



Originally published as:

Boxberger, T., Picozzi, M., Parolai, S. (2011): Shallow geology characterization using Rayleigh and love wave dispersion curves derived by seismic noise array measurements. - Journal of Applied Geophysics, 75, 2, 345-354

DOI: [10.1016/j.jappgeo.2011.06.032](https://doi.org/10.1016/j.jappgeo.2011.06.032)

1 **Shallow geology characterization using Rayleigh and Love wave dispersion**  
2 **curves derived from seismic noise array measurements**

3 T. Boxberger, M. Picozzi\*, S. Parolai

4 *Helmholtz Centre Potsdam GFZ German Research Centre for Geosciences, Germany*

5 \*picoz@gfz-potsdam.de

6  
7

8 **Abstract**

9 The local geology and shallow S-wave velocity structure of a site are recognized to be key factors for  
10 the increase in the damaging potential of seismic waves. Indeed, seismic amplitudes may be  
11 amplified in frequency ranges unfavorable for building stock by the presence of soft sedimentary  
12 covers over lying hard bedrock. Hence, microzonation activities, which aim at assessing the site  
13 response as accurately as possible, have become a fundamental task for the seismic risk reduction of  
14 urbanized areas. Methods based on the measurement of seismic noise, which typically are fast, non-  
15 invasive, and low cost, have become a very attractive option in microzonation studies.

16 Using observations derived from seismic noise recordings collected by two-dimensional arrays of  
17 seismic stations, we present a novel joint inversion scheme for surface wave curves. In particular, the  
18 Love wave, the Rayleigh wave dispersion and the HVSR curves are innovatively combined in a joint  
19 inversion procedure carried out following a global search approach (i.e., the Genetic Algorithm).

20 The procedure is tested using a data set of seismic noise recordings collected at the Bevagna (Italy)  
21 test-site. The results of the novel inversion scheme are compared with the inversion scheme  
22 proposed by Parolai et al. (2005), where only Rayleigh wave dispersion and HVSR curves are used,  
23 and with a cross-hole survey.

24  
25  
26  
27  
28  
29

30 Abbreviations:

31 ESAC - extended spatial autocorrelation method  
32 GA - genetic algorithm  
33 HVSR - horizontal-to-vertical spectral ratio  
34 INGV-Mi - Istituto Nazionale di Geofisica e Vulcanologia - Sezione di Milano  
35 RMS - root mean square  
36 SPAC - spatial autocorrelation method

37

38

39

40

41 **1. Introduction**

42 It has been recognized for some time that the local geology is a key factor when considering the  
43 damaging potential of seismic waves. Seismic amplitudes may in fact be amplified in frequency  
44 ranges unfavorable for building stock by the presence of sedimentary covers. Hence, microzonation  
45 activities, which aim at assessing a site's response as accurately as possible, have become a  
46 fundamental task for the seismic risk reduction of urbanized areas.

47 The variation of the S-wave velocity with depth and the total thickness of the sedimentary cover are  
48 the most important physical parameters controlling the amplification of seismic waves. For these  
49 reasons, during the last decades, several approaches have been developed that aim to provide in-situ  
50 measurements of S-wave velocity variations with depth. The usage of very accurate techniques such  
51 as downhole or cross-hole methods, however, are often too expensive for microzonation studies,  
52 where a large territorial coverage is mandatory. Therefore, methods based on the measurement of  
53 seismic noise, which typically are fast, non-invasive, and low cost, have become a very attractive  
54 option.

55 The dispersion curves of surface waves are very sensitive to S-wave velocity variations with depth  
56 (Zhang *et al*, 1996). Therefore, several approaches have been proposed in the literature to extract  
57 from seismic noise recordings information about the velocity of seismic wave propagation in the  
58 sedimentary cover. In general, all methods for retrieving surface wave dispersion curves are based  
59 on phase-coherency measurements between pairs (at least two) of signals. Aki (1957) proposed the  
60 spatial autocorrelation (SPAC) method, which recently has been generalized in the extended spatial  
61 autocorrelation (ESAC) method by Ohori *et al.* (2002), to extract both the Love and Rayleigh wave  
62 dispersion curves. Similarly, frequency-wavenumber approaches, including the beam-forming (BFM;  
63 Lacoss *et al*, 1969) and the maximum likelihood (MLM; Capon, 1969) methods, together with their  
64 variants, can also be successfully used for the same purpose.

65 Similarly, Scherbaum *et al.* (2003), and Arai and Tokimastu (2004) showed that Horizontal-to-Vertical  
66 Spectral Ratio (HVSr) curves from seismic noise recordings, originally proposed by Nogoshi & Igarashi

67 (1971), are sensitive to the shallow S-wave velocity structure of a site, and can be profitably inverted  
68 to obtain the structure.

69 Once the surface wave dispersion or HVSR curves of a site are available, from their inversion the S-  
70 wave velocity profile of a site can be obtained. In order to overcome the difficulties related to the  
71 non-linear nature of this inverse problem, several inversion approaches have been tested over the  
72 last few years. Among others, Parolai et al. (2006) discussed the pro and cons of both linearized and  
73 global inversion methods when applied to the inversion of Rayleigh wave dispersion curves.  
74 Scherbaum et al. (2003) showed that the independent inversions of Rayleigh wave dispersion and  
75 HVSR curves are inexorably affected by the trade-off between S-wave velocity and the thickness of  
76 the sedimentary cover. For this reason, Parolai et al. (2005) and Arai and Tokimatsu (2005) proposed  
77 the joint inversion of Rayleigh wave dispersion and HVSR curves, where in each case surface wave  
78 higher modes are included in the analysis. Similarly, Köhler et al. (2007) presented a combined  
79 inversion of Love and Rayleigh dispersion curves that takes advantage of the different sensitivity of  
80 Love and Rayleigh waves to the S-wave velocity structure.

81 In this work we present a joint inversion of Rayleigh and Love dispersion curves together with the  
82 HVSR curve (hereafter, LRHV inversion). This novel inversion approach is applied to seismic noise  
83 data recorded by a 2D array of seismic stations at the Bevagna (Italy) test site, where independent  
84 geological and geophysical information was available. The Rayleigh and Love dispersion curves have  
85 been estimated using the ESAC method, and specifically using the procedure proposed by Metaxian  
86 et al. (1997) for the analysis of the horizontal components of ground motion. The two dispersion  
87 curves are then combined with HVSR curves representative of the whole array into a joint inversion,  
88 based on a genetic algorithm (Parolai et al., 2005). In order to evaluate the advantages related to the  
89 proposed inversion scheme, we compared the resultant S-wave velocity profile with the one  
90 obtained by the classical joint inversion approach of Parolai et al. (2005), where only the Rayleigh  
91 dispersion and HVSR curves are considered (hereafter, RHV inversion). Moreover, we compared the  
92 inversion results with the S-wave profile obtained by a cross-hole survey made until a depth of 40 m.

93 Finally, both the performance of the RHV and LRHV inversion schemes and the different sensitivities  
94 of the Love, Rayleigh dispersion and HVSR curves with respect to the S-wave velocity variation with  
95 depth are discussed by comparing the resolution and Jacobian matrixes.

96

## 97 **2. Methods**

98 In this section we present a summary of the methods used for deriving the Rayleigh and Love waves  
99 dispersion curves, the HVSR curves, as well as the Genetic Algorithm used in the joint inversion  
100 analysis for the estimation of the S-wave velocity profile. A comprehensive literature review on the  
101 topic is reported by Foti et al. (2011).

102

### 103 **2.1 SPatial Auto-Correlation (SPAC) and Extended Spatial Auto-Correlation (ESAC) for the vertical** 104 **component of ground motion**

105 The SPAC method was originally proposed by Aki (1957, 1965) as a statistical tool for the extraction  
106 from seismic noise of information dealing with the surface wave phase velocities in sedimentary  
107 layers. The method is based on the assumption that the seismic noise represents the sum of waves  
108 propagating without attenuation in a horizontal plane in different directions with different powers,  
109 but with the same phase velocity for a given frequency. Moreover, it is also assumed that waves with  
110 different propagation directions and different frequencies are statistically independent. As discussed  
111 in Ohori et al. (2002), when these assumption are verified, the space-correlation function for one  
112 angular frequency  $\omega_0$ , normalized to the power spectrum, can be expressed in the form

$$113 \quad \bar{\rho}(r, \omega_0) = J_0 \left( \frac{\omega_0}{c(\omega_0)} r \right), \quad (1)$$

114 where,  $c(\omega)$  is the frequency-dependent phase velocity,  $r$  is the interstation distance, and  $J_0$  is the  
115 zero order Bessel function. Eq. 1 provides a link between the Rayleigh wave phase velocity and the  
116 theoretical space cross-correlation function values. Considering that experimental values of the  
117 azimuthally averaged spatial correlation function  $\bar{\rho}(r, \omega_0)$  can be obtained from seismic noise

118 measurements carried out with 2D arrays of seismic stations, by fitting these latter values to the  
 119 theoretical Bessel function values, the phase velocity  $c(\omega)$  can be retrieved.  
 120 Typically in the SPAC method, a fixed value of  $r$  is used. However, Okada (2003) and Ohori et al.  
 121 (2002) showed that, since  $c(\omega)$  is a function of frequency  $\omega$ , better phase-velocity estimates are  
 122 achieved by fitting the spatial-correlation function at each frequency to a Bessel function, which  
 123 depends on the inter-station distances (extended spatial autocorrelation, ESAC). For every pair of  
 124 stations, the function  $\phi(\omega)$  can be calculated in the frequency domain by means of (Malagnini et  
 125 al., 1993; Ohori et al., 2002; Okada, 2003):

$$126 \quad \bar{\rho}_{jn}(f) = \frac{\frac{1}{M} \sum_{l=1}^M \text{Re}({}_m S_{jn}(\omega))}{\sqrt{\frac{1}{M^2} \sum_{l=1}^M {}_m S_{jj}(\omega) \sum_{l=1}^M {}_m S_{nn}(\omega)}} , \quad (2)$$

127 where  ${}_m S_{jn}$  is the cross-spectrum for the  $m$ th segment of data between the  $j$ th and the  $n$ th  
 128 stations and  $M$  is the total number of used segments. The power spectra of the  $m$ th segments at  
 129 stations  $j$  and  $n$  are therefore expressed as  ${}_m S_{jj}$  and  ${}_m S_{nn}$ , respectively. Hence, the experimental  
 130 space-correlation values from eq. (2) are plotted for every frequency as a function of distance, and  
 131 an iterative grid-search procedure can be then performed using eq. (1) in order to find the value of  
 132  $c(\omega)$  that gives the best fit to the data.

133

## 134 **2.2 ESAC for the horizontal components of ground motion**

135 As discussed by Aki (1965), the SPAC equations derived for the analysis of the vertical component of  
 136 ground motion can in principle be adapted for the analysis of the horizontal components, with the  
 137 aim of extracting the phase velocities of Rayleigh and Love waves. In fact, in the case of Rayleigh  
 138 waves, when they are polarized parallel to the propagation direction, equations similar to Eq. 1 can  
 139 be derived for both the radial and tangential components of motion. In particular, the azimuthally

140 averaged correlation coefficient for the radial  $\bar{\rho}_r(r, \omega_0)$  and tangential  $\bar{\rho}_t(r, \omega_0)$  components are  
 141 expressed by

$$142 \quad \bar{\rho}_r(r, \omega_0) = J_0\left(\frac{\omega_0}{c_R(\omega_0)} r\right) - J_2\left(\frac{\omega_0}{c_R(\omega_0)} r\right), \quad (3)$$

143 and

$$144 \quad \bar{\rho}_t(r, \omega_0) = J_0\left(\frac{\omega_0}{c_R(\omega_0)} r\right) + J_2\left(\frac{\omega_0}{c_R(\omega_0)} r\right), \quad (4)$$

145 where  $J_2$  is the second-order Bessel function.

146 Similarly, for Love waves, when they are perpendicularly polarized to the direction of propagation,  
 147  $\bar{\rho}_r(r, \omega_0)$  and  $\bar{\rho}_t(r, \omega_0)$  are expressed by

$$148 \quad \bar{\rho}_r(r, \omega_0) = J_0\left(\frac{\omega_0}{c_L(\omega_0)} r\right) + J_2\left(\frac{\omega_0}{c_L(\omega_0)} r\right), \quad (5)$$

149 and

$$150 \quad \bar{\rho}_t(r, \omega_0) = J_0\left(\frac{\omega_0}{c_L(\omega_0)} r\right) - J_2\left(\frac{\omega_0}{c_L(\omega_0)} r\right), \quad (6)$$

151 where  $c_L(\omega_0)$  is the phase velocity of the Love waves.

152 However, during real surveys, the seismic noise in the horizontal components of ground motion is  
 153 characterized by the superposition of both Rayleigh and Love waves. Thus, the extraction of phase-  
 154 velocity information for the single phases is not straightforward.

155 For this reason, Metaxian et al. (1997) proposed, under the assumption that the contribution of both  
 156 Rayleigh and Love waves is statistically independent, to adopt the equation

$$157 \quad \bar{\rho}_r(r, \omega_0) = \alpha \left[ J_0\left(\frac{\omega_0}{c_R(\omega_0)} r\right) - J_2\left(\frac{\omega_0}{c_R(\omega_0)} r\right) \right] + (1 - \alpha) \left[ J_0\left(\frac{\omega_0}{c_L(\omega_0)} r\right) + J_2\left(\frac{\omega_0}{c_L(\omega_0)} r\right) \right], \quad (7)$$

158 and

$$\begin{aligned}
\bar{\rho}_i(r, \omega_0) = & \alpha \left[ J_0 \left( \frac{\omega_0}{c_R(\omega_0)} r \right) + J_2 \left( \frac{\omega_0}{c_R(\omega_0)} r \right) \right] \\
& + (1 - \alpha) \left[ J_0 \left( \frac{\omega_0}{c_L(\omega_0)} r \right) - J_2 \left( \frac{\omega_0}{c_L(\omega_0)} r \right) \right], \tag{8}
\end{aligned}$$

160 where  $\alpha(\omega)$  represents the proportion of Rayleigh and Love waves in the wave field energy . In  
161 particular, for  $\alpha = 1$ , the wavefield is dominated by Rayleigh waves, while for  $\alpha = 0$ , only Love  
162 waves exist. Eqs. (7) and (8) can therefore be exploited to retrieve Love wave phase velocity  
163 estimates. In particular, once the Rayleigh wave phase velocities are constrained by the analysis of  
164 the vertical component of motion, a similar iterative procedure can be implemented on the  
165 horizontal components of motion for the estimation of Love wave phase velocities, with only the  
166 addition of a loop accounting for the variation of the parameter  $\alpha$ .

167

### 168 **2.3 Horizontal-to-Vertical Spectra Ratios (HVSr)**

169 Nakamura (1989) revised the Horizontal-to-Vertical Spectral Ratio (HVSr) technique, based on  
170 seismic noise recordings, that was first proposed by Nogoshi and Igarashi (1970, 1971). Since then,  
171 the HVSr technique has been used in site effect estimation in a large number of studies (e.g., Field  
172 and Jacob, 1993; Lermo and Chavez-Garcia, 1994; Mucciarelli, 1998; Bard, 1998; Parolai et al., 2001),  
173 especially due to its very low cost of execution. Similarly to surface wave dispersion curves, HVSr  
174 also provides information on the subsoil structure of a site. Specifically, HVSr curves are strongly  
175 conditioned by the properties (depth and S-wave velocity contrast) of the interface between the soft  
176 sediment and bedrock. Arai and Tokimatsu (2004) proposed the inversion of HVSr curves for the  
177 estimation of the S-wave velocity profile, while Parolai et al. (2005) and Arai and Tokimatsu (2005)  
178 first proposed using HVSr curves together with Rayleigh wave dispersion curves in a joint inversion  
179 scheme. Computing the HVSr corresponds to the normalization of the horizontal spectral amplitude  
180 with respect to the vertical one.



181 In practice, the method for evaluating HVSR curves consists of (a) merging the two Fourier spectra of  
182 the two horizontal components,  $X(\omega)$  and  $Y(\omega)$ , of motion to obtain a single combined horizontal  
183  $H(\omega)$  component, whose modulus is given by

$$184 \quad H(\omega) = \sqrt{(X(\omega))^2 + Y(\omega)^2} \quad (9)$$

185 and (b) computing the ratio between these  $H(\omega)$  spectra with the one from the vertical  
186 component.

187

## 188 **2.4 Genetic Algorithm (GA) Inversion**

189 As discussed before, both the surface wave dispersion and HVSR curves provide the necessary  
190 information for S-wave velocity estimation. However, the relationship between the surface wave  
191 velocities and HVSR curves and the S-wave velocity and sediment thickness is not linear.

192 Until only a few years ago, most works dealing with estimating the S-wave velocity profile of a site  
193 using seismic noise recordings focused only on the separate inversion of either the Rayleigh wave  
194 dispersion curve or the HVSR curves. However, as shown by Scherbaum et al. (2003), the HVSR ratio  
195 and the Rayleigh wave dispersion curve display different sensitivities to the S-wave velocity and  
196 thickness of the sedimentary layers. In particular, they showed that when these curves are inverted  
197 singularly, there is an un-resolvable trade-off between the model parameters that hampers the  
198 analysis results. Therefore, to overcome this drawback and the non-linear nature of the inverse  
199 problem, Parolai et al. (2005) proposed a joint inversion of Rayleigh wave phase velocity and HVSR  
200 curves using a genetic algorithm (GA) while also considering also the higher modes. They showed  
201 that with this approach, the trade-off between the model parameters can be reduced and a reliable  
202 evaluation of the local S-wave velocity structure can be obtained. In fact, the dispersion curve  
203 provides a constraint on the S-wave velocity of the soft sediments, while the fundamental frequency,  
204  $f_0$  estimated from the HVSR ratio peak, represents a constraint for the total sedimentary-cover  
205 thickness.

206 Parolai et al. (2005) proposed to perform the joint inversion of the Rayleigh wave dispersion and  
 207 HVSR curves using the genetic algorithm (GA) code presented by Yamanaka and Ishida (1996).  
 208 Indeed, GAs (Goldberg, 1989) belong to the class of evolutionary algorithms that generate solutions  
 209 for the optimization of non-linear inverse problems by the analysis of thousands of possible models  
 210 and by the usage of techniques inspired by the natural evolution theory of Darwin, such as  
 211 inheritance, mutation, selection and crossover.  
 212 Parolai et al. (2005) showed the efficiency of joint inversion strategies in generating models near the  
 213 global optimal solution (i.e., shear-wave velocity models able to justify both the observed Rayleigh  
 214 wave dispersion and HVSR curves) using a cost function defined as

$$\begin{aligned}
 \text{cost} = & [(1-p)N + pK] \left\{ \frac{1-p}{N} \left[ \sum_{j=1}^N \left( \frac{c_{R_o}(f) - c(f)}{c_{R_o}(f)} \right)^2 \right] \right. \\
 & \left. + \frac{p}{K} \left[ \sum_{j=1}^K \left( \frac{hv_o(f) - c(f)}{hv_o(f)} \right)^2 \right] \right\}
 \end{aligned} \tag{10}$$

216 where the subscript  $o$  indicates the observed data, and  $N$ , and  $K$  are the number of data points for  
 217 the Rayleigh wave dispersion and HVSR curves, respectively.

218 In this study, we slightly modified the Eq. (10) in order to include within the joint inversion scheme  
 219 the Love wave dispersion curve. Hence, a new cost function was defined as

$$\begin{aligned}
 \text{cost} = & [(1-p-q)N + qM + pK] \left\{ \frac{1-p-q}{N} \left[ \sum_{j=1}^N \left( \frac{c_{R_o}(f) - c_R(f)}{c_{R_o}(f)} \right)^2 \right] \right. \\
 & \left. + \frac{q}{M} \left[ \sum_{j=1}^M \left( \frac{c_{L_o}(f) - c_L(f)}{c_{L_o}(f)} \right)^2 \right] + \frac{p}{K} \left[ \sum_{j=1}^K \left( \frac{hv_o(f) - hv(f)}{hv_o(f)} \right)^2 \right] \right\}
 \end{aligned} \tag{11}$$

221 where  $M$  is the number of points in the Love wave dispersion curve. The relative influence of the  
 222 data sets in Eqs. (10) and (11) is controlled by the parameters  $p$  and  $q$ , where  $p$  controls the  
 223 relative importance of the HVSR curve and  $q$  of the Love wave dispersion curve. That is to say, if  $p$  or  
 224  $q$  are set equal to zero, then the HVSR or Love wave dispersion curves are not considered during the  
 225 inversion.

226 Finally, the forward modeling of Rayleigh and Love wave phase velocities and of HVSR curves was  
227 performed using an improved Thomson-Haskell method proposed by Wang (1999). This included the  
228 influence of higher modes of surface waves in both the dispersion and HVSR curves following the  
229 equations proposed by Tokimatsu et al. (1992) and Arai and Tokimatsu (2004), under the assumption  
230 of vertically heterogeneous 1D earth models.

231

### 232 **3 The Bevagna test site**

#### 233 **3.1 Data acquisition**

234 In September 2007, an array of 15 seismic stations was installed in the village of Bevagna (Italy) by  
235 INGV-MI near the BVG accelerometric station of the Italian Accelerometric Network. The same data  
236 set has been used by Puglia et al. (2011) for a study of site effects, based on the separate inversion of  
237 Rayleigh and Love wave dispersion curves. All stations were equipped with Lennartz LE-3D/5s three  
238 component sensors using a Reftek 130 digitizer. Figure (1) shows the selected array geometry, which  
239 is characterized by side lengths of 170 and 85 m, and a minimum inter-station distance of 10 m. The  
240 selected range of station inter-distances allows for an optimal compromise between resolution at  
241 shallow depths and the maximum depth of investigation. Seismic noise was recorded at 500 Hz  
242 sampling rate for over 3 hours. Moreover, cross-hole measurements of the S-wave velocity down to  
243 40 m depth have been carried out within the framework of the DPC-INGV S6 project in 2006  
244 (Deliverable, 2009), making the shallow structure of the site quite well constrained.

245

#### 246 **3.2 ESAC on vertical and horizontal components**

247 The Rayleigh wave phase velocities were computed by analyzing the seismic noise recorded on the  
248 vertical component, using 300 windows of signal of 30 seconds length. In order to reduce leakage  
249 problems, each signal window was tapered for 5 per cent of its length using a cosine function.

250 Figure (2) shows an example of the application of the ESAC analysis for 4 frequencies. The space-  
251 correlation values for every frequency are plotted as a function of distance, and an iterative grid-

252 search procedure was performed using Eq. (1) in order to find the Rayleigh phase velocity value that  
253 gives the best fit to the observed data. The tentative phase velocity was generally varied over large  
254 intervals (e.g., between 100 and 3000 m/s) in small steps (i.e., 1 m/s). The best fit is achieved by  
255 minimizing the root mean square (RMS) of the differences between the values calculated using Eqs.  
256 (1) and (2). Following Parolai et al., (2006), three grid-search iterations were performed, and data  
257 points, which differ by more than two standard deviations from the value obtained with the  
258 minimum-misfit velocity, were removed before the next iteration. Then, using the Rayleigh wave  
259 phase velocities constrained so far, the procedure detailed in section 2.2 was performed using the  
260 horizontal components of the seismic noise ground motion to estimate the Love wave dispersion  
261 curve. With respect to the analysis carried out on the vertical component, Eqs. (7) and (8) were used,  
262 and only an additional loop was added to the iterative procedure to take into account variations of  
263 the parameter  $\alpha$  (i.e., between 0 and 1 with steps of 0.1), which measures the proportion of Rayleigh  
264 and Love waves within the seismic noise (Metaxian et al. 1997). The number of seismic noise  
265 windows considered, the window length, and the phase velocity range used in the grid search  
266 procedure were the same as for vertical component.

267 Figure (3) shows the contour plot of the RMS values derived using Eqs. (7) and (8) and the  
268 experimental spatial correlation coefficients when  $\alpha$  is varied between 0 and 1 for each frequency.  
269 Interestingly, a rather stable trend in the proportion of Rayleigh and Love waves within the seismic  
270 noise wavefield is observed in the frequency range 1.5 to 2.5 Hz, where Rayleigh and Love waves  
271 constitute about 30% and 70% of the seismic noise, respectively. However, we observed that the  
272 proportion of Rayleigh wave content increases with increasing frequency, until a value of between  
273 70- 90 % at 4 Hz. Figure (4) shows the results of the ESAC procedure when applied to the horizontal  
274 component of ground motion for the same frequencies as in Figure (2). In this case, the minima  
275 observed in the different RMS functions indicates the phase velocity of Love waves.

276 Finally, Figure (5) shows for the frequency range 1.3 Hz to about 5 Hz the comparison between the  
277 Rayleigh and Love wave dispersion curves together with the associated uncertainties. Unfortunately,

278 outside of this frequency range, it was not possible to obtain reliable estimates of the phase velocity.  
279 The two dispersion curves show a similar trend for the frequency range 3.4 to 5 Hz, which  
280 corresponds to the shallower portion of the subsurface. By contrast, for frequencies below 3 Hz, the  
281 two dispersion curves diverge considerably.

282

### 283 **3.3 HVSR**

284 The HVSR were computed for all the stations of the array, again using 300 windows of 30 seconds  
285 length, tapered with a cosine function for 5 per cent of their length. Figure (6) shows that most of the  
286 stations are characterized by a predominant peak at 1.3 Hz with amplitudes of around 6. Only for  
287 three stations located in the northern part of the array (BE01, BE09, and BE13) does the 1.3 Hz peak  
288 display a broader shape, probably due to the presence of spurious signals generated by a nearby  
289 small road and an irrigation canal.

290 Some of the HVSR curves (e.g., BE01, BE03, BE07, and BE15 in Figure 6) showed a further secondary  
291 peak at frequency around 0.3-0.4 Hz. Unfortunately, most of the stations show for frequencies below  
292 0.5 Hz an anomalous increasing trend in the HVSR curve, probably due to tilt effects affecting the  
293 horizontal components of the sensors (Forbriger, 2006). Therefore, in order to avoid introducing to  
294 the inversion analysis observations that might be biased by systematic errors, we decided to neglect  
295 the HVSR peak at frequency between 0.3-0.4 Hz. Moreover, we selected the HVSR from station BE14  
296 (Figure 4, subplot), as being representative of the majority of the curves to be included in the joint  
297 inversion scheme, since it shows a clear peak at 1.3 Hz, but is not affected by any artificial trend in  
298 the lower frequencies.

299

### 300 **3.4 Inversion analysis**

301 In this section we present the S-wave velocity results obtained by the inversion analyses. Following  
302 the procedure described in section 2.4, we first show the RHV inversion results and then compare  
303 them to those obtained by the LRHV inversion scheme.

304

### 305 **3.4.1 Joint inversion of Rayleigh wave dispersion and HVSR curves (RHV)**

306 As discussed in the previous section, the inversion based on GA does not employ any explicit starting  
307 models, but requires suitable parameter limits for each layer of the model to be defined. In this case,  
308 we inverted for the parameters S-wave velocity ( $V_s$ ) and thickness ( $H$ ), which as shown by Arai and  
309 Tokimatsu (2004) are the parameters that most influence the propagation of surface waves into the  
310 ground. Therefore, during the inversion analysis, the thickness and S-wave velocity of each layer  
311 could vary within pre-defined ranges, while the density ( $d$ ) and the P-wave velocity ( $V_p$ ) were  
312 constrained. In particular, the density was set by selecting values from the literature in agreement  
313 with *a priori* geological information, while the  $V_p$  were related to the S-wave velocity using the  
314 relationship of Kitsunozaki *et al.* (1990)  $V_p = 1.1 \cdot V_s + 1290$ , where both  $V_s$  and  $V_p$  are expressed  
315 in m/s.

316 Different parameterizations of the model were tested, and finally we selected one consisting of 4  
317 layers. In order to avoid over-parameterization, the range of thicknesses explored by the GA were  
318 increased with depth. In table (1) the tested parameter ranges of the S-wave velocity and layer  
319 thicknesses are presented.

320 The GA inversion consisted of 150 generations of a population of 50 models. Moreover, the inversion  
321 was repeated starting from 5 different seed numbers, i.e., from a different population of initial  
322 models, with the aim of increasing the exploration of the model parameters space and thus  
323 increasing the probability of converging towards the global minimum of the inverse problem.

324 Following Parolai *et al.* (2005), within the cost function (Eq. 13) we weighted the two data-sets using  
325 a value of 0.9 for the Rayleigh wave dispersion curves and 0.1 for the HVSR curves. The inversion was  
326 performed over the frequency band 1.25-5.3 Hz for the Rayleigh wave dispersion curves and 0.5-5.3  
327 Hz for the HVSR curves.

328 Figure (7) shows the results of the GA inversion procedure. The best-fit S-wave velocity profile is  
329 characterized by a significant impedance contrast at about 16 m depth, and by a second deeper

330 impedance contrast at about 85 meters. Using the best-fit S-wave velocity profile, and computing the  
331 average velocity ( $\overline{V_s}$ ) at the different impedance contrasts, we estimated the theoretical  
332 fundamental resonance frequency for SH-waves [i.e.,  $f_0 = \frac{\overline{V_s}}{4H}$ ] for the different layers.  
333 Interestingly, we found that a value of  $f_0$  comparable to the maximum in the HVSR curve is found for  
334 the impedance contrast at 16m of depth. On the contrary, for the impedance contrast at 85 m, a  
335 value of  $f_0$  around 0.6 Hz is found. In light of these observations, the HVSR and Rayleigh wave  
336 dispersion curves cooperate in constraining the impedance contrast at 16 m, while the deeper  
337 contrast is constrained only by the Rayleigh waves phase velocity.

338 The misfit function (Figure 7b) of the seed number leading to the best fit model shows a consistent  
339 decreasing trend during the first 20 generations, and then a rather stable trend with continuously  
340 decreasing misfit with increasing numbers of generations. It worth noting that the theoretical  
341 Rayleigh wave dispersion and HVSR curves computed for the best-fit model are in good agreement  
342 with the observed curves, and lay inside their uncertainty bounds. These observations suggest that  
343 the GA found a solution close to the global minimum of the inverse problem.

344 The distribution of the family of models with misfits within 10% RMS of the best-fit model (hereafter,  
345 models-10%) around the best-fit model itself is a further indication of the quality of the solution.  
346 Focusing on this feature within Figure (7a), we observe that for the 1<sup>st</sup> and 3<sup>rd</sup> layers, the models-10%  
347 are very close to the best-fit model, indicating that these two layers are well constrained. On the  
348 other hand, for the 2<sup>nd</sup> and 4<sup>th</sup> layers, the models-10% display that a range of other possible  $\overline{V_s}$  and  
349 H values could justify, in terms of the chosen misfit functions, in a very similar way the observations.

350 Therefore, despite the very good quality of the inversion results, it seems that the Rayleigh wave  
351 dispersion and the HVSR curves are not able to constrain with the same robustness the different  
352 portions of the model. This is consistent with previous studies (see among the others Scherbaum et  
353 al., 2003; Köhler et al. 2007), that showed the different sensitivities of HVSR, Rayleigh wave and Love  
354 wave dispersion curves to the S-wave velocity profile of the site. For this reason, in order to improve

355 our estimates of the S-wave velocity profiles from seismic noise recordings, in a further test we also  
356 included the Love wave dispersion curves within the joint inversion scheme.

357

### 358 **3.4.2 Joint inversion of Love and Rayleigh wave dispersion and HVSR curves (LRHV)**

359 The joint inversion of Love and Rayleigh wave dispersion and HVSR curves was carried out following  
360 the same strategy described in section 3.4.1. That is to say, the same number of layers and range of  
361 values for the model parameters, the same number of generations and seed numbers, as well as the  
362 same information in terms of Rayleigh wave dispersion and HVSR curves were used. The only  
363 differences we introduced into this new inversion scheme were the inclusion of the Love wave  
364 dispersion curve shown in Figure (5) and the use of the new cost function expressed by Eq. (14). In  
365 particular, after tests carried out on synthetic data sets, we selected the weights  $p$  and  $q$  of Eq. (14)  
366 to equal 0.1 and 0.3, respectively, as the optimal values for combining the information given by the  
367 different curves during the inversion, that is to say to obtain an optimal fit of all the data sets used in  
368 the inversion.

369 Figure (8) shows the results of the novel GA joint inversion. Interestingly, despite the best-fit model  
370 of the LRHV inversion showing general characteristics very similar to those of the best-fit model from  
371 the RHV inversion, it presents some important new features. In fact, the velocity variation between  
372 the 1<sup>st</sup> and 2<sup>nd</sup> layer of the model hints at an inversion of the S-wave velocity with depth (Figure 8a).  
373 Moreover, we observe that in the present inversion results, especially in the range 15-25m depth,  
374 the LRHV models-10% are less dispersed around the best-fit model than the RHV ones (Figure 7).  
375 Furthermore, the LRHV results appear to better constrain the depth and value of the main  
376 impedance contrasts at around 18 and 95m depth.

377 Also for the LRHV analysis, the misfit values reached a stable minimum plateau with the increasing  
378 generation number (Figure 8b), suggesting that no improvement can be expected by continuing the  
379 process and that the final solution is likely to lie in the vicinity of the global minimum of the inversion



380 problem. Moreover, Figures (8b and c) show that the best-fit model allows us to reproduce very well  
381 the trends in all three experimental curves.

382

### 383 **3.5 Quantitative comparison of the RHV and LRHV results**

384 The results presented in Figure (8) already suggested qualitatively that including the Love wave  
385 dispersion curves together with the Rayleigh wave dispersion and HVSR curves within a joint  
386 inversion scheme allows us to better constrain the S-wave velocity profile. However, in order to  
387 quantify the improvement in the solution, the inversion results are also compared using different  
388 criteria.

389 The first one is based on calculating for the two best-fit models of the LRHV and RHV inversions the  
390 model resolution matrices (Menke, 1989) while considering the experimental derived surface wave  
391 dispersion and HVSR curves. Although the problem is non-linear, we expect that the obtained final  
392 models will lie close to the global minimum of the solution. Picozzi and Albarello (2007) showed that  
393 in such cases, a linear inversion can be carried out to refine the velocity models. Here, we only take  
394 advantage of the linearization of the problem to study how the procedure and data can be reliably  
395 constrained through the analysis of the model and data resolution matrices. In fact, focusing on the  
396 diagonal elements of the resolution matrixes, which represent how much each of the model  
397 parameters is resolved by the experimental observations, Figure (9) shows that including the Love  
398 wave dispersion curves improves the resolution of the first two layers of the model. The reason for  
399 the improvements in the results from the LRHV inversion can be understood by analysing the  
400 Jacobian matrixes in Figure (10). Each Jacobian matrix is defined as the matrix of the first-order  
401 partial derivative, here numerically calculated, of one of the observations (Love wave dispersion,  
402 Rayleigh wave dispersion, or HVSR) with respect to the model parameters (in this case, the S-wave  
403 velocity at different depth). The distribution of maxima within the Jacobian matrix in Figure (10)  
404 indicates for which change in velocity at depth is each frequency of the Rayleigh, Love wave  
405 dispersion and HVSR curves more sensitive.

406 The comparison of Figures (10a and b) shows that in the frequency range 2 Hz to 5 Hz, Love waves  
407 are much more sensitive to the very shallow material properties (i.e., less than 10 m), while Rayleigh  
408 waves have a higher sensitivity at a depth of around 10 m. For frequencies less than 2 Hz, both Love  
409 and Rayleigh waves tend to be sensitive to the model characteristics over the depth range 20 m to 50  
410 m. Finally, the HVSR curve shows for frequencies around the peak at about 1.3 Hz a high sensitivity to  
411 the model characteristics at a depth of around 20 m, which corresponds to the impedance contrast  
412 between the 2<sup>nd</sup> and 3<sup>rd</sup> layers in Figure (8a). These observations confirm that the main HVSR peak is  
413 constrained by the impedance contrast between these layers, while the second deeper impedance  
414 contrast at around 90 m is constrained by the surface wave dispersion curves. The evidence of  
415 different sensitivities with respect to different model parameters explains well why the joint  
416 inversion performed including all the surface wave derived curves, that is to say Love, Rayleigh wave  
417 dispersion and HVSR curves, can provide a better constrained S-wave velocity model.

418 Figure (11) shows the comparison of both the best-fit models and models-10% with the S-wave  
419 velocity profile obtained by cross-hole (CH) measurements. In general, both the best-fit S-wave  
420 velocity profiles RHV and LRHV are in good agreement with the S-wave CH profile. However,  
421 interestingly, this comparison confirmed that by including the Love wave dispersion curve in the  
422 inversion scheme (Figure 11b) the S-wave velocities in the shallow layers can be better retrieved, as  
423 well as the S-wave velocity increase at about 16 m depth. Moreover, as also observed in Figure (8),  
424 when the Love wave dispersion curve is included in the analysis, the models-10% are considerably  
425 closer to the relevant best-fit model, indicating a higher capacity of the inversion scheme to  
426 constrain the optimal solution.

427

#### 428 **4 Conclusion**

429 In this work we introduce a joint inversion of Love wave dispersion, Rayleigh wave dispersion and  
430 HVSR curves using a data set of seismic noise recordings collected by a 2D array of seismic stations at  
431 the Bevagna (Italy) test-site.

432 In the first part of the work, we summarized the methods used for the estimation of the different  
433 surface wave derived curves, as well as for their inversion. In particular, we focused on the  
434 estimation of the Love wave dispersion curve using the ESAC approach on the horizontal component  
435 of the seismic noise ground motion.

436 In the second part of the work, we focused on the joint inversion analysis of the observed surface  
437 wave curves. In particular, we performed the inversion of real data using the joint inversion scheme  
438 proposed by Parolai et al. (2005), where only Rayleigh wave dispersion and HVSR are included, and  
439 by novel joint inversion scheme where the Love wave dispersion curves are also included. The  
440 comparison of the results from the two inversion analyses highlights that including the Love wave  
441 dispersion curves allows for the shallower portion of the S-wave velocity model to be better  
442 resolved. This point is confirmed by several observations. For example, when the joint inversion  
443 includes the Love wave dispersion curve, the family of models with misfit values close to the best-fit  
444 model of the inversion (i.e., the minimum plus the 10%) are distributed within a very narrow domain  
445 around the best-fit model itself, indicating that the solution of the inverse problem is very well  
446 constrained. Moreover, comparing the resolution matrixes for the final best fit models obtained by  
447 the RHV and LRHV schemes showed that when Love waves are considered, the shallower layers of  
448 the model can be better resolved. Finally, comparing the S-wave velocity models estimated following  
449 the two approaches with a-priori information from a cross-hole survey confirmed that the joint  
450 inversion of Love wave dispersion, Rayleigh wave dispersion, and HVSR curves can help to better  
451 constrain the S-wave velocity soil structure.

452 In fact, the Jacobian matrixes estimated with respect to the different observations clearly show how  
453 each of the inverted surface wave curves has a high sensitivity to a specific range of model  
454 parameters. In particular, Love waves provide more information about the shallower parts of the  
455 model, the Rayleigh waves allow the investigation of a deeper portion of the S-wave velocity profile,  
456 while the HVSR peak constrains the main impedance contrast. In conclusion, the three surface wave  
457 related curves appear to complement each other very well, and we believe that their joint inversion

458 within the framework of microzonation studies potentially allows a more detailed characterization of  
459 the subsoil structure.  
460

461 **Acknowledgments:**

462 We would like to thank Editor-in-chief K. Holliger D. Albarello and one anonymous reviewer for their  
463 comments and suggestions that allowed us to significantly improve the manuscript.

464 The data set used in this work is part of the INGV-DPC S4 Research Project “Italian Strong Motion  
465 Data Base”, promoted by the Istituto Nazionale di Geofisica e Vulcanologia (INGV) and funded by the  
466 Dipartimento della Protezione Civile (DPC) of the Italian Government (Agreement INGV-DPC 2007-  
467 2009). The authors wish to thank the DPC and the project team coordinators, Dr. Francesca Pacor  
468 and Prof. Roberto Paolucci, and Dr. R. Puglia for provided the data set. The INGV-Mi team (composed  
469 of Dr. Dino Bindi, Ezio D’Alema, Sara Lovati and Dr. Marco Massa) who previously executed the in situ  
470 microtremor measurements is also warmly acknowledged. K Fleming kindly improved the English.  
471

472  
473  
474  
475  
476  
477  
478  
479  
480  
481  
482  
483  
484  
485  
486  
487  
488  
489  
490  
491  
492  
493  
494  
495  
496  
497  
498  
499  
500  
501  
502  
503  
504  
505  
506  
507  
508  
509  
510  
511  
512  
513  
514  
515  
516  
517  
518  
519  
520

## REFERENCES

- Aki, K. (1957). Space and time spectra of stationary stochastic waves, with special reference to microtremors. *Bulletin of the Earthquake Research Institute* 35, 415–456.
- Aki, K. (1965). A note on the use of microseisms in determining the shallow structures of the earth's crust. *Geophysics* 29, 665–666.
- Arai, H. and K. Tokimatsu (2004). S-wave velocity profiling by inversion of microtremor h/v spectrum. *Bulletin of the Seismological Society of America* 94, 53–63.
- Arai, H. and K. Tokimatsu (2005). S-wave velocity profiling by joint inversion of microtremor dispersion curve and horizontal-to-vertical (h/v) spectrum. *Bulletin of the Seismological Society of America* 95, 1766–1778.
- Bard, P. Y. (1998). Microtremor measurement: a tool for site effect estimation? *Second International Symposium on the Effects of the Surface Geology on Seismic Motion ESG98, Japan*.
- Capon, J. (1969). High-resolution frequency-wavenumber spectrum analysis. *Proceedings of the IEEE* 57(8), 1408–1418.
- Chávez-García, F. J. and J. Lermo (1994). Are microtremors useful in site response evaluation? *Bulletin of the Seismological Society of America* 84, No. 5, pp. 1350-1364
- Deliverable, (2009). *Responsibles: Pacor, F. Paolucci, R.; edited by: Foti, S., Parolai, S. Albarello, D.; contributors: Albarello, D., Comina, C., Foti, S., Maraschini, M., Parolai, S., Picozzi, M., Puglia, R. Tokeshi, K. Technical report, of DPC-INGV S4 Project 2007-2009.*
- Field, E. H., and K. Jacob (1993). The theoretical response of sedimentary layers to ambient seismic noise, *Geophysical Research Letters* . 20–24, 2925–2928.
- Forbriger, T. (2006). Low frequency limit for H/V studies due to tilt, *Arbeitsgruppe-Seismology Session, 4–6 October, Haidhof*.
- Foti, S. Parolai, D. Albarello, and M. Picozzi (2011). Application of surface-wave methods for seismic site characterization. 2011, *Surv Geophys*. DOI 10.1007/s10712-011-9134-2.
- Goldberg, D.E. (1989) .*Genetic Algorithms in Search, Optimization and Machine Learning*, Addison-Wesley, Reading, MA.
- Kitsunezaki, C., N. Goto, Y. Kobayashi, T. Ikawa , M. Horike, T. Saito, T. Kurota, K. Yamane and K. Okuzumi (1990). Estimation of P- and S- wave velocities in deep soil deposits for evaluating ground vibrations in earthquake. *J. JSNDS*,9, 1-17.
- Köhler, A., M. Ohrnberger, F. Scherbaum, M. Wathelet, and C. Cornou (2007). Assessing the reliability of the modified three-component spatial autocorrelation technique. *Geophysical Journal International* . 168, 779–796
- Lacoss, R., E. Kelly, and M. Toksöz (1969). Estimation of seismic noise structure using arrays. *Geophysics* 34, 21–38.

521 Malagnini, L., A. Rovelli, S. Hough, and L. Seeber (1993). Site amplification estimates in the garigliano  
522 valley, central italy, based on dense array measurements of ambient noise. Bulletin of the  
523 Seismological Society of America 83, 1744–1755.  
524

525 Menke, W., (1989). Geophysical data analysis: Discrete inverse theory, Academic Press,  
526 ISBN 0-12-490921-3.  
527

528 Metaxian, J.-P., P. Leesa, and J. Dorel (1997). Permanent tremor of masaya volcano, nicaragua:Wave  
529 field analysis and source location. Journal of Geophysical Research 102,NO. B10, 529–545.66  
530

531 Mucciarelli, M. (1998). Reliability and applicability of Nakamura's technique using microtremors: an  
532 experimental approach, Journal of Earthquake Engineering. 2, 625-638.

533 Nakamura, Y. (1989). A method for dynamic characteristics estimations of subsurface using  
534 microtremors on the ground surface. Q. Rept. RTRI Jpn. 30, 25–33.

535 Nogoshi, M. and T. Igarashi (1971). On the amplitude characteristics of microtremor (part2).Journal  
536 Seism. Soc. Jpn. 24, 26–40.  
537

538 Ohori, M., A. Nobata, and K. Wakamatsu (2002). A comparison of esac and fk methods of estimating  
539 phase velocity using arbitrarily shaped microtremor arrays. Bulletin of the Seismological Society of  
540 America 92(6), 2323–2332.  
541

542 Okada, H. (2003). The Microtremor Survey Method. Geophysical monograph series 12.  
543

544 Parolai, S., P. Bormann, and C. Milkereit. (2001). Assessment of the natural frequency of the  
545 sedimentary cover in the Cologne area (Germany) using noise measurements, Journal of Earthquake  
546 Engineering. 5, 541–564.  
547

548 Parolai, S., M. Picozzi, S. M. Richwalski, and C. Milkereit (2005). Joint inversion of  
549 phase velocity dispersion and h/v ratio curves from seismic noise recordings using a genetic  
550 algorithm, considering higher modes. Geophysical Research Letters 32, L01303.  
551

552 Parolai, S., S. M. Richwalski, C. Milkereit, and D. Fäh (2006). S-wave velocity profiles for earthquake  
553 engineering purposes for the cologne area (germany). Bulletin of Earthquake Engineering 4, 65–94.  
554

555 Picozzi, M., and D. Albarello (2007). Combining Genetic and Linearized algorithms for a two-step joint  
556 inversion of Rayleigh wave dispersion and  $H/V$  spectral ratio curves. *Geophysical Journal  
557 International Vol. 169 Issue 1 Page 189 April 2007*

558 Puglia R., K. Tokeshi, M. Picozzi, E. D’Alema, S. Parolai, S. Foti (2011). Interpretation of microtremor  
559 2D array data using Rayleigh and Love waves: the case study of Bevagna (central Italy). Submitted to  
560 Near Surface Geophysics  
561

562 Scherbaum, F., K.-G. Hinzen, and M. Ohrnberger (2003). Determination of shallow shearwave velocity  
563 profiles in cologne, germany area using ambient vibrations. Geophysical Journal International 152,  
564 597–612.  
565

566 Tokimatsu, K., S. Tamura, and H. Kojima (1992). Effects of multiple modes on Rayleigh wave  
567 dispersion characteristics. Journal of Geotechnical Engineering 118, 1529–1543.  
568

569 Wang, R. (1999). A simple orthonormalization method for stable and efficient computation of  
570 Green's functions. Bulletin of the Seismological Society of America. 89 733–41  
571  
572 Yamanaka, H. and H. Ishida (1996). Application of generic algorithms to an inversion of surface-wave  
573 dispersion data. Bulletin of Earthquake Engineering 86, 436–444.  
574  
575 Zhang B, M Yu, Lan C Q and Xiong W 1996 Elastic waves and excitation mechanism of surface waves  
576 in multilayered media J. Acoust. Soc. Am. 100 3527–38  
577  
578



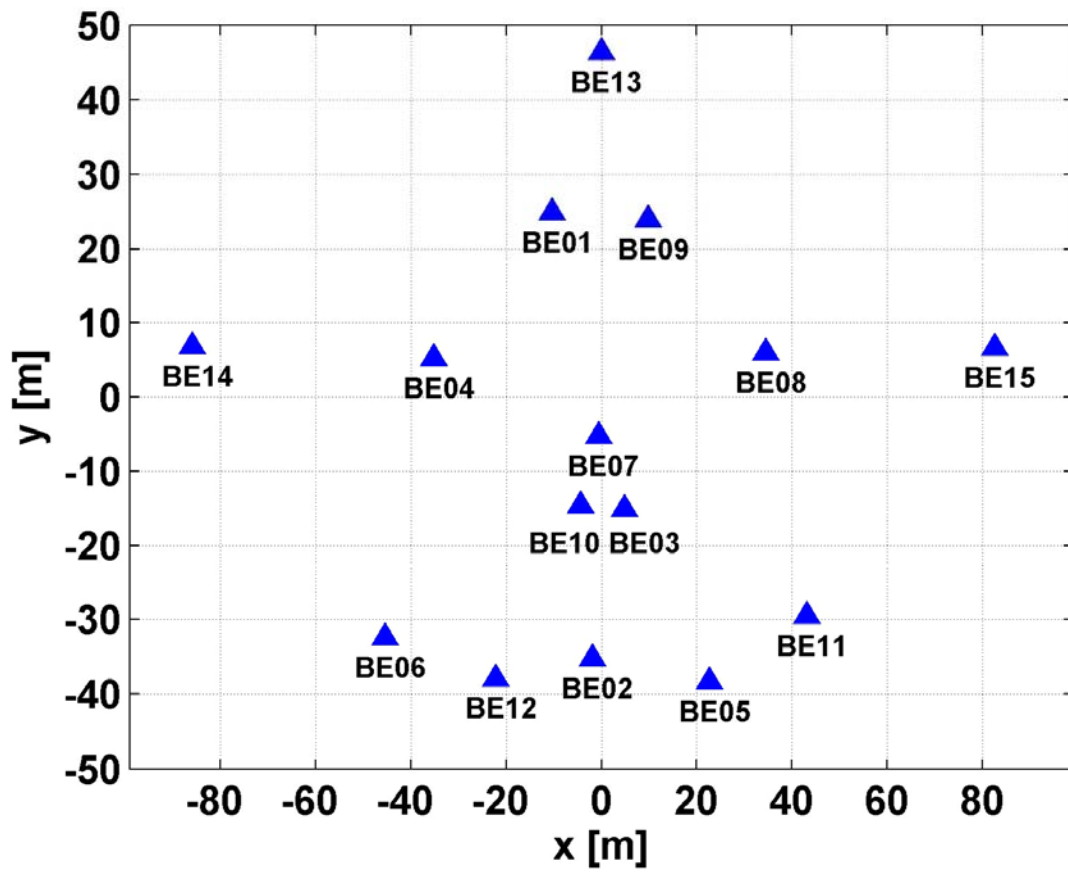
579  
580 **Tables**  
581

layer	thickness[m]		vs [m/s]	
	min	max	min	max
1	5.00	20.00	70.00	230.00
2	5.00	20.00	100.00	300.00
3	30.00	80.00	200.00	600.00
4	10.00	40.00	250.00	650.00

582

583 **Table 1:** Parameter range of depth [m] and S-wave velocity [m/s] for the joint inversion of the  
584 Bevagna measurement considering 4 different layers

585  
586  
587  
588  
589  
590  
591  
592  
593  
594  
595



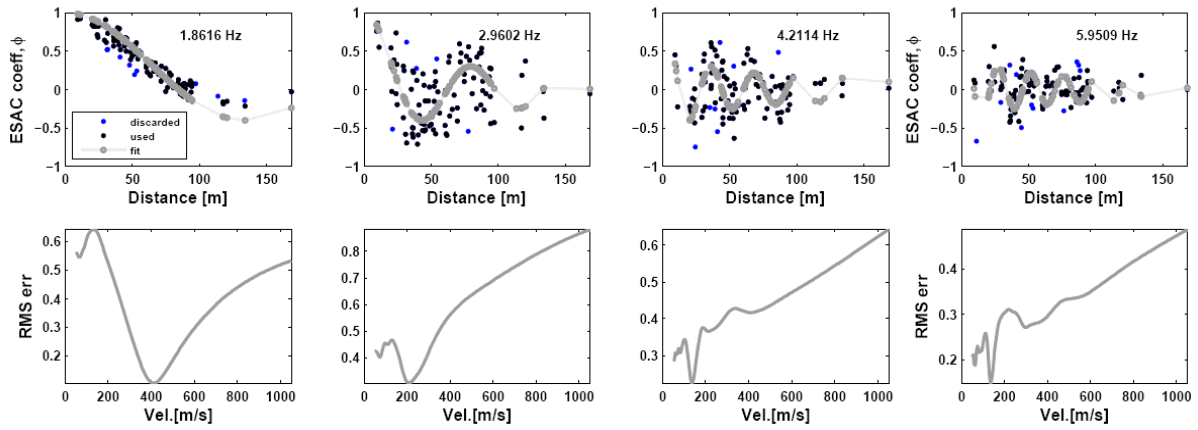
597  
598

599 **Figure 1:** Array geometry of the Bevagna test site measurement.  
600

601

602

603

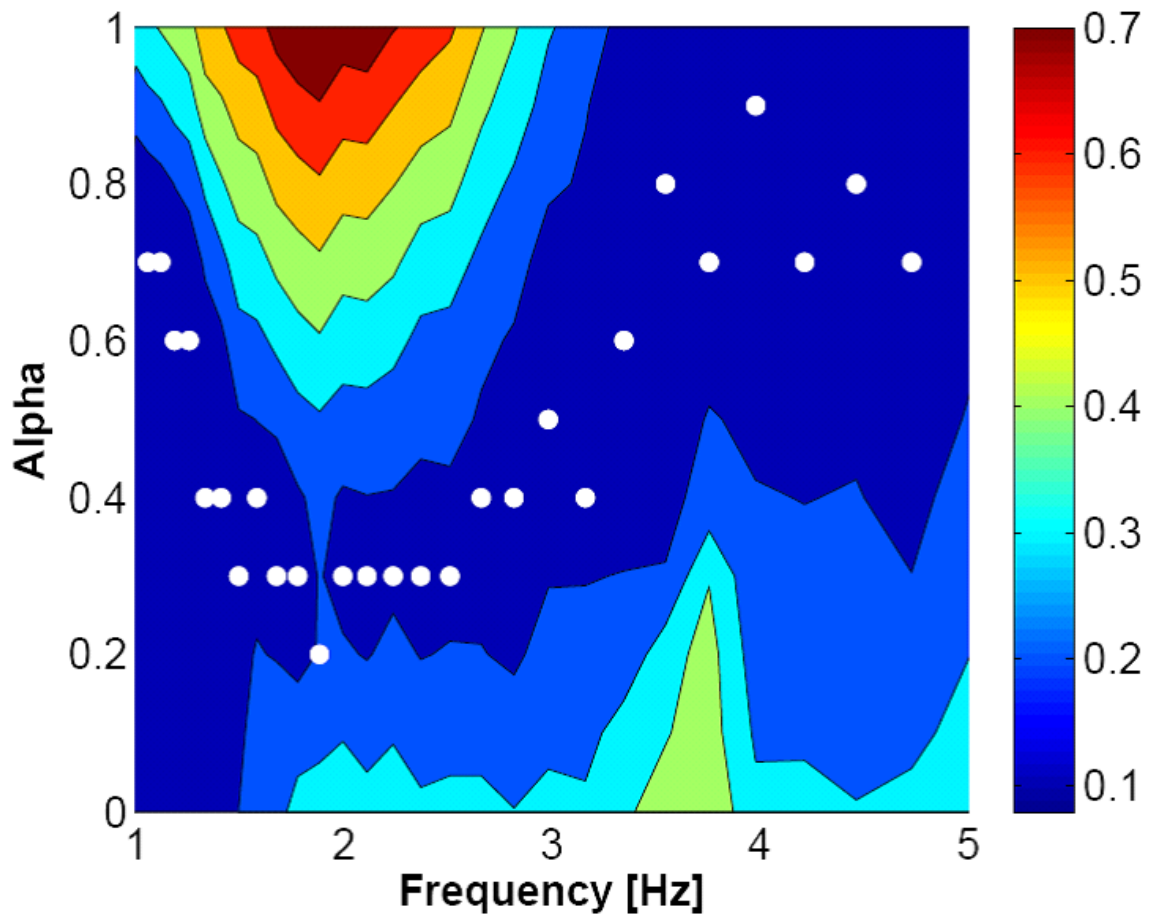


604  
605

606 **Figure 2:** Top row: ESAC function values (black circles) for different frequencies and the vertical  
607 component. The blue circles indicate the discarded values. The grey lines depict the best-fitting  
608 Bessel function. Bottom row: RMS error versus phase velocity curves.

609

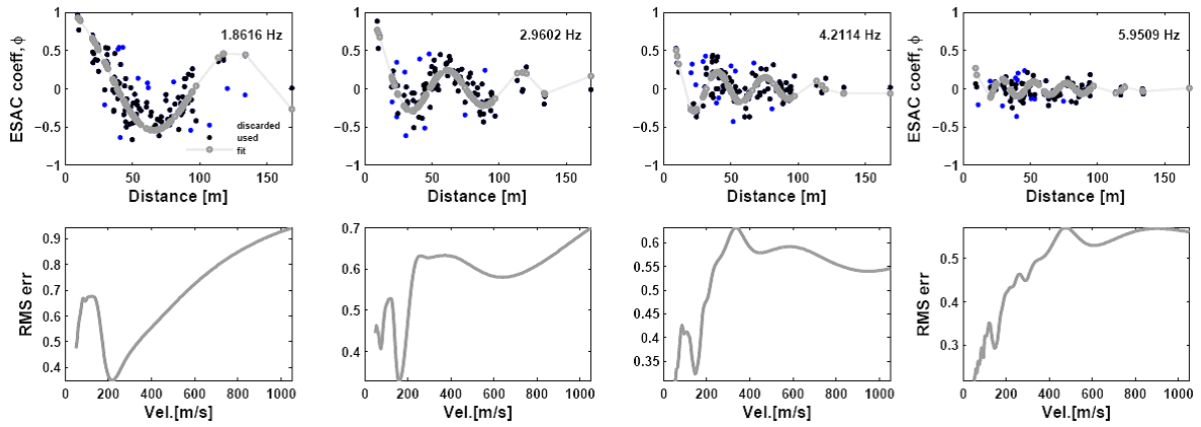
610



611  
612

613 **Figure 3:** Distribution of  $\alpha$  values (white dots) for different frequencies and the associated RMS  
614 error.  
615

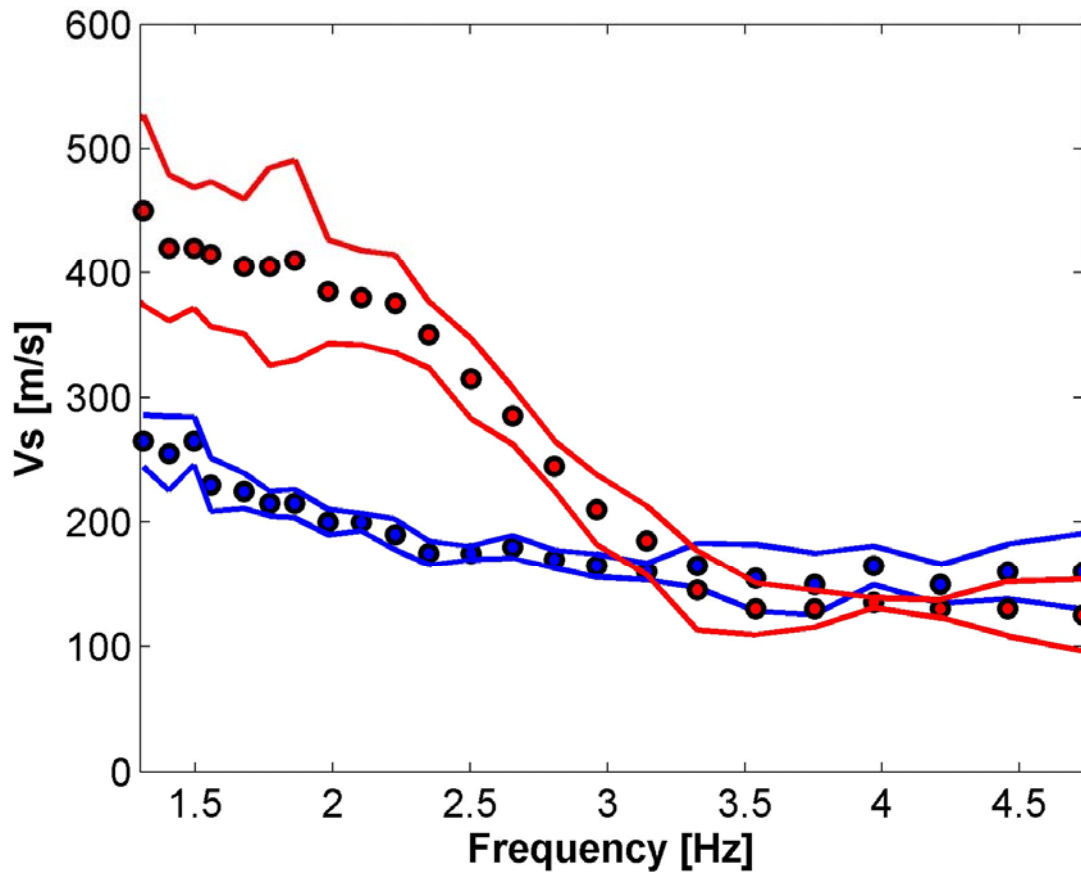
616



617  
618

619 **Figure 4:** Top row: ESAC function values (black circles) for different frequencies and the horizontal  
620 components. The blue circles indicate the discarded values. The grey lines depict the best-fitting  
621 Bessel function. Bottom row: RMS error versus phase velocity curves.  
622

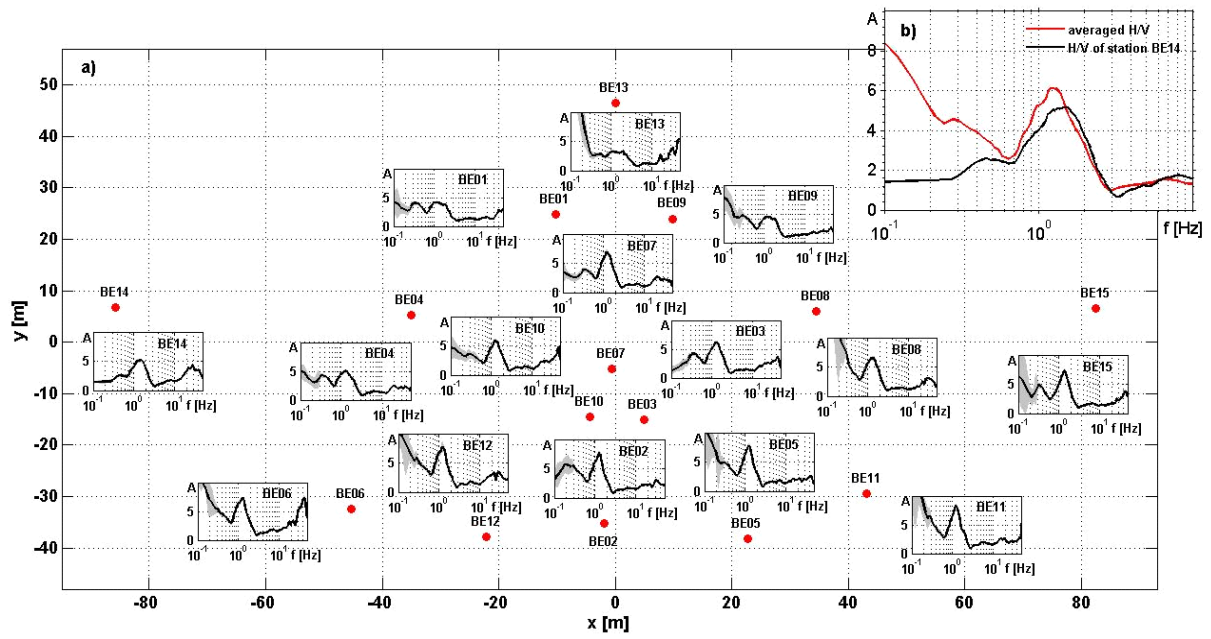
623



624  
625

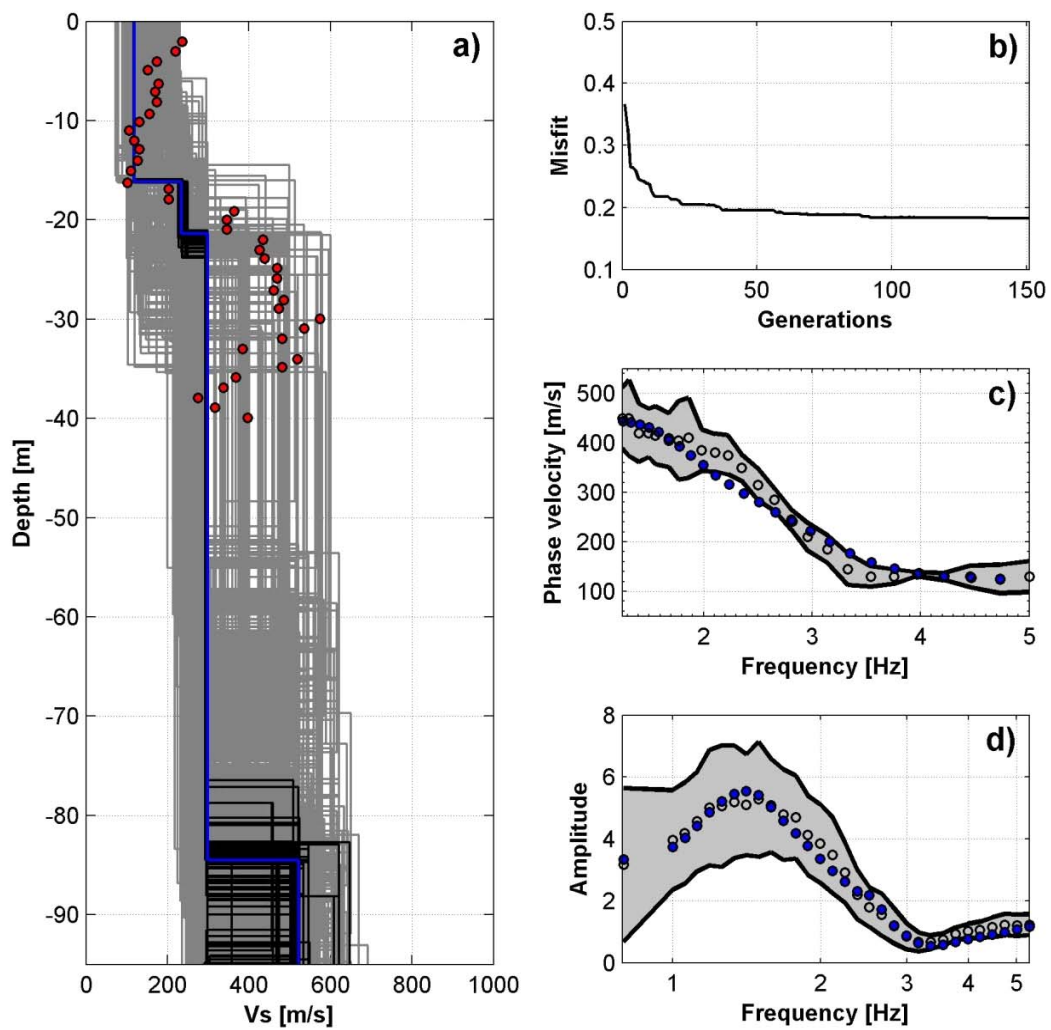
626 **Figure 5:** Rayleigh (red) and Love (blue) wave dispersion curves.  
627

628



629  
630

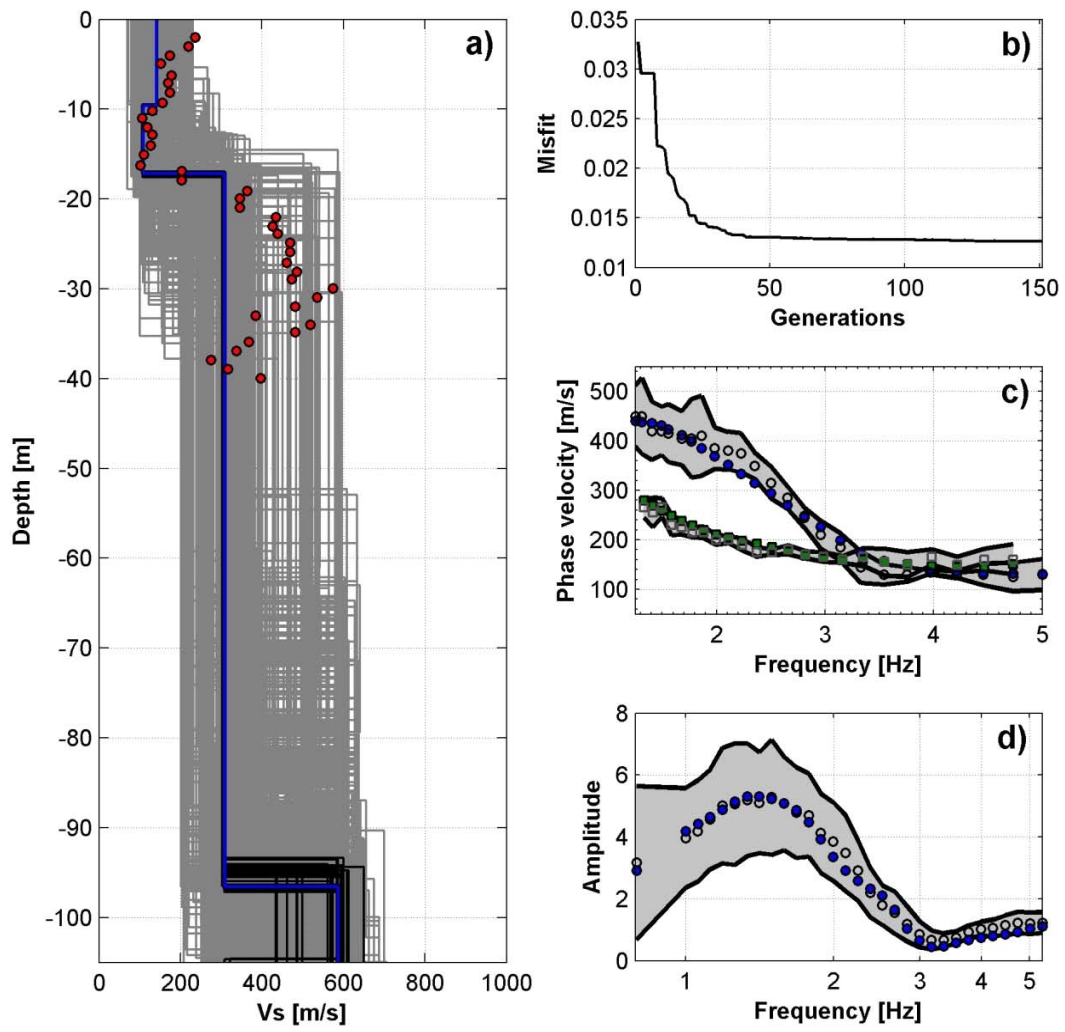
631 **Figure 6:** H/V spectral ratio curves of the array microtremor measurements from Bevagna.  
632 a) H/V curves of all stations together with their positions. b) H/V curve of station BE14 compared  
633 with the averaged H/V curve of all stations.



634  
635

636 **Figure 7:** a) S-wave velocity profile from the RHV joint inversion; best fit model (blue line), models in  
 637 a range of the best fit model + 10 % (black lines), all models (gray lines) b) misfit for all 150  
 638 generations c) observed (open circles) and calculated (blue circles) Rayleigh wave dispersion curve,  
 639 standard deviation values (grey lines) of the observed dispersion curve d) observed (open circles) and  
 640 calculated (blue circles) H/V spectral ratio, standard deviation values (grey lines) of the observed H/V  
 641 curve.

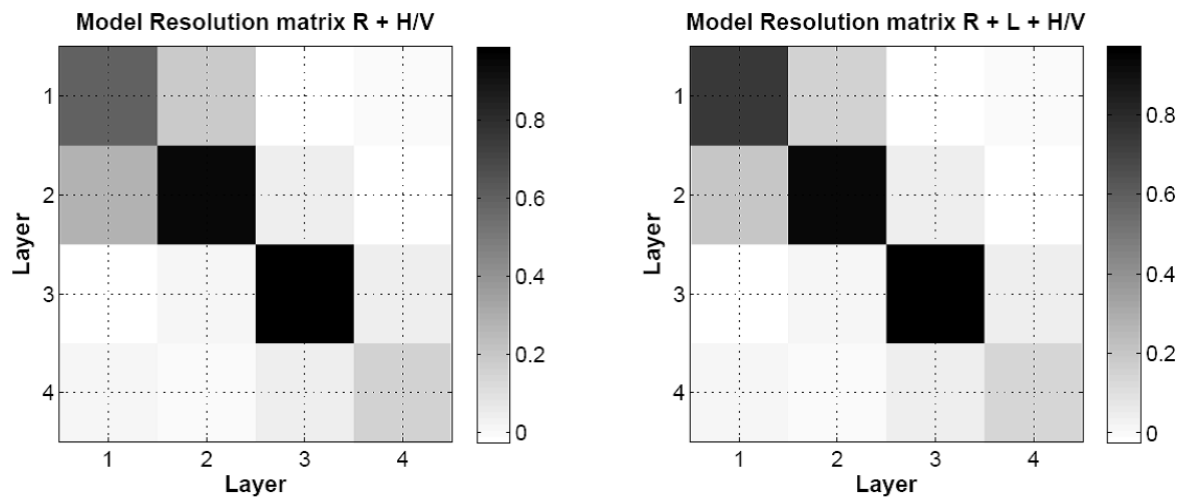




642  
643

644 **Figure 8:** see Figure 7; additional c) observed (gray squares) and calculated (green squares) Love  
645 wave dispersion curve.  
646

647



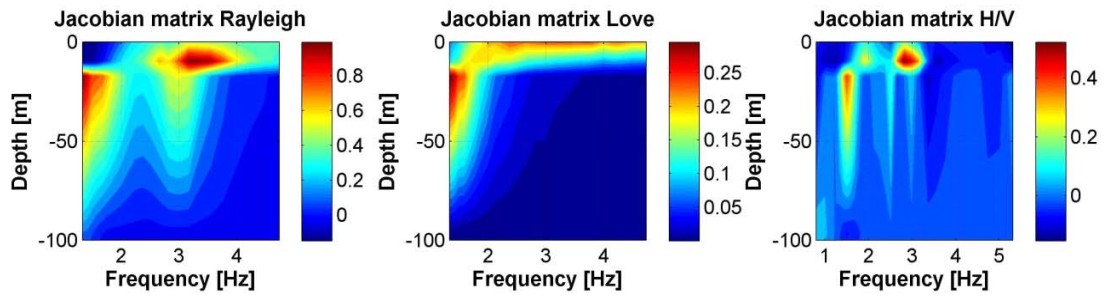
648

649

Figure 9: Model resolution matrices: left) RHV inversion, right) RLHV inversion.

650

651



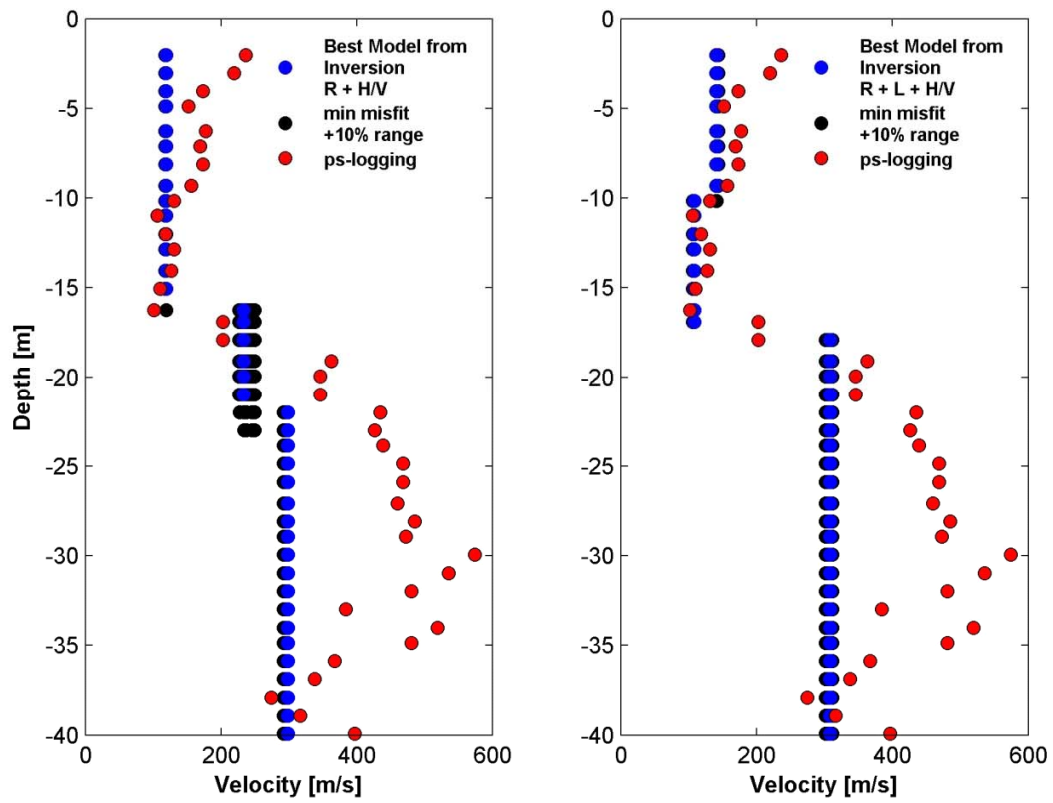
652

653

**Figure 10:** Jacobian matrix for Rayleigh and Love wave dispersion curve and H/V curve.

654

655



656

657

**Figure 11:** Best fit models and models in a range of 10% of the best fit model in comparison to CH data for RHV (left) and RLHV (right) inversion.

658

659

660