ~800 km beneath Yellowstone which is wider than our array could detect. This  
957 suggests that a thermal instability from this deeper body could leak melt blobs into the transition  
958 zone.

959 One of those buds could have turned into the Yellowstone plume. The other buds may  
960 not develop further since the heat was effectively transported away via the plume conduit.  
961 However it is not clear how leaky thermal instability hypothesis fits in with the history of the  
962 Yellowstone plume. For example, the buds may be of younger origin than the overall plume as  
963 melt derivatives that leak off the main plume stem.

964

#### 965 **4.4 The Source of Plume-Plate Magma Plumbing**

966 An important Yellowstone hotspot problem is understanding the connection between the  
967 Yellowstone crustal magma body and the mantle plume source. This results from the lack of  
968 knowledge of the seismic structure of the lower crust and upper-most mantle and hence the  
969 mantle-crustal plumbing system from current technology. Crustal seismic refraction and  
970 regional earthquake studies and the local earthquake tomography do not have sufficient spatial  
971 resolution due to: 1) the lack of lower-crustal penetrating horizontal-propagating head waves  
972 and the lack of sufficient regional earthquakes, recorded from 200 km to 1000 km distances,  
973 required in these studies, 2) the lack of resolution in this depth range for local-earthquake  
974 tomography because of the lack of local earthquake sources deeper than the mid-crust, ~ 16 km  
975 maximum depth, 3) the lack of resolution for teleseismic tomography because of insufficiently  
976 wide angle crossing rays, and 4) the lack of resolution of surface waves due to their inherent  
977 longer periods and hence longer wavelength resolving kernels.

978           However some observations on this problem come from a study using controlled source  
979 refraction studies (Lehman et al., 1982). Lehman et al. (1982) recognized that the lower crust of  
980 Yellowstone was similar to that of the volcanically unmodified lower crust of surrounding  
981 regions. They evaluated the wide-angle reflector from the Moho boundary and found that the  
982 travel-times for equivalent ray paths, inside and outside the Yellowstone caldera, are essentially  
983 the same. This suggested that: (1) the lower crust was homogenous between the active area of  
984 Yellowstone volcanism and the surrounding Archean Rocky Mountain crust, and (2) while  
985 recognizing that magmas must have propagated through the Yellowstone lower crust, it did not  
986 alter the velocity structure and thus the lower crust did not contain large bodies of remnant melt  
987 such as the well-resolved in the middle crust (Husen et al., 2004).

988           A mid crustal sill at the base of the SRP upper crust was identified by seismic refraction  
989 and receiver function studies (Braile et al., 1982; Sparlin et al., 1982, Peng and Humphreys,  
990 1998) and interpreted to be a mafic remnant of the fractionation process that differentiates the  
991 Yellowstone magma system into the bimodal basaltic-rhyolitic volcanic rocks of the YSRP. It  
992 was modeled as composed of a series of gabbroic lenses inter-fingering with the granitic upper  
993 crust (see details of the SRP mid crustal sill gravity and seismic modeling in this volume by  
994 DeNosaquo et al., 2008). This geometry and modeling of the velocity-density relationships  
995 yields a bulk composition comparable to diorite and a density of  $2900 \text{ kg/m}^3$ . The high-density  
996 mid crustal sill varies from 4 to 11 km in thickness, resulting in a series of SE-NW trending  
997 gravity highs observed along the axis of the SRP. The sill extends up to 20 km southeast of the  
998 volcanic field, causing the asymmetry of the gravity field southeast of the SRP. This suggests  
999 that basaltic mantle magma ascends buoyantly to mid-crustal depths, where it attains neutral  
1000 buoyancy spreads out as a province-wide sill. Hildreth et al. (1991) suggested that isotopically

1001 that Yellowstone rhyolites are derived from the melting of young underplated mafic crust rather  
1002 than felsic upper crust magmas, possibly from fractionated and re-melted basalts. This  
1003 petrologic model is consistent with the seismic and gravity models.

1004         Seismic refraction studies of the YSRP (Schilly et al., 1982; Lehmann et al., 1982;  
1005 Braile et al., 1982; Schutt et al., 2008) provide information on the lower crustal structure of the  
1006 YSRP and reveal a relatively high-velocity layer beneath Yellowstone but lower velocity beneath  
1007 the Snake River Plain. On the basis of these generalized data, crustal magma intrusion is  
1008 characterized by magmas rising buoyantly from the plume source at ~70 km depth through a  
1009 vertically oriented plexus of basaltic dikes in the upper 30 km of the mantle. These dikes  
1010 possibly pond at the Moho to produce underplating, but then ascending through the lower crust  
1011 via another plexus of dikes. As the mafic magma passes through the lower crust, it promotes  
1012 remelting of the surrounding silicic, granulitic host rock, which also ascends. The magma then  
1013 differentiates into a mid crustal magma body composed of basalt and rhyolitic melts. Whether  
1014 the lower crust flows laterally to accommodate the mantle magmas or whether there is a balance  
1015 between the eruptive volumes and intruded volumes is problematical given the lack of  
1016 quantitative relationships between the parental and erupted magma volumes, especially those of  
1017 rhyolitic-basaltic mixes. The seismic data of Husen et al. (2004) suggest that this body is as  
1018 shallow as 6 km beneath the resurgent domes and extending to at least 16 km, the maximum  
1019 depth of the resolving paths for local earthquake tomography. Between the crust and mantle  
1020 plume, 40 to 70 km, the seismic velocity data are ambiguous for the reasons stated above.

1021

## 1022 **5. Geodynamic Plume Modeling**

1023 Our tomographic images provide a background model for dynamic modeling of the  
1024 Yellowstone plume. The geodynamic plume models were analyzed for seismic velocity  
1025 anomalies including the effects of temperature, anisotropy and composition, and also the  
1026 presence of water or melt (Fig. 20). According to Karato (1993), temperature is the main source  
1027 of seismic velocity anomalies if no heterogeneities of the chemical composition are present.

1028 While anisotropy can have a significant effect on seismic travel times, it is neglected in  
1029 this study since olivine fast axis orientation is approximately unidirectional (Waite et al., 2005).  
1030 In this case, all compressional P-wave travel times from any one event are equally affected and  
1031 the homogeneous imprint of anisotropy is removed by calculating relative travel time residuals.  
1032 Therefore we concentrate on temperature and composition in terms of the presence of water and  
1033 melt.

1034 The presence of a thermal anomaly in the 410–660 km depth range, as allowed in one of  
1035 the models by Humphreys et al. (2000), would be expected to be accompanied by thinning of the  
1036 transition zone through depression of the 410 km discontinuity and elevation of the 660 km  
1037 discontinuity (Bina and Helffrich, 1994). The depths to these discontinuities have been studied  
1038 using receiver functions for earthquakes recorded on a 500-km-long profile traversing the eastern  
1039 Snake River Plain and flanking swell at a distance of 200 km southwest of Yellowstone (Dueker  
1040 and Sheehan, 1997). The results provide no evidence for a coherent thermal anomaly extending  
1041 throughout the transition zone under the plain, and the results are confirmed by migration  
1042 reanalysis of the data by using different techniques (Beucler et al., 1999).

1043 Beneath the profile, transition-zone thickness was found to vary by 30–35 km, but the  
1044 topographies of the two discontinuities are uncorrelated. The most significant feature is a  
1045 deepening of the 410 km discontinuity by 20 km from the northwest margin of the plain to the

1046 eastern boundary of the Basin-Range. It is interpreted as a result of elevated mantle temperature  
1047 with excess partial melt in the shallow upper mantle (e.g., Bina and Hellfrich 1994). The change  
1048 in discontinuity depth implies uncorrelated lateral temperature variations of up to 250 K across  
1049 the Snake River Plain and flanking swell, with maximum temperatures at 400 km depth at a  
1050 location 150 km southeast of the plain.

1051 Possible non-thermal explanations for the observed topography of the mantle  
1052 discontinuities include the combined effects of garnet pyroxene phase transformations, chemical  
1053 layering, and variations in mantle hydration (Dueker and Sheehan, 1997). The fact that the 660  
1054 km discontinuity seems to be unaffected by the mantle plume could be also be explained by the  
1055 discontinuity being equally effected by a widespread thermal body underlying the 660 km  
1056 discontinuity in the area of our tomographic model.

1057 To assess the dynamic properties of the imaged plume we employed the thermodynamic  
1058 methodology of Cammarano et al. (2003) and Cammarano and Romanowicz (2007) with the  
1059 constraints of our velocity and attenuation models. We realize that our geodynamic plume  
1060 model is for a one-dimensional case, whereas we have parameterized data from a 3D structure,  
1061 but this model should not affect the overall dynamic processes.

1062 Following Cammarano et al. (2003), we derive models for attenuation and temperatures  
1063 in a two-step procedure for dry regimes first. We first derive an anelasticity model in terms of  
1064 depth-dependent  $Q_s$ -values from the melting curves and temperatures along the geotherm. These  
1065  $Q_s$ -values then are transformed into  $Q_p$ -values (Anderson and Given, 1982) based on the  
1066 comparison of compressional and shear wave velocity contrasts by Waite et al. (2006). In the  
1067 next step we calculate the partial derivatives, following the work of Karato (1993) on olivine  
1068 polycrystals at upper mantle temperatures and pressures. These are used to estimate the changes

1069 in velocity with temperature for a given attenuation or quality factor, respectively, so that the  
1070 velocity changes can be forward modeled and excess temperatures derived as in Fig. 20.

1071 For the wet regime, we followed Karato and Jung's (1998) explanation of the effect of  
1072 water on the seismic velocities by enhancing anelastic relaxation and by lowering the melting  
1073 point. Transition zone minerals can dissolve 2-3% water (Karato and Jung, 1998). Hence we  
1074 examine the influence on the excess temperatures and attenuation properties that can produce the  
1075 observed reduction in P-wave speed. Considering the imaged low velocity zone as a plume  
1076 consisting of an upwelling of hotter mantle originating in the transition zone, the plume will  
1077 carry water up into the upper mantle. The effects are decreased viscosity, lowered melting point  
1078 and, when the solidus is reached, partial melting. This process will remove water from the solid  
1079 minerals and therefore increase properties like shear modulus and seismic velocities (Karato and  
1080 Jung, 1998). Dehydration will reduce the negative effect of increased temperature on the seismic  
1081 velocities.

1082 Since S-waves are more affected than P-waves, the dehydration may even compensate for  
1083 the temperature effect on S-wave velocities. Waite et al. (2006) observe such a "hole" in their  
1084 low velocity zone at about 200 km depth in their S-wave but not their P-wave model. This is  
1085 consistent with findings from Kawamoto and Holloway (1997). Partial melting and dehydration  
1086 may be responsible for the rapid increase in size of the Yellowstone plume above 200-250 km  
1087 depth. We estimate the attenuation in terms of Q-values for this case following Karato and Jung  
1088 (1998) who find  $Q_{\text{wet}}=2.5 \cdot Q_{\text{dry}}$ , based on the enhanced creep rate and frequency dependence.  
1089 This is consistent with empirical results from Jackson et al., (1992) for dunite composition.

1090 Also we calculated the corresponding partial derivatives and estimate excess  
1091 temperatures. Ignoring the effect of the partial melting process, we find excess temperatures of

1092 145°-168°K in the uppermost mantle, 60-72 K in the lower upper mantle, and 78-85 K in the  
1093 transition zone (see Fig. 20). However, 1% partial melt can lower the compressional wave speed  
1094 between 1.8% (Faul et al., 1994) and 3.6% (Hammond and Humphreys, 2000) per 1% partial  
1095 melt. Consequently, the amount of partial melt in the uppermost mantle has to be far less than  
1096 1%. Assuming 0.5% melt and the relation of Hammond and Humphreys (2000), this will  
1097 account for 1.8% velocity reduction leaving -1.2% as a temperature effect.

1098         These observations and models agree well with models derived by Farnetani and Samuel  
1099 (2004) for so-called “spout” plumes, predicting widely spread ponding beneath the transition  
1100 zone and only a narrow tail with 120-180 km diameter and 100°-150°K excess temperatures.  
1101 Moreover this model also matches the global tomography model by Montelli et al. (2004) and  
1102 could explain our observations of blobs in the transition zone. However, our estimated excess  
1103 temperatures (dry and wet) are considerably lower than the 200 K determined by Fee and Dueker  
1104 (2004) from deflections of the 410 km discontinuity.

1105

## 1106 **6. Deflection of the Yellowstone Plume in Large-Scale Mantle Flow**

1107         Guided by the tomographic images of the tilted upper mantle body and geodynamic  
1108 models of mantle properties, we can evaluate the effect of mantle flow on the orientation of the  
1109 hypothesized Yellowstone plume.

1110         Based on theoretical, experimental, and numerical results (e.g., Whitehead and Luther,  
1111 1975; Olson and Singer, 1985; Griffiths and Campbell, 1990) the following standard view of a  
1112 mantle plume has evolved: initially, a plume head rises from a source region in a thermal  
1113 boundary layer (often assumed above the core-mantle boundary, but in the case of Yellowstone  
1114 perhaps more appropriately at the boundary between upper and lower mantle at a depth of 660



1115 km). It remains connected with the source region through a conduit and molten material  
1116 continues to flow through the conduit to the base of the lithosphere. Volcanism may occur at the  
1117 surface above the plume-plate interaction, and when volcanic products are carried away with the  
1118 lithosphere moving over it, a hotspot track is created.

1119         If there is large-scale flow in the mantle this plume conduit will “blow in the wind” and  
1120 become tilted (Richards and Griffiths, 1988). The tilt will depend on the large-scale mantle flow  
1121 and buoyant rising speed of material within the conduit, and can thus be computed based on  
1122 models for both. Comparison of the computed conduit shape with observations obtained through  
1123 seismic tomography, and of the predicted hotspot track with the observed distribution of  
1124 volcanism in space and time, can thus provide important insights regarding mantle flow and  
1125 plume conduit rising speed. More generally, such analysis determines whether a mantle plume is  
1126 an appropriate explanation for a particular intraplate volcanic center. Locations of intraplate  
1127 volcanism, such as in Yellowstone, are often attributed to mantle plumes (Wilson, 1963;  
1128 Morgan, 1972), but other upper mantle-lithosphere interactions have been proposed as the source  
1129 of Yellowstone volcanism as described in section 2 (Smith, 1977; Christiansen and McKee,  
1130 1978; King and Anderson, 1995; Humphreys et al., 2000).

1131         Here we contrast the predicted plume conduit shapes for various modeling assumptions  
1132 with a tomography model of the Yellowstone plume in the upper mantle. We also compare the  
1133 predicted hotspot track with geometry and age progression of volcanism along the Snake River  
1134 Plain, the presumed track of the Yellowstone hotspot (e.g., Pierce and Morgan, 1992; Smith and  
1135 Braile, 1993). The general features of our plume model have been explained by Steinberger and  
1136 O'Connell (1998). Regarding specific parameters and characteristics, we will mostly follow the  
1137 work of Steinberger and Antretter (2006), which has been extended to 44 hotspots (including

1138 Yellowstone) by Boschi et al. (2007). While the full model description is given in these papers,  
1139 we are here mostly interested in the plume conduit in the upper mantle, and thus give a  
1140 simplified description.

1141 If we disregard time dependence, lateral variations, and the vertical components of large-  
1142 scale flow, we expect that the conduit becomes tilted if the horizontal mantle flow velocity at  
1143 depth  $z$ ,  $v(z)$ , differs from flow  $v(z_0)$  at source depth  $z_0$ . More specifically, if conduit rising speed  
1144 at depth  $z$  is  $v_r(z)$ , the conduit takes a time  $dt = dz/v_r(z)$  to rise through a layer of thickness  $dz$ .  
1145 During this time, it will get displaced relative to the source by an amount  $dx = (v(z) - v(z_0))dt =$   
1146  $(v(z) - v(z_0))/v_r(z)dz$ . Integrating from depth  $z_0$  to depth  $z$  thus yields a total displacement

1147

$$x(z) = \int_{z_0}^z (v(z) - v(z_0))/v_r(z) dz \quad (1)$$

1148

1149

1150 For a source depth at the upper-lower mantle boundary, this implies that conduit tilt in the  
1151 upper mantle should be in the direction of upper mantle flow, relative to flow at source depth,  
1152 and that the tangent of conduit tilt should be the same order of magnitude as the ratio of  
1153 horizontal upper mantle flow, relative to flow at source depth, to buoyant rising speed. Under  
1154 these simplifying assumptions, the shape of the conduit remains constant, but the conduit moves  
1155 with the flow at source depth. The predicted azimuth and age progression along the hotspot  
1156 track thus depends on the difference vector between plate motion and flow at source depth.

1157 However, if we initiate the computation with a vertical conduit, the effect of the conduit  
1158 being progressively tilted by mantle flow introduces a further component of hotspot motion in  
1159 the direction of upper mantle flow until steady state is reached. If a deeper source depth such the  
1160 core-mantle boundary is assumed, then steady-state is not reached and the conduit will

1161 experience tilting in both the lower and upper mantle, with different directions and magnitudes  
1162 of tilt at depth depending on differences in mantle flow, thus contributing an additional  
1163 component to hotspot motion.

1164         There are various parameters influencing flow in the mantle, but the largest uncertainties  
1165 arise from variations in the mantle density models derived from seismic tomography, subduction  
1166 history, and viscosity structure. We will use different models to obtain a realistic range of flow  
1167 and plume conduit shape predictions, and consider several models of plate motion for hotspot  
1168 track predictions.

1169         Computation of large-scale mantle flow is done with the method of Hager and O'Connell  
1170 (1979, 1981), employing prescribed plate motions as surface boundary conditions and internal  
1171 density heterogeneities, both expanded in spherical harmonics, to compute flow. Density  
1172 variations are converted from global tomography models following Steinberger and Calderwood  
1173 (2006) or inferred from subduction history (Steinberger, 2000) and are named Smean (Becker  
1174 and Boschi, 2002), SAW24B16 (Megnin and Romanowicz, 2000) and TX2007 (Simmons et al.,  
1175 2006)

1176         An example of a flow model is shown in Fig. 21. Computed upper mantle flow in the  
1177 vicinity of Yellowstone is eastward. This eastward flow is part of a large-scale convection cell,  
1178 from an upwelling underneath the Pacific towards downward flow due to the subducted Farallon  
1179 slab beneath central and eastern North America. The viscosity model primarily used in our  
1180 models (VM1) was derived by Steinberger and Calderwood (2006) and was based on fitting the  
1181 geoid and other observational constraints, and consistent with mineral physics. VM2 is a simpler  
1182 model (Becker et al., 2006) also used in our mantle flow computations. Specifics and parameters  
1183 of the flow model follow Steinberger and Antretter (2006). This eastward flow in the upper

1184 mantle gives a first indication that we should expect an eastward-tilted Yellowstone plume  
1185 conduit (i.e., coming up from the west).

1186 Fig. 22 shows that this eastward flow component is strongest in the upper mantle and  
1187 transition zone but decreases with depth until the bottom of the transition zone at 660 km. The  
1188 flow direction and depth dependence are general features common to a large number of models.  
1189 However there are variations among the models, with flow direction ranging between  
1190 southeastward and northeastward, and variable flow speeds.

1191 For the preferred model of Steinberger and Antretter (2006), plume conduit diameter in  
1192 the upper mantle and transition zone is about 100 km, and buoyant rising speed increases from  
1193 about 2 cm/yr at a depth of 660 km to 10 cm/yr at 400 km and remains approximately constant  
1194 through the upper mantle. Total rise time from a depth of 660 km is about 12 Ma. Given typical  
1195 upper mantle horizontal flow speeds (relative to flow at 660 km) of a few cm/yr (see Fig.22), we  
1196 expect a conduit tilted by a few hundred kilometers, and that the surface plume position is  
1197 displaced relative to the position at a depth of 660 km approximately towards the east, between  
1198 southeast and northeast.

1199 The comparison of observations with actual computations more or less confirms this  
1200 expectation: Fig. 23a shows results for the case of plumes coming from 660 km depth with no  
1201 assumption about the initial conduit made, i.e., all conduit elements originate at depth 660 km  
1202 and the plume conduit is already tilted when the plume first reaches the surface. Tilts are in  
1203 directions similar to upper mantle flow, and amounts of tilt vary between less than 100 and ~250  
1204 km. Differences between predicted hotspot tracks and the corresponding fixed-hotspot track,  
1205 shown in Fig. 22, approximately correspond to the amount of plume source displacement due to  
1206 the horizontal flow component at a depth of 660 km. Differences between this case and the

1207 simplified model discussed in the introduction are due to time dependence and the vertical  
1208 component of the flow field

1209 Computed tilts are somewhat steeper (~150 to 400 km), but generally in the same direction in  
1210 the case shown in Fig. 23b where plumes rise from the lowermost mantle with an initially  
1211 vertical conduit (at 15 Ma). This larger tilt can be attributed to the cumulative effect of tilting in  
1212 the lower mantle added to tilting in the upper mantle. For the plume model based on tomography  
1213 model SAW24B16, both direction and amount of predicted tilt approximately agrees with the  
1214 tomographic observations reported here (Figs. 17 and 18 ). Computed hotspot tracks for the case  
1215 of a whole-mantle plume tend to be longer (i.e. with the predicted 15 Ma location further from  
1216 Yellowstone) than in Fig. 23a. This is due to the effect of the initially vertical conduit becoming  
1217 tilted in the upper and lower mantle.

1218 For models with an initially vertical conduit but plume initiation ages older than 15 Ma,  
1219 the predicted age progressions become more similar to that shown in Fig. 23a as the “blowing  
1220 over” effect of upper mantle flow causes the conduit shape to converge with the initially tilted  
1221 plume models. In the case of plumes rising from the lowermost mantle, the predicted conduit tilt  
1222 becomes stronger with greater age, as tilting in the lower mantle contributes to the total tilt.  
1223 Predicted tracks are very similar in Fig. 23a and Fig. 23b, because in both cases, the effect of the  
1224 initially vertical conduit becoming “blown over” by the upper mantle “wind” is important.

1225 We find the best agreement between predicted and observed azimuth and age progression  
1226 of the hotspot track can be achieved with eastward flow in the upper mantle. The best agreement  
1227 between predicted and observed conduit shape can be achieved with southeast flow, as with  
1228 models TX2007 and SAW24B16. Both are within the range of our flow models. Moreover the  
1229 amount of tomographically observed tilt can be better matched with our models by having the

1230 plume originate in the lower mantle. However, a larger tilt can also result from a slower buoyant  
1231 rising speed and smaller conduit diameter. Increasing temperature dependence of viscosity could  
1232 decrease the conduit rising speed and diameter. Such a stronger temperature dependence would  
1233 for example result if a linear stress-strain relationship (diffusion creep) was assumed in the upper  
1234 mantle rather than dislocation creep with a non-linear stress-strain relationship with stress  
1235 exponent  $n = 3.5$  (as done by Steinberger and Antretter, 2006, whose model we adopt here).

1236         To summarize the effect of the mantle flow on plume conduit tilt, we plot in Fig. 24 the  
1237 flow and conduit for a lowermost mantle plume source from an earlier model superimposed on a  
1238 background of S-wave velocity structure (Steinberger, 2000). For this model the plume origin  
1239 would be at longitude  $120^\circ\text{W}$ , beneath the Oregon Coast and displaced 1100 km west at  
1240 Yellowstone. The modeled plume location at  $\sim 660$  km depth is at  $\sim 115^\circ\text{W}$ , beneath the Oregon  
1241 High Lava Plain and the Columbia Plateau basalt field. If Yellowstone plume volcanism  
1242 initiated with the Columbia Plateau flood basalts, then this coincidence of plume position at the  
1243 base of the transition zone implies an initially vertical plume in the upper mantle.

1244

## 1245 **7. Yellowstone Geoid Anomaly**

1246         Just as the earth's topographic field responds to crustal loads, the long-wavelength gravity  
1247 field and topography generally reflects deeper mantle sources. To analyze this feature we  
1248 examine the Earth's geoid field. Most of the local geoid features are due to topographic  
1249 variation, but deep density variations form an important source of the Yellowstone anomaly.

1250

1251         The large-scale isostatic properties of the YSRP can be seen in the GEOID03 model for  
1252 the U.S. (Fig. 25). The model was constructed from a combination of gravity data and

1253 orthometric heights determined by geodetic surveys, with the resulting equipotential surface  
1254 reflecting an amalgam of topographic relief and density variation within the earth (see Milbert,  
1255 1991 for explanation).

1256         To model the Yellowstone geoid signal we assumed the geoid height was due to the static  
1257 uplift due to density variations. We recognized that dynamic processes could also contribute to  
1258 the signal but that it was negligible given the weak buoyancy flux that we determined for the  
1259 Yellowstone plume model. Using the observed geoid height for the Yellowstone anomaly, we  
1260 calculated the B-value, or relation between seismic velocity perturbations and mantle density  
1261 variations ( $\otimes V_p/\otimes \rho$ ) (e.g., Birch, 1961, Lees and VanDecar, 1991) at different depths of the  
1262 plume conduit by the forward modelling method developed by Emile Klingele (2006, personal  
1263 communication) based on Tsuboi, 1954. The resulting B-values are then used to interpret the  
1264 composition, temperature, pressure, and melt within the plume. For example Schmitz et al.  
1265 (1997) found a correlation between B-values and degree of partial melt. The overall objective of  
1266 our modelling was to evaluate the contribution of the Yellowstone plume to the geoid anomaly  
1267 relative to other local and regional tectonic and topographic sources and determine its melt  
1268 percentage.

1269         We parameterized the geometry and density of the Yellowstone geoid model by  
1270 converting the velocity perturbation structure of Yellowstone plume, described above, to density  
1271 variations as a starting model. While the separate derivation of density structure from either  
1272 velocity perturbations or geoid data is highly non-unique, the combination of both leads to a  
1273 more constrained solution.

1274         Forward modelling of the density variation was done on a two-dimensional profile, A-A'  
1275 crossing the Yellowstone hotspot from NW to SE (Fig. 25a). The density model was initially

1276 divided into nine bodies extending from 14 km to 660 km in depth, but only the top four  
1277 segments were found to contribute significantly to the model (Fig. 25b). Each body was given a  
1278 velocity perturbation based on the results of the reconstruction tests from the tomographic  
1279 inversion. Since the true background velocities of the tomographic model are not known, the  
1280 absolute velocity perturbations are based on the whole earth velocity IASP91 P-wave velocity  
1281 model. 80 000 forward models were run with varying B-value combinations.

1282         Since both the reference values for the geoid data and the velocity and density models are  
1283 not well resolved, the observed and the forward calculated geoid data have to be compared on a  
1284 relative basis. To do so, we shift the forward modelled geoid profile so that its maximum  
1285 amplitude matches the peak of the observed geoid profile (Fig. 25c). This does not affect the  
1286 lateral position or the shape of the geoid.

1287         The geoid data were obtained from the U.S. National Oceanic and Atmospheric  
1288 Administration website ([www.ngs.noaa.gov/GEOID](http://www.ngs.noaa.gov/GEOID)) along the profile at a 5 km spacing. The  
1289 data are plotted in Fig. 25a. The profile A-A' coordinates were adapted to the profile from the  
1290 seismic tomography model. The origin of the Yellowstone tomography coordinate system is  
1291 located at kilometer 700 along the profile. Since the geoid signal shows high frequency  
1292 components that are not associated with a deep plume model and are likely are of local tectonic  
1293 or topographic origin, we applied a one-dimensional filter to the geoid data using an optimum  
1294 window length of 300 km to remove short wavelength crustal contributions. Since regional  
1295 effects are obvious in the data but are not contained in the model, we only calculated the misfit  
1296 between 250 km and 1000 km along the profile (Fig. 25c). Only the plume effect is calculated  
1297 and the surrounding velocity and density variations are not taken into account.



1298 Density variations were forward calculated from the velocity perturbations assuming  
1299 initial B-values between 2 and 6 for each body. The B-value, can be an indicator of the presence  
1300 of partial melt. The effect on the geoid of the resulting density structure was calculated  
1301 separately for each body based on the algorithms of Tsuboi (1954). The results were then  
1302 superimposed to give the full geoid signal. The B-values of bodies 5 to 9 were fixed to a median  
1303 B-value of 3.0, since the corresponding variation between the minimum and maximum effect on  
1304 the geoid surface expression is about 0.3 m for body 5 and even less for the deeper bodies.

1305 To assess the modelling result, we calculated the RMS misfit between the observed and  
1306 the modelled geoid data for each of the forward models (Fig. 25c). The models were sorted so  
1307 that the minimum misfit occurs in model number 1 and the maximum misfit occurs in model 80  
1308 000. The resulting model of density variations for bodies 1-4 is given in Table 1, together with  
1309 the corresponding B-values for selected optimal solutions. The comparison between the B-  
1310 values for bodies 1 and 2 shows that their mean value is approximately constant at 3.5. This  
1311 implies that the modelling cannot distinguish between the geoid signals caused by body 1 and  
1312 body 2. The optimum B-values for bodies 3 and 4 are small, implying a larger density variation  
1313  $\Delta\rho$  per velocity contrast  $\Delta V_p$  than for bodies 1 and 2. We interpret that the B-values to thus  
1314 reflect notable density decreases of 1.3% to 3.6% in the upper part of the plume relative to the  
1315 commonly assumed density of 3400 kg/m<sup>3</sup> of the upper mantle. The largest density anomaly is  
1316 in the upper mantle at depths of 150 km to 170 km, the same depth range as the velocity  
1317 anomaly.

1318 As the temperature reaches the melting point, the seismic velocities decrease rapidly,  
1319 while the densities only decrease slowly, resulting in increased B-values (Schmitz et al., 1997).  
1320 We interpret the models to indicate that partial melt is present in bodies 1 and 2, but this

1321 interpretation is less justified for bodies 3 and 4. Since the absolute velocities and densities are  
1322 not known we do not determine the exact percentage of partial melt here. However, the  
1323 modelling results of relatively high B-values in bodies 1 and 2 (14 km to 110 km depth) and  
1324 consistently lower B-values in bodies 3 and 4 (110 km to 285 km depth) indicate significant  
1325 differences between those two regions within the plume. This correlates well with our  
1326 interpretation of a plume in a wet state, where partial melt is present in the uppermost part of the  
1327 plume. Below that, the plume was dehydrated by the melting process, reducing the negative  
1328 effect on the amplitudes of seismic velocity perturbations but only slightly changing the density  
1329 variations. This results in smaller B-values suggesting possible dehydration already below 110  
1330 km depth.

1331

## 1332 **8. Yellowstone-Snake River Plain Kinematics and Dynamics**

1333 To evaluate the large-scale effects of the Yellowstone hotspot on the western U.S., we  
1334 have determined the contemporary velocity field from over 2100 GPS measurements and 245  
1335 fault-slip rates. These data were input into the dynamic modeling codes of Haines and Holt  
1336 (1993) and Haines et al. (1998). The GPS velocities were compiled from 22 studies across the  
1337 western U. S., while fault-slip rates greater than 0.2 mm/yr were obtained from the USGS  
1338 Quaternary Fault and Fold Database and other sources (Haller et al., 2002; Chang and Smith,  
1339 2002; McCalpin and Nishenko, 1996). The data were then interpolated to a grid using a bi-cubic  
1340 spline interpolation (Fig. 26a).

1341 The resulting model reveals a generalized clockwise rotation in the direction of crustal  
1342 motions (Fig. 26a). Southwest extension across the Yellowstone caldera at up to  $4.3 \pm 0.2$   
1343 mm/yr drives southwest motion of the Snake River Plain at  $2.1 \pm 0.2$  mm/yr (Puskas et al.,

1344 2007a). To the south, the direction of extension rotates from southwest to west in the eastern  
1345 Basin-Range. The western Basin-Range marks a transition to shear deformation that is driven by  
1346 relative shear between the North America and Pacific plates (Thatcher, 2003; Meade and Hager,  
1347 2005; Puskas et al., 2007b). In the western Basin-Range and Pacific Northwest, GPS ground  
1348 deformation rates decrease from  $14.6 \pm 0.1$  mm/yr in the Sierra Nevada to  $2.9 \pm 0.2$  mm/yr in  
1349 northeast Oregon, and the direction of deformation rotates from northwest to east (Puskas et al.,  
1350 2008). Velocities decrease in northern Idaho and western Montana to 1 mm/yr or less, and  
1351 deformation in this transition zone becomes difficult to resolve with available GPS data. The  
1352 observed rotation of velocities requires shearing and/or rotation of the Idaho Batholith to  
1353 accommodate the pattern of rotation.

1354         Horizontal deviatoric stresses were calculated from the potential energy of the western  
1355 U.S. lithosphere, which in turn was dependent on elevation and density structure (Flesch et al.,  
1356 2000). In order to improve the detail of the model and resolve the effects of the Yellowstone  
1357 hotspot, the standard global crustal density-thickness model (Bassin et al., 2000) was rescaled for  
1358 the western U.S. using Moho depths from receiver functions from EarthScope (Crotwell and  
1359 Owens, 2005). The Yellowstone crustal density structure, based on gravity-density modeling  
1360 (Denasquo et al., 2008, this volume), was added to the model to account for volcanic reworking  
1361 of the crust in the YSRP. With the addition of topographic and isostatic corrections, these steps  
1362 produced a lithospheric density structure model that accurately reflected the tectonic provinces of  
1363 the western U.S. (Puskas et al., 2007b).

1364         The deviatoric stress modeling largely corroborates the observed velocity field (Fig. 26).  
1365 Internal stresses arising from lateral mass variations in the lithosphere show a notable rotation of  
1366 tension directions centered on the Yellowstone hotspot. Northeast-southwest tension at the

1367 Yellowstone Plateau rotates to east-west tension in the eastern Basin-Range. The Basin-Range  
1368 experiences primarily east-west uniaxial tensional stress with shear stress to the south and west.  
1369 Stress orientations in the Basin-Range are strongly affected by the boundary conditions of the  
1370 stress model, which incorporates the kinematic data to constrain the stress tensor orientations at  
1371 the model boundaries and hence includes relative North America-Pacific plate motions (Puskas  
1372 et al., 2007b). The Pacific Northwest is a region of compression and shear (Zoback and Zoback,  
1373 1991; McCaffrey et al., 2000; McCaffrey et al., 2007) associated with the subduction of the Juan  
1374 de Fuca plate. Our stress model predicts compression and shear in Oregon and Washington but  
1375 finds tensional stresses in northern Idaho and the Idaho Batholith. The high stresses at the Idaho  
1376 Batholith do not correspond to high deformation rates. Both GPS measurements and kinematic  
1377 models support low deformation rates of less than 1 mm/yr. The discrepancy can be accounted  
1378 for by postulating a strong lithosphere, so that high stresses will result in very little strain (Puskas  
1379 et al., 2007b).

1380

## 1381 **9. Effects of Mantle on the Overriding Lithosphere**

1382 On a global scale, we compare our hypothesized Yellowstone plume with other hotspots  
1383 by computing the buoyancy flux using properties derived from tomographic models after the  
1384 method of Ritter (2004) (Fig. 27). The buoyancy flux is estimated from the width and elevation  
1385 of the hotspot topographic or bathymetric anomalies, velocity of the overriding plate, and excess  
1386 plume temperature (e.g., Davies, 1988; Sleep, 1990). Given the small ( $< 150$  K) excess  
1387 temperatures predicted for a dry mantle with low  $Q$  (estimated jointly from the  $V_P$  and  $V_S$   
1388 models), the Yellowstone buoyancy flux is at least one order of magnitude lower than previous  
1389 estimates (Sleep, 1990; Smith and Braile, 1994; Lowry et al., 2000; Nolet et al., 2007; Schutt et

1390 al., 2008; Stachnik et al., 2008). Likewise the Yellowstone buoyancy flux is estimated to be an  
1391 order of magnitude lower than its oceanic counterparts below Iceland and Hawaii, which have  
1392 fluxes of 1.4 Mg/s and 8.7 Mg/s, respectively (Sleep, 1990). At 0.25 Mg/s, Yellowstone is  
1393 comparable to the other continental hotspots with weak flux, calculated by Ritter (2004) to be  
1394 0.09 Mg/s at Eifel and 0.7 Mg/s at the Massif Central. A main consequence of such a weak flux  
1395 is that the low volume of ascending magma and reduced excess temperature together produce  
1396 less melting. The results are lower plume buoyancy and reduced impact on lithosphere uplift and  
1397 magmatic volume (Waite, 2004).

1398         If the anomaly is assumed to be a plume, there are two peculiarities that need to be  
1399 addressed. One is the much larger volume and amplitude of the anomaly in the upper 200 km  
1400 versus the lower 200 km of the upper mantle. The second is the northwest tilt of the anomaly  
1401 with depth. The tilt is difficult to reconcile with plate motion and mantle flow models.  
1402 Southeastward uppermost mantle flow can result in a plume dipping down to the northwest but  
1403 predicts hotspot motion to the southeast. However, when combined with the plate motion vector,  
1404 the base of the plume will be off to the NW of Yellowstone instead of along the Snake River  
1405 Plain (Figure 29).

1406         An alternate interpretation that cannot be ruled out is that the upper mantle velocity  
1407 anomaly may be caused by magma rising along in a weak or thinned lithosphere. If lithosphere  
1408 has been eroded along a preexisting structural weakness, then the upwelling could follow the  
1409 path of least resistance to the surface. This idea has been invoked by some researchers to explain  
1410 the dynamics of the Yellowstone system without a plume (e.g., Smith and Sbar, 1974; Favela  
1411 and Anderson, 2000; Dueker et al., 2001; Christiansen et al., 2002).

1412           The velocity anomaly in the upper 200 km is much larger than the anomaly in the range  
1413 of 200 to 400 km depth. This type of contrast between the upper and lower 200 km of the upper  
1414 mantle is not seen at Iceland, which is the best-imaged plume. Similarly, the Eifel plume has  
1415 strong low  $V_p$  anomaly to at least 400 km depth. It is possible that small-scale convection in the  
1416 uppermost mantle may draw material up from greater depths below the YSRP. In fact, it may be  
1417 necessary to draw material up, since the melt-depleted residuum is not expected to cycle back  
1418 through the convection cell. It spreads laterally instead, making room for more mantle material  
1419 to be drawn up. The extension of the Basin-Range enhances the effect and creates space for  
1420 mantle to ascend.

1421           Analog models of plumes by Whitehead (1982) showed that in a viscoelastic media  
1422 plumes rise vertically. However, if the ascending material were bent over by more than  $30^\circ$  from  
1423 the vertical, then the plume would break off from the original source, leaving a single, tilted  
1424 feature that Whitehead (1982) called a “plumelet”. Such a scenario is consistent with mantle  
1425 flow models and the geometry of the tomographically imaged Yellowstone plume. Steinberger  
1426 and O’Connell (2000) constructed global maps of hotspot tracks in laterally varying large-scale  
1427 mantle convection models and found that at transition zone depths beneath the western U.S.  
1428 interior, the ascending flow geometry would have exceeded  $30^\circ$  tilt from the vertical, cutting off  
1429 the heads of any pre-existing plumes. Thus Yellowstone may be a beheaded remnant of a  
1430 stronger plume that could have originated at the core-mantle boundary. Such a feature would  
1431 have originated in the lower mantle but was cut off by the high angle of tilt, leaving melt from a  
1432 more limited thermal source in the transition zone. The remaining material would have a low  
1433 buoyancy flux characteristic of a weaker plume.

1434           The small buoyancy flux of the modeled Yellowstone plume is not enough to totally  
1435 support the topography (e.g., Lowry et al., 2000). Compared with other hotspots, the eruption  
1436 rate at Yellowstone is much too large to be attributed to its small buoyancy flux of 0.25 Mg/s.  
1437 The same small-scale convection processes have been proposed for such features as  
1438 Yellowstone, the St. George, UT, and Jemez, NM, volcanic trends, which are all parallel and  
1439 trend to the northeast (Hernlund et al., 2008).

1440           However, the buoyancy flux of Yellowstone is much larger than St. George and Jemez, as  
1441 evidenced by the topographic features, eruptive volumes, and geoid anomaly. This suggests  
1442 something different about Yellowstone, which can be satisfied by the plume component in the  
1443 plume-convection hybrid model. The high  $^3\text{He}/^4\text{He}$  ratio, deep seismic anomaly, and thin  
1444 transition zone are also satisfied by the plume component. The persistence of basaltic  
1445 magmatism along the SRP for hundreds of kilometers from Yellowstone may be attributed to  
1446 continued convection millions of years after the plate has passed the plume. A weak lineament  
1447 in the lithosphere can help explain the apparent deflection of the plume, although we consider  
1448 mantle flow a more plausible cause of deflection.

1449           A model of plume-fed upper mantle convection generally agrees with our observations.  
1450 However, other models argue for buoyant decompression melting instabilities in an extending  
1451 lithosphere above regions of partially molten upper mantle (Lowry et al., 2000; Hernlund et al.,  
1452 2008). These models have been proposed to account for some characteristics of intraplate  
1453 volcanism in extensional lithospheric regimes including Yellowstone. Such models do not  
1454 require spatially and temporally correlated volcanism and extension and may account for  
1455 localized volcanic activity following Basin and Range extension in the western United States.  
1456 We suggest that our seismic images of a conduit of melt from ~660 km argues for a plume

1457 geometry, not a shallow planar zone of decompression melting.

1458         The well-known  $^{87}\text{Sr}/^{86}\text{Sr}=0.706$  boundary (Farmer et al., 1983) separates accreted  
1459 oceanic lithosphere to the west from continental cratonic lithosphere to the east. This boundary  
1460 is also marked by sharp decreases in the normalized isotope ratios  $\Sigma\text{Nd}$  and  $\Sigma\text{Hf}$  found in  
1461 Yellowstone silicic magmas, indicating a decrease in the mantle component of erupted materials  
1462 (Nash et al., 2006). That is, as hotspot volcanism progressed from accreted to cratonic terrain,  
1463 there was a fundamental change in magma composition, eruptive frequency, and temperature in  
1464 association with the change in overriding lithosphere (Perkins and Nash, 2002). The  
1465 configuration of subducting slab and thin oceanic lithosphere and thick continental lithosphere in  
1466 the Pacific Northwest has important implications for the evolutions of Yellowstone hotspot  
1467 volcanism.

1468         In Fig. 28 we illustrate the development of the YSRPN in terms of the plume-convection  
1469 model. The original voluminous plume head was vertical, rising from the deep mantle only to be  
1470 entrained in westward return flow driven by the eastward subduction of the Juan de Fuca plate.  
1471 The relatively weaker and thinner oceanic lithosphere allowed the plume head to spread out and  
1472 protected the plume from the eastward currents that dominated upper mantle convective return  
1473 flow below the thicker continental lithosphere to the east. As the North America plate  
1474 progressed southwest it encountered the much thicker continental lithosphere and lost the  
1475 protection of the back-arc geometry from large-scale mantle flow. Nash et al. (2006) showed  
1476 that, based on geochemical data, the transition from accreted to cratonic lithosphere and a shift  
1477 from westward to eastward mantle flow occurred at the Oregon-Idaho border. Here a plume with  
1478 a conduit diameter of 70 km became caught in the mantle return flow, tilting it and smearing out



1479 the magma against the overriding lithosphere. This process was responsible for the YSRP  
1480 hotspot track.

1481 A further postulation of this model would be the southward offset of volcanism over time  
1482 relative to the initial plume head position. The original Yellowstone-related silicic centers of  
1483 volcanism proposed by several authors (Pierce and Morgan, 1992; Smith and Braile, 1994;  
1484 Jordan et al., 2005) were associated with the McDermitt lava field of northern Nevada. Recent  
1485 studies by Nash et al. (2006) and Camp and Ross (2004) argue for a more widely distributed area  
1486 of initial silicic volcanism over southeastern Oregon. That is, later Yellowstone volcanism was  
1487 offset to the south of initial volcanism in eastern Oregon. Alternately, if we assume a linear  
1488 track for the base of the plume at 660 as well as the top of the plume, then the trace of the mantle  
1489 source follows a southwest trend beneath the northern Rockies and the Idaho Batholith, ending  
1490 below the western Snake River Plain, notably ~150 km north of the originally defined beginning  
1491 of the YSRP, at the McDermitt, NV, volcanic field (Figure 29).

1492 Our data and resulting model of the Yellowstone plume (Smith et al. 2003; Smith et al.,  
1493 2005) is consistent with the area hypothesized by Camp and Ross (2004) to be affected by the  
1494 plume head. This area encompasses much of the silicic volcanic area of the Oregon High Lava  
1495 Plains and the southern part of the Columbia Plateau basalt field. This suggests that Columbia  
1496 Plateau basalt outpouring that began at ~17 Ma may have had a common mantle source with the  
1497 YSRP, i.e., the Yellowstone plume. This concept is also corroborated by the geochemical  
1498 analysis of Hanan et al. (2008) who noted that the Steens basalt eruptive center may have been  
1499 an early eruptive phase of the Columbia River basalts and is also located near our modeled  
1500 location of the Yellowstone plume head at 17 Ma.

1501

1502 **10. Concluding Remarks**

1503           The Yellowstone hotspot is a profound tectono-magmatic feature of the western U.S. that  
1504 results from interaction of a mantle plume with the overriding North America plate. The hot,  
1505 low-density Yellowstone plume head has sufficient buoyancy to induce a large topographic swell  
1506 over the continental part of the hotspot track, now the Yellowstone swell. The high swell  
1507 elevation imparts a high potential energy that causes southwest downhill flow of the lithosphere  
1508 from Yellowstone driving compression along the eastern Snake River Plain and adding to the  
1509 westward extension of the Basin-Range. Kinematically, the plume magma is sheared to the  
1510 southwest against the southwest moving North American plate (Lowry et al., 2000) producing an  
1511 elongate plume head beneath the SRP and Yellowstone. Regionally, lithospheric extension  
1512 drives SW motion of the YSRP and is part of a larger kinematic pattern of clockwise rotation of  
1513 the western U.S. whose motion is partially driven by the potential energy of the topographically  
1514 high swell.

1515           The southwest motion of the YSRP is one element of the gyre of clockwise rotation of  
1516 deformation direction in the western U.S. The lithospheric extension associated with the Basin-  
1517 Range tectonics further reduces the horizontal confining stress, amplifying magma ascent from  
1518 plume to surface. Negative loading of the Snake River Plain by its mid crustal high density sill  
1519 also produces flexure in the crust extending ~ 30 km SE beyond the SRP boundary. Along with  
1520 lithospheric cooling, this contributes to systematic decrease in elevation with increasing age of  
1521 silicic volcanism in a trend that extends from Yellowstone back to the hotspot origin in eastern  
1522 Oregon.

1523           Our results confirm that the Yellowstone volcanic field is the locus of the highest level of  
1524 seismicity in the Rocky Mountains. Moreover it has experienced the largest historic earthquake,

1525 a1959 M7.5 event, of the Intermountain region. Local earthquake tomography images of  
1526 Yellowstone confirm a low- $V_p$  magma body beneath the caldera at 8-16 km, with 8-15% melts,  
1527 i.e. the Yellowstone magma chamber. Heat flow of 2000 mW/m<sup>2</sup> and Yellowstone's Quaternary  
1528 dominantly silicic volcanism from this crustal magma system drives Yellowstone's hydrothermal  
1529 features.

1530           Moreover, contemporary deformation of Yellowstone from geodetic measurements  
1531 reveals an energetic system dominated by lithospheric extension of up to 4 mm/yr with  
1532 superimposed volcanic uplift and subsidence with average rates of ~2 cm/yr. But the caldera has  
1533 experienced an unexpected episode of accelerated uplift from 2004 to 2007 at up to 7 cm/yr that  
1534 is attributed to magmatic recharge of the crustal magma system

1535           Teleseismic tomography employing  $V_p$  inversion imaged a P-wave low-velocity body  
1536 from 80 to 250 km directly beneath Yellowstone, but continuing at a tilt of 60° northwest to the  
1537 bottom of the transition zone at 660 km, 150 km west of Yellowstone. We interpret this body to  
1538 be the Yellowstone plume. Dynamics of the plume reveal excess temperature of 85°-120°K and  
1539 up to 1.5% melt with a relatively weak buoyancy flux of ~0.25 Mg/s that is several times smaller  
1540 than oceanic plumes.

1541           Employing the inclined plume-geometry and plate motion history, we extrapolate the  
1542 Yellowstone mantle-source southwestward to its initial position at 17 million years beneath  
1543 eastern Oregon and the southern part of the Columbia Plateau basalt field, which suggests a  
1544 common source for the entire YSRPN system (Fig. 29). Our model is consistent with the  
1545 original plume head rising vertically behind the subducting Juan de Fuca plate, but at ~12 million  
1546 years it encountered cooler continental lithosphere and horizontal mantle flow, imparting the  
1547 observed westward tilt.

1548           Although this paper synthesizes a great deal of research to produce an integrated analysis  
1549 of the Yellowstone plume and volcanic history of the Yellowstone hotspot, several issues still  
1550 need to be addressed. The dynamics and mechanics of magma generation in the upper mantle  
1551 and its subsequent transport from the tilted plume to the lithosphere and upper crust have not  
1552 been resolved. The mantle plume contributes to the Yellowstone topographic swell and geoid  
1553 anomaly, but the processes through which plume contributes regional extension at the surface  
1554 and depth to allow magma emplacement requires further study. Hydrothermal activity is driven  
1555 by heat from the crustal magma chamber, but the processes involved are poorly understood.  
1556 More work needs to be done to understand the spatial distributions of thermal basins in the  
1557 caldera, the changes in thermal activity that occur over time, and the thermal and mechanical  
1558 connections between surface features and crustal magma. Changes in the caldera magmatic  
1559 system through fluid transport (i.e., magma intrusion, dike injection, escape of volatiles, etc.)  
1560 contribute to local deformation, and the resulting stress changes are hypothesized to affect  
1561 nearby normal faults. Similarly, tectonic loading of Late Quaternary Yellowstone faults can  
1562 affect the volcanic system. Stress interactions between faults and the volcanic system at short  
1563 and long time scales remain poorly understood. As discussed in Section 4.4, the structure and  
1564 thermal state of the lower crust is not known. All these topics are at present unknown, but they  
1565 provide opportunities for investigation and quantitative modeling of the geochemical and  
1566 geophysical characteristics of the Yellowstone hotspot.

1567           The Yellowstone plume thus has had a profound effect on the western U.S. interior with  
1568 hotspot-driven Cenozoic volcanism affecting lithospheric structure, stress state, deformation, and  
1569 topography. Hotspot volcanism has produced the geology and environment that is used as the  
1570 basis for designating the world's first national park. The outstanding physical features of

1571 Yellowstone National Park include its world-renowned hot springs and geysers. These features  
1572 are thermal phenomena that are driven by Yellowstone’s extraordinarily high heat flow, which in  
1573 turn is caused by its active magmatic sources. For this reason we commonly say that “heat  
1574 drives it all” in Yellowstone. Moreover, we believe that our results demonstrate the dynamic  
1575 properties of the Yellowstone hotspot. Yellowstone caldera deformation and intense earthquake  
1576 activity denote “a living, breathing, shaking” caldera. In conclusion the contemporary volcanic  
1577 and tectonic processes of Yellowstone demonstrate that it is truly “A Window into the Earth’s  
1578 Interior”.

1579

#### 1580 **Acknowledgements**

1581 Collaborators in the Yellowstone Geodynamic Project included Eugene Humphreys, Paul  
1582 Tackley, Paul Hernlund, and Ken Dueker. Suzette Payne of the Idaho National Laboratory and  
1583 Ronald Harris of Brigham Young University and their staffs, assisted with temporary GPS  
1584 deployments. Our project materially benefited with discussions and assistance from David  
1585 Drobeck, Henry Heasler, Jake Lowenstern, David Lageson, Dan Dzurisin, Anthony Lowry,  
1586 Barbara Nash, Michael Perkins, John Shervais, Lisa Morgan, Kenneth Pierce, Suzette Payne,  
1587 David Rodgers, and Rick Hutchinson. We appreciate the support with data acquisition of our  
1588 Yellowstone projects from the National Science Foundation, the U.S. Geological Survey,  
1589 Yellowstone and Grand Teton National Parks, the Idaho National Laboratory (INL), the  
1590 Yellowstone Volcano Observatory and the Plate Boundary Observatory. The University of Utah  
1591 provided computational and technical assistance and the University of Utah Seismograph  
1592 Stations supplied engineering support for seismic and GPS monitoring. The IRIS-PASSCAL  
1593 project portable seismographs and UNAVCO provided GPS instruments. Our main Yellowstone

1594 hotspot project was primarily supported by the National Science Foundation Continental  
1595 Dynamics program grants #EAR 9725431, 0314298, 9725431, and 9316289. Funding was also  
1596 provided by William and Sue Carrico Foundation and by The Brinson Foundation.

1597 **References**

- 1598 Allen, R.M., Nolet, G., Morgan, W.J., Vogfjord, K., Bergsson, B.H., Erlendsson, P., Foulger,  
1599 G.R., Jakobsdottir, S., Julian, B.R., Pritchard, M., Ragnarsson, S., Stefansson, R. 2002.  
1600 Imaging the mantle beneath Iceland using integrated seismological techniques, *J.*  
1601 *Geophys. Res.* 107: doi:10.1029/2001JB000595.
- 1602
- 1603 Anderson, D.L., J. L. Given, 1982. Absorption band Q model for the earth. *J. Geophys. Res.* 87:  
1604 3893-3904.
- 1605
- 1606 Anders M. H., Sleep N. H., 1992. Magmatism and extension: The thermal and mechanical  
1607 effects of the Yellowstone hotspot. *J. Geophys. Res.* 97: 15,379–15,393.
- 1608
- 1609 Annen, C., Blundy, J. D., Sparks, R. S., 2006. The genesis of intermediate and silica magmas in  
1610 deep crustal hot zones. *J. Pet.* 47: 505-539.
- 1611
- 1612 Armstrong, R. L., Leeman, W.P., Malde, H.E., 1975. K-Ar dating, Quaternary and Neogene  
1613 volcanic rocks of the Snake River Plain, Idaho. *Am. J. Sci.* 275: 225-251.
- 1614
- 1615 Barth, A., Jordan, M., Ritter, J., 2007. Crustal and upper mantle structure of the French Massif  
1616 Central plume. In: J.R.R. Ritter, U.R. Christensen, (Editors), *Mantle Plumes - A*  
1617 *Multidisciplinary Approach*. Springer Verlag, Heidelberg, pp. 159-184.
- 1618
- 1619 Bassin, C., Laske, G., Masters, G., 2000. The current limits of resolution for surface wave  
1620 tomography in North America. *Eos Trans. AGU* 81(48): F897.
- 1621
- 1622 Becker, T. W., Boschi, L., 2002. A comparison of tomographic and geodynamic mantle models.  
1623 *Geochem. Geophys. Geosys.* 3(1): 1003, doi:10.1029/2001GC000168.
- 1624
- 1625 Becker, T.W., Chevrot, S., Schulte-Pelkum, V. Blackman, D.K., 2006. Statistical properties of  
1626 seismic anisotropy predicted by upper mantle geodynamic models. *J. Geophys. Res.* 111:  
1627 B08309, doi:10.1029/2005JB004095.
- 1628
- 1629 Bennett, R. A., Davis, J. L., Normandeau, J. E., Wernicke, B. P., 2001. Space geodetic  
1630 measurements of plate boundary deformation in the western U.S. Cordillera. In: S.  
1631 Stein, J. Freymueller (Editors), *Plate Boundary Zones*. AGU, Washington, D.C., pp. 27-  
1632 55.
- 1633
- 1634 Beucler, E., Chevrot, S., Montagner, J. P. 1999. The Snake River Plain experiment revisited.  
1635 Relationships between a Farallon plate fragment and the transition zone. *Geophys. Res.*  
1636 *Lett.* 26: 2673-2676.

1637  
1638 Bina C. R., Helffrich, G., 1994. Phase transition Clapeyron slopes and transition zone seismic  
1639 discontinuity topography. *J. Geophys. Res.* 99: 15,853-15,860.  
1640  
1641 Birch, F., 1961. The Velocity of compressional waves in rocks to 10 kilobars, Part 2. *J.*  
1642 *Geophys. Res.* 66(7): 2199-2224.  
1643  
1644 Blackwell, D. D., Negraru, P. T., Richards, M. C., 2006. Assessment of the enhanced  
1645 geothermal system resource base of the United States. *Nat. Resources Res.* 15(4): 283-  
1646 308, doi:10.1007/s11053-007-9028-7.  
1647  
1648 Boschi, L., Becker, T. W., Steinberger, B., 2007. Mantle plumes: Dynamic models and seismic  
1649 images. *Geochem. Geophys. Geosys.* 8: Q10006, doi:10.1029/2007GC001733.  
1650  
1651 Braile, L. W., Smith, R.B., Anson, J., Baker, M. R. , Sparlin, M. A., Prodehl, C. M. Schilly,  
1652 M.M., Healy, J. H. Mueller, S. Olsen, K. H. 1982. The Yellowstone Snake River Plain  
1653 seismic profiling experiment: Crustal structure of the eastern Snake River Plain. *J.*  
1654 *Geophys. Res.*, 84: 2597-2610.  
1655  
1656 Burke, K., 1996. The African Plate. *S. African J. Geol.* 99: 339-409.  
1657  
1658 Burdick, S., Li, C., Martynov, V., Cox, T., Eakins, J., Mulder, T., Astiz, L., Vernon, F.L., Pavlis,  
1659 G.L., van der Hilst, R., 2008. Upper mantle heterogeneity beneath North America from  
1660 travel time tomography with global and USArray Transportable array data. *Seismol. Res.*  
1661 *Lettr.* 79(3): 384-392.  
1662  
1663 Cammarano F., Goes S., Vacher P., Giardini D., 2003. Inferring upper mantle temperatures from  
1664 seismic velocities. *Phys. Earth Plan. Int.*, 138: 197-222.  
1665  
1666 Cammarano F., Romanowicz, B., 2007. Insights into the nature of the transition zone from  
1667 physically constrained inversion of long-period seismic data. *Proceedings Natl. Acad.*  
1668 *Sci. U.S.A.* 104(22): 9139-9144.  
1669  
1670 Camp, V. E., Ross, M. E., 2004. Mantle dynamics and genesis of mafic magmatism in the  
1671 intermontane Pacific Northwest. *J. Geophys. Res.* 109: B08204,  
1672 doi:10.1029/2003JB002838.  
1673  
1674 Carlson, R. W., Hart, W. K., 1988. Flood basalt volcanism in the northwestern United States.  
1675 In: J. D. Macdougall (Editor), *Continental Flood Basalts*. Kluwer Academic Publishers,  
1676 Dordrecht, The Netherlands, p. 35-61.  
1677  
1678 Chang, W. L., Smith, R. B., 2002. Integrated seismic-hazard analysis of the Wasatch Front,  
1679 Utah. *Bull. Seis. Soc. Am.* 92(5): 1904-1922.  
1680

1681 Chang, W. L., Smith, R. B., 2005, Lithospheric rheology from postseismic deformation of a  
1682 M=7.5 normal-faulting earthquake with implications for continental kinematics. 2005  
1683 Salt Lake City Annual Meeting, Geol. Soc. Amer., Abs. 225-4.  
1684

1685 Chang, W. L., Smith, R. B., 2008. Lithospheric rheology from postseismic deformation of a  
1686 M=7.5 normal-faulting earthquake with implications for continental kinematics. J.  
1687 Geophys. Res. in press.  
1688

1689 Chang, W., Smith, R.B., Wicks, C., Puskas, C., Farrell, J., 2007. Accelerated uplift and source  
1690 models of the Yellowstone caldera, 2004-2006, from GPS and InSAR observations,  
1691 Science 318(5852): 952-956, doi:10.1126/science.1146842.  
1692

1693 Christiansen, R. L., 2001. The Quaternary and Pliocene Yellowstone Plateau volcanic field of  
1694 Wyoming, Idaho, and Montana. U. S. Geol. Surv. Prof. Pap. 729-G, U. S. Geol. Surv.,  
1695 Denver, CO, 120 pp.  
1696

1697 Christiansen R. L., Foulger G. R., Evans J. R., 2002. Upper-mantle origin of the Yellowstone  
1698 hotspot. GSA Bull. 114: 1245–1256.  
1699

1700 Christiansen, R.L., McKee, E.H., 1978. Late Cenozoic volcanic and tectonic evolution of the  
1701 Great Basin and Columbia intermontane regions. In: R. B. Smith, and G. P. Eaton  
1702 (Editors), Cenozoic Tectonics and Regional Geophysics of the Western Cordillera:  
1703 Geological Society of America Memoir 152. GSA, Boulder, CO, pp. 283-311.  
1704

1705 Clawson, S. R., Smith, R. B. Benz, H.M., 1989. P-wave attenuation of the Yellowstone caldera  
1706 from three-dimensional inversion of spectral decay using explosion source seismic data.  
1707 J. Geophys. Res. 94: 7205-7222.  
1708

1709 Craig, H., Lupton, J. E., Welhan, J. A., Poreda, R., 1978. Helium isotope ratios in Yellowstone  
1710 and Lassen Park volcanic gases. Geophys. Res. Lett. 5(11): 897-900.  
1711

1712 Crampin S., Chastin S., 2003. A review of shear wave splitting in the crack-critical crust.  
1713 Geophys. J. Int. 155: 221-240.  
1714

1715 Crotwell, H. P., Owens, T. J., 2005. Automated receiver function processing. Seis. Res. Lett.  
1716 76(6): 702-713.  
1717

1718 Crough, S.T., 1978. Thermal origin of mid-plate hot-spot swells. Geophys. J. Royal Astron.  
1719 Soc. 55: 451-469.  
1720

1721 Davies, G. F., 1988. Dynamic Earth: Plates, Plumes and Mantle Convection. Cambridge  
1722 University Press, Cambridge, 458 pp.  
1723

1724 DeNosaquo, K., Smith, R.B., Lowry, A.R., 2008. Density and lithospheric strength models of  
1725 the Yellowstone-Snake River Plain volcanic system from gravity and heat flow data, J.  
1726 Vol. Geotherm. Res. (this volume).



1727  
1728 Dietz, R.S., Holden, J. C., 1970. Reconstruction of Pangaea: Breakup and dispersion of  
1729 continents, Permian to present. *J. Geophys. Res.* 75: 4939-4956  
1730  
1731 Doe, B. R., Leeman, W. P., Christiansen, R. L., Hedge, C. E., 1982. Lead and strontium isotopes  
1732 and related trace elements as genetic tracers in the upper Cenozoic rhyolite-basalt  
1733 association of the Yellowstone plateau volcanic field. *J. Geophys. Res.* 87(B6): 4785-  
1734 4806.  
1735  
1736 Doser, D. I., 1985. Source parameters and faulting processes of the 1959 Hebgen Lake,  
1737 Montana, earthquake sequence. *J. Geophys. Res.* 90: 4537-4555.  
1738  
1739 Dueker, K. G., Sheehan, A. F., 1997. Mantle discontinuity structure from mid-point stacks of  
1740 converted P to S waves across the Yellowstone Hotspot Track. *J. Geophys. Res.*  
1741 102(B4): 8313-8327.  
1742  
1743 Dueker, K., Yuan, H., Zurek, B., 2001. Thick-structured Proterozoic lithosphere of the Rocky  
1744 Mountain region. *GSA Today* 11: 4-9.  
1745  
1746 Dzurisin, D., Yamashita K. M., 1987. Vertical surface displacements at Yellowstone caldera,  
1747 Wyoming, 1976-1986. *J. Geophys. Res.* 92: 13,753-13,766.  
1748  
1749 Evans, J.R., Achauer, U., 1993. Teleseismic velocity tomography using the ACH-method.  
1750 Theory and application to continental scale studies. In: H.M. Iyer and K. Hirahara  
1751 (Editors), *Seismic Tomography Theory and Practice*. Chapman and Hall, London, pp.  
1752 319–360.  
1753  
1754 Farmer, G., DePaolo, D. 1983. Origin of Mesozoic and Tertiary Granite in the Western United  
1755 States and Implications for Pre-Mesozoic Crustal Structure 1. Nd and Sr Isotopic Studies  
1756 in the Geocline of the Northern Great Basin. *J. Geophys. Res.* 88(B4): 3379-3401.  
1757  
1758 Farnetani, C. G., Samuel, H., 2005. Beyond the thermal plume paradigm. *Geophys. Res. Lett.*  
1759 32: L07311, doi:10.1029/2005GL022360.  
1760  
1761 Farrell, J. M., 2007. Space-time seismicity and development of a geographical information  
1762 system database with interactive graphics for the Yellowstone region. Masters Thesis,  
1763 University of Utah, Salt Lake City, Utah.  
1764  
1765 Favela, J., Anderson, D. L., 2000. Extensional tectonics and global volcanism. In: E. Boschi,  
1766 G. Ekstrom, A. Morelli, (Editors), *Problems in geophysics for the new millennium*.  
1767 Editrice Compositori, Bologna, pp. 463–498.  
1768  
1769 Fee, D., Dueker K., 2004. Mantle transition zone topography and structure beneath the  
1770 Yellowstone hotspot. *Geophys. Res. Lett.* 31: L18603, doi:10.1029/2004GL020636.  
1771

- 1772 Faul, U. H., Toomey, D. R., Waft, H. S., 1994. Late granular basaltic melt is distributed in thin,  
1773 elongated inclusions. *Geophys. Res. Lett.* 21: 29-32.  
1774
- 1775 Flesch, L. M., Holt, W. E., Haines, A. J., Shen-Tu, B., 2000. Dynamics of the □Pacific-North  
1776 American plate boundary zone in the western United □States. *Science* 287: 834-836.  
1777
- 1778 Fournier, R. O., 1989. Geochemistry and dynamics of the Yellowstone National Park  
1779 hydrothermal system. *Ann. Rev. Earth Plan. Sci.* 17: 13-53.  
1780
- 1781 Fournier, R. O., Pitt, A. M., 1985. The Yellowstone magmatic-hydrothermal system, U.S.A. In:  
1782 C. Stone (Editor), 1985 International Symposium on Geothermal Energy. Geothermal  
1783 Resource Center, Davis, CA, pp. 319-327.  
1784
- 1785 Geist, D., Richards, M., 1993. Origin of the Columbia Plateau and Snake River Plain: Deflection  
1786 of the Yellowstone plume. *Geology* 21: 789-792.  
1787
- 1788 Grand, S. P., van der Hilst, R. D., Widiyantoro, S., 1997. Global seismic tomography: A  
1789 snapshot of convection in the Earth. *GSA Today* 7: 1-7.  
1790
- 1791 Griffiths, R. W., Campbell, I. H., 1990. Stirring and structure in mantle starting plumes. *Earth.*  
1792 *Plan. Sci. Lett.* 99(1-2): 67-78.  
1793
- 1794 Gripp, A. E., Gordon, R. G., 2002. Young tracks of hotspots and current plate velocities.  
1795 *Geophys. J. Int.* 150: 321-361.  
1796
- 1797 Hager, B. H., O'Connell, R. J., 1979. Kinematic models of large-scale flow in the earth's  
1798 mantle. *J. Geophys. Res.* 84(B3): 1031-1048.  
1799
- 1800 Hager, B. H., O'Connell, R. J., 1981. A simple global model of plate dynamics and mantle  
1801 convection. *J. Geophys. Res.* 86(B6): 6003-6015.  
1802
- 1803 Haines, A. J., Holt, W. E., 1993. A procedure for obtaining the complete horizontal motions  
1804 within zones of distributed deformation from the inversion of strain rate data. *J.*  
1805 *Geophys. Res.* 98(B7): 12,057-12,082.  
1806
- 1807 Haines, A. J., Jackson, J. A., Holt, W. E., Agnew, D. C., 1998. Representing distributed  
1808 deformation by continuous velocity fields. *Sci. Rept.* 98/5. Inst. Of Geol. and Nucl. Sci.,  
1809 Wellington, New Zealand.  
1810
- 1811 Haller, K. M., Wheeler, R. L., Rukstales, K. S., 2002. Documentation of changes in fault  
1812 parameters for the 2002 National Seismic Hazard Maps-Conterminous United States  
1813 except California, U. S. Geol. Surv. Open-File Rep. 02-467. U. S. Geol. Surv., Denver,  
1814 CO, 34 pp.  
1815
- 1816 Hammond, W.C., Humphreys, E.D., 2000. Upper mantle seismic wave velocity: Effects of  
1817 realistic partial melt geometries. *J. Geophys. Res.* 105: 10,975-10,986.

1818  
1819 Hanan, B. B., Shervais, J. W., Vetter, S. K., 2008. Yellowstone plume-continental lithosphere  
1820 interaction beneath the Snake River Plain. *Geology* 36: 51-54.  
1821  
1822 Hart, W.K., Carlson, R.W., 1987. Tectonic controls on magma genesis and evolution in the  
1823 northwestern United States. *J. Volc. Geotherm. Res.* 32: 119–135.  
1824  
1825 Hernlund, J.W., Tackley, P.J., Stevenson, D.J., 2008. Bouyant melting instabilities beneath  
1826 extending lithosphere:1 Numerical models. *J. Geophys. Res.* 113: B04405,  
1827 doi:10.1029/2006JB004862.  
1828  
1829 Hill, D. P., 1992. Temperatures at the base of the seismogenic crust beneath the Long Valley  
1830 caldera, California, and the Phlegrean fields caldera, Italy. In: P. Gasparini (Editor),  
1831 *Proceedings in Volcanology: Volcanic Seismology Vol. 3.* Springer-Verlag, Berlin, pp.  
1832 432-460.  
1833  
1834 Hill, D. P., Reasenber, P. A., Michael, A., Arabasz, W. J., Beroza, G., Brumbaugh, D., Brune, J.  
1835 N., Castro, R., Davis, S., dePolo, D., Ellsworth, W. L., Gombert, J., Harmsen, S., House,  
1836 L., Jackson, S. M., Johnston, M. J. S., Jones, L., Keller, R., Malone, S., Munguia, L.,  
1837 Nava, S., Pechmann, J. C., Sanford, A., Simpson, R. W., Smith, R. B., Stark, M.,  
1838 Stickney M., Vidal, A., Walter, S., Wong, V., Zollweg, J., 1993. Seismicity Remotely  
1839 Triggered by the Magnitude 7.3 Landers, California, Earthquake. *Science*, 260(5114):  
1840 1617-1623, doi:10.1126/science.260.5114.1617.  
1841  
1842 Hildreth, W., Halliday, A.N. and Christiansen, R.L., 1991. Isotopic and chemical evidence  
1843 concerning the genesis and contamination of basaltic and rhyolitic magma beneath the  
1844 Yellowstone Plateau volcananic field. *J. Pet.* 32: 63-138.  
1845  
1846 Holdahl, S. R., Dzurisin, D., 1991. Time-dependent models of vertical deformation for the  
1847 Yellowstone-Hebgen Lake region, 1923-1987. *J. Geophys. Res.* 96(B2): 2465-2483.  
1848  
1849 Humphreys, E. D., Dueker, K. G., Schutt, D. L., Smith, R. B., 2000. Beneath Yellowstone:  
1850 Evaluating plume and nonplume models using teleseismic images of the upper mantle.  
1851 *GSA Today* 10(12): 1-7.  
1852  
1853 Husen, S., Smith, R. B., 2004. Probabilistic earthquake relocation in three-dimensional velocity  
1854 models for the Yellowstone National Park Region, Wyoming. *Bull. Seis. Soc. Am.*  
1855 94(3): 880-896.  
1856  
1857 Husen, S., Smith, R. B., Waite, G. P., 2004. Evidence for gas and magmatic sources beneath the  
1858 Yellowstone volcanic field from seismic tomographic imaging. *J. Volc. and Geotherm.*  
1859 *Res.* 131: 397-410, doi:10.1016/S0377-0273(03)00416-5.  
1860  
1861 Husen, S., Taylor, R., Smith, R. B., Healsler, H., 2004. Changes in geyser behavior and remotely  
1862 triggered seismicity in Yellowstone National Park produced by the 2002 M7.9 Denali  
1863 fault earthquake. *Geology* 32: 537-540.

- 1864  
1865 Husen, S., Weimer, S., Smith, R. B., 2004. Remotely triggered seismicity in the Yellowstone  
1866 National Park region by the 2002 MW 7.9 Denali Fault Earthquake, Alaska. *Bull. Seism.*  
1867 *Soc. Am.* 94(6B): S317-S331.  
1868  
1869 Ito, G., van Keken, P.E., 2007. Hotspots and melting anomalies. In: D. Bercovici (Editor),  
1870 *Mantle Dynamics, Treatise on Geophysics v. 7.* Elsevier Press, Amsterdam, The  
1871 Netherlands.  
1872  
1873 Iyer, H.M., Evans, J. R., Zandt, G., Stewart, R. M., Coakley, J. M., Roloff, J. N., 1981. A deep  
1874 low-velocity body under the Yellowstone caldera, Wyoming: Delineation using  
1875 teleseismic P-wave residuals and tectonic interpretation: Summary. *Geol. Soc. of Am.*  
1876 *Bull.*, 92,(11): 792-798.  
1877  
1878 Jackson, I., Paterson, M. S., Fitzgerald, J. D., 1992. Seismic wave dispersion and attenuation in  
1879 Aheim dunite: an experimental study. *Geophys. J. Int.* 108: 517-534.  
1880  
1881 Jordan, M., 2003. JI-3D A new approach to high resolution regional seismic tomography: theory  
1882 and applications. PhD thesis, University of Göttingen, Göttingen, Germany.  
1883  
1884 Jordan, M., Smith, R. B., Waite, G. P., 2004. Tomographic Images of the Yellowstone Hotspot  
1885 Structure. *Eos Tran. AGU* 85(47), Fall Meet. Suppl., Abstract V51B-0556.  
1886  
1887 Jordan, M., Smith, R. B., Puskas, C., Farrell, J., Waite, G., 2005. The Yellowstone hotspot and  
1888 related plume: volcano-tectonics, tomography, kinematics and mantle flow. *Eos Trans.*  
1889 *AGU*, 86(52), Fall Meet. Suppl., Abstract T51D-1388.  
1890  
1891 Karato, S., 1993. Importance of anelasticity in the interpretation of seismic tomography.  
1892 *Geophys. Res. Lett.* 20: 1623-1626.  
1893  
1894 Karato S., Jung, H., 1998. Water, partial melting and the origin of the seismic low velocity and  
1895 high attenuation zone in the upper mantle. *Earth Planet. Sci. Lett.* 157: 193–207.  
1896  
1897 Kawamoto T., Holloway, J., 1997. Melting temperature and partial melt chemistry of H<sub>2</sub>O–  
1898 saturated mantle peridotite to 11 gigapascals. *Science* 276: 240–243.  
1899  
1900 Kennedy, B.M., Lynch, M.A., Reynolds, J.H., Smith, S.P., 1985. Intensive sampling of noble  
1901 gases in fluids at Yellowstone: I. Early overview of the data; regional patterns. *Geochim.*  
1902 *Cosmochim. Acta* 49: 1251–1261.  
1903  
1904 Kennett, B. L. N. Engdahl, E. R., 1991. Traveltimes for global earthquake location and phase  
1905 identification. *Geophys. J. Int.* 105(2): 429–465, doi:10.1111/j.1365-  
1906 246X.1991.tb06724.x.  
1907  
1908 King, S. D., Anderson, D. L., 1995. An alternative mechanism of flood basalt formation. *Earth*  
1909 *and Plan. Sci. Lett.* 136(3-4): 269-279.

1910

1911 Leeman, W.P., 1982. Development of the Snake River Plain-Yellowstone Plateau province,  
 1912 Idaho and Wyoming: An overview and petrologic model. In: B. Bonnichsen and R.M.  
 1913 Breckenridge (Editors), Cenozoic Geology of Idaho: Idaho Bureau of Mines and Geology  
 1914 Bulletin 26. Idaho Bureau of Mines and Geology, Moscow, ID, pp. 155-178.

1915

1916 Lees, J., VanDecar, J., 1991. Seismic tomography constrained by bouguer gravity anomalies:  
 1917 Applications in western Washington, Pure App. Geophys. 135: 31-52.

1918

1919 Lehman, J. A., Smith, R.B., Schilly, M.M, Braile, L.W., 1982. Crustal structure of the  
 1920 Yellowstone caldera from delay-time analyses and correlation with gravity data. J.  
 1921 Geophys. Res. 84: 2713-2730.

1922

1923 Lowenstern, J. B., Hurwitz, S., 2008. Monitoring a supervolcano in repose: Heat and volatile  
 1924 flux at the Yellowstone caldera, Elements, DOI: 10.2113/GSELEMENTS.4.1.35.

1925

1926 Lynch, D., Smith, R. B., Benz, H. M., 1997. Three-dimensional tomographic inversion of crust  
 1927 and upper mantle structure of the eastern Basin Range-Rocky Mountain transition from  
 1928 earthquake and regional refraction data. Abstracts from the 9th Annual IRIS Workshop,  
 1929 IRIS Consortium, Breckenridge, CO.

1930

1931 Lowry, A. R., Ribe, N. M., Smith, R. B., 2000. Dynamic elevation of the Cordillera, western  
 1932 United States. J. Geophys. Res. 105(B10): 23,371-23,390.

1933

1934 McCaffrey, R., 2000. Rotation and plate locking at the southern Cascadia subduction zone.  
 1935 Geophys. Res. Lett. 27(19): 3117-3120.

1936

1937 McCaffrey, R., Qamar, A. I., King, R. W., Wells, R., Ning, Z., Williams, C. A., Stevens, C. W.,  
 1938 Vollick, J. J., Zwick, P. C., 2007. Plate coupling, block rotation and crustal deformation  
 1939 in the Pacific Northwest. Geophys. J. Int., in press.

1940

1941 McCaIpin, J. P., Nishenko, S. P., 1996. Holocene paleoseismicity, temporal clustering, and  
 1942 probabilities of future large ( $M>7$ ) earthquakes on the Wasatch fault zone, Utah. J.  
 1943 Geophys. Res. 101: 6233-6253.

1944

1945 Meade, B. J., Hager, B. H., 2005. Block models of crustal motion in southern California  
 1946 constrained by GPS measurements. J. Geophys. Res. 110: B03403,  
 1947 doi:10.1029/2004JB003209.

1948

1949 Megnin, C., Romanowicz, B., 2000. The three-dimensional shear velocity structure of the  
 1950 mantle from the inversion of body, surface and higher-mode waveforms. Geophys. J. Int.  
 1951 143(3): 709-728, doi:10.1046/j.1365-246X.2000.00298.x.

1952

1953 Milbert, D.G., 1991. Computing GPS-derived orthometric heights with the GEOID90 geoid  
 1954 height model. Technical Papers of the 1991 ACSM-ASPRS Fall Convention. American  
 1955 Congress on Surveying and Mapping, Washington, D.C., pp. A46-55.

1956

1957 Miller, D. S., Smith, R. B., 1999. P and S velocity structure of the Yellowstone volcanic field

1958 from local earthquake and controlled-source tomography. *J. Geophys. Res.* 104: 15,105-

1959 15,121.

1960

1961 Montelli, R., Nolet, G., Dahlen, F., Masters, G., Engdahl, E. R. & Hung, S. H., 2004. Finite-

1962 frequency tomography reveals a variety of plumes in the mantle. *Science* 303: 338-343,

1963 doi:10.1126/science.1092485.

1964

1965 Morgan, W.J., 1972. Plate motions and deep mantle convection. *Geol. Soc. Am. Memoir* 132: 7-

1966 22.

1967

1968 Morgan, J. P., Morgan, W. J., Price, E., 1995. Hotspot melting generates both hotspot volcanism

1969 and a hotspot swell? *J. Geophys. Res.* 100: 8045-8062.

1970

1971 Morgan, P. Blackwell, ., D. D., Spafford R. E., Smith, R.B., 1977. Heat flow measurements in

1972 Yellowstone Lake and the thermal structure of the Yellowstone Caldera. *J. Geophys.*

1973 *Res.* 82: 379-3732.

1974

1975 Mueller, I.I., 1991. The International GPS Geodynamics Service. *GPS Bull.* 4: 7-16.

1976

1977 Nabelek, J., Xia, G., 1995. Moment-tensor analysis using regional data: application to the 25

1978 March, 1993, Scotts Mills, Oregon, earthquake. *Geophys. Res. Lett.* 22(1): 13–16.

1979

1980 Nash, B. P., Perkins, M. E., Christensen, J. N., Lee, D. C., Halliday, A., 2006. The Yellowstone

1981 hotspot in space and time: Nd and Hf isotopes in silicic magmas. *Earth and Plan. Sci.*

1982 *Lett.* 247(1-2): 143-156.

1983

1984 Nolet, G. R. Allen, D. Zhao, D., 2007. Mantle plume tomography. *Geochem. Geol.* 241: 248-

1985 263, doi: 10.1016/j.chemgeo.2007.01.022.

1986

1987 Nolet, G., Karato. S., Montelli, R., 2006. Plume fluxes from seismic tomography. *Earth and*

1988 *Plan. Sci. Lett.* 248: 685-699.

1989

1990 Olson, P., Singer, H., 1985. Creeping plumes. *J. Fluid Mech.* 158: 511-531.

1991

1992 Parsons, T., Thompson, G. A., Smith, R. P., 1998. More than one way to stretch: a tectonic

1993 model for extension along the plume track of the Yellowstone hotspot and adjacent Basin

1994 and Range Province. *Tect.* 17(2): 221-234.

1995

1996 Pelton, J.R., Smith, R.B., 1982. Contemporary vertical surface displacements in Yellowstone

1997 National Park. *J. Geophys. Res.* 87: 2745-2761.

1998

1999 Peng, X. and Humphreys, E.D., 1998. Crustal velocity structure across the eastern Snake River

2000 Plain and the Yellowstone swell. *J. Geophys. Res.*, 103: 7171-7186.

2001

- 2002 Perkins, M. E., Nash, B. P., 2002. Explosive silicic volcanism of the Yellowstone hotspot: the  
 2003 ash fall tuff record. *Geol. Soc. Am. Bull.* 114(3): 367-381.  
 2004
- 2005 Pierce, K. L., Morgan, L. A., 1992. The track of the Yellowstone hot spot: Volcanism, faulting,  
 2006 and uplift. In: P.K. Link, M.A. Kuntz, L.B. Platt (Editors), *Regional Geology of Eastern*  
 2007 *Idaho and Western Wyoming: Geological Society of America Memoir 179.* Geological  
 2008 Society of America, Boulder, CO, pp. 1-53.  
 2009
- 2010 Pitt, A. M., Weaver, C. S., Spence, W., 1979. The Yellowstone Park earthquake of June 30,  
 2011 1975. *Bull. Seismol. Soc. Am.* 69: 187-205.  
 2012
- 2013 Priestly, K., Orcutt, J., 1982. Extremal travel time inversion of explosion seismology data from  
 2014 the eastern Snake River Plain, Idaho. *J. Geophys. Res.* 87: 2634-2642.  
 2015
- 2016 Puskas, C. M., Smith, R. B., Meertens, C. M., Chang, W. L., 2007. Crustal deformation of the  
 2017 Yellowstone-Snake River Plain volcanic system: campaign and continuous GPS  
 2018 observations, 1987-2004. *J. Geophys. Res.* 112: B03401, doi:10.1029/2006JB004325.  
 2019
- 2020 Puskas, C. M., Smith, R. B., Flesch, L. M., Settles, K., 2007. Effects of the Yellowstone Hotspot  
 2021 on western U.S. Stress and Deformation. *Eos Trans. AGU* 88(52), Fall Meet. Suppl.,  
 2022 Abstract V51F-04.  
 2023
- 2024 Puskas, C. M., Smith, R. B., 2008. Intraplate Deformation and Microplate Tectonics of the  
 2025 Yellowstone Hotspot. *Earth Plan. Sci. Lett.*, (submitted).  
 2026
- 2027 Richards, M. A., Griffiths, R. W., 1988. Deflection of plumes by mantle shear flow:  
 2028 experimental results and a simple theory. *Geophys. J. Int.* 94(3): 367-376,  
 2029 doi:10.1111/j.1365-246X.1988.tb02260.x.  
 2030
- 2031 Ritter, J. R. R., 2004. Small-scale mantle plumes: imaging and geodynamic aspects. In: K.  
 2032 Fuchs, F. Wenzel (Editors), *Challenges for Earth Sciences in the 21st Century.* Springer  
 2033 Verlag, Berlin.  
 2034
- 2035 Rodgers, D. W., Hackett, W. R., Ore, H. T., 1990. Extension of the Yellowstone Plateau, eastern  
 2036 Snake River Plain, and Owyhee plateau. *Geol.* 18: 1138-1141.  
 2037
- 2038 Ruppel, E.T., 1972. Geology of pre-Tertiary rocks in the northern part of Yellowstone National  
 2039 Park, Wyoming, with a section on Tertiary laccoliths, sills, and stocks in and near the  
 2040 Gallatin Range, Yellowstone National Park. *Geology of Yellowstone National Park:*  
 2041 *U.S. Geol. Surv. Prof. Pap. 729-A, U.S. Geolo. Surv., Denver, CO, 66 pp.*  
 2042
- 2043 Saltzer, R. L., Humphreys, E. D., 1997. Upper mantle P wave velocity structure of the eastern  
 2044 Snake River Plain and its relationship to geodynamic models of the region. *J. Geophys.*  
 2045 *Res.* 102: 11,829-11,841.  
 2046

2047 Schmitz, M., Heinshohn, W. D., Schilling, F. R., 1997. Seismic, gravity and petrological  
2048 evidence for partial melt beneath the thickened Central Andean crust (21-23°S).  
2049 *Tectonophys.* 270(3-4): 313-326, doi:10.1016/S0040-1951(96)00217-X.  
2050

2051 Schutt D. L., K. Dueker, H. Yuan, 2008. Crust and upper mantle velocity structure of the  
2052 Yellowstone hot spot and surroundings. *J. Geophys. Res.* 113: B03310,  
2053 doi:10.1029/2007JB005109.  
2054

2055 Schutt, D.L., Humphrey, E.D., 2004. P and S wave velocity and V-P/V-S in the wake of the  
2056 Yellowstone hot spot. *J. Geophys. Res.* 109: B01305, doi:10.1029/2003JB002442.  
2057

2058 Shervais, J.W., Vetter, S.K. and Hanan, B.B., 2006. Layered mafic sill complex beneath the  
2059 eastern Snake River Plain: Evidence from cyclic geochemical variations in basalt.  
2060 *Geology*, 34: 365-368.  
2061

2062 Sigloch, K., McQuarrie, N., Nolet, G., 2008. Two-stage subduction history under N. America  
2063 inferred from multiple-frequency tomography. *Nature Geoscience*, (in press).  
2064

2065 Sillard, P., Altamimi, Z., Boucher, C., 1998. The ITRF96 realization and its associated velocity  
2066 field. *Geophys. Res. Lett.* 25: 3223-3226, doi: 10.1029/98GL52489.  
2067

2068 Simmons, N. A., Forte, A. M., Grand, S. P., 2006. Constraining mantle flow with seismic and  
2069 geodynamic data: a joint approach. *Earth Plan. Sci. Lett.* 46: 109-124.  
2070

2071 Sleep, N. H., 1990. Hotspots and mantle plumes: some phenomenology. *J. Geophys. Res.*  
2072 95(B5): 6715-6736.  
2073

2074 Smith, R. B., 1977. Intraplate tectonics of the Western North American Plate. *Tectonophys.* 37:  
2075 323-336.  
2076

2077 Smith, R. B., Braile, L. W., Schilly, M. M., Ansorge, J., Prodehl, C., Baker, M. Healey, J. H.  
2078 Mueller, S. Greensfelder, R., 1982. The Yellowstone- eastern Snake River Plain seismic  
2079 profiling experiment: Crustal structure of Yellowstone, *J. Geophys. Res.*, 84, 2583-2596.  
2080

2081 Smith, R. B., Arabasz, W. J., 1991. Seismicity of the Intermountain Seismic Belt. In: D.B.  
2082 Slemmons, E.R. Engdahl, M.L. Zoback, and D.D. Blackwell (Editors), *Neotectonics of*  
2083 *North America*. Geological Society of America, Boulder, CO, pp. 185-228.  
2084

2085 Smith, R. B., Braile, L. W. 1993. Topographic signature, space-time evolution, and physical  
2086 properties of the Yellowstone-Snake River Plain volcanic system: the Yellowstone  
2087 hotspot, In: A. W. Snoke, J. Steidtmann, S. M. Roberts (Editors), *Geology of Wyoming:*  
2088 *Geological Survey of Wyoming Memoir No. 5*, Wyoming State Geological Survey,  
2089 Laramie, WY, p. 694-754.  
2090

2091 Smith, R. B., Braile, L. W., 1994. The Yellowstone hotspot. *J. Volc. and Geotherm. Res.* 61:  
2092 121-187.



2093  
2094 Smith, R. B., Bruhn, R. L., 1984. Intraplate extensional tectonics of the eastern Basin-Range:  
2095 inferences on structural style from seismic reflection data, regional tectonics, and  
2096 thermal-mechanical models of brittle-ductile deformation. *J. Geophys. Res.* 89(B7):  
2097 5733-5762.  
2098  
2099 Smith, R.B., Siegel, L., 2000. *Windows into the Earth's interior; The geologic story of*  
2100 *Yellowstone and Grand Teton National Parks*, Oxford University Press, 242 pp.  
2101  
2102 Smith, R. B., Jordan, M., Puskas, C., Waite, G. Farrell, J., 2005. Geodynamic models of the  
2103 Yellowstone Hotspot constrained by seismic and GPS imaging and volcano-tectonic data.  
2104 2005 Salt Lake City Annual Meeting, Geol. Soc. Amer., Abstract 54-2.  
2105  
2106 Smith, R.B. Sbar, M., 1974. Contemporary tectonics and seismicity of the Western United  
2107 States with emphasis on the Intermountain Seismic Belt. *Geol. Soc. Am. Bull.* 85: 1205-  
2108 1218.  
2109  
2110 Smith, R. B., Blackwell, D.D., 2000. Heat flow and energetics of Yellowstone Lake  
2111 hydrothermal systems. *Eos Trans. AGU* 81(48), Fall Meet. Suppl., Abstract V22F-17.  
2112  
2113 Smith, R. B., G. P. Waite, C. M. Puskas, D.L. Shut and E.D. Humphreys, 2003, Dynamic and  
2114 kinematic models of the Yellowstone Hotspot constrained by seismic anisotropy, GPS  
2115 measurements and fault slip rates. *Eos. Trans. AGU* 84(86), Fall Meet. Suppl., Abstract  
2116 T51G-05.  
2117  
2118 Smith, R. P., Jackson, S. M. and Hackett W. R., 1996. Paleoseismology and seismic hazards  
2119 evaluations in extensional volcanic terrains. *J. Geophys. Res.* 101(B3): 6277–6292.  
2120  
2121 Sparlin, M. A., Braile, L. W., Smith, R. B., 1982. Crustal structure of the eastern Snake River  
2122 Plain determined from ray trace modeling of seismic refraction data. *J. Geophys. Res.*  
2123 87(B4): 2619-2633.  
2124  
2125 Steck, L.K., Prothero, W.A., 1991. A 3-D raytracer for teleseismic body-wave arrival times.  
2126 *Bull. Seis. Soc. Am.* 81: 1332–1339.  
2127  
2128 Steinberger, B., 2000. Plumes in a convecting mantle: models and observations from individual  
2129 hotspots. *J. Geophys. Res.* 105: 11,127-11,152.  
2130  
2131 Steinberger, B., Antretter, A., 2006. Conduit diameter and buoyant rising speed of mantle  
2132 plumes: implications for the motion of hot spots and shape of plume conduits. *Geochem.*  
2133 *Geophys. Geosys.* 7: Q11018, doi:10.1029/2006GC001409.  
2134  
2135 Steinberger, B., Calderwood, A. R., 2006. Models of large-scale viscous flow in the Earth's  
2136 mantle with constraints from mineral physics and surface observations. *Geophys. J. Int.*  
2137 167: 1461-1481.  
2138

- 2139 Steinberger, B., O'Connell, R.J., 1998. Advection of plumes in mantle flow; implications for hot  
2140 spot motion, mantle viscosity and plume distribution. *Geophys. J. Int.* 132: 412-434.  
2141
- 2142 Steinberger, B., O'Connell, R. J., 2000. Effects of mantle flow on hotspot motion. *Geophys.*  
2143 *Mono.* 121: 377-398  
2144
- 2145 Steinberger, B., Sutherland, R., O'Connell, R. J., 2004. Prediction of Emperor-Hawaii seamount  
2146 locations from a revised model of global plate motion and mantle flow. *Nature* 430:  
2147 167-173.  
2148
- 2149 Tarantola, A., Valette, B., 1982. Generalized nonlinear inverse problems solved using the least  
2150 squares criterion. *Rev. Geophys. Space Phys.* 20: 219–232.  
2151
- 2152 Tapley, B., Ries, J., Bettadpur, S., Chambers, D., Cheng, M., Condi, F., Gunter, B., Kang, Z.,  
2153 Nagel, P., Pastor, R., Pekker, T., Poole, S., Wang, F., 2005. GGM02 – An improved  
2154 Earth gravity field model from GRACE. *J. Geod.* 79: 467-478, doi:10.1007/s00190-005-  
2155 0480-z.  
2156
- 2157 Thatcher, W., 2003. GPS constraints on the kinematics of continental deformation. *Int. Geol.*  
2158 *Rev.* 45: 191-212.  
2159
- 2160 Tsuboi, C., 1954. A new and simple method for calculating the deflections of the vertical from  
2161 gravity anomalies with the aid of BESSEL FOURIER series, *Proceedings of the Japan*  
2162 *Academy* 30(6): 461-466.  
2163
- 2164 van der Hilst, R.D., De Hoop, M.V., 2005. Banana-doughnut kernels and mantle tomography.  
2165 *Geophys. J. Int.* 163: 956– 961.  
2166
- 2167 Vasco, D.W., Johnson, L. R., Goldstein, N. E., 1988. Using surface displacement and strain  
2168 observations to determine deformation at depth, with an application to Long Valley  
2169 Caldera, California. *J. Geophys. Res.* 93: 3232-3242.  
2170
- 2171 Vasco, D. W., Puskas, C. M., Smith, R. B., Meertens, C. M., 2007. Crustal deformation and  
2172 source models of the Yellowstone volcanic field from geodetic data. *J. Geophys. Res.*  
2173 112: B07402, doi:10.1029/2006JB004641.  
2174
- 2175 Vasco, D. W., Smith, R., Taylor, C., 1990. Inversion of Yellowstone vertical displacements and  
2176 gravity changes, 1923 to 1975-1977 to 1986. *J. Geophys. Res.* 95: 19,839-19,856.  
2177
- 2178 Waite, G. P., 1999. Seismicity of the Yellowstone Plateau: Space-time patterns and stresses  
2179 from focal mechanism inversion. M.S. thesis, University of Utah, Salt Lake City, UT.  
2180
- 2181 Waite, G. P., 2004. Upper mantle structure of the Yellowstone hotspot from teleseismic body-  
2182 wave velocity tomography and shear-wave anisotropy. Ph.D dissertation, University of  
2183 Utah, Salt Lake City, Utah.  
2184

- 2185 Waite, G. P., Chang, W., 2007. Shear-wave splitting from local earthquakes as an indicator of  
2186 crustal stress at Yellowstone. *Eos Trans. AGU* 88(52), Fall Meet. Suppl., Abstract V3B-  
2187 1319.
- 2188
- 2189 Waite, G. P., Schutt, D. L., Smith R. B., 2005. Models of lithosphere and asthenosphere  
2190 anisotropic structure of the Yellowstone hot spot from shear wave splitting. *J. Geophys.*  
2191 *Res.* 110: B11304, doi:1029/2004JB003501.
- 2192
- 2193 Waite, G. P., Smith, R. B., 2002. Seismic evidence for fluid migration accompanying  
2194 subsidence of the Yellowstone caldera. *J. Geophys. Res.* 107(B9): 2177,  
2195 doi:10.1029/2001JB000586.
- 2196
- 2197 Waite, G. P., Smith R. B., 2004. Seismotectonics and stress field of the Yellowstone volcanic  
2198 plateau from earthquake first motions and other indicators. *J. Geophys. Res.* 109:  
2199 B02301, doi:1029/2003JB002675.
- 2200
- 2201 Waite, G. P., Smith, R. B., Allen, R. M., 2006. Vp and Vs structure of the Yellowstone hot spot:  
2202 evidence for an upper mantle plume. *J. Geophys. Res.* 111(B4): B04303,  
2203 doi:10.1029/2005JB003867.
- 2204
- 2205 Wells D. L. Coppersmith K. J., 1994. New empirical relationships among magnitude, rupture  
2206 length, rupture width, rupture area, and surface displacement. *Bull. Seis. Soc. Am.* 84:  
2207 974-1002.
- 2208
- 2209 White, B. J. P., Smith, R. B., Farrell, J., Husen, S., Wong, I., 2008. Seismicity and earthquake  
2210 hazard analysis of the Teton-Yellowstone region, Wyoming. *J. Vol. Geotherm. Res.*,  
2211 (this volume).
- 2212
- 2213 Whitehead, J. A., 1982. Instabilities of fluid conduits in a flowing Earth: are plates lubricated by  
2214 the asthenosphere? *Geophys. J. R. Astr. Soc.* 70: 415–433.
- 2215
- 2216 Whitehead, J. A., Luther, D. S., 1975. Dynamics of laboratory diapir and plume models. *J.*  
2217 *Geophys. Res.* 80(B5): 705-717.
- 2218
- 2219 Wicks, C., Thatcher, W., Dzurisin, D., 1998. Migration of fluids beneath Yellowstone caldera  
2220 inferred from satellite radar interferometry. *Science* 282(5388): 458-462,  
2221 doi:10.1126/science.282.5388.458.
- 2222
- 2223 Wicks, C., Thatcher, W., Dzurisin, D., Svarc, J., 2006. Uplift, thermal unrest and magma  
2224 intrusion at Yellowstone caldera. *Nature* 440: 72-75, doi:10.1038/nature04507.
- 2225
- 2226 Wilson, J. T., 1963. A possible origin of the Hawaiian Islands. *Can. J. Phys.* 41: 863-870.
- 2227
- 2228 Wolfe, C. J., Solomon, S. C., Silver, P. G., VanDecar, J. C. and Russo, R. M., 2002. Inversion of  
2229 body-wave delay times for mantle structure beneath the Hawaiian islands: results from  
2230 the PELENET experiment. *Earth Planet. Sci. Lett.* 198: 129-145.

- 2231  
2232 Wüllner, U., Christensen, U. R., M. Jordan, 2006. Joint geodynamical and seismic modeling of  
2233 the Eifel plume. *Geophys. J. Int.* 165(1): 357–372, doi:10.1111/j.1365-  
2234 246X.2006.02906.x  
2235
- 2236 Xue, M., R.M. Allen, 2007. The fate of the Juan de Fuca plate: Implications for a Yellowstone  
2237 plume head. *Earth Planet. Sci. Lett.*, 264(1): 266-276.  
2238
- 2239 Yuan, H., Dueker, K., 2005. Teleseismic P-wave tomogram of the Yellowstone plume.  
2240 *Geophys. Res. Lett.* 32: L07304, doi:10.1029/2004FL022056.  
2241
- 2242 Zeyen, H., Achauer, U., 1997. Joint inversion of teleseismic delay times and gravity anomaly  
2243 data for regional structures: theory and synthetic examples. In: K. Fuchs (Editor), *Upper*  
2244 *Mantle Heterogeneities from Active and Passive Seismology*. Kluwer Academic  
2245 Publishers, Dordrecht, Netherlands, pp. 155-168.  
2246
- 2247 Zoback, M. D., Zoback, M. L., 1991. Tectonic stress field of North America and relative plate  
2248 motions. In: D. B. Slemmons, E. R. Engdahl, M. D. Zoback, D. D. Blackwell (Editors),  
2249 *Neotectonics of North America*. Geological Society of America, Boulder, CO, p. 339-  
2250 366.
- 2251 **Table Captions**
- 2252 **Table 1.** Table of the B-values and density variations based on the geoid modelling and the  
2253 results of the optimized tomographic inversion. The B-values and density deviations are given  
2254 for the optimum result (No.1), and further examples of plausible results at calculation numbers  
2255 500 and 1000.

2256 **Figure Captions**

2257 Fig. 1. Global signatures of the Yellowstone hotspot. (a) Global free-air gravity anomaly of  
2258 some notable hotspots including Hawaii, Iceland, and Yellowstone in Mgal. These anomalies  
2259 have wavelengths the order of 1000 km long and relative amplitudes of +20 to +40 Mgal; and (b)  
2260 North America geoid map showing the Yellowstone -7 geoid anomaly. Yellowstone has a  
2261 ~1000 km wide topographic swell.

2262  
2263 Fig. 2. Track of the Yellowstone hotspot (Y) showing the relative motion of age-transgressive  
2264 SRP silicic volcanic centers at 180° to the direction of the direction North American plate  
2265 motion. The topographically low area occupied by the Snake River Plain is outlined in green.  
2266 Centers of post-17 Ma silicic volcanism (in yellow) contain multiple caldera-forming eruptions.  
2267 Red dots are historic earthquake epicenters, taken from compilations of the University of Utah  
2268 and the USGS of M1.5 – 7.5 earthquakes. Late Quaternary faults are shown by black lines;  
2269 Cenozoic basaltic dikes (age in Ma) are shown in yellow and orange; <sup>87</sup>Sr/<sup>86</sup>Sr isotope boundary  
2270 for the 0.706 value is shown as a black-dashed line; and the YSRP tectonic parabola is defined  
2271 by the bow-shaped pattern of high topography and seismicity surrounding the YSRP (yellow  
2272 dashed lines). The relative motion vector of the North American Plate over the mantle at 2.2  
2273 cm/yr is noted as a large arrow. The Newberry-Oregon trend of silicic volcanism extends NW  
2274 across southeast Oregon to the Newberry caldera (N).

2275  
2276 Fig. 3. Space view of Grand Teton and Yellowstone National Parks from Landsat satellite  
2277 images overlain on digital elevation data. The 2.4 km-high Yellowstone caldera was produced  
2278 by a giant volcanic eruption 630 000 years ago. The caldera occupies a 60 by 40-km-wide area  
2279 of central Yellowstone. The Teton fault bounds the east side of the Teton Range and raised the  
2280 mountains high above Jackson Hole's valley floor. From Smith and Siegel (2000).

2281  
2282 Fig. 4. Map of the seismic and GPS stations deployed for the 1999-2003 Yellowstone Hotspot  
2283 Geodynamic project. The networks contain 166 seismic stations (broadband and short-period),  
2284 15 permanent GPS and 150 campaign GPS stations. Note the linear distribution of stations in  
2285 600 km long arrays with a NW azimuth designed to best record earthquakes at teleseismic  
2286 distances (>1000 km) from the major seismic belts of the western Pacific and South America.

2287  
2288 Fig. 5. Volcanic and tectonic features of Yellowstone and surrounding area. Yellowstone  
2289 calderas I (2.1 Ma), II (1.2 Ma) and III (0.64 Ma) are shown as black lines and labeled. The two  
2290 resurgent domes Mallard Lake (ML) and Sour Creek (SC) are shown with dashed black lines.  
2291 Yellow stars mark post-caldera volcanic vents of 640 000 to 70 000 years in age. Late  
2292 Quaternary faults are heavy black lines with ticks on downthrown side. Fault abbreviations are  
2293 EGF=eastern Gallatin fault, HLF=Hebgen Lake fault, MF=Madison fault, CF=Centennial fault,  
2294 TF=Teton fault, MSF=Mount Sheridan fault, YLF=Yellowstone Lake fault, BFF=Buffalo Fork  
2295 fault, UYF=Upper Yellowstone Valley fault. Areas of hydrothermal features, including geysers,  
2296 fumaroles, and hot springs, are shown in orange.

2297  
2298 Fig. 6. (a) Heatflow of the Yellowstone and the Snake River Plain, with heatflow of the SRP  
2299 averaging ~ 150 mWm<sup>-2</sup>; (b) Yellowstone Plateau averaged heat flow ~2000 mWm<sup>-2</sup>; and (c)  
2300 very high heatflow of Yellowstone Lake ranges from ~100 mWm<sup>-2</sup> to extraordinarily high 30  
2301 000 mWm<sup>-2</sup> (after Blackwell and Richards, 2006).

2302

2303 Fig. 7. Photographs of prominent hydrothermal features of Yellowstone. (a) Aerial view of a  
2304 rare eruption of Steamboat Geyser in Norris Geyser Basin, July 6, 1984. It the largest geyser in  
2305 the world, sending water up to 250 m high. Steamboat Geyser erupts sporadically at decadal  
2306 scales but has erupted 7 times in 2006-2007. (b) The Hot Springs Basin group of fumaroles and  
2307 hot springs is the largest hydrothermal area of Yellowstone and is located 15 km north of the  
2308 caldera boundary, between the Sour Creek dome and the Mirror Plateau. Pictures by Robert B.  
2309 Smith.

2310

2311 Fig. 8. Station map of Yellowstone National Park and surrounding area showing seismograph,  
2312 GPS, and borehole strainmeter networks. Seismic stations consist of broadband and short-period  
2313 seismometers operated by the University of Utah. GPS sites include 26 permanent and 90  
2314 temporarily occupied (campaign) sites operated by the University of Utah and EarthScope PBO.  
2315 Borehole strainmeter sites also contain downhole seismometers and are operated by the  
2316 EarthScope PBO project. All data from these instruments are available in real-time and online.

2317

2318 Fig. 9. Seismicity of the Yellowstone Plateau (1975-2007). Epicenters are located by employing  
2319 a three-dimensional P-wave velocity model of Husen and Smith (2004). Locations of the M7.5  
2320 1959 Hebgen Lake Mt and M6.1 Norris Jct are highlighted as a large red star and large circle,  
2321 respectively.

2322

2323 Fig. 10. Photographs of active faults of the Yellowstone Plateau: (a) the Hebgen Lake fault  
2324 broke vertically during the M7.5 Hebgen Lake, MT, earthquake with as much as 5.7 m of  
2325 vertical offset; and (b) caldera-boundary faults on the Mirror Plateau, northeast Yellowstone,  
2326 with a maximum offset of ~30 on the left, SW-facing fault. Note three antithetic NW-facing  
2327 faults to right, creating a graben occupied by Mirror Lake. Pictures by Robert B. Smith.

2328

2329 Fig. 11. Maximum focal depths of the Hebgen Lake-Yellowstone area serve as a proxy for  
2330 conductive temperature: (a) well-located hypocenters (red dots) along 10-km wide windows  
2331 corresponding to profiles in map view (b). The 80<sup>th</sup> percentile maximum focal depth is marked  
2332 by the dashed line. This depth is interpreted as the brittle-ductile transition at ~400°C. (b)  
2333 Contoured map of focal depths showing the very shallow focal depths in the caldera produced by  
2334 high temperatures. The thin seismogenic layer of the caldera, ~5 km thick, restricts the  
2335 maximum earthquake magnitude to Mw 6.5.

2336

2337 Fig. 12. Isosurfaces of anomalously low P-wave bodies are determined from local earthquake  
2338 tomography of the Yellowstone caldera and reveal the Yellowstone magma chamber. The  
2339 shallow anomaly plotted in blue is interpreted to be a gas-saturated body. The red anomaly is  
2340 interpreted to be 5% to 15% partial melt corresponding to a crystallizing magma body that feeds  
2341 the surface silicic and basaltic magmatism of Yellowstone (from Husen and Smith, 2004).

2342

2343 Fig. 13. Crustal deformation of the Yellowstone Plateau deformation from leveling and GPS  
2344 observations (after Pelton and Smith, 1982; Puskas et al., 2007a; Vasco et al., 2007). Color  
2345 backgrounds represent vertical motion measured from (a) leveling surveys between 1923 and  
2346 1987 and (b)-(d) GPS campaigns between 1987 and 2003. Red circles represent campaign GPS

2347 sites, yellow circles represent permanent GPS stations, and arrows are the direction of motion.  
2348 Time windows correspond to the distinct periods of caldera uplift and subsidence.

2349  
2350 Fig. 14. Stress field of Yellowstone and the Snake River Plain illustrating dominant lithospheric  
2351 NE-SW extension: (a) directions of stress in the YSRP from focal mechanism T axes, minimum  
2352 horizontal principals stresses ( $\sigma_3$ ), slip directions of normal faults ( $\phi$ ), post-caldera volcanic vent  
2353 alignments, and GPS- derived strain tensors; and b) similar stress directions of the Yellowstone  
2354 Plateau (Waite and Smith, 2004).

2355  
2356 Fig. 15. Temporal history of deformation and earthquakes of Yellowstone. Earthquakes are  
2357 sorted by date into quarters to obtain the total number of earthquakes per three-month period.  
2358 Specific leveling and GPS surveys are shown as black squares and white circles, respectively.  
2359 Dates when InSAR images were taken are shown as gray diamonds. Deformation rates are  
2360 averaged for periods of homogeneous deformation, either uplift or subsidence (from Chang et al.,  
2361 2007).

2362  
2363 Fig. 16. Unprecedented uplift of the Yellowstone caldera revealed by GPS and InSAR data  
2364 (2004-2007) modified from Chang et al. (2007). (a) Map view of the uplift with GPS vertical  
2365 and horizontal vectors and background showing line of sight (nearly vertical) deformation in 28  
2366 mm displacement bands. Note the maximum 7 cm/yr of uplift of the caldera compared to up to  
2367 1.5 cm/yr of subsidence of the Norris Geyser basin area. (b) Cross section of modeled 10° SE-  
2368 dipping sill that is interpreted to be inflating at 0.1 km<sup>3</sup> per year, consistent with the modeled rate  
2369 of inflation from the heatflow and geochemical data. Color contours are Coulomb stress increase  
2370 (red) or decrease (blue) caused by inflation of the sill. Hypocenters of earthquakes that occurred  
2371 during the period of accelerated uplift are shown as black dots.

2372  
2373 Fig. 17. P-wave velocity slices (km/s) from tomographic inversion of teleseismic data for the  
2374 Yellowstone hotspot (after Jordan et al., 2005; Waite et al., 2006). Data consisted of P-wave  
2375 arrivals of 115 earthquakes recorded at 86 stations with 3399 P and 380 PKIPK arrivals. Maps  
2376 are horizontal slices of P-wave velocity at selected depths with corresponding relative decrease  
2377 (red) and increase (blue) velocities. Note the low-velocity anomaly beneath Yellowstone is  
2378 displaced to the west as depth increases. A high-velocity zone is located to the east of  
2379 Yellowstone. Profile lines for Fig. 18 are shown in the 30 km and 330 km depth slices.

2380  
2381 Fig. 18. Two-dimensional cross sections of the Yellowstone P-wave low velocity anomalies  
2382 corresponding to Fig. 17 (Jordan et al., 2005). (a) NW-SE cross-section across western Montana  
2383 and western Wyoming, and b) NE-SW profile along the YSRP. Significantly these profiles  
2384 reveal a 60° west-dipping low-velocity anomaly of up to -1.5%. The anomaly extends to 660 km  
2385 in the NW-SE profile, but does not extend deeper than 200 km beneath Yellowstone in the NE-  
2386 SW profile.

2387  
2388 Fig. 19. Seismic image of the Yellowstone plume as a 60° west-dipping, rising column of  
2389 molten rock of up to -1.5% melt originating in the mantle transition zone. The plume is  
2390 represented by a three-dimensional P-wave velocity isosurface of the ~-1% value. The top of  
2391 this upper mantle plume underlies Yellowstone to depths of ~250 km, but the deeper part to the  
2392 northwest is at a depth of ~650 km, at the bottom of the mantle transition zone. (Also see 3D

2393 animation of the University of Utah Yellowstone hotspot project results: earthquakes, geysers,  
2394 faults, GPS sites, volcanoes, calderas, topography and other geo-located geologic features of the  
2395 Yellowstone-Teton region <http://siovizcenter.ucsd.edu/library/objects/detail.php?ID=210>. The  
2396 iView viewer is free and can be downloaded from:  
2397 [http://www.ivs3d.com/download/iview3d\\_download.html](http://www.ivs3d.com/download/iview3d_download.html) )  
2398

2399 Fig. 20. One-dimensional geodynamic model of the Yellowstone plume plotted in two  
2400 dimensions for realism. Effect of two-dimensional extrapolation is negligible for the low relative  
2401 velocities. Model of excess temperature ( $^{\circ}\text{K}$ ) is constrained by seismic P-wave velocity and  
2402 relative attenuation of P-waves ( $Q_p$ ) in  $Q^{-1}$ , for wet and dry models of Cammarano et al. (2003,  
2403 2007).  $Q^{-1}$  is the relative attenuation of seismic waves for P- and S-waves modified for  
2404 Yellowstone by Clawson et al., (1989) and converted to  $Q_s$  by standard methods.  
2405

2406 Fig. 21. Computed regional mantle density structure and flow beneath North America and the  
2407 northeastern Pacific for tomography model mean (mean shear wave) with viscosity structure  
2408 VM1 from Fig. 22. (a) Mantle cross-section showing density structure along the line shown in  
2409 the bottom panel. The relatively dense subducted Farallon plate is located beneath the eastern  
2410 U.S. at depths of 1000 to 1900 km. (b) Map view of upper mantle flow at 359 km depth.  
2411 Vectors represent horizontal components of flow and color background represents vertical flow.  
2412

2413 Fig. 22. Computed horizontal upper mantle flow in the vicinity of the Yellowstone hotspot at  
2414 various depths. Viscosity models VM1 and VM2, used as the basis of the mantle flow models,  
2415 are shown in the upper right panel. Model VM1 constrains the tomography model of Fig. 21.  
2416 Bottom left panel also includes fixed-hotspot tracks (0-15 Ma) for four different models of  
2417 North American “absolute” plate motion.  
2418

2419 Fig. 23. Models of hotspot plume conduits for (a) plumes ascending from the mantle transition  
2420 zone (660 km), and (b) initially vertical plumes ascending from the core-mantle boundary. The  
2421 upper panels in (a) and (b) show the surface hotspot tracks and the progression of surface  
2422 volcanism over time (0-15 Ma) from Steinberger et al. (2004). Gray shaded areas represent  
2423 provinces of basaltic volcanism in the Columbia Plateau and eastern Snake River Plain. The  
2424 lower panels show the projection of the plume conduits into map view and plume displacement  
2425 with depth. Colored lines represent plume conduit models corresponding to mantle flow models  
2426 from Fig. 22.  
2427

2428 Fig. 24. Cross-section of western North America mantle S-wave velocity structure from mantle  
2429 tomography (Grand, 1997). Mantle flow directions represented by vectors. Yellowstone mantle  
2430 plume is superimposed as thick orange line, with hypothesized lower mantle extension shown as  
2431 a dashed line.  
2432

2433 Fig. 25. (a) Geoid map of Western U.S. from GEOID2003 model with profile line AA', (b)  
2434 parameterized Yellowstone plume model, and (c) forward models of the geoid based on plume  
2435 parameterizations (blue lines) compared to the filtered geoid (red lines) in the area of the  
2436 Yellowstone hotspot swell. Geoid data were shifted relative to background geoid values so that  
2437 modeled heights ranged from +7 to +30 m. The black line represents the unfiltered geoid and the  
2438 white lines are selected models discussed in Section 8. The plume in panel (b) was initially



2439 parameterized into nine sections, but only the uppermost four segments contributed to the  
2440 solution. Density perturbations and P-wave velocities for the best-fit model are included in the  
2441 plot.

2442  
2443 Fig. 26. (a) Velocity field based on kinematic modeling, and (b) total stress field of western U.S.  
2444 from dynamic modeling of the lithosphere. The velocity field is interpolated from GPS  
2445 velocities and fault-slip rates. The stress field is calculated from a detailed lithosphere density  
2446 structure model that includes the YSRP. Boundary stresses are constrained to match the strain  
2447 tensors from kinematic modeling.

2448  
2449 Fig. 27. Comparison of buoyancy flux estimates for oceanic and continental hotspots plotted as  
2450 a function of plume radii and excess temperature. White circles approximate the uncertainty.  
2451 Radius and excess temperature estimates of Eifel (E), Massif Central (MC), and Yellowstone (Y)  
2452 from tomographic images are used to estimate the buoyancy flux using the method of Ritter  
2453 (2004). The buoyancy fluxes for Iceland (I) and Hawaii (H) is taken from Sleep (1990). The  
2454 buoyancy flux is calculated for a potential mantle temperature of 1300° C. If the mantle beneath  
2455 Yellowstone were 200° C hotter, then it would yield a buoyancy flux 10 times larger.

2456  
2457 Fig. 28. Schematic diagram of the Yellowstone plume progression. (a) Behind-arc plume-head  
2458 phase located beneath the accreted oceanic plate of the Columbia Plateau and behind the  
2459 descending Juan de Fuca plate, and (b) sheared and tilted plume head entrained in mantle flow.  
2460 Note the depleted upper-mantle residuum body above the plume interacting with a continental  
2461 lithosphere.

2462  
2463 Fig. 29. Track of the Yellowstone plume tail originating at ~650 km depth and 150 km west of  
2464 Yellowstone. At its origin at 15 Ma, the plume had a vertical ascending path beneath the  
2465 Columbia Plateau to the west of the Sr 0.706 boundary and coincident with the implied outline of  
2466 the plume head by Camp (2004). The plume was tilted 60° to the SE by mantle flow, so that the  
2467 plume base (red circles) was offset from surface silicic volcanic centers (yellow circles). Ages of  
2468 silicic and basaltic volcanic centers are also shown on the map. The initial plume head spread  
2469 out beneath the thin oceanic lithosphere to the west of the Sr 0.706, and the extend of spreading  
2470 at 15 Ma and present day is shown by shaded areas.

2471  
2472 **Supplementary figures**

2473 Fig. S1 – Time series for station WLWY (White Lake, WY) in Yellowstone National Park.

2474  
2475 Fig. S2 – Resolution test for a -1% reduction in the upper mantle.

2476  
2477 Fig. S3 – Resolution test for a -3% reduction in the upper mantle.

2478  
2479 Fig. S4 – Resolution test for a -1% reduction in the transition zone.

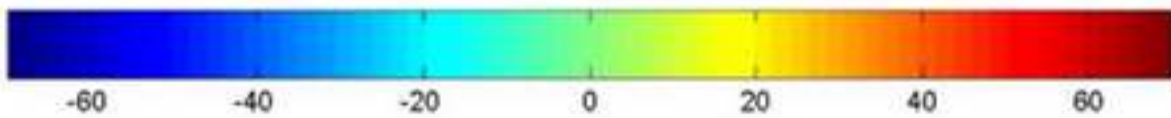
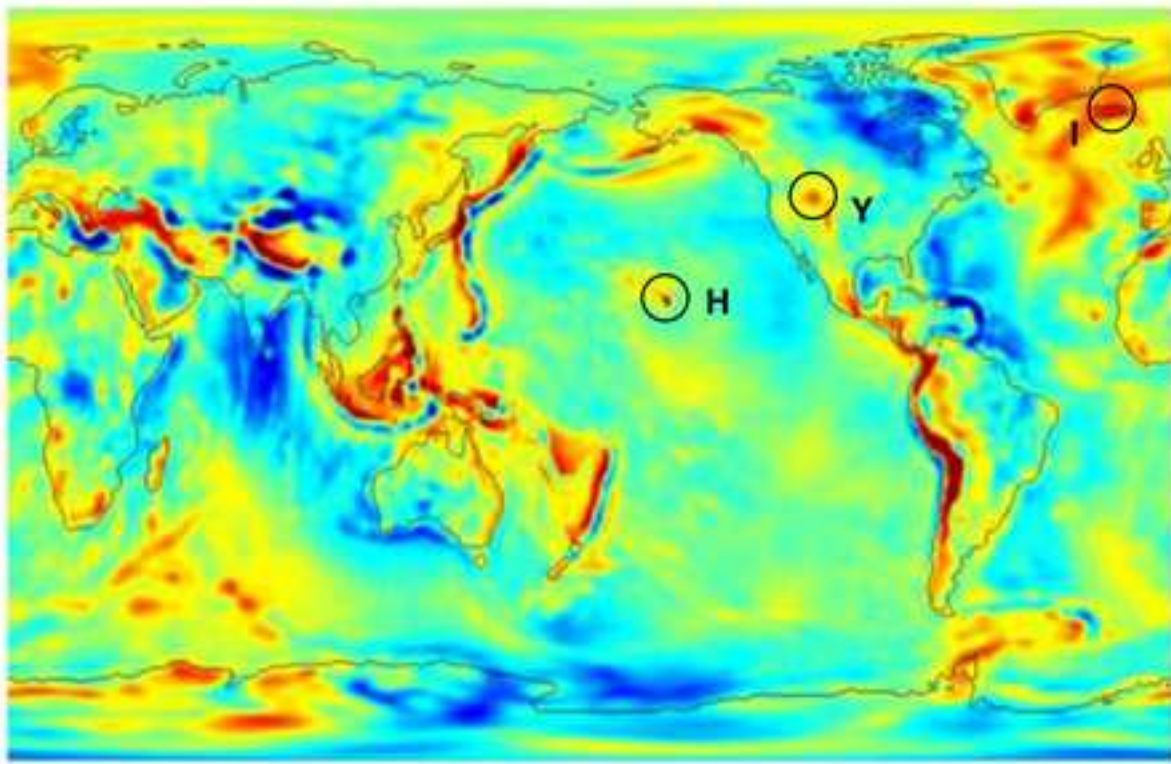
2480  
2481 Fig. S5 – Resolution test for a -1% reduction in the sub transition zone.

2482  
2483 Fig. S6 – The parameterization scheme of the optimized tomography.

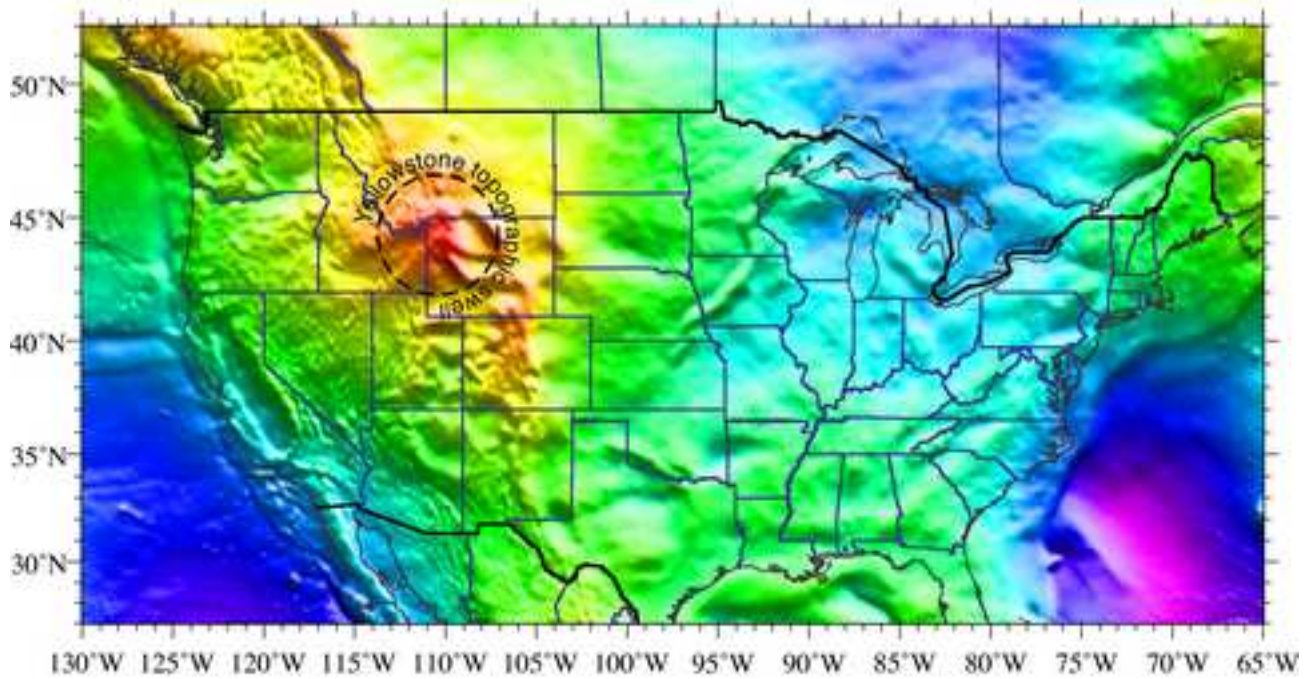
Figure 01

[Click here to download high resolution image](#)

a)



b)



Geoid Height above NAD83 Ellipsoid (m)

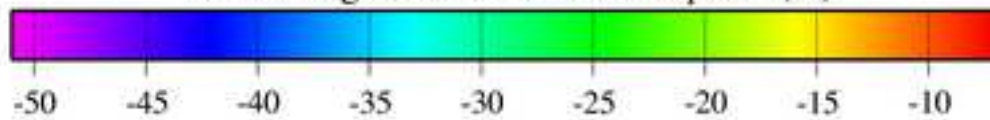




Figure 02  
[Click here to download high resolution image](#)

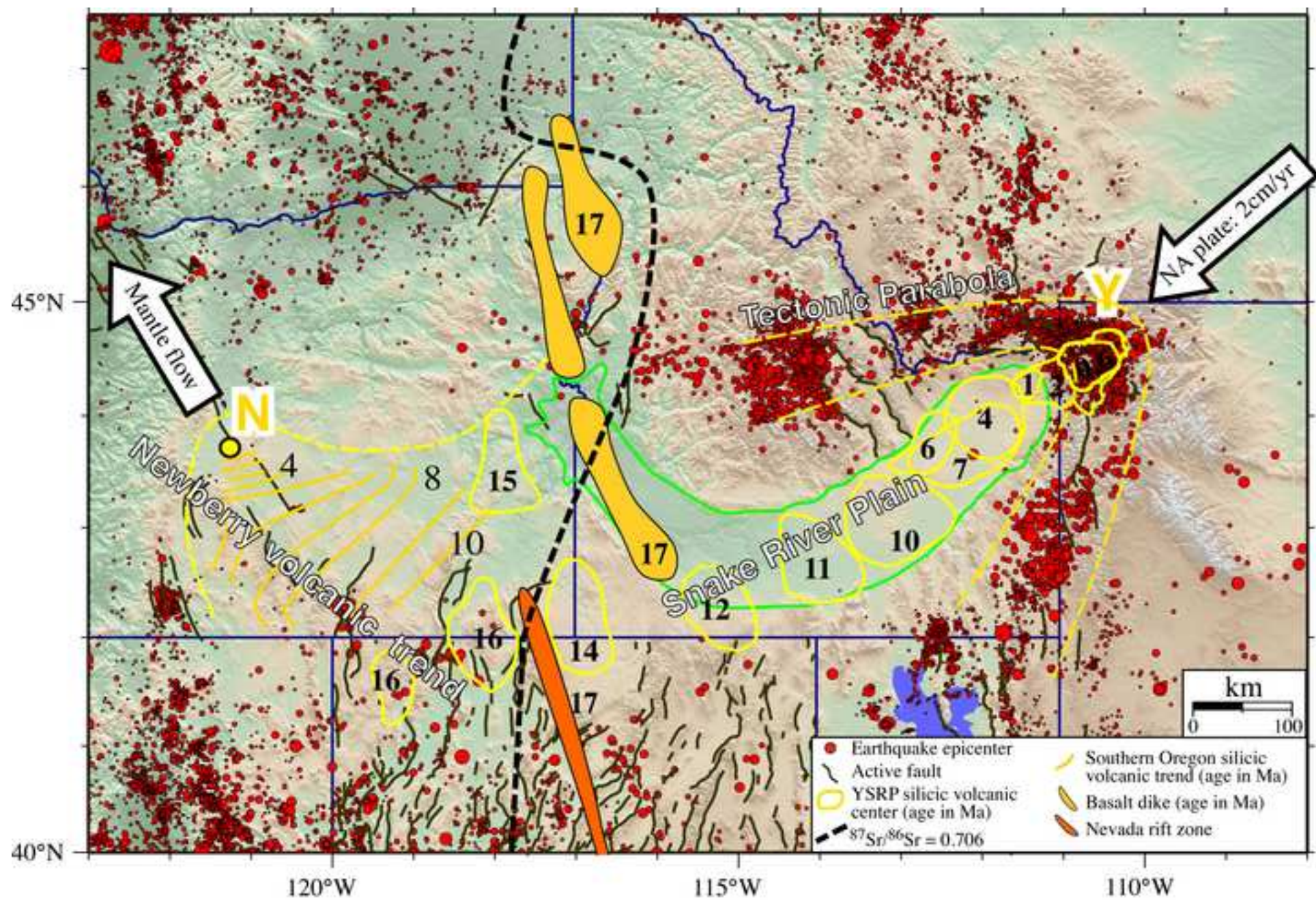




Figure 03  
[Click here to download high resolution image](#)

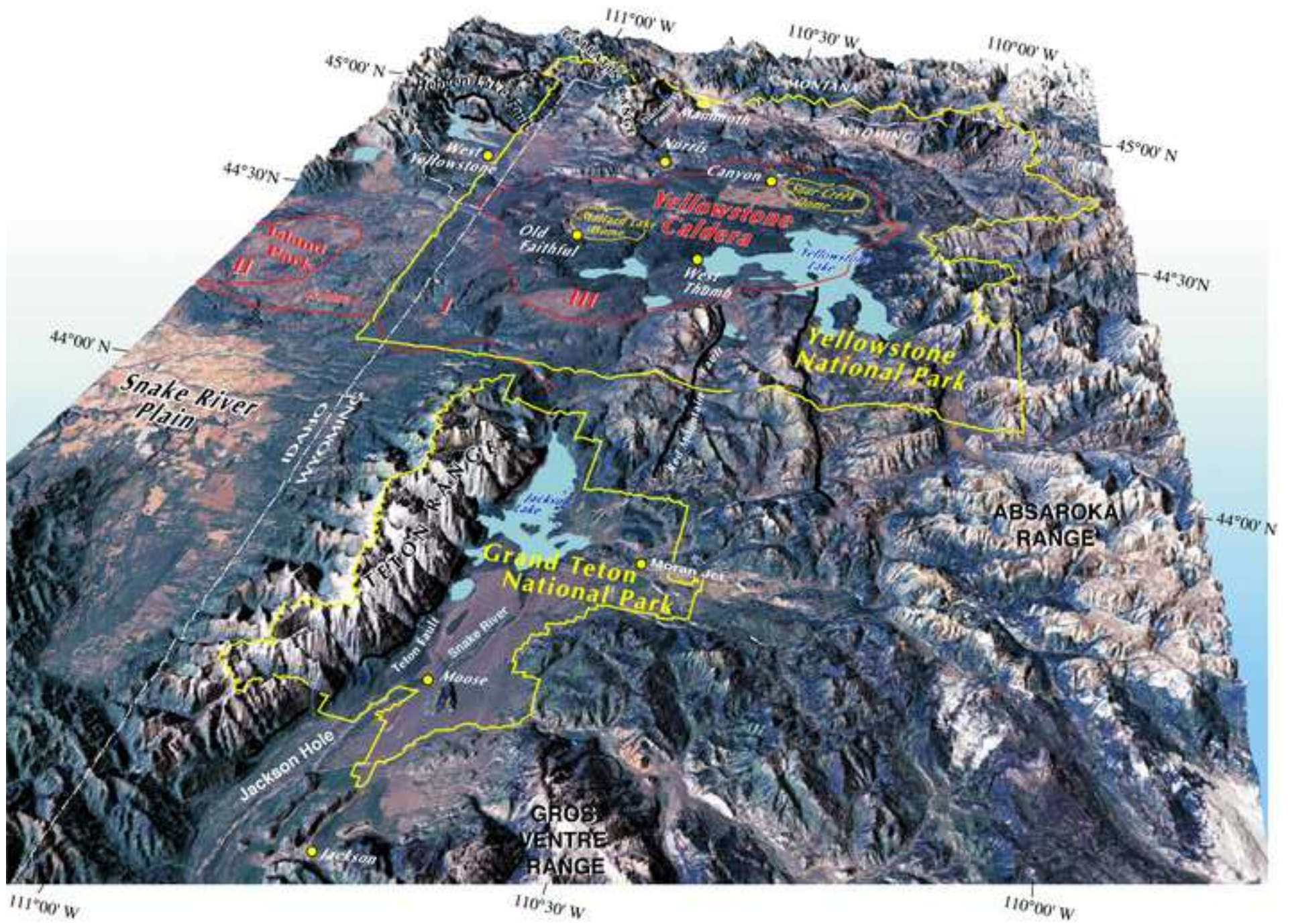


Figure 04

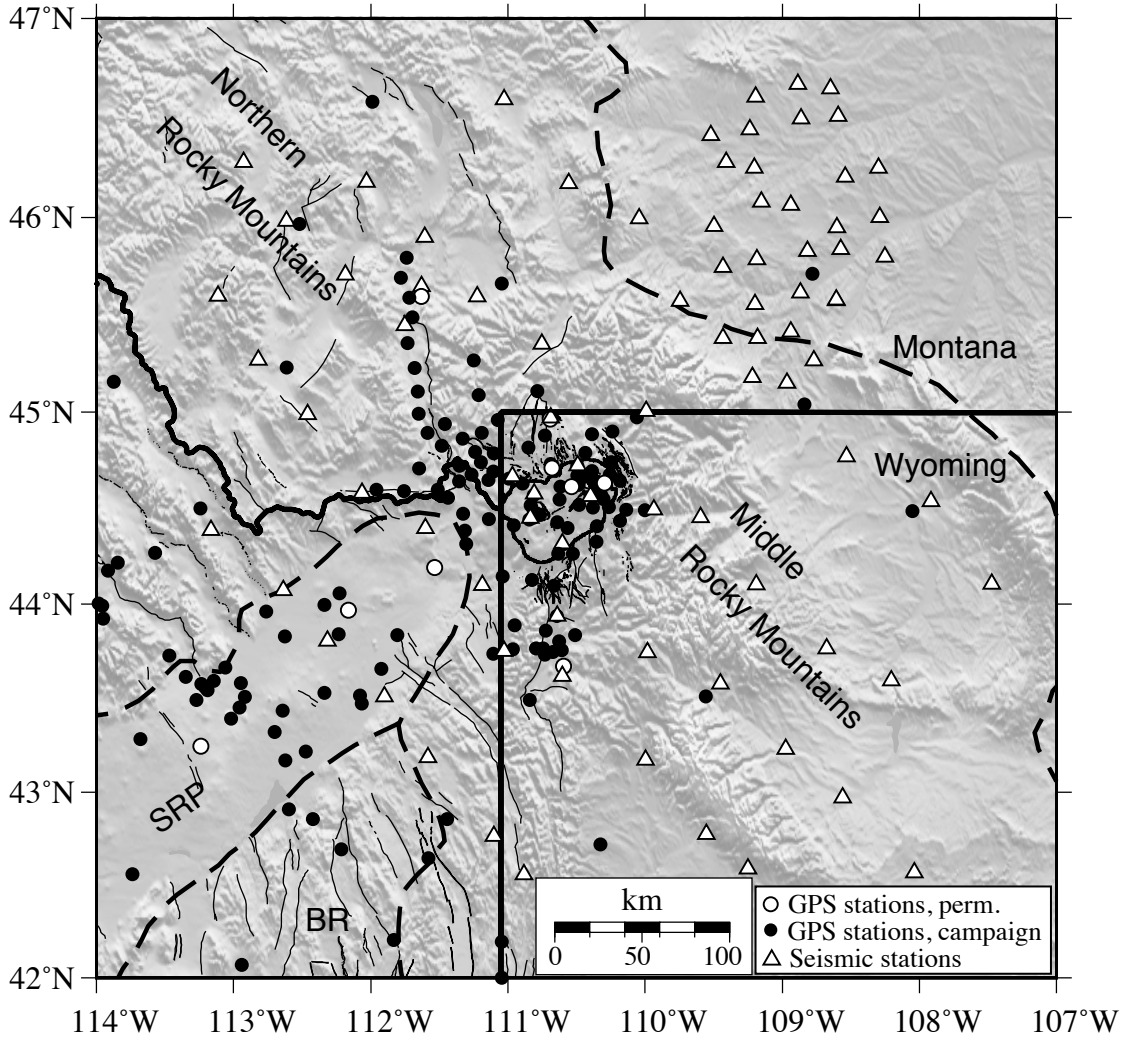




Figure 05  
[Click here to download high resolution image](#)

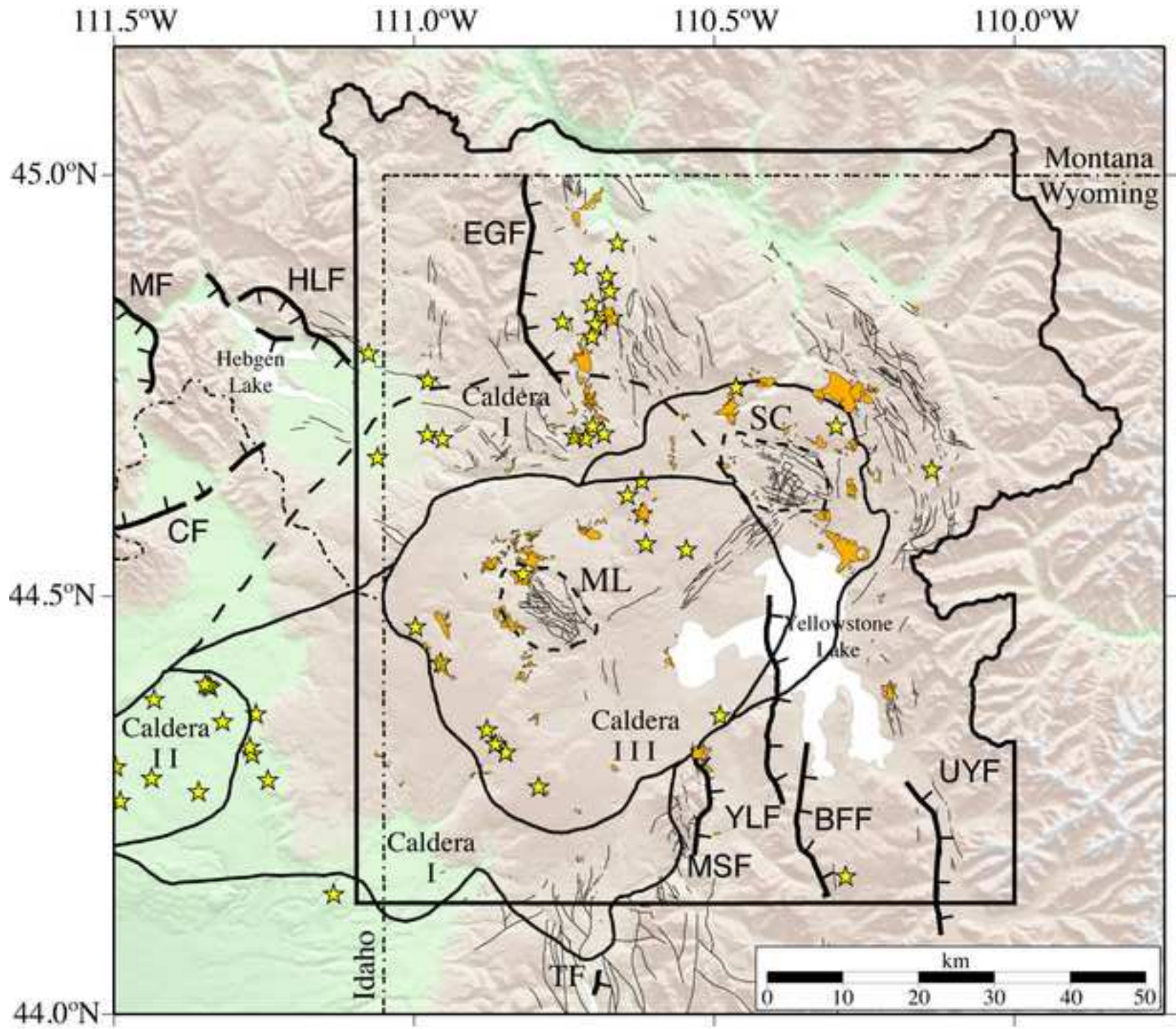
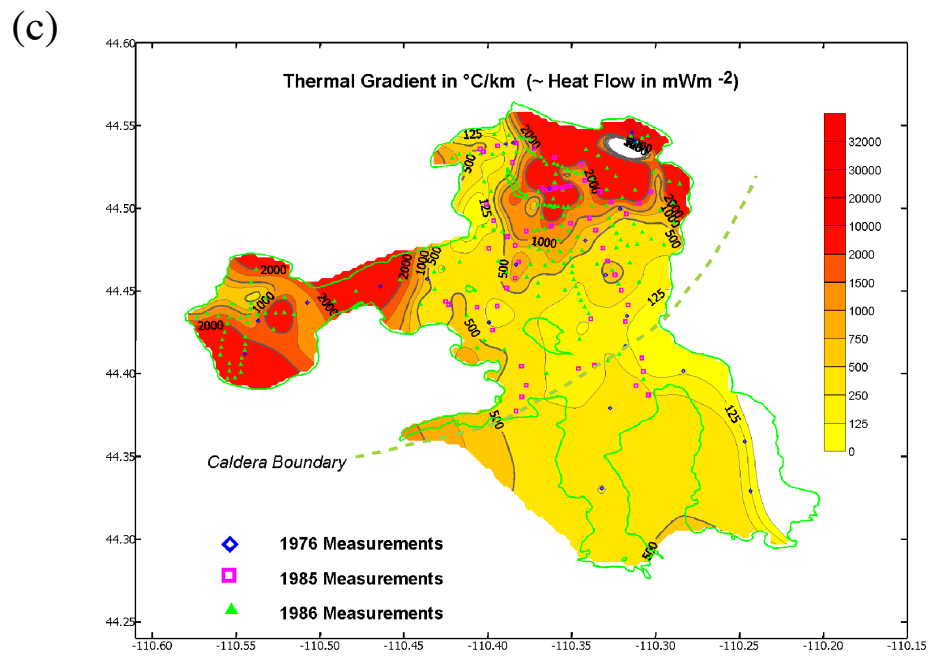
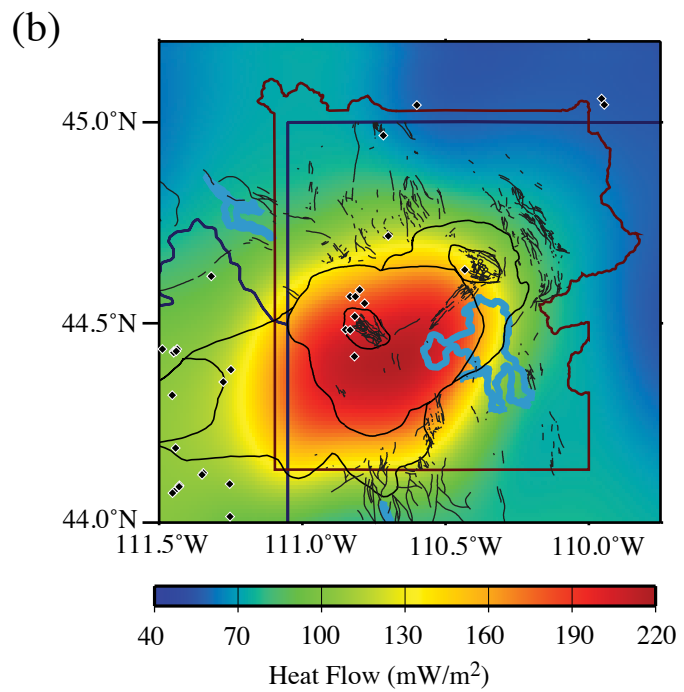
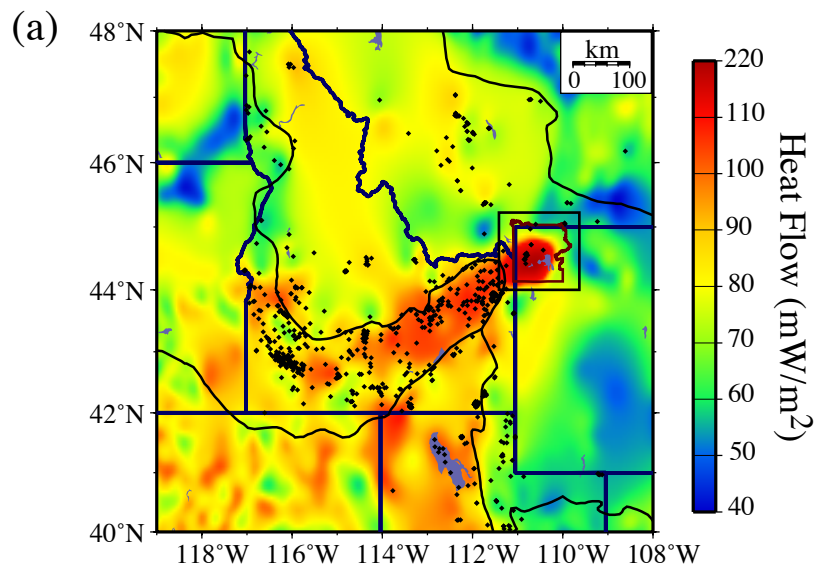


Figure 06



a)



b)





Figure 08  
[Click here to download high resolution image](#)

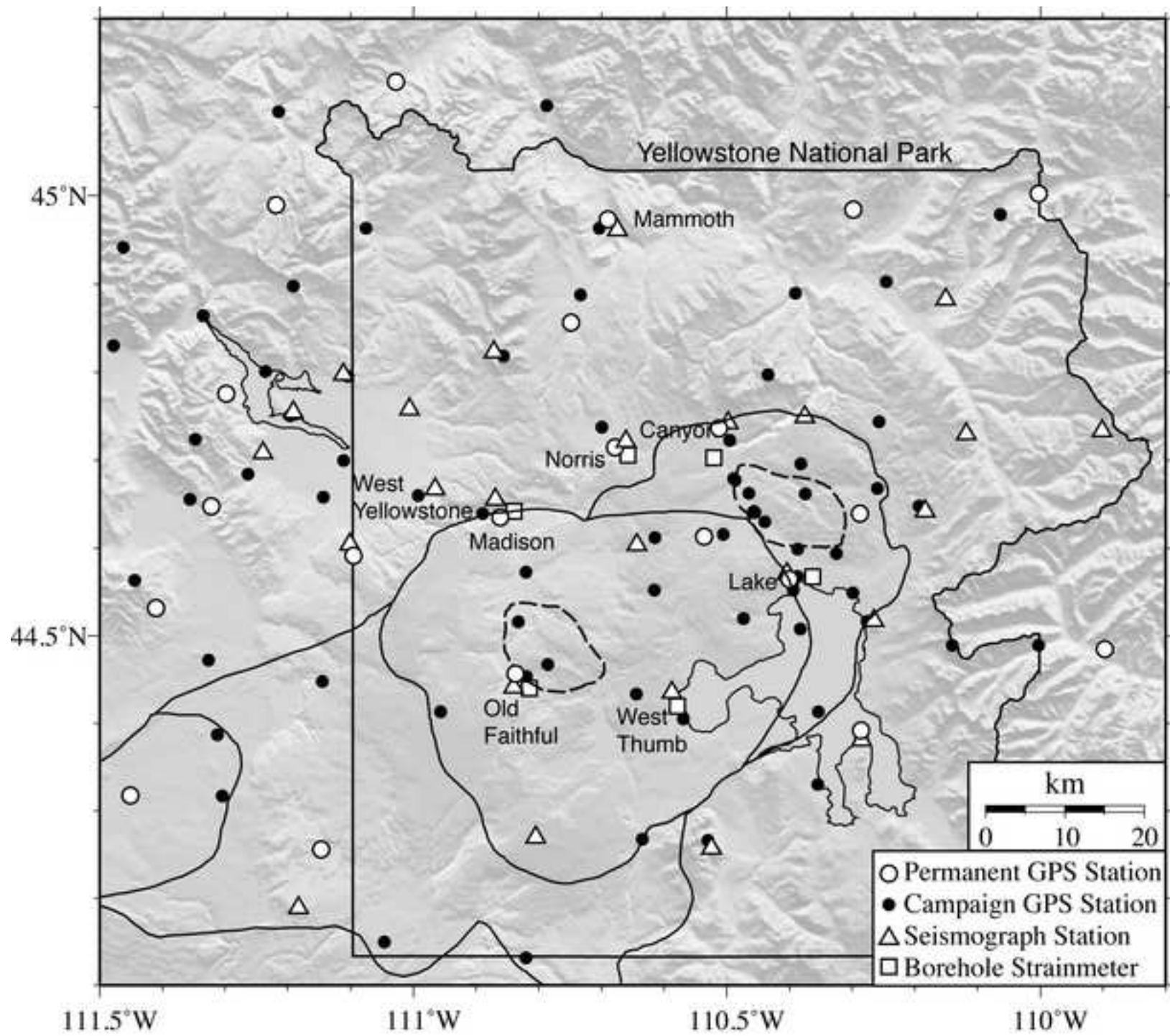




Figure 09  
[Click here to download high resolution image](#)

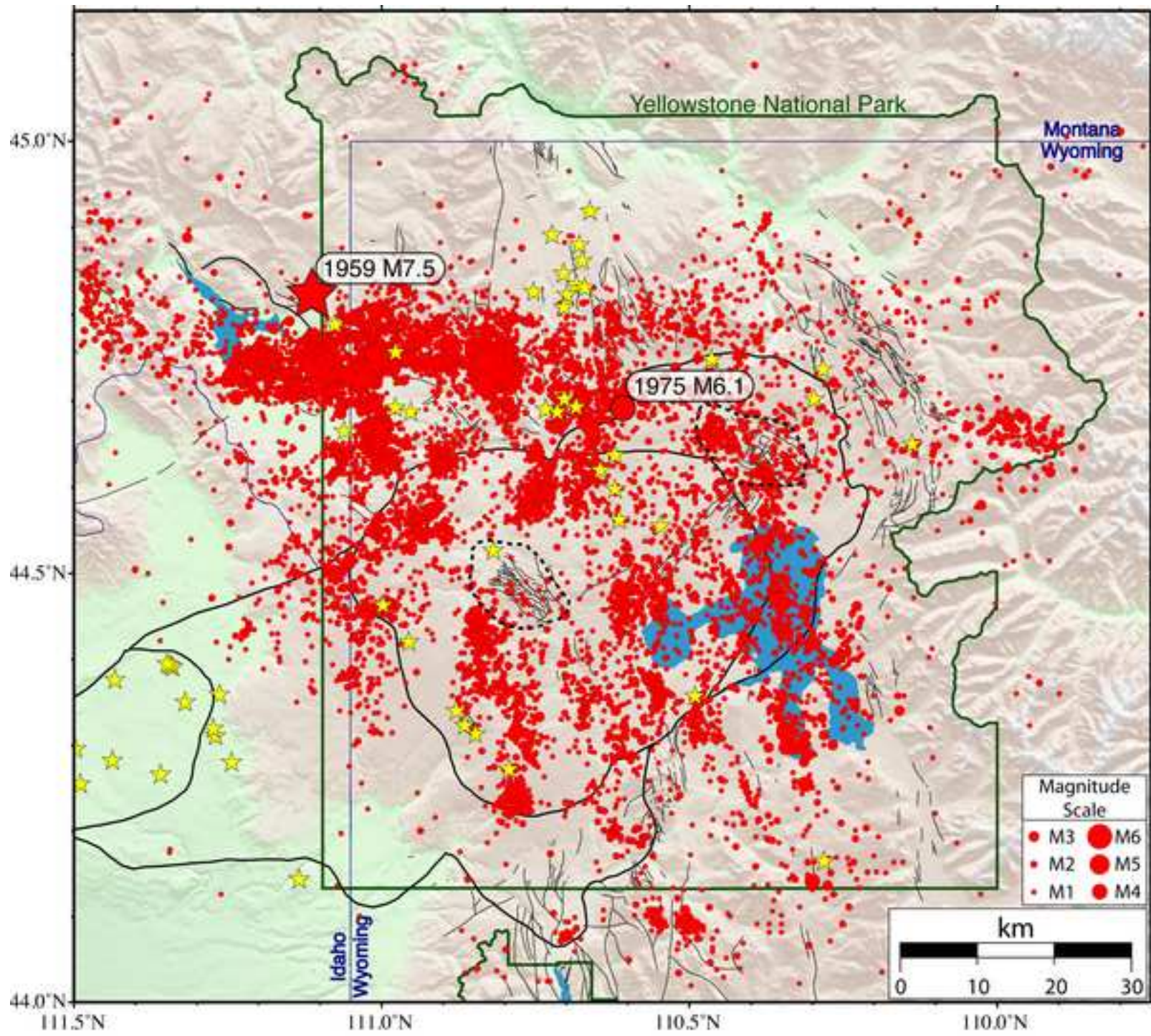




Figure 10  
[Click here to download high resolution image](#)

a)



b)



Figure 11

[Click here to download high resolution image](#)

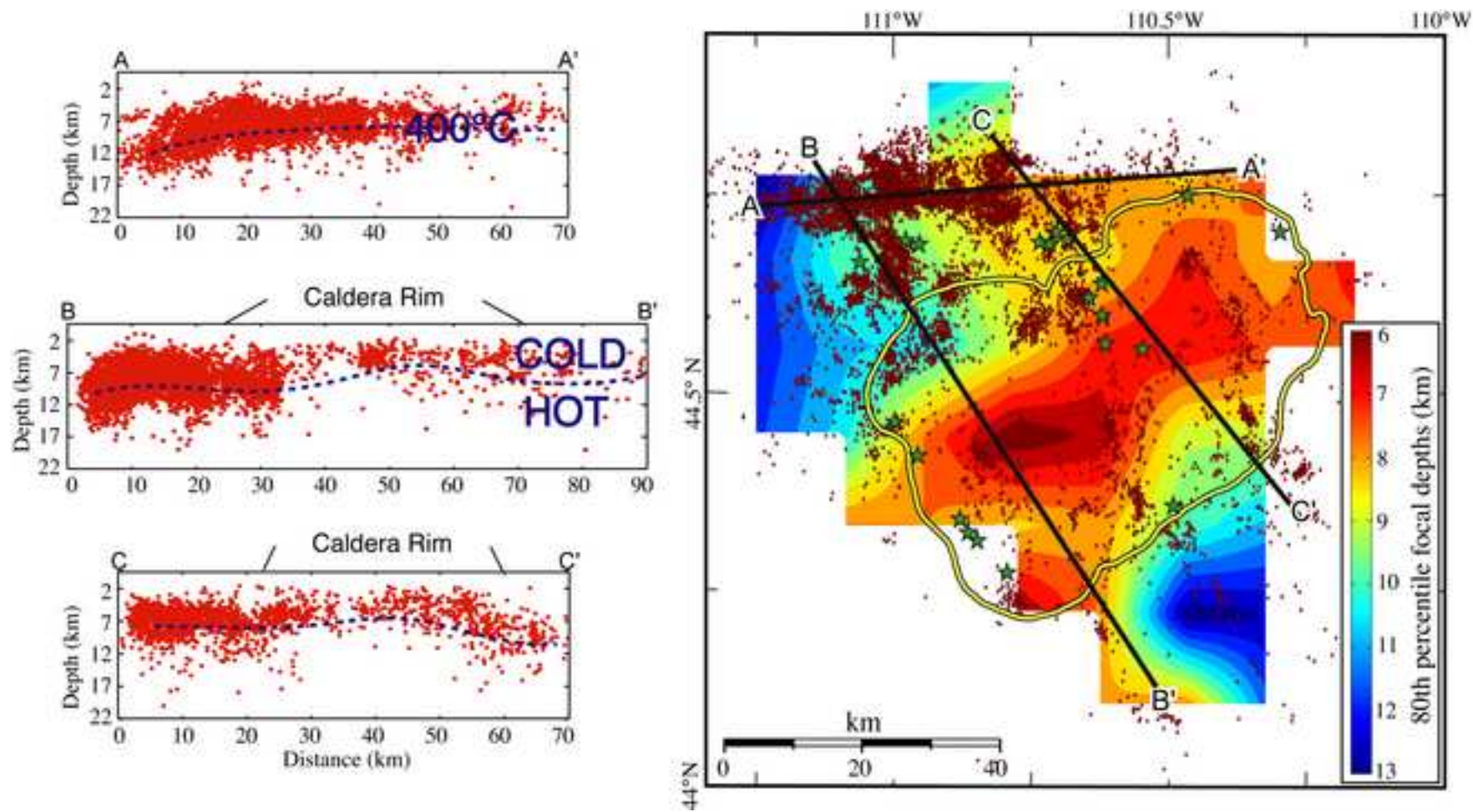
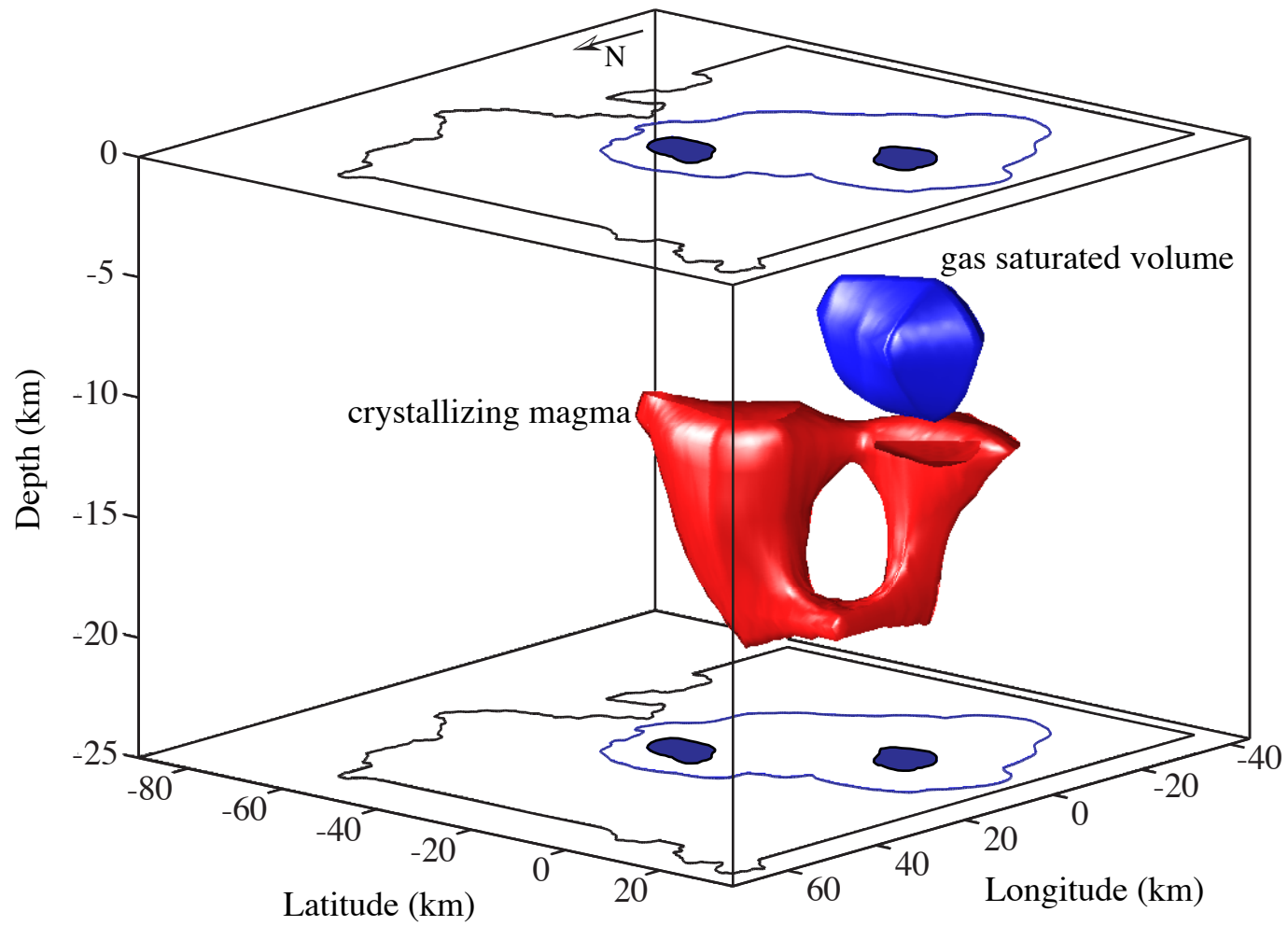


Figure 12



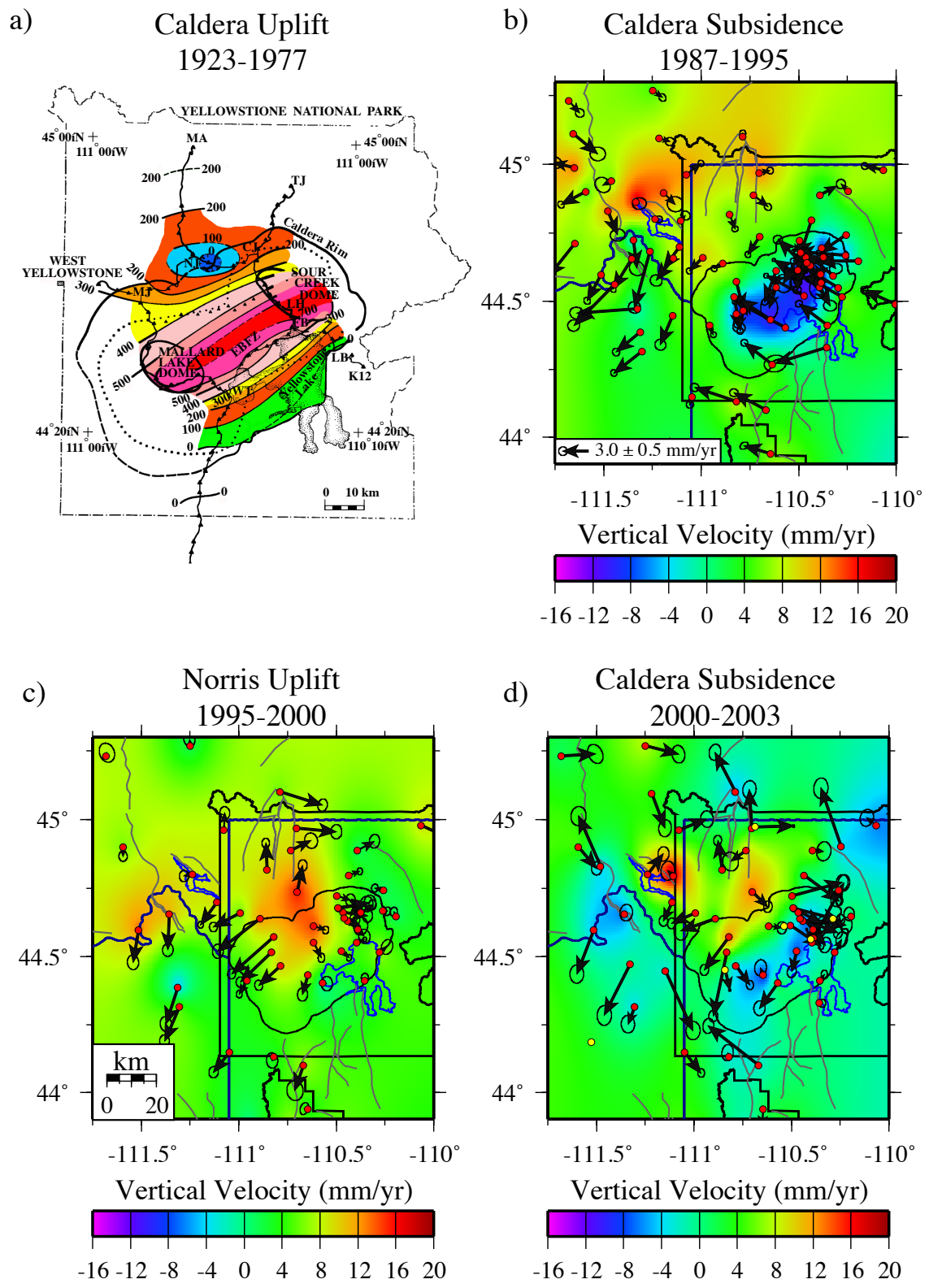




Figure 14  
[Click here to download high resolution image](#)

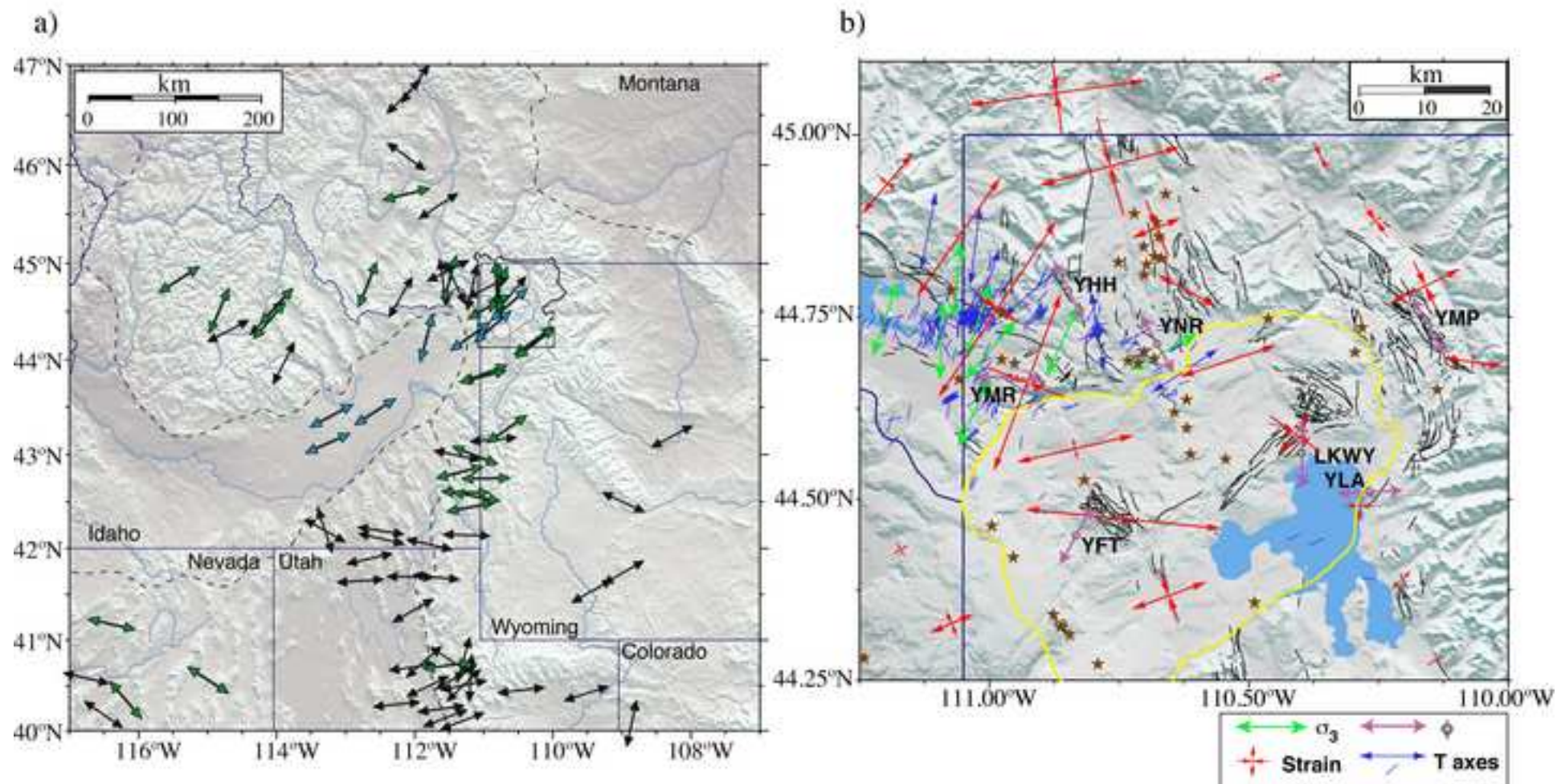


Figure 15

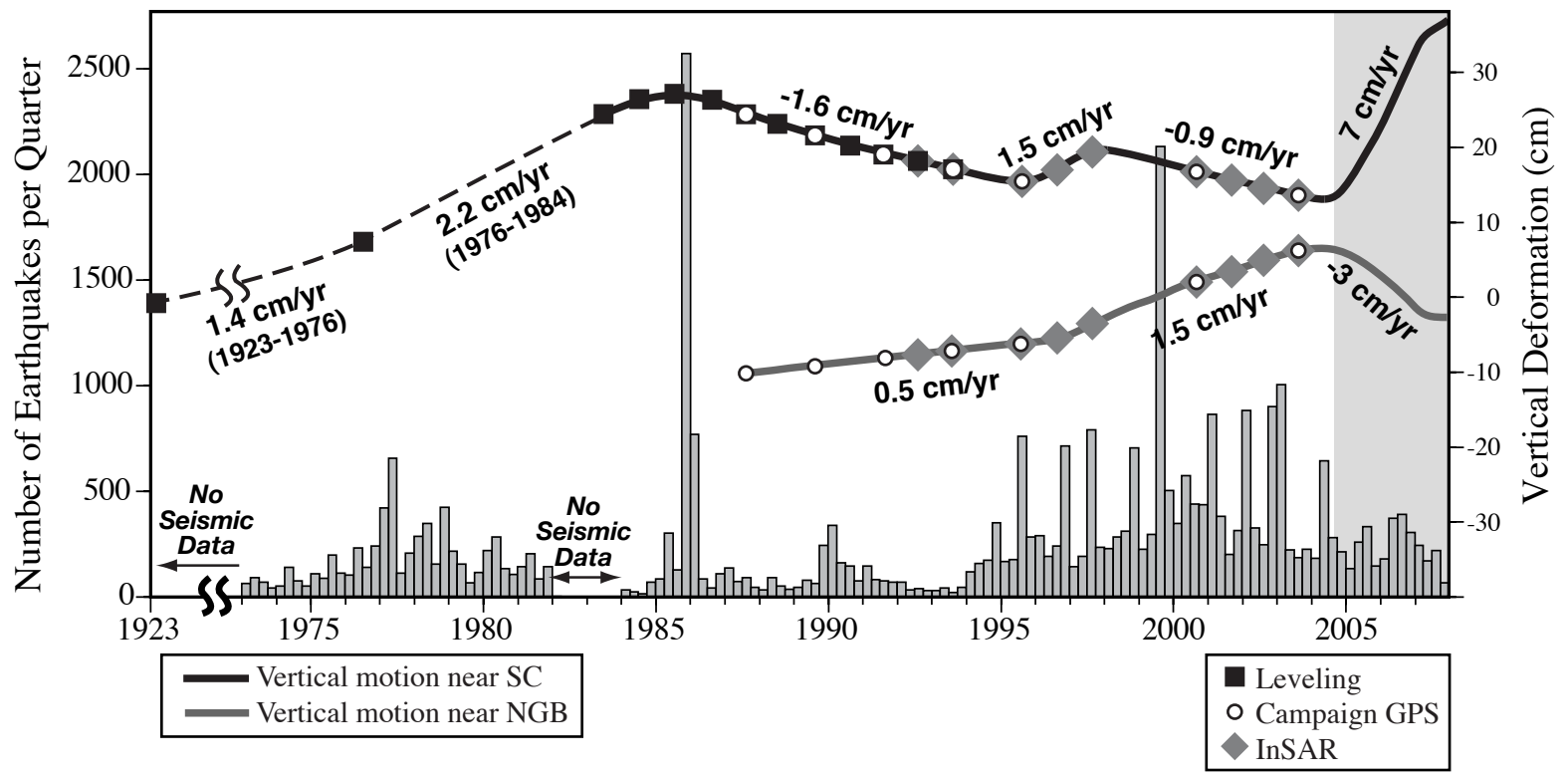




Figure 16  
[Click here to download high resolution image](#)

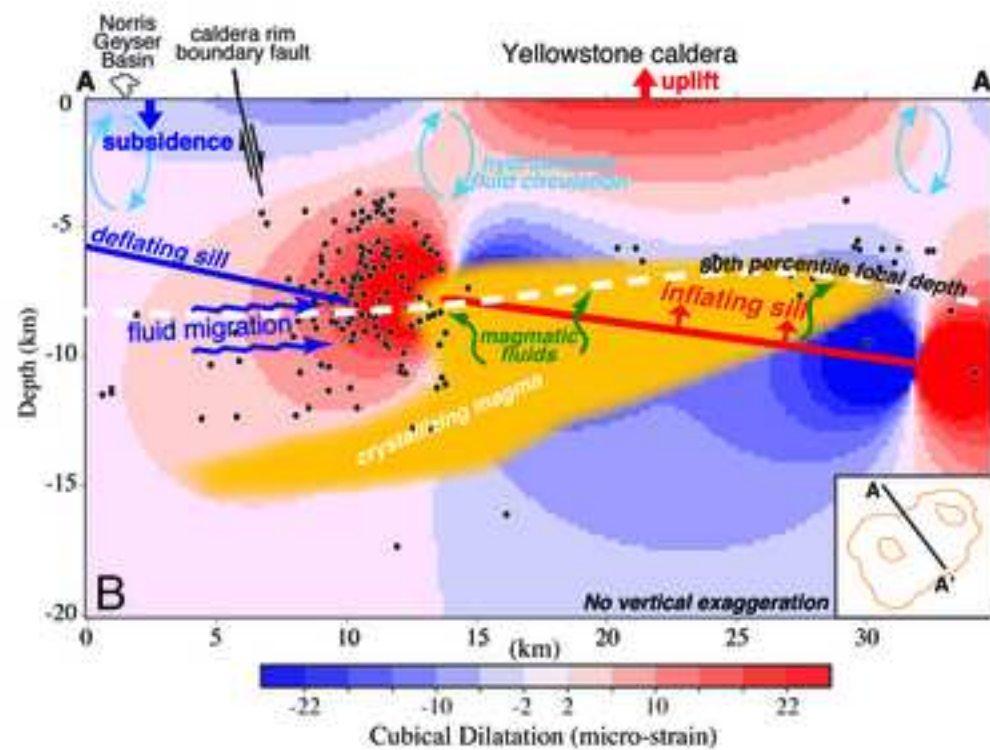
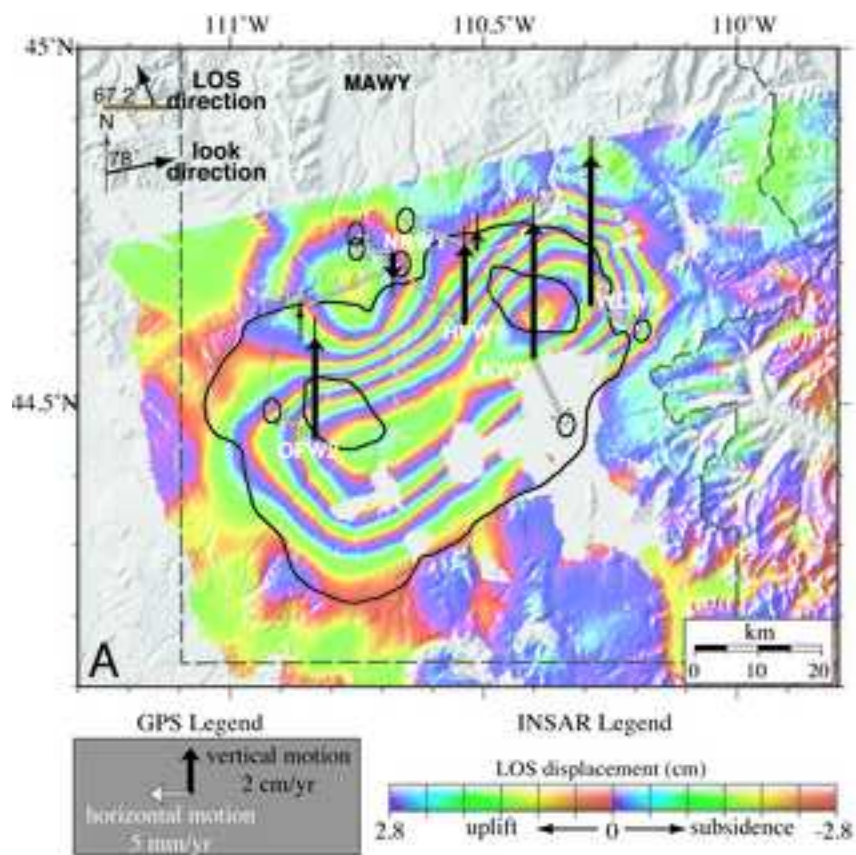


Figure 17

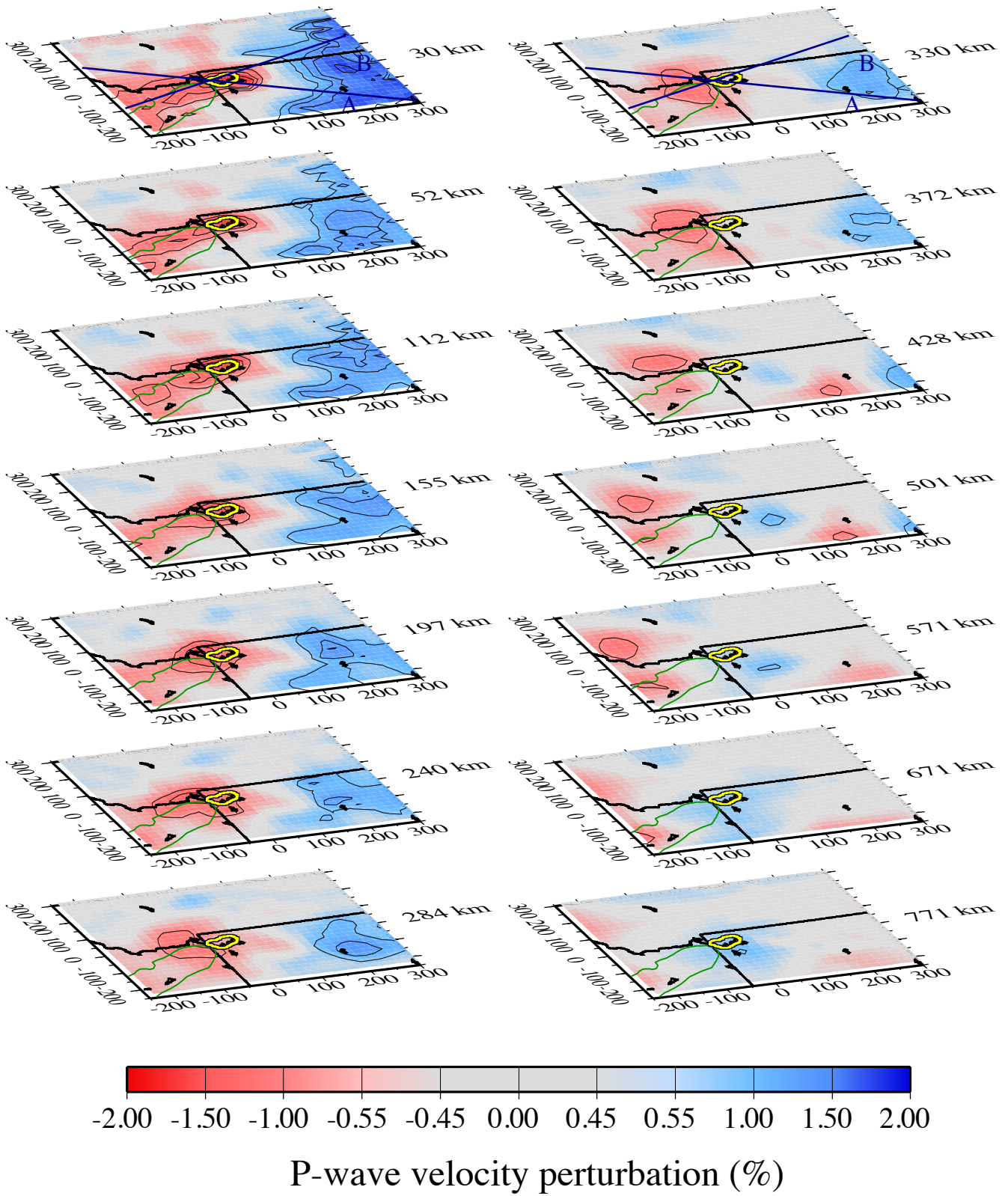


Figure 18

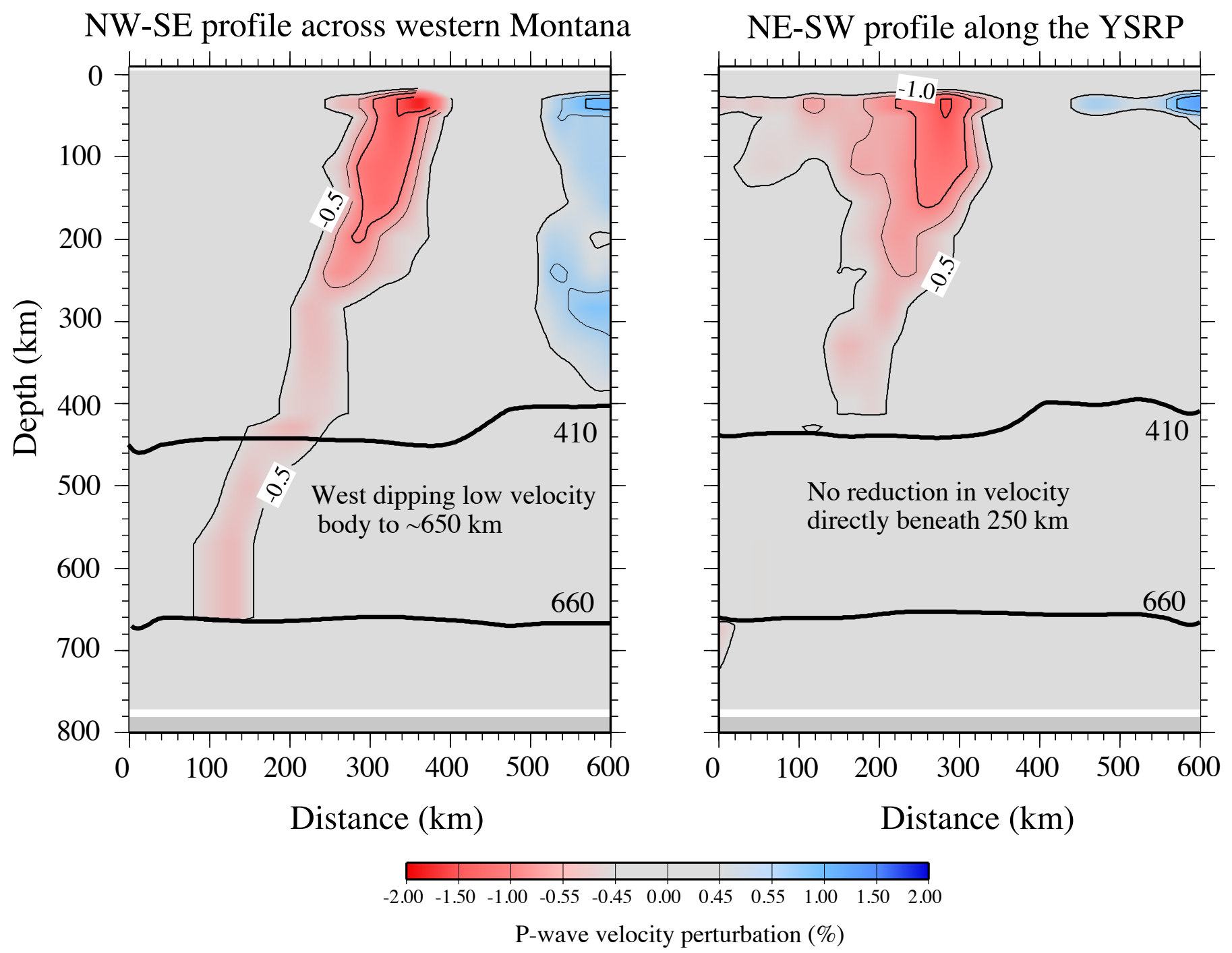


Figure 19

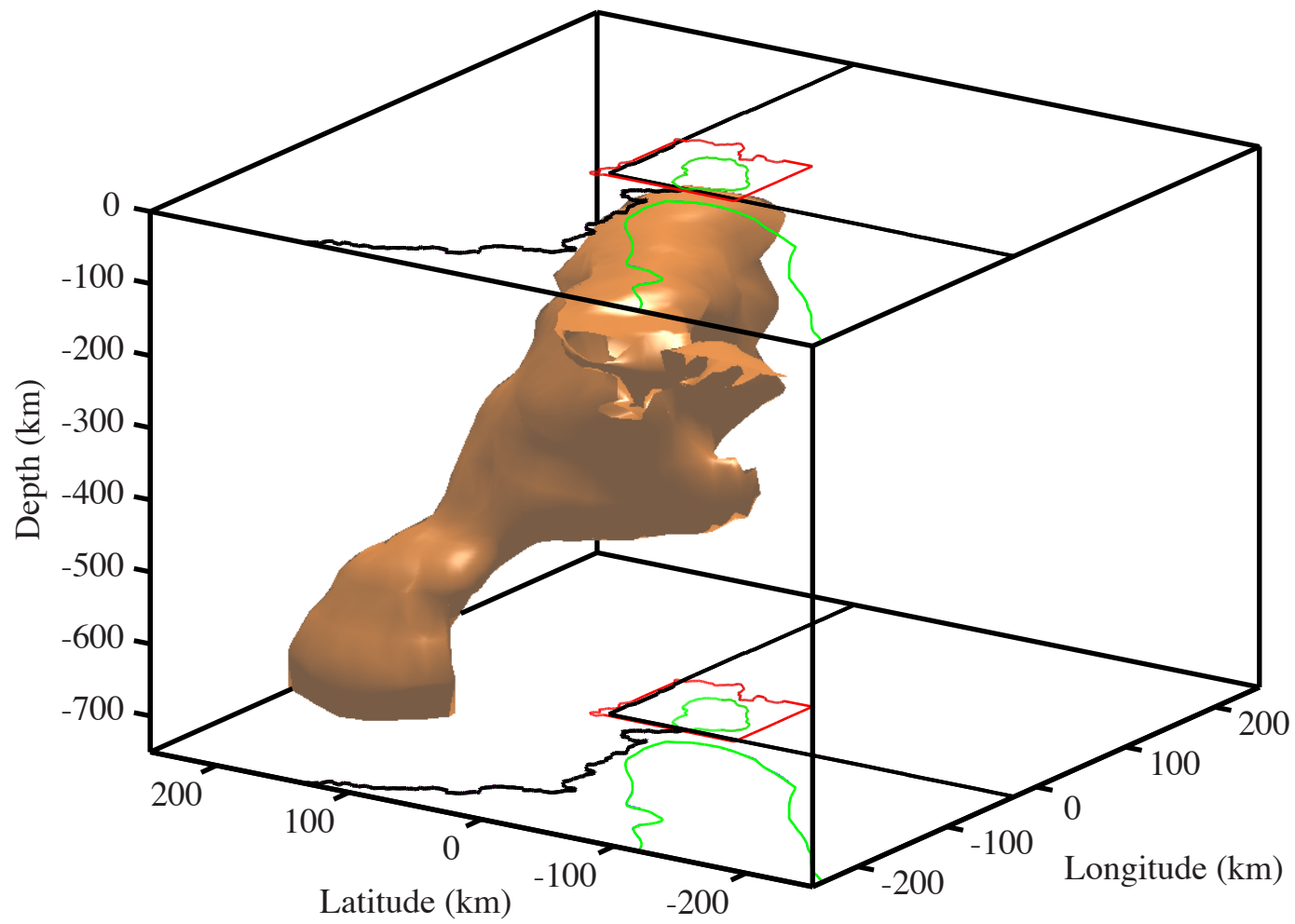


Figure 20

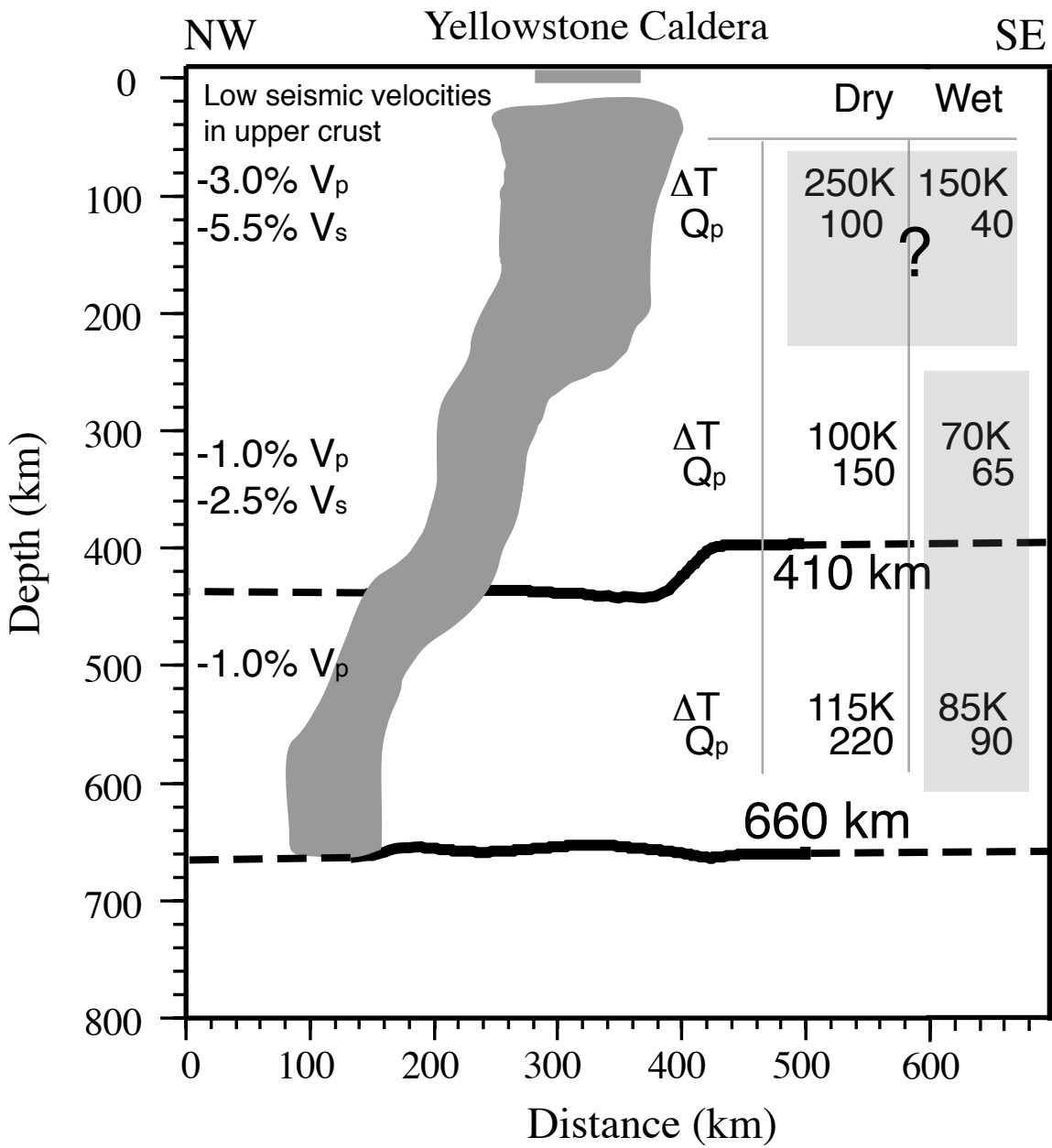
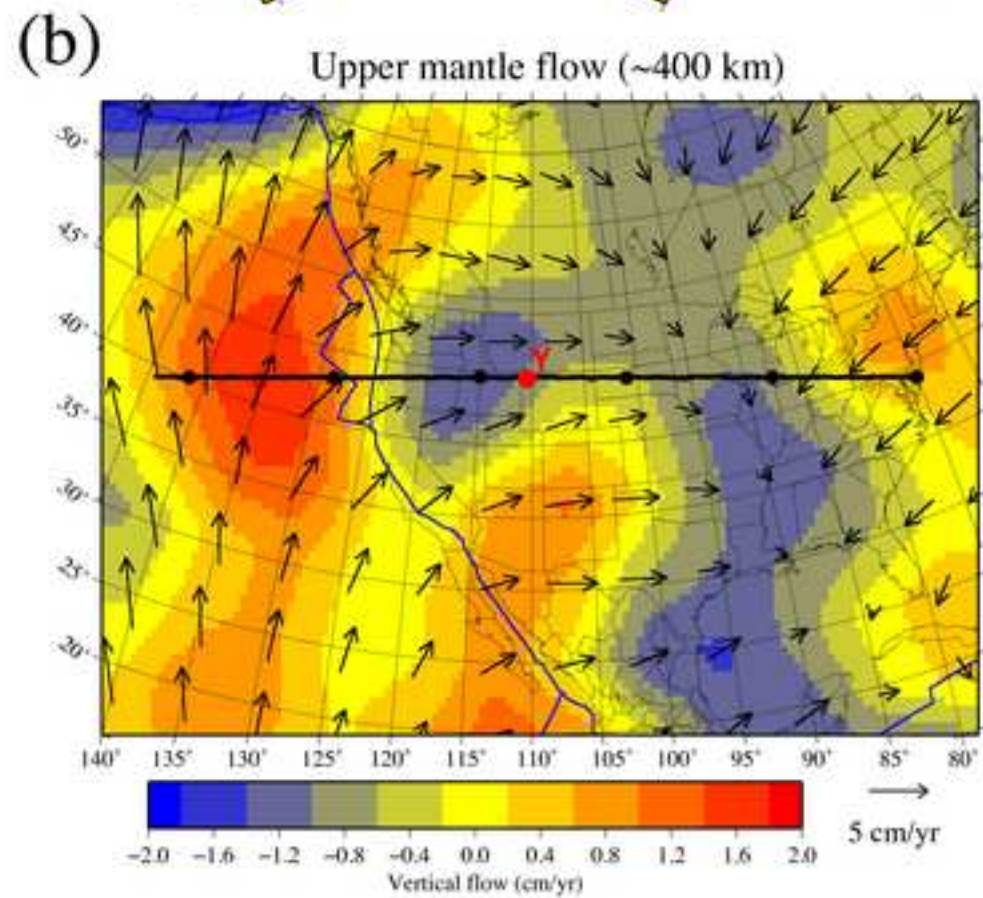
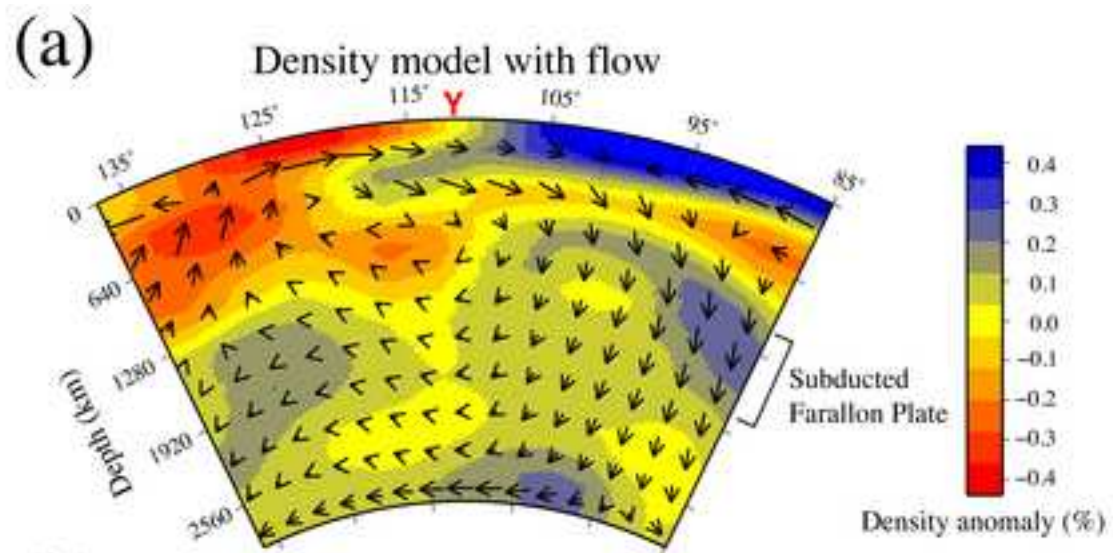


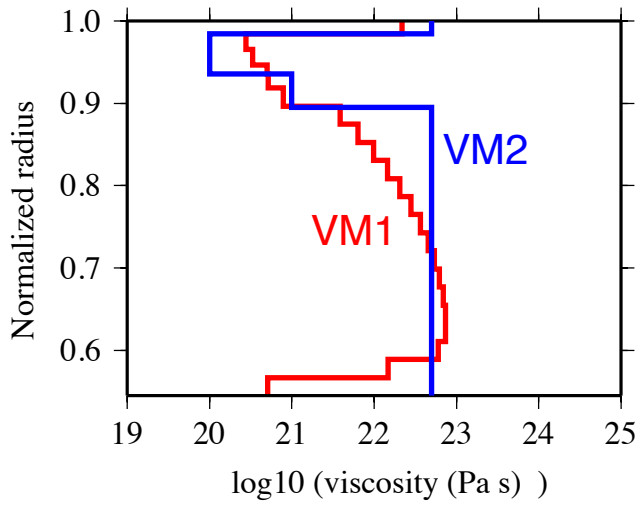


Figure 21

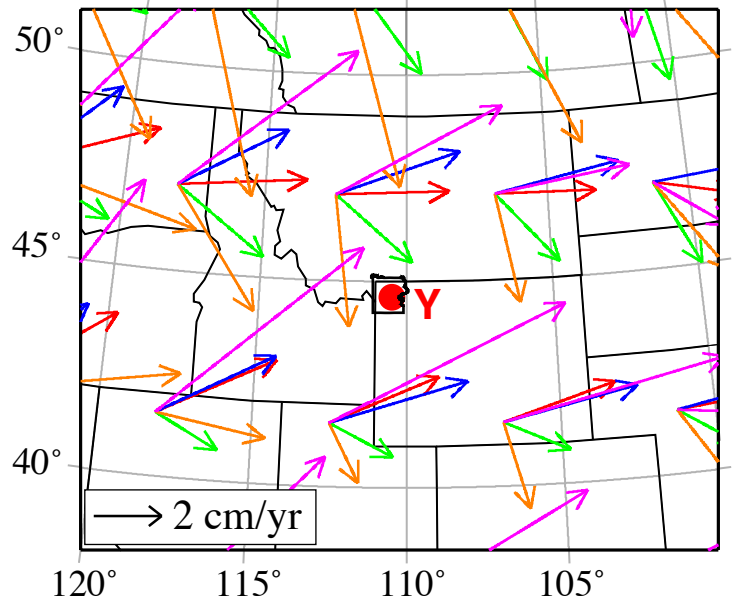
[Click here to download high resolution image](#)



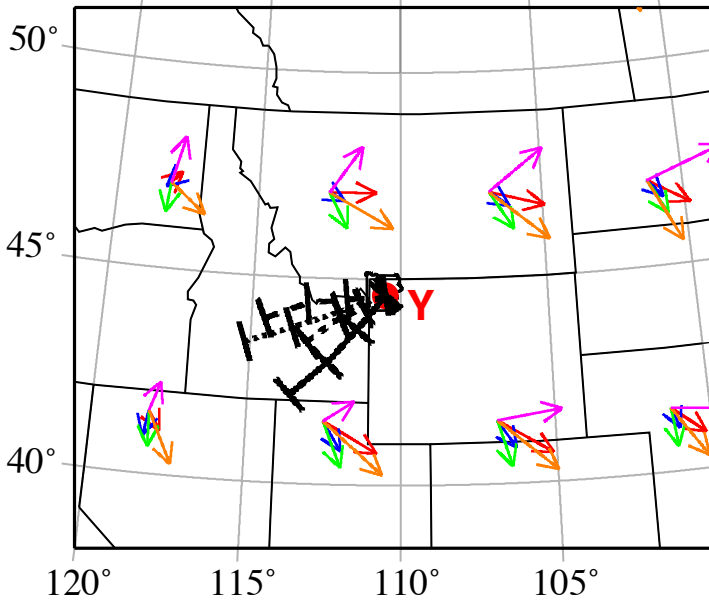
Mantle Viscosity Structure



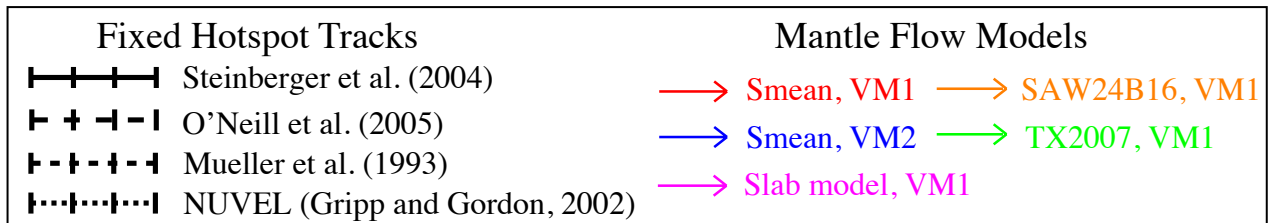
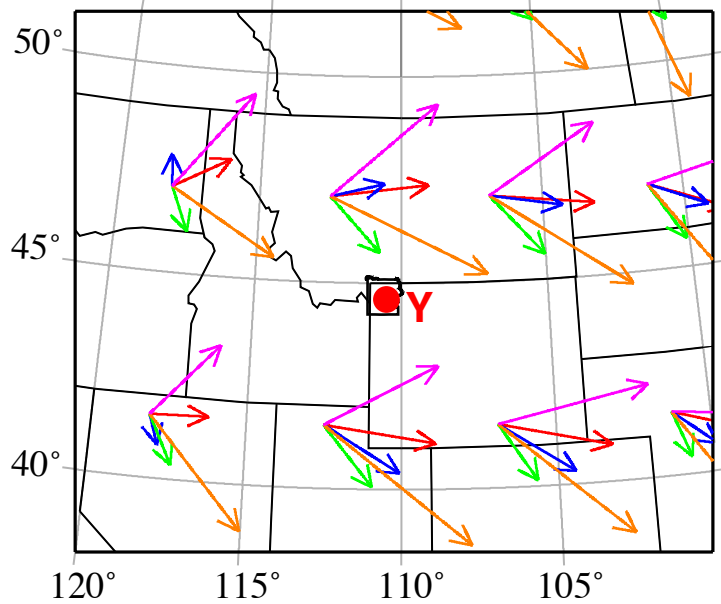
Upper mantle flow



Flow at ~670 km depth

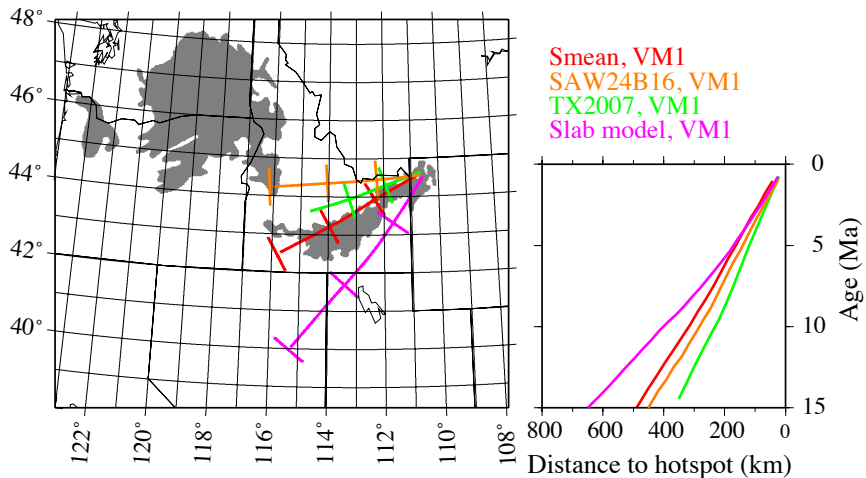


Transition zone flow



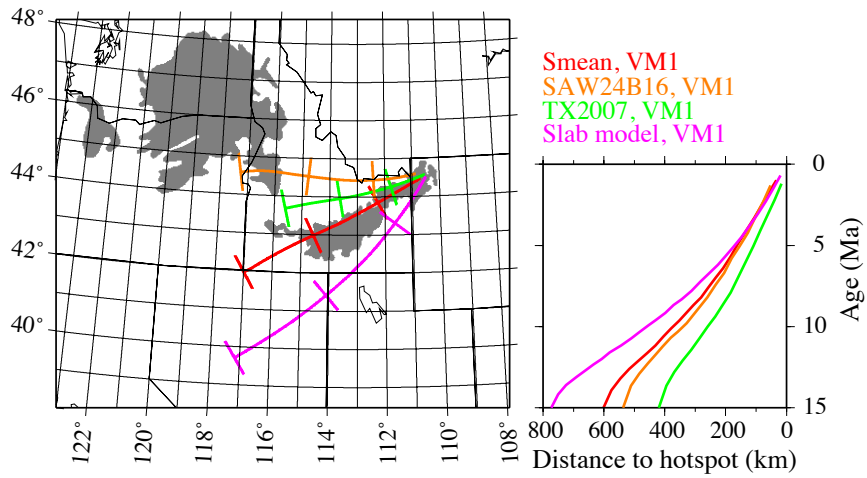
(a) Plume from 660 km depth

Hotspot track (0-15 Ma) for Steinberger et al. (2004) absolute plate motion

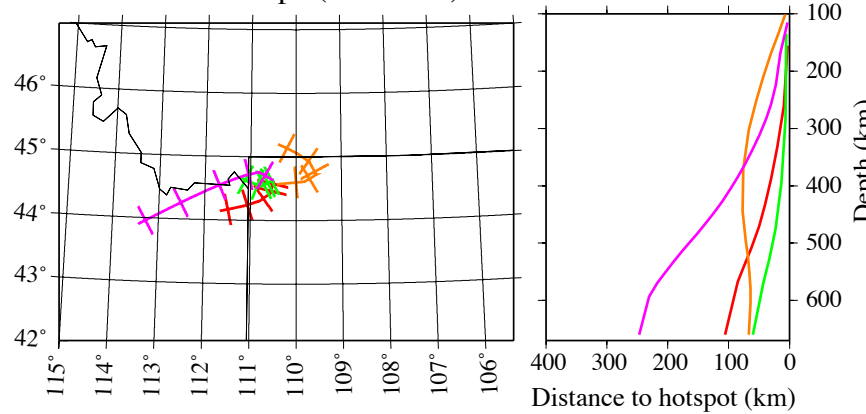


(b) Initially vertical plume from lowermost mantle

Hotspot track (0-15 Ma) for Steinberger et al. (2004) absolute plate motion



Conduit shape (0-660 km)



Conduit shape (0-660 km)

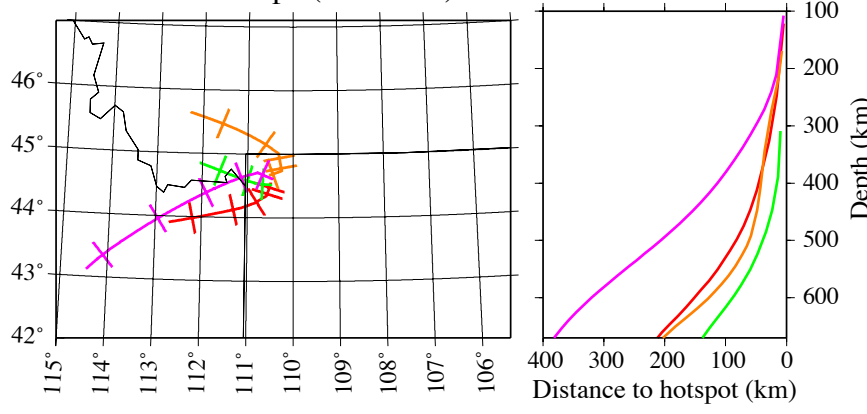




Figure 24

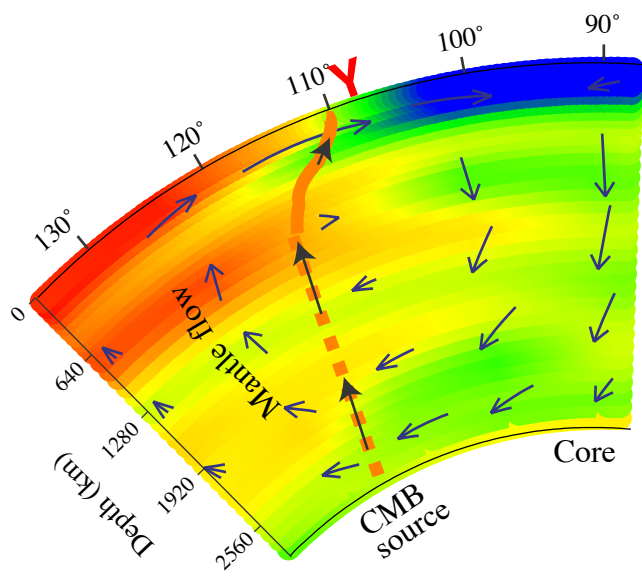


Figure 25

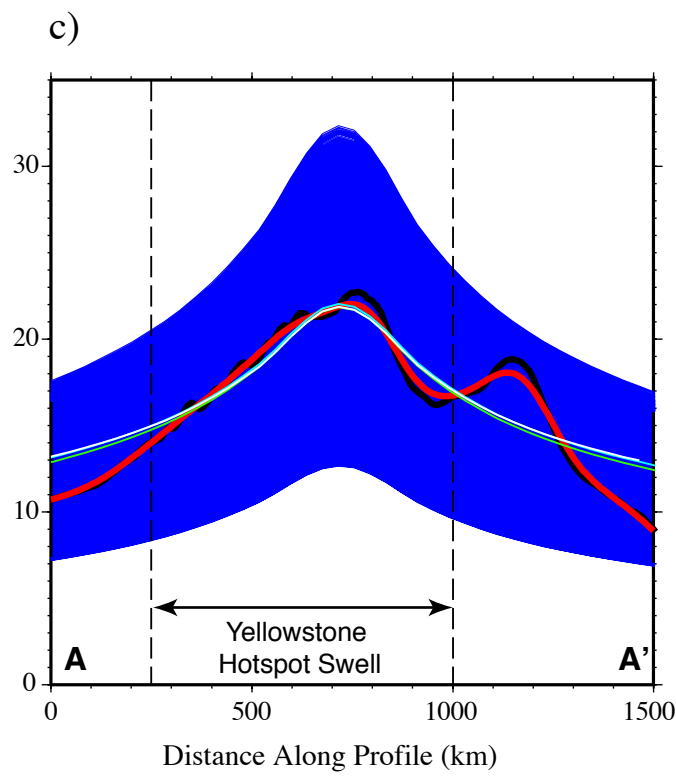
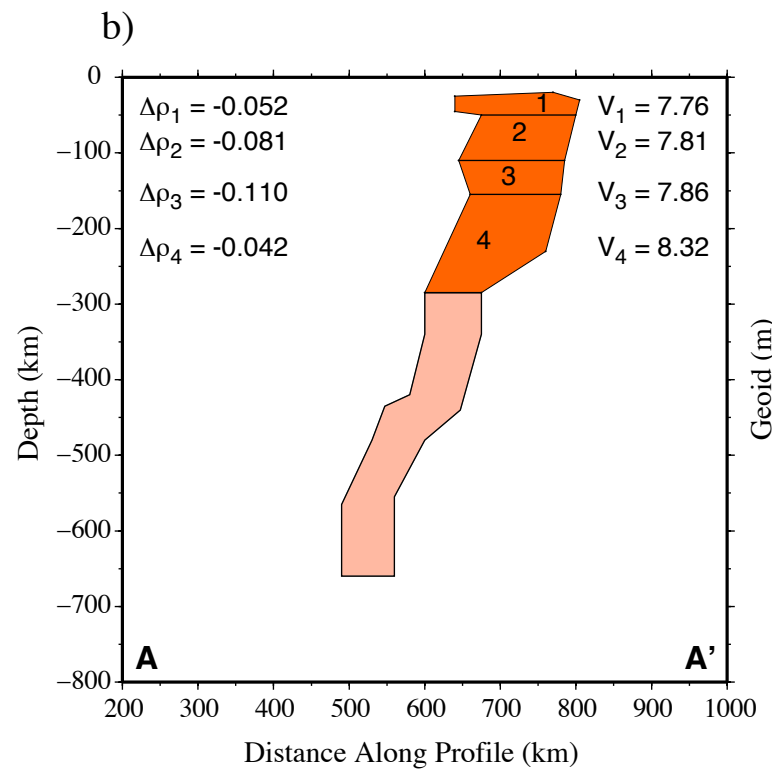
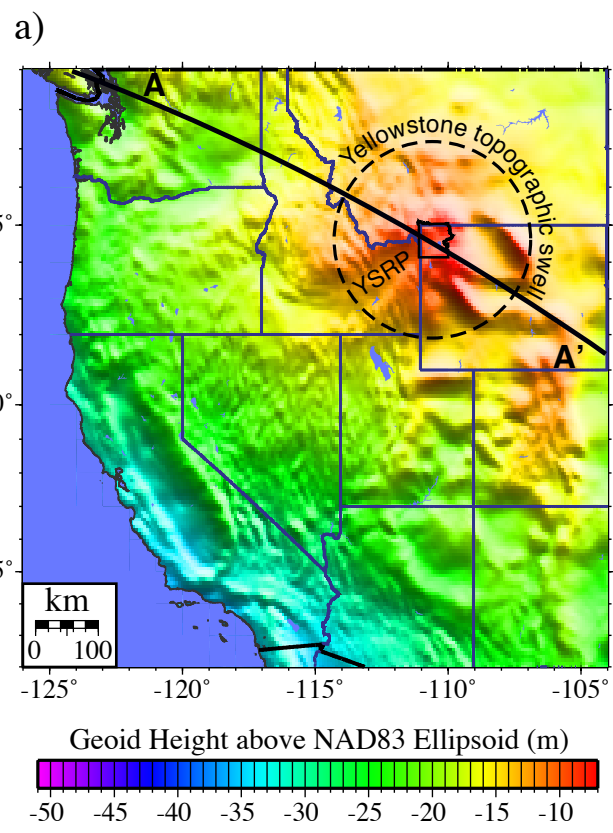


Figure 26  
[Click here to download high resolution image](#)

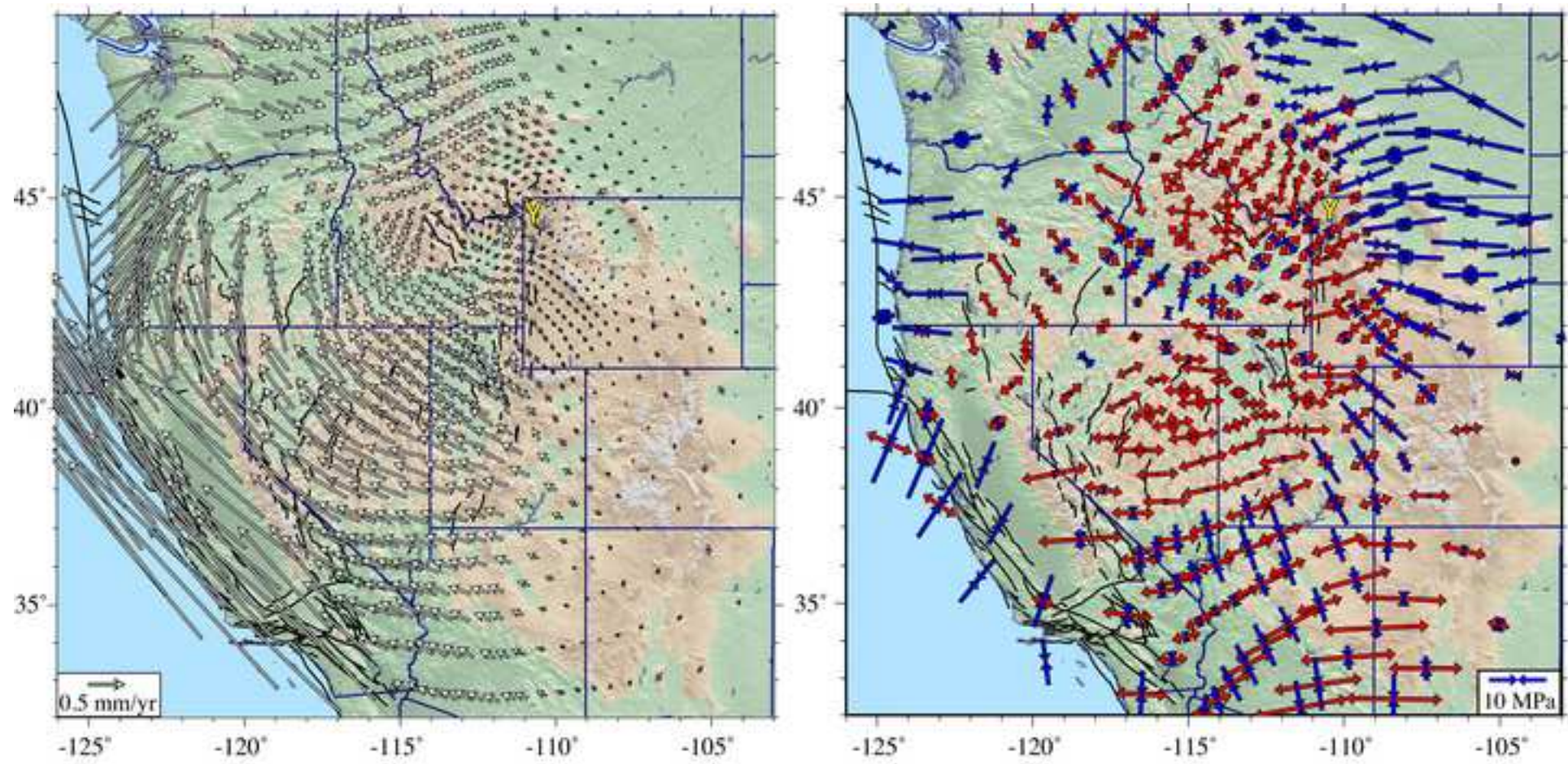


Figure 27

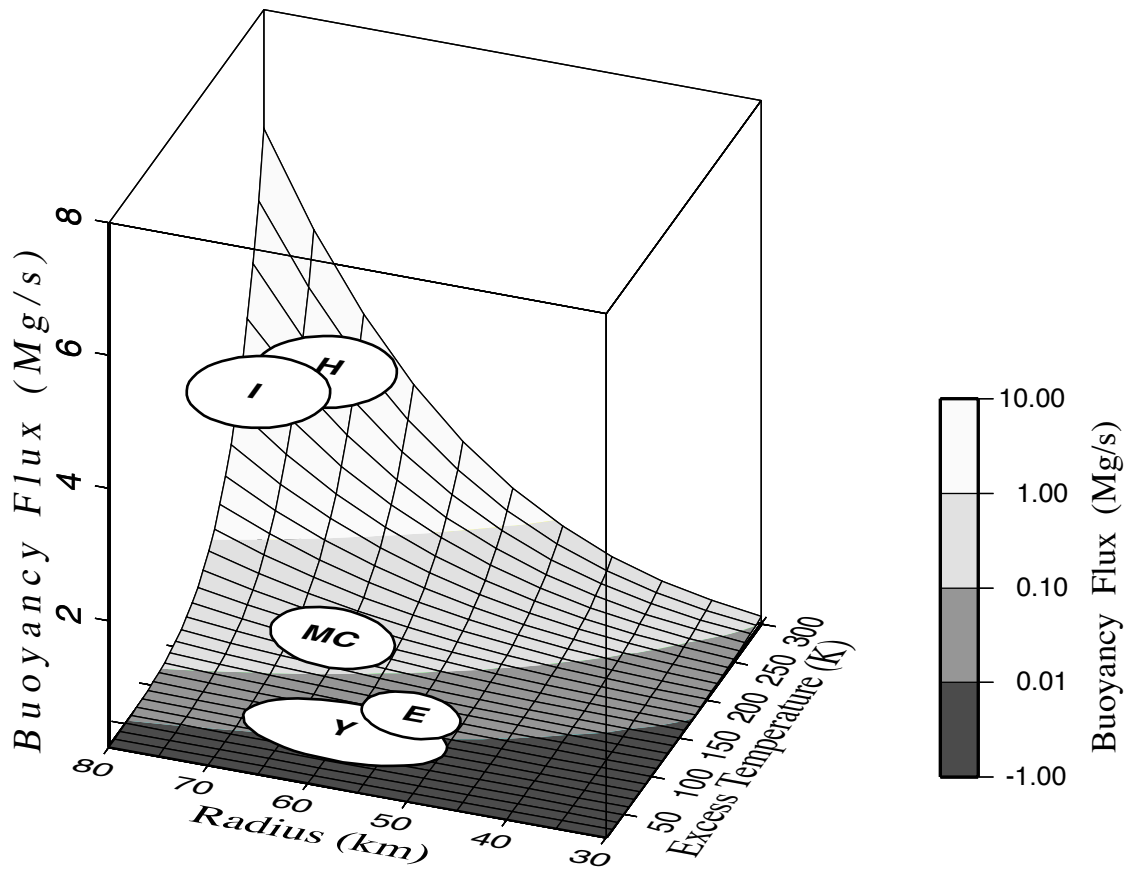
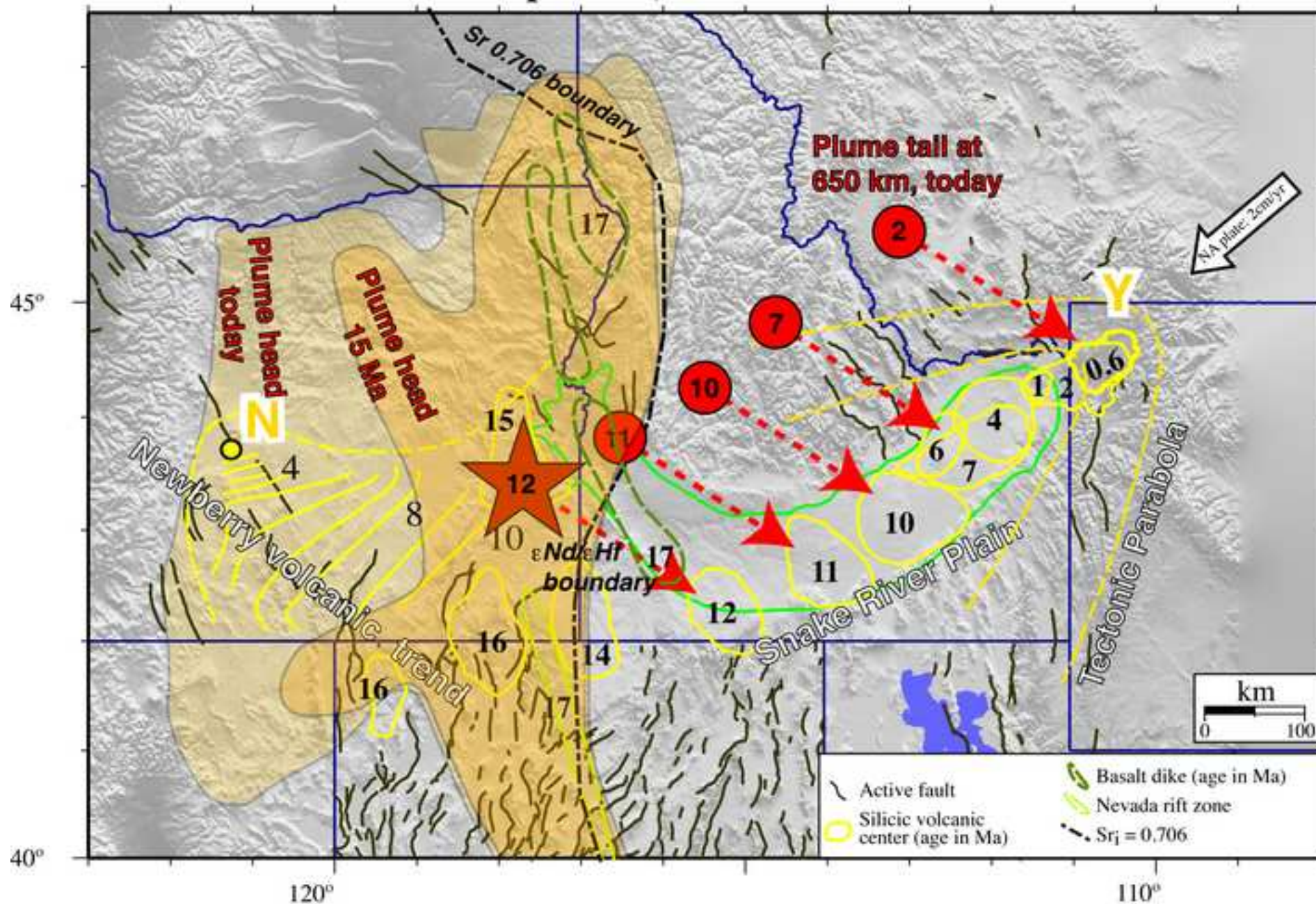






Figure 29  
[Click here to download high resolution image](#)

## Yellowstone Hotspot Volcanism, 16 Ma to present, 650 km to the surface



**Table 1.**

<b>Body #</b>	<b>depth min [km]</b>	<b>depth max [km]</b>	<b>B Model 1</b>	<b><math>\Delta\rho</math> [kg/dm<sup>3</sup>] Model 1</b>	<b>B Model 500</b>	<b><math>\Delta\rho</math> [kg/dm<sup>3</sup>] model 500</b>	<b>B Model 1000</b>	<b><math>\Delta\rho</math> [kg/dm<sup>3</sup>] Model 10000</b>
1	25	50	4.6	-0.052	4.0	-0.060	3.4	-0.071
2	50	110	3.0	-0.081	3.2	-0.076	3.8	-0.064
3	110	155	2.2	-0.110	2.6	-0.093	3.0	-0.081
4	155	285	2.0	-0.042	2.2	-0.038	3.2	-0.026

Figure s1

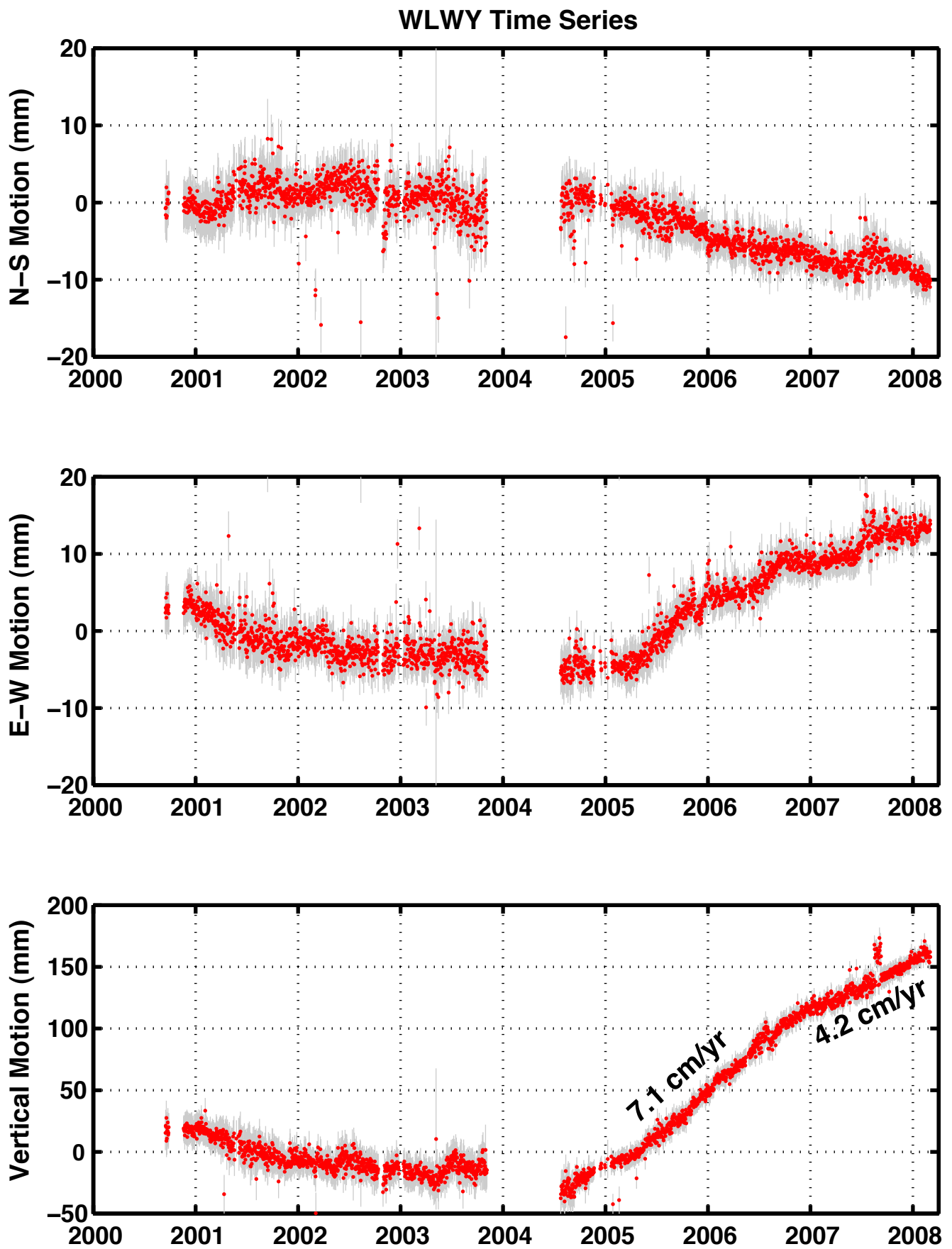




Figure s2

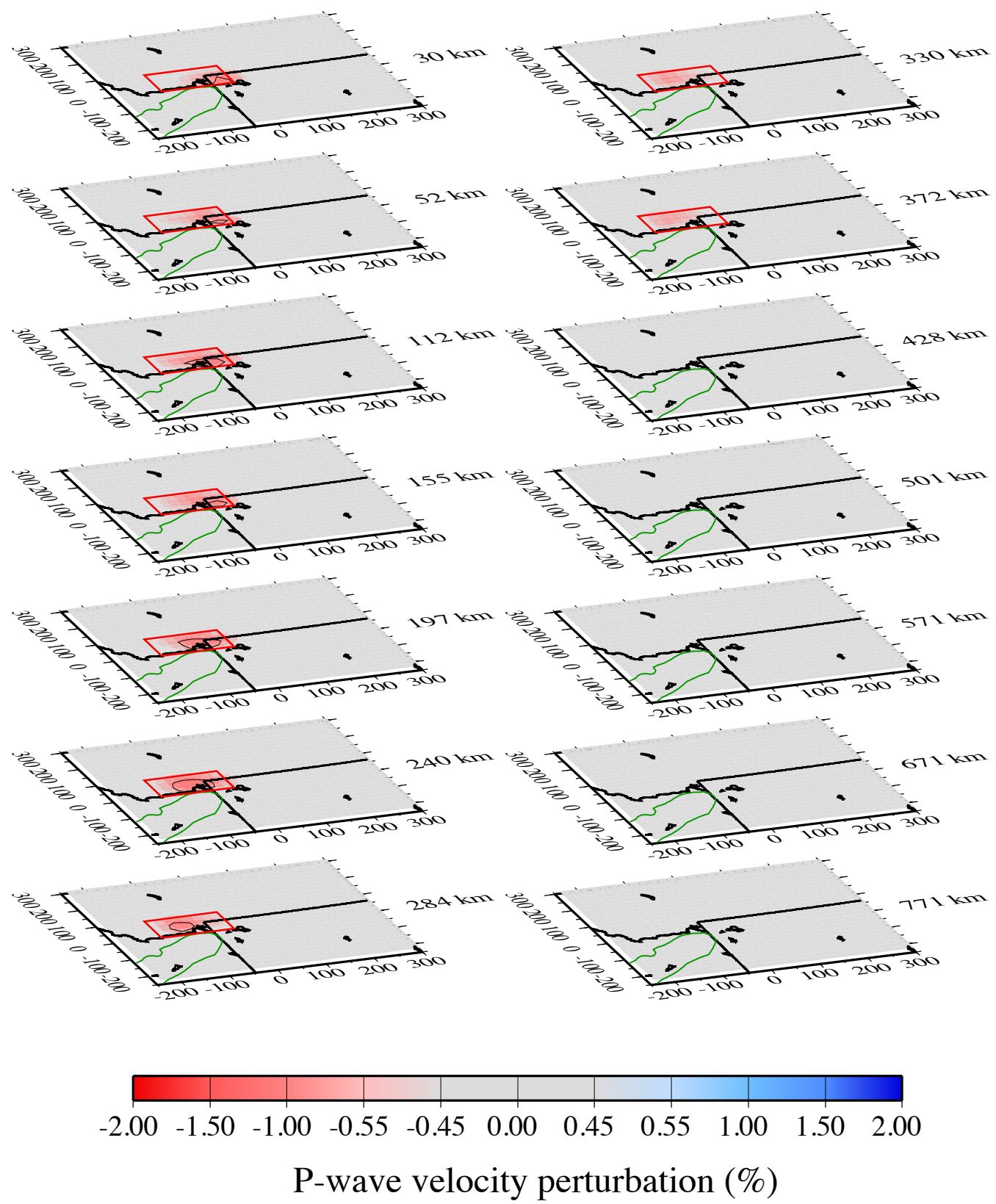


Figure s3

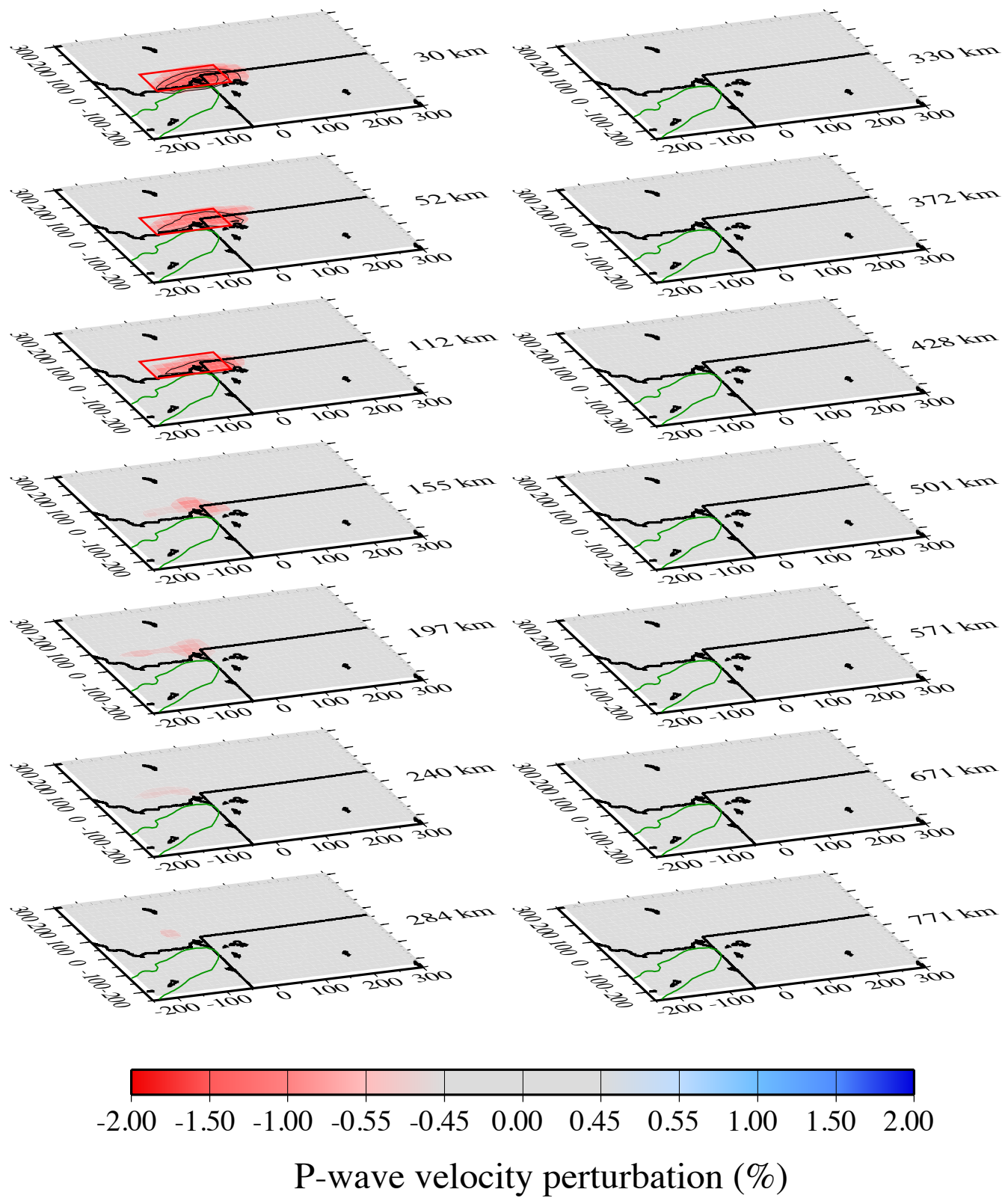


Figure s4

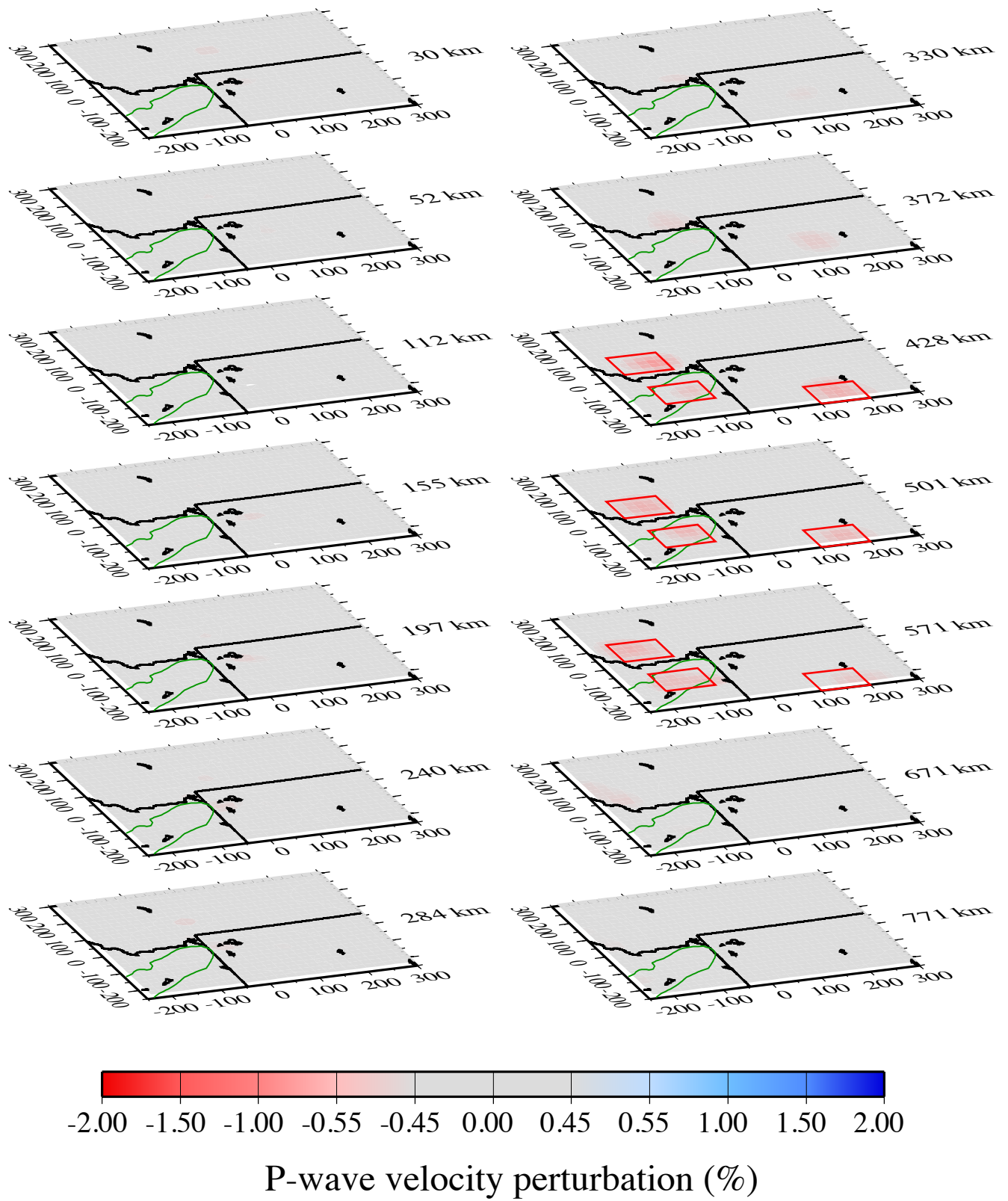
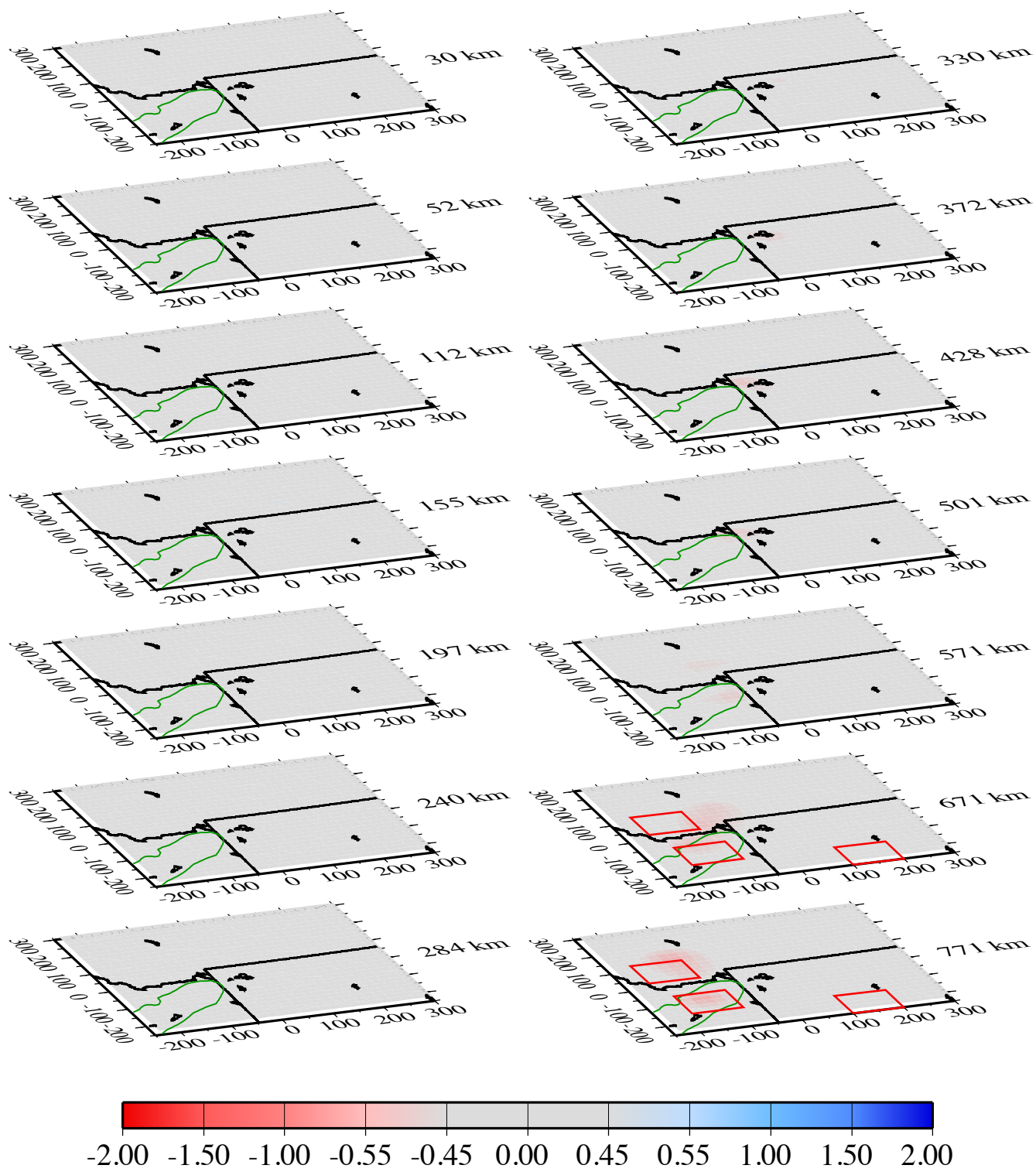


Figure s5



P-wave velocity perturbation (%)

Figure s6

