



Originally published as:

Matmon, A., Fink, D., Davis, M., Niedermann, S., Rood, D., Frumkin, A. (2014): Unraveling rift margin evolution and escarpment development ages along the Dead Sea fault using cosmogenic burial ages. - *Quaternary Research*, 82, 1, p. 281-295.

DOI: <http://doi.org/10.1016/j.yqres.2014.04.008>

# Unraveling rift margin evolution and escarpment development ages along the Dead Sea fault using cosmogenic burial ages

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## Abstract

The Dead Sea fault (DSF) is one of the most active plate boundaries in the world. Understanding the Quaternary history and sediments of the DSF requires investigation into the Neogene development of this plate boundary. DSF lateral motion preceded significant extension and rift morphology by ~10 Ma. Sediments of the Sedom Formation, dated here between  $5.0 \pm 0.5$  Ma and  $6.2_{-2.1}^{+inf}$  Ma, yielded extremely low  $^{10}\text{Be}$  concentrations and  $^{26}\text{Al}$  is absent. These reflect the antiquity of the sediments, deposited in the Sedom Lagoon, which evolved in a subdued landscape and was connected to the Mediterranean Sea. The base of the overlying Amora Formation, deposited in the terminal Amora Lake which developed under increasing relief that promoted escarpment incision, was dated at  $3.3_{-0.8}^{+0.9}$  Ma. Burial ages of fluvial sediments within caves ( $3.4 \pm 0.2$  Ma and  $3.6 \pm 0.4$  Ma) represent the timing of initial incision. Initial DSF topography coincides with the earliest Red Sea MORB's and the East Anatolian fault initiation. These suggest a change in the relative Arabian-African plate motion. This change introduced the rifting component to the DSF followed by significant subsidence, margin uplift, and a reorganization of relief and drainage pattern in the region resulting in the topographic framework observed today.

## Introduction

Temporal frameworks of rift morphology initiation and formation of rift-margin escarpments are difficult to establish. Sometimes, these can be provided based on seismic stratigraphy,

35 radiometric dating, lithostratigraphy and/or biostratigraphy (e.g. Summerfield, 1991a,b; Ollier,  
36 1995; Rosendahl, 1987; Ebinger, 1989; Ebinger et al., 1987). One of the youngest, most active,  
37 and studied rifts in the world is located along the Dead Sea fault (DSF) which is part of the  
38 Arabian-African plate boundary. The DSF, defined geographically from the Gulf of Aqaba to the  
39 Lebanon Mountains, connects the spreading zones of the Red Sea and Gulf of Aden with the  
40 convergence belts along the northern Mediterranean Sea, Turkey, and Iran (Fig. 1). The  
41 evolution of the Arabian plate boundaries is inherently related to the way it translates relative  
42 to the African plate. As such, the DSF provides the opportunity to understand the temporal  
43 relations between tectonic activity (i.e. motion along faults) and the development of  
44 topography. To understand the present relations between tectonic activity and morphology  
45 and the way they developed throughout the Quaternary it is essential to go back several million  
46 years, describe the major phases in the tectonic activity of the DSF, and determine the time and  
47 processes that triggered major morphologic development. Specifically, the Quaternary  
48 landscape evolution of the western margin of the DSF stems from the Neogene tectonic  
49 evolution of the Dead Sea plate boundary and the Pleistocene-Holocene development of  
50 drainage systems can be understood only in the light of the preceding tectonic processes.

51

### 52 ***Sediments and Paleohydrology of the Dead Sea basin***

53

54 The sediments deposited in the water bodies of the Dead Sea basin include coarse to fine  
55 clastic sediments, dolomites, and evaporates such as aragonite, gypsum, and halite (Zak, 1967;  
56 Begin et al., 1974; Ken-Tor et al., 2001a,b). These sediments reflect the environmental  
57 conditions that existed during the life span of each lake and comprise the Dead Sea Group (Zak,  
58 1967), which includes, from oldest to most recent, the Sedom, Amora, Lisan, and Ze'elim  
59 Formations (Fig. 2). The earliest known water body in the DSF is the Sedom Lagoon, an  
60 elongated gulf connected to the Mediterranean Sea through the Yizre'el Valley (Neev and  
61 Emery, 1967; Zak, 1967) (Fig. 3). The Sedom Formation, deposited in the Sedom Lagoon,  
62 consists of a series of 1500-2000 m thick evaporites, shales, dolomites, and sand. Since the end  
63 of the Sedom lagoonal stage, the Dead Sea rift valley has been occupied by terminal lakes of  
64 which Lake Amora, Lake Lisan and the Dead Sea are the better-studied ones (Katz and Starinsky,

65 2009).

66 In the Miocene, during the initial stages of Dead Sea fault activity, extension was not significant  
67 enough to promote development of a deep rift valley with uplifted rift shoulders. The shallow  
68 valley which then existed along the Dead Sea fault accommodated intermittent lagoons that  
69 extended to the Mediterranean Sea (Katz and Starinsky, 2009). Nevertheless, in some places  
70 rivers that originated hundreds of kilometers east of the plate boundary flowed to the west and  
71 drained into the Mediterranean (Neev, 1960; Garfunkel and Horowitz, 1966; Calvo and Bartov,  
72 2001; Calvo, 2002). Overall, the landscape and drainage systems were continuous across the  
73 Dead Sea plate boundary. The morphological barriers of the valley and the main water divide  
74 between the Dead Sea and the Mediterranean Sea did not exist at these times. At some point  
75 during the Plio-Pleistocene the scale of rift subsidence combined with uplift along the western  
76 margin was sufficient to cause collapse of the contemporary drainage pattern and create a  
77 continuous water divide between the evolving rift and the Mediterranean Sea (Picard, 1943;  
78 Garfunkel and Horowitz, 1966; Zilberman et al., 1996; Wdowinski and Zilberman, 1997; Avni et  
79 al., 2000). Drainage systems developed on both sides of this established water divide:  
80 westward-flowing streams that flowed to the Mediterranean Sea, and a system of eastward-  
81 flowing streams that incised deeply in response to significant subsidence in the Dead Sea area.

82

### 83 ***Geochronology of the Dead Sea basin***

84

85 Our general understanding of the evolution of the Dead Sea fault shows that lateral motion  
86 initiated about 18 – 14 Ma (Bartov et al, 1980) but that subsequent development of rift  
87 morphology and relief formation occurred many millions of years later. Earlier studies suggest  
88 that the establishment of the present relief and drainage systems along the western margin of  
89 the DSF occurred more recently during the early Pliocene and into the Quaternary (e.g.  
90 Matmon et al., 1999, 2000; Guralnik et al., 2010). This long-scale chronology is based on  
91 regional geologic correlations (Zak, 1967; Shaliv, 1991) and few radiometric ages (Steinitz and  
92 Bartov, 1991) of basalts that limit the time of formation of presently existing canyons. During  
93 the Pliocene, increased extension along the DSF resulted in development of a deep inland  
94 drainage basin (Garfunkel, 1981). The initiation of major relief formation after millions of years

95 of dominant lateral slip may be an expression of significant changes in the relative motion  
96 between the Arabian and African plates. Similar geomorphic processes of basin evolution have  
97 been reported from other extensional provinces around the world (e.g. Summerfield, 1991a,b;  
98 Ollier, 1995; Zelilidis, 2000; Mack et al., 2006).

99 Previous age estimates of the Sedom Formation range between 12 and 3 Ma, based on  
100 stratigraphical-paleogeographical and geochemical considerations (Zak, 1967; Agnon, 1983a,  
101 1993; Shaliv, 1989; Shaliv, 1991; Steinitz and Bartov, 1991; Stein et al., 1994; Stein et al., 2000;  
102 Horowitz, 2001; Stein and Agnon, 2007; Torfstein, 2008; Ryb et al., 2009a). The fossil  
103 assemblage in the formation does not enable its dating (Zak, 1967). Current efforts using  
104 meteoric <sup>10</sup>Be concentrations in the Sedom Formation halite place it in the early to middle  
105 Pliocene (Belmaker et al., 2013).

106 The the Amora Formation, includes the lacustrine sequence overlying the Sedom Formation  
107 (Torfstein et al., 2009; Zak, 1967). The Amora Formation exposure at the eastern flanks of Mt.  
108 Sedom is ~300 m thick and consists of alternating aragonite–detritus laminae (Marco, 1996),  
109 halite, and clastic sediments. Age estimates based on different techniques for the Amora  
110 Formation outcrop are all <1 Ma (Zak, 1967; Kaufman, 1971; Torfstein et al., 2009). The  
111 unexposed base of the Amora Formation and the age of the Sedom-Amora contact have yet to  
112 be dated. Overall, the lithology of the Amora section is similar to that of overlying sediments  
113 comprising the Pleistocene Lake Lisan and Holocene Dead Sea (e.g. Katz and Starinsky, 2009).  
114 Based on geological considerations and field relation, Zak (1967) estimated the age of the  
115 Amora Formation to be between 1 Ma and ~100 ka. Kaufman (1971) provided three U–Th ages  
116 that range between 380 ka and 190 ka. Torfstein et al. (2009) dated the exposed upper part of  
117 the Amora Formation using the U-disequilibrium dating method to range between ~500 ka  
118 (their stratigraphically lowest sample) and ~150 ka (their stratigraphically highest sample).

119 The time at which the Sedom Formation ceased sedimentation and the Amora Formation  
120 started being deposited highlights a special and significant point in the morphotectonic  
121 evolution of the Dead Sea plate boundary. Until then, sedimentation in the basins competed  
122 with subsidence, Mediterranean waters could flow into the DSF valley through structural and  
123 topographic lows, and some major rivers continued flowing across the fault westward (Neev,  
124 1960; Zilberman, 1993). From that time on, the rate of subsidence outpaced the rate of

125 sedimentation, forcing the Dead Sea valley to become an inland endorheic basin whose base  
126 level directed river flow into it, effectively disconnecting transfer between the Mediterranean  
127 Sea and regions east of the plate boundary (Katz and Starinsky, 2009). This change set the scene  
128 for the Quaternary evolution of DSF morphology and to the establishment of the current  
129 drainage pattern along the margins of the rift. Determining when this shift occurred is needed  
130 for the overall understanding of the evolution of this major plate boundary.

131

### 132 ***Cosmogenic burial dating***

133

134 The application of terrestrial cosmogenic nuclides (TCN) provides insight into the important  
135 evolutionary stage in the DSF history that occurred during the Pliocene and controlled the  
136 Quaternary landscape evolution of this region. Cosmogenic radionuclides such as  $^{26}\text{Al}$  and  $^{10}\text{Be}$   
137 (0.705 and 1.39 Ma half lives, respectively; Chmeleff et al., 2010; Korschinek et al., 2010;  
138 Nishiizumi, 2004) and stable  $^{21}\text{Ne}$  are produced in-situ within minerals at the Earth's surface  
139 (e.g. Gosse and Phillips, 2001; Niedermann, 2002). After sufficient exposure, production  
140 balances loss via the sum of decay and erosion, whereby radionuclide concentrations reach a  
141 steady-state or secular equilibrium value which is effectively a function of the erosion rate (e.g.  
142 Granger, 2006):

$$143 \quad N_i = P_i / \left[ \left( \frac{1}{\tau_i} \right) + \frac{\rho \varepsilon}{\Lambda} \right] \quad (1)$$

144 where  $P_i$  is the production rate of nuclide  $i$  at the site of interest (atoms  $\text{yr}^{-1} \text{g}^{-1}$  quartz),  $N_i$  is the  
145 concentration (atoms  $\text{g}^{-1}$  quartz),  $\tau_i$  is the mean life (yr),  $\varepsilon$  is the erosion rate ( $\text{cm yr}^{-1}$ ),  $\Lambda$  is the  
146 attenuation length ( $\text{g cm}^{-2}$ ), and  $\rho$  is the rock density ( $\text{g cm}^{-3}$ ).

147 Cosmogenic burial dating employs the measurement of paired terrestrial cosmogenic nuclides  
148 in buried sedimentary quartz grains. When quartz grains that are dosed with cosmogenic  
149 nuclides are transported and buried to a sufficient depth (generally more than several tens of  
150 meters) that shields cosmic radiation, production ceases and only radioactive decay controls  
151 their concentration ratio. Thus, due to the shorter half-life of  $^{26}\text{Al}$  compared to  $^{10}\text{Be}$ ,  
152 the  $^{26}\text{Al}/^{10}\text{Be}$  ratio decreases as a function of burial time. The  $^{26}\text{Al}/^{10}\text{Be}$  surface production rate  
153 ratio is 6.75 (Balco and Shuster, 2009). For most environments, the concentration ratio  
154 of  $^{26}\text{Al}/^{10}\text{Be}$  is identical to the production rate ratio. Hence, for most cases we can assume that

155 the initial  $^{26}\text{Al}/^{10}\text{Be}$  ratio is  $\sim 6.75$ . It should be noted that for quartz sourced from very slow  
 156 eroding terrains or extended ( $>1\text{Ma}$ ) continuous exposure, the  $^{26}\text{Al}/^{10}\text{Be}$  initial ratio can  
 157 decrease to as low as  $\sim 3.45$ . Nevertheless, in most cases measured  $^{26}\text{Al}/^{10}\text{Be}$  ratios that are  
 158 lower than 6.75 suggest long periods of burial. The time dependent  $^{26}\text{Al}/^{10}\text{Be}$  ratio during burial  
 159 can be described by the equation (Granger et al., 1997):

$$160 \left(\frac{N_{26}}{N_{10}}\right)_{t_b} = \left(\frac{N_{26}}{N_{10}}\right)_0 e^{-t_{burial}\left(\frac{1}{\tau_{26}} - \frac{1}{\tau_{10}}\right)} \quad (2)$$

161 where  $\left(\frac{N_{26}}{N_{10}}\right)_{t_b}$  is the ratio of  $^{26}\text{Al}$  and  $^{10}\text{Be}$  after a burial period of time  $t_b$ ,  $\left(N_{26}/N_{10}\right)_0$  is the  
 162 initial  $^{26}\text{Al}/^{10}\text{Be}$  ratio at time of burial,  $t_b$  is the time since burial, and  $\tau_{26}$  and  $\tau_{10}$  are the mean  
 163 lives of  $^{26}\text{Al}$  and  $^{10}\text{Be}$  ( $\tau_{26}=1.01\pm 0.04\text{ Ma}$  and  $\tau_{10}=2.00\pm 0.08\text{ Ma}$ ), respectively. Equations 1 and 2  
 164 can be solved iteratively to yield the burial age and the source erosion rate (Granger et al.,  
 165 1997). The  $^{21}\text{Ne}/^{10}\text{Be}$  production ratio is 4.08, and the  $^{21}\text{Ne}/^{26}\text{Al}$  production ratio is 0.606 (Balco  
 166 and Shuster, 2009). These ratios also change during burial since  $^{21}\text{Ne}$  does not decay while the  
 167 two other nuclides do. In this paper we use cosmogenic isotope burial dating to explore one of  
 168 the major stages in the evolution of the DSF: the stage at which the rift became an inland base  
 169 level disconnected from the Mediterranean and topography initiated along its margins.

170 We present TCN burial ages for three groups of samples: 1) samples from the top of the Sedom  
 171 Formation, representing the final stages of deposition from water bodies that were connected  
 172 to the Mediterranean Sea (labeled as SDM samples), 2) samples from the base of the Amora  
 173 Formation, representing the initial stages of deposition from water bodies that were not  
 174 connected to the Mediterranean Sea and were confined to the Dead Sea rift valley (labeled as  
 175 AMZ samples), and 3) samples from two caves. The Cave of the Letters (labeled as COL samples)  
 176 is close to the top of the canyon cliffs of Wadi Hever, a major canyon flowing into the Dead Sea.  
 177 The Masada cave (labeled as the MZ sample) is located close to the top of the escarpment  
 178 flanking the Dead Sea on the west. Buried sediments within these caves should post-date SDM  
 179 samples and be similar to AMZ samples, as they represent the initial stages of relief forming  
 180 and incision along the Dead Sea escarpment, that followed subsidence of the Dead Sea basin  
 181 and the uplift of its western margin. In the following sections we compare burial ages using  
 182 different isotopic pairs ( $^{26}\text{Al}/^{10}\text{Be}$ ,  $^{26}\text{Al}/^{21}\text{Ne}$ ,  $^{10}\text{Be}/^{21}\text{Ne}$ ) and highlight the difficulties when  
 183 interpreting such cosmogenic nuclide data pertaining to ancient landscapes and sediments with

184 burial ages in the Plio-Pleistocene age range.

185

### 186 **Site description and sampling strategy**

187

188 The Sedom Formation is exposed in Mount Sedom, a rising salt diapir located within the Dead  
189 Sea rift valley along the southwestern shores of the Dead Sea, Israel (Weinberger et al., 2006)  
190 (Figs. 3,4). The diapir is rising at a rapid rate of 5-9 mm yr<sup>-1</sup> (Frumkin, 2009; Pe'eri et al., 2004).  
191 Four samples (SDM 2,3,4,5; Fig. 4) were collected from a terrigenous quartz-rich sand layer,  
192 located about 250 m below the Sedom and Amora Formation boundary.

193 The samples were collected from a dry ravine incised ~20 m into the soft sediments. To further  
194 eliminate the interference of recent cosmogenic nuclide production, loose streambed material  
195 was removed and an additional 0.5-1 m were dug into the soft sediment before the quartz was  
196 collected. Some of the samples were dug from the base of small knickpoints (1-2 m) increasing  
197 the total overburden.

198 The Amiaz-1 deep borehole is located west of Mt. Sedom (Figs. 3,4) and penetrates 3.4 km of  
199 basin fill (Kashai and Croker, 1987). Three medium-coarse quartz sand samples (AMZ 1,2,3) of  
200 fluvial origin were obtained for TCN burial dating. The sampled section is attributed to the base  
201 of the Amora Formation (Kashai and Croker, 1987) and lies immediately above evaporates of  
202 the Sedom Formation. TCN burial ages of the Amiaz-1 borehole samples provide an age for the  
203 base of the Amora Formation and a minimum age constraint for the top of the Sedom  
204 Formation. Sedom Formation sand beds, such as those found in the Benot Lot member, were  
205 not available from this core.

206 The Cave of the Letters is located 160 m above sea level, 3 km west of the Dead Sea at the  
207 northern wall of Wadi Hever (Figs. 3,5), a deep canyon draining into the Dead Sea. The thalweg  
208 of Wadi Hever canyon lies 230 m below the Cave of the Letters and the top of the escarpment  
209 is 70 m above the cave. The Cave of the Letters is located within Turonian limestone, and cave  
210 sediments consist of autogenic dolomite, collapsed rocks and sorted fine-grained clastics  
211 (Frumkin, 2001) (Fig. 5) which are moderately bedded to laminated, with cross bedding  
212 structures. Individual sand grains are well-rounded, indicating transport by water. The cave  
213 origin is hypogene without a vadose stream, hence fluvial deposits within the cave must



214 originate from sediments transported by Wadi Hever as it incised and truncated the pre-  
215 existing cave (Frumkin, 2001). Biogenic activity in the form of wasp drill holes (ranging in  
216 diameter from a few mm to 1 cm) is apparent in the soft sediment in the cave (Fig. 5) as well as  
217 in the exposed sediments along the wall of the canyon. Although this disturbance is recognized,  
218 we collected samples from this cave for cosmogenic burial dating. Five samples were collected,  
219 four from the top of the sedimentary column, deep within the cave passages, 10-21 m from the  
220 entrance (COL 1,2,5,7), and one from the face of exposed sediment along the external wall of  
221 the canyon (COL 3). We avoided sampling sections that showed clear evidence of wasp drilling.  
222 However, we assume that such biogenic activity has continued since the sediment was  
223 excavated to allow easier passage into the back of the cave, ~2000 yr ago, and therefore,  
224 consider all our results from the Cave of the Letters as minimum ages because the ancient  
225 buried sediment may have been mixed with modern silt.

226 The Masada cave is located on the main escarpment of the Dead Sea below the ruins of the  
227 archeological site of Masada (Fig. 6). The cave is located in Turonian carbonate rocks, 10 m  
228 above sea level, ~20 m below the upper desert surface and about 400 m above the Dead Sea.  
229 Similar to the COL setting, the Masada cave was exposed during the initial stages of Dead Sea  
230 escarpment formation. Shoreline sediments were washed into the cave when it was at lake  
231 level and since then have been buried within the cave and preserved from erosion. The burial  
232 age of these shoreline sediments correlates to the initial stages of Dead Sea escarpment  
233 formation. We collected one sediment sample (M22) from a sandy lens, composed mostly of  
234 chert fragments from the back of the cave (Fig. 6) where the sediment is well shielded from the  
235 opening of the cave entrance and its overburden is the thickest.

236 All samples were sieved and the 250-850  $\mu\text{m}$  size fraction was further processed. Quartz was  
237 separated after carbonate dissolution in warm HCl (18%) and magnetic separation. Quartz was  
238 further etched three times in a 2.5% HF:1% HNO<sub>3</sub> solution. Extraction of Al and Be followed  
239 standard techniques described in Bierman and Caffee (2001). Three samples from the Sedom  
240 Formation were analyzed for <sup>10</sup>Be/<sup>9</sup>Be and <sup>26</sup>Al/<sup>27</sup>Al ratios at the ANTARES Accelerator Mass  
241 Spectrometry (AMS) facility at the Australian Nuclear Science and Technology Organization  
242 (ANSTO) (Fink and Smith, 2007). All other samples were analyzed at the Lawrence Livermore  
243 National Laboratory (LLNL) (Rood et al., 2010). Two Amora Formation samples (AMZ-1 and

244 AMZ-3) and two Cave of the Letter samples (COL-1 and COL-2) were analyzed for  $^{21}\text{Ne}$  at  
245 Deutsches GeoForschungsZentrum (GFZ) in Potsdam (Niedermann et al., 1997). Stable Al was  
246 measured using an ICP-OES at the Hebrew University.

247

## 248 **Results**

249

250 In the upper part of the Sedom Formation,  $^{10}\text{Be}$  concentrations for SDM 3,4,5 are low and range  
251 between  $1.1\pm 0.2\times 10^4$  and  $1.7\pm 0.5\times 10^4$  atoms  $\text{g}^{-1}$  quartz (Table 1).  $^{26}\text{Al}/^{27}\text{Al}$  ratios for these 3  
252 samples are effectively indistinguishable from the  $^{26}\text{Al}/^{27}\text{Al}$  ratio measured in the blank. Thus,  
253 their  $^{26}\text{Al}/^{10}\text{Be}$  ratios are below current AMS detection limits indicating the extreme antiquity of  
254 these sediments. Sample SDM-2 yielded a  $^{26}\text{Al}/^{27}\text{Al}$  ratio marginally above background, albeit  
255 with a large uncertainty. This ratio corresponds to a  $^{26}\text{Al}$  concentration of  $0.5\pm 0.5\times 10^4$  atoms  $\text{g}^{-1}$   
256 quartz. Unfortunately, the  $^{10}\text{Be}/^9\text{Be}$  ratio could not be measured for this sample due to a  
257 laboratory problem.

258  $^{10}\text{Be}$  concentrations in the Amiaz-1 borehole samples range between  $1.0\pm 0.1\times 10^4$  and  
259  $1.5\pm 0.2\times 10^4$  atoms  $\text{g}^{-1}$  quartz.  $^{26}\text{Al}$  concentrations are also very low, although in contrast to  
260 results for SDM samples, provide measurable  $^{26}\text{Al}$  distinguishable from background, and range  
261 between  $2.0\pm 0.7\times 10^4$  and  $2.8\pm 0.7\times 10^4$  atoms  $\text{g}^{-1}$  quartz (Table 1). Most of the uncertainty is  
262 attributed to  $^{26}\text{Al}/^{27}\text{Al}$  analytical uncertainties that reach 30%.  $^{26}\text{Al}/^{10}\text{Be}$  ratios range between  
263  $1.44\pm 0.45$  and  $2.09\pm 0.80$ .

264 Cosmogenic  $^{21}\text{Ne}$  concentrations in two of the Amiaz-1 borehole samples (AMZ-1, AMZ-3) are  
265  $6.7\pm 0.7\times 10^6$  atoms  $\text{g}^{-1}$  quartz and  $5.3_{-1.2}^{+1.5}\times 10^6$  atoms  $\text{g}^{-1}$  quartz (Table 1). The uncertainties  
266 in  $^{21}\text{Ne}$  concentrations are large (between 10% and 28%) and are due to large amounts of  
267 trapped atmosphere-like Ne (high  $^{20}\text{Ne}$  concentrations), and subsequently to rather small  
268 excesses in  $^{21}\text{Ne}/^{20}\text{Ne}$  resulting from cosmic ray production. In the Ne three-isotope diagram of  
269 the AMZ samples (Fig. 8a), the isotope ratios for extractions up to  $800^\circ\text{C}$  follow a line parallel to  
270 the expected spallation line (i.e., the mixing line of atmospheric and cosmogenic Ne;  
271 Niedermann et al., 1993) through the data point representing the crushing extraction of AMZ-1,  
272 except for the  $400^\circ\text{C}$  step of AMZ1 which contributes only  $\sim 2\%$  to the total  $^{21}\text{Ne}$  excess up to  
273  $800^\circ\text{C}$  for this sample. These data are thus consistent with a two-component mixture of

274 fractionated atmospheric and cosmogenic Ne. The 1200°C ratios, however, plot above that  
275 mixing line. Because no cosmogenic Ne is released from quartz above 800°C (Niedermann,  
276 2002), these data indicate a contribution of nucleogenic  $^{21}\text{Ne}$  and  $^{22}\text{Ne}$  from ( $\alpha,n$ ) reactions  
277 on  $^{18}\text{O}$  and  $^{19}\text{F}$ , respectively. In absolute terms, the 1200°C steps released  $3.3\pm 0.7\times 10^6$  and  
278  $1.9\pm 0.8\times 10^6$  atoms  $\text{g}^{-1}$  of excess  $^{21}\text{Ne}$  for AMZ-1 and 3, respectively, which is another 50% and  
279 35% of the excess  $^{21}\text{Ne}$  released up to the 800°C step. Some contribution of nucleogenic Ne in  
280 the lower temperature steps can therefore not be excluded, and the cosmogenic  $^{21}\text{Ne}$   
281 concentrations given above must be considered as maximum values.

282 Higher  $^{10}\text{Be}$  concentrations occur in the Cave of the Letters samples. They range between  
283  $12.0\pm 0.4\times 10^4$  and  $57.5\pm 1.6\times 10^4$  atoms  $\text{g}^{-1}$  quartz.  $^{26}\text{Al}$  concentrations range between  
284  $49.5\pm 7.2\times 10^4$  and  $69.4\pm 4.2\times 10^4$  atoms  $\text{g}^{-1}$  quartz (Table 1).  $^{26}\text{Al}/^{10}\text{Be}$  ratios range between  
285  $1.11\pm 0.10$  and  $4.32\pm 0.28$ . These ratios correspond to simple burial ages that range between  
286  $0.9\pm 0.1$  Ma (COL2) and  $3.4\pm 0.2$  Ma (COL1) (Fig. 7). Age calculations assume negligible post  
287 burial production due to the massive shielding of 70 m of dense dolomite over the cave.  
288 Cosmogenic  $^{21}\text{Ne}$  concentrations in two of the COL samples (COL-1, COL-2) are  $4.3\pm 1.2\times 10^6$   
289 atoms  $\text{g}^{-1}$  quartz and  $4.05\pm 0.88\times 10^6$  atoms  $\text{g}^{-1}$  quartz (Table 1). In this case, the 400-800°C data  
290 are consistent with a two-component mixture of unfractionated atmospheric and cosmogenic  
291 Ne (Fig. 8b), while the 1200°C data again indicate some contribution of nucleogenic Ne, which is  
292 however smaller than for the AMZ samples.

293 The  $^{10}\text{Be}$  concentration in the Masada cave sample (M22) is  $18.1\pm 0.5\times 10^4$  atoms  $\text{g}^{-1}$  quartz,  
294 its  $^{26}\text{Al}$  concentration is  $19.7\pm 7.5\times 10^4$  atoms  $\text{g}^{-1}$  quartz (Table 1) and the  $^{26}\text{Al}/^{10}\text{Be}$  ratio is  
295  $1.09\pm 0.42$ . This ratio corresponds to a simple burial age of  $3.6\pm 0.4$  Ma (Fig. 7). This age  
296 calculation assumes negligible post burial production even though the present geometry of the  
297 site suggests only 12-20 m of rock shielding.

298

## 299 **Discussion**

300

### 301 ***The age of the Amora and Sedom Formations***

302

303 TCN results and their corresponding burial ages of the Amora Formation and the top of the

304 Sedom Formation are in agreement with their stratigraphic position. AMZ samples provide an  
305 age range for the base of the Amora Formation between  $2.7_{-0.6}^{+0.7}$  Ma (AMZ-1) and  $3.3_{-0.8}^{+0.9}$  Ma  
306 (AMZ-3) (Fig. 7). Age calculations safely assume negligible post burial production due to the  
307 high sedimentation regime and the extraction of the samples directly from a deep borehole,  
308 >1000 m below the surface.

309  $^{21}\text{Ne}$ - $^{10}\text{Be}$  and  $^{21}\text{Ne}$ - $^{26}\text{Al}$  ratios of the AMZ samples yield apparent burial ages of  $\sim 8.5$  Ma and  $\sim 6$   
310 Ma, respectively. These ages are significantly older than their  $^{10}\text{Be}$ - $^{26}\text{Al}$  counterpart burial ages.  
311 Some contribution of nucleogenic Ne in the lower temperature steps cannot be excluded, and  
312 the cosmogenic  $^{21}\text{Ne}$  concentrations given above must be considered as maximum values.  
313 However, being a stable nuclide,  $^{21}\text{Ne}$  can be retained in the quartz grains from previous cycles  
314 of exposure. Since Amora Formation sediments include recycled sand from the Miocene Hazeva  
315 Formation which by itself consists of recycled Nubian Lower Cretaceous and Lower Paleozoic  
316 sands (Sa'ar, 1985), the excess  $^{21}\text{Ne}$  in samples AMZ is not surprising.

317 Dealing with Miocene sediments, which are at the very limit of the cosmogenic burial dating  
318 range, places a challenge in determining the  $^{26}\text{Al}$  cosmogenic isotope concentrations by AMS.  
319 The antiquity of such sediments prevents precise and unambiguous measurements of the  $^{26}\text{Al}$   
320 concentrations in the SDM samples. The absence of  $^{26}\text{Al}$  in all four SDM samples suggests three  
321 rather surprising conclusions when one considers that currently they are exposed surface  
322 samples: 1) total burial of the Sedom Formation sediments has been sufficiently long to allow  
323 the decay of essentially all the  $^{26}\text{Al}$  that was in the quartz at the time of deposition. Assuming a  
324 reasonable initial  $^{26}\text{Al}$  concentration of  $\sim 10^6$  atoms  $\text{g}^{-1}$  (as suggested by sample MK22W;  
325 Fruchter et al., 2011) and an AMS minimum detection limit of  $\sim 10^4$  atoms  $\text{g}^{-1}$  (Table 1; sample  
326 SDM-2), at least 6-7  $^{26}\text{Al}$  half lives have elapsed since sediment burial, which is equivalent to 4.2  
327 – 4.9 Ma, 2) the reduction in shielding as the salt diapir was uplifted and exhumed was rapid,  
328 and 3) any recent production of  $^{26}\text{Al}$  has been negligible. With a site production rate of  $\sim 18$   
329 atoms  $\text{g}^{-1} \text{yr}^{-1}$ , accumulation of  $^{26}\text{Al}$  would reach the detection limit of  $\sim 10^4$  atoms  $\text{g}^{-1}$  (Table 1  
330 sample SDM-2) within 500-1000 yr, assuming instantaneous exposure. Since erosion and  
331 surface change on Mt. Sedom are extremely rapid, and noticeable changes, such as gully  
332 incision to a depth of several meters, occur in a life time, a recent exposure of ancient deposits  
333 for only a few centuries may not be all that surprising. Hence, we consider that recent

334 accumulation of post-burial produced  $^{26}\text{Al}$  that would exceed the AMS detection limit to be  
335 unlikely.

336 If the calculation above for  $^{26}\text{Al}$  yields a burial age of 4.2 – 4.9 Ma (6-7 half lives), then for the  
337 same sand grain and same initial  $^{26}\text{Al}/^{10}\text{Be}$  input at  $t=0$  we should get the same burial age  
338 estimate from the  $^{10}\text{Be}$  concentration. Indeed, the initial  $^{10}\text{Be}$  concentration would have been  
339  $1.5 \times 10^5$   $^{10}\text{Be}$  atoms  $\text{g}^{-1}$ . This concentration stems from the initial  $^{26}\text{Al}$  concentration (as shown  
340 above) divided by 6.75 (which is the production rate ratio). From Table 1 we can see that the  
341 mean  $^{10}\text{Be}$  concentration for SDM samples is  $1.3 \times 10^4$  atoms  $\text{g}^{-1}$ , indicating that  $\sim 3$ -4 half-lives  
342 (equal to  $\sim 4$ -6 Ma) must have elapsed since burial.

343 In addition to constraining the minimum age of the SDM samples by the absence of  
344 measurable  $^{26}\text{Al}$ , we suggest two approaches to better approximate the limits for the Sedom  
345 Formation burial ages. First, assuming that all SDM samples were deposited within a relatively  
346 short time, such that within the resolution of burial dating they are of similar age, we can  
347 combine the  $^{26}\text{Al}$  concentration of sample SDM-2 together with the range in  $^{10}\text{Be}$   
348 concentrations measured in samples SDM-3, SDM-4, and SDM-5 to estimate  
349 representative  $^{26}\text{Al}/^{10}\text{Be}$  ratios for the SDM samples in order to calculate minimum burial ages.  
350 These calculations yielded  $^{26}\text{Al}/^{10}\text{Be}$  ratios in the range of 0.35-0.52 (see Table 1) which  
351 correspond to burial ages in the range of  $5.3_{-1.8}^{+inf}$  Ma and  $6.2_{-2.1}^{+inf}$  Ma for the top of the Sedom  
352 Formation (Fig. 7).

353 Another independent estimate of the age of the Sedom Formation uses a previously published  
354 measurement of the  $^{10}\text{Be}$  concentration from presently exposed Hazeva Formation sand  
355 (Fruchter et al., 2011). We consider this result to be a reasonable estimate of the initial pre-  
356 burial concentration ( $N_0$ ). Using this value we can calculate the time required to reach the  
357 measured  $^{10}\text{Be}$  concentrations due to decay. A provenance study of DSF sediments (Sa'ar, 1985)  
358 suggests that sands imbedded in the Sedom Formation consist of continental sediments that  
359 were transported by an extensive drainage system that flowed from the Trans-Jordanian  
360 plateau across the Dead Sea transform to the Mediterranean Sea. Such Miocene sands that  
361 accumulated in the DSF and in the Negev Desert belong to the Hazeva Formation (Calvo, 2000).  
362 Age estimates for the base of the Hazeva Formation range between 25 Ma and 20 Ma (Steinitz  
363 et al., 1978; Calvo, 2002). The age of the top of the Hazeva Formation is older than 8-6 Ma

364 (Steinitz et al., 2000; Steinitz and Bartov, 1991). Based on mineralogical assemblages, Agnon  
365 (1983a,b) estimated that the Sedom Formation was deposited synchronously with the later part  
366 of the Hazeva Formation, thus enabling the inclusion of Hazeva Formation sand in the Sedom  
367 Formation. If we assume that the concentration of  $^{10}\text{Be}$  in these sediments was mostly  
368 controlled by the rate at which the Miocene sands were exhumed and transported into the  
369 Sedom Lagoon and that such exhumation has been operating continuously,  $^{10}\text{Be}$  concentrations  
370 in the present outcrops of the Miocene sands should be representative of the sands that were  
371 deposited in the Sedom Lagoon. Today, Hazeva Formation sediments are preserved mainly in  
372 local structural lows such as synclines and grabens located along the western margin of the DSF.  
373 One such location is the Yamin Plane,  $\sim 30$  km SW of the study site (Fig. 3), a syncline overlain by  
374 Hazeva Formation sediments (Hirsch, 1995) and presently drained to the southern end of the  
375 Dead Sea. A sand sample (MK22W) collected from a small streambed draining the Yamin Plane  
376 was analyzed for  $^{26}\text{Al}$  ( $95.6 \pm 9.9 \times 10^4$  atoms  $\text{g}^{-1}$  quartz) and  $^{10}\text{Be}$  ( $16.8 \pm 0.2 \times 10^4$  atoms  $\text{g}^{-1}$  quartz)  
377 by Fruchter et al. (2011).

378 The results of these analyses show that this sample is positioned at the beginning point of a  
379 decay line passing through the SDM samples (Fig. 7). If the sample from the Yamin Plane is in  
380 fact an appropriate analog for  $N_0$ , then the age of the SDM samples can be based solely on  
381 the  $^{10}\text{Be}$  measurements. This procedure yields an average  $^{10}\text{Be}$  burial age of  $5.0 \pm 0.5$  Ma for the  
382 top of Sedom Formation. This age approximation agrees with the other age estimation  
383 presented above and supports a late Miocene age for the upper part of the Sedom Formation.  
384 Since an additional 250 m of Sedom Formation sediments overlie our SDM samples, the actual  
385 termination of sediment deposition by the Sedom Lagoon is somewhat younger than the age  
386 attributed here. It is interesting to note that Steinitz and Bartov (1991) provided similar age  
387 constraints for the Sedom Formation based on a broad examination of the paleogeographical  
388 evolution of the DSF by means of radiometric dating of plateau pre-rift basalts and basalts  
389 which flowed through incised canyons towards the Dead Sea.

390 The rise of the Sedom diapir has been examined in numerous multi-disciplinary studies and is  
391 reviewed by Weinberger et al. (2006). They conclude that Mt. Sedom breached the surface  
392 about 14 ka and very rapid erosion of the soft sediment followed the rise of the mountain. The  
393 absence of  $^{26}\text{Al}$  and the low levels of  $^{10}\text{Be}$  in all the SDM samples forces us to assume that our

394 sampling locations must have been shielded by at least tens of meters of sediment until very  
395 recently. Nevertheless, it is instructive to consider the possible effect of post burial production  
396 of  $^{26}\text{Al}$  on the SDM samples as its absence constrains possible scenarios of diapir emergence. In  
397 the context of post-burial production, the exposure history is simplified and divided into two  
398 stages: 1) rise of the diapir to the surface through the basin sediments. During this stage, the  
399 overburden is well in excess of hundreds to thousands of meters. The contribution of post-  
400 burial production of cosmogenic nuclides during the first stage of the diapir uplift through the  
401 basin sediments (until 14 ka) is found to be insignificant at a total of  $\sim 200$   $^{26}\text{Al}$  atoms  $\text{g}^{-1}$  quartz,  
402 considering a rising rate of  $5 \text{ mm yr}^{-1}$  and a minimum of 20 m of final overburden, and 2)  
403 breaching of the surface and the formation of Mt. Sedom. During this stage TCN production can  
404 occur. The transition between the two stages took place after the retreat of Lake Lisan at  $\sim 14$   
405 ka as both alluvial and lake sediments are found on top of Mt. Sedom, disconnected from their  
406 surroundings (Fig. 4; Weinberger et al., 2006). The overburden above the SDM samples at the  
407 time of the diapir emergence is considered to be at least 50 m: 20 m of current stream incision  
408 into the Sedom Formation and  $\sim 30$  m of overlying Lake Lisan sediments (Weinberger et al.,  
409 2007; Fig. 4).

410 Post-burial production during the second stage could have a noticeable effect on burial ages,  
411 depending on the exposure history (Fig. 9). Although erosion of this overburden began  
412 immediately after the retreat of Lake Lisan at 14 ka, major stream incision started later as a  
413 result of relief development and rapid dissolution of halite within the Sedom Formation.  $^{14}\text{C}$   
414 dating of cave passages at different depths places the beginning of incision at  $\sim 3.3$  ka for the  
415 north-eastern boundary of Mt. Sedom (Frumkin, 1996). If erosion of the overburden and late  
416 incision (i.e. since 3.3 ka) are assumed as suggested by dated cave passages (Frumkin, 1996),  
417 then post-burial production of  $^{26}\text{Al}$  may amount to no more than 2000 atoms  $\text{g}^{-1}$  quartz (if initial  
418 overburden after 14 ka was 40 m and final overburden was 1 m; Fig. 9), and no more than 3000  
419 atoms  $\text{g}^{-1}$  quartz (if initial overburden after 14 ka was 20 m and final overburden was only 0.5  
420 m; Fig. 9). In both these cases the TCN ages presented above may be underestimated by only  
421 0.7-1.3 million yr. If, however, erosion of the overburden and incision started immediately after  
422 emergence of the diapir ( $\sim 14$  ka), then post burial production of  $^{26}\text{Al}$  may amount to as much as  
423 10,000 atoms  $\text{g}^{-1}$  quartz (Fig. 9) and the TCN ages presented above may be extremely

424 underestimated. This scenario is not supported by the results of Frumkin (1996), by field  
425 evidence that indicates very rapid erosion and incision of the mountain, and by the absence of  
426 measurable  $^{26}\text{Al}$ .

427 It should be noted here that we concentrated our discussion concerning post-burial production  
428 on  $^{26}\text{Al}$  because it is produced much faster than  $^{10}\text{Be}$ . Total post burial production of  $^{10}\text{Be}$  would  
429 range between 300 and 450 atoms  $\text{g}^{-1}$  quartz, if incision started at 3.3 ka. This amount is within  
430 the measuring error of  $^{10}\text{Be}$  and would not contribute to any change in the calculated  $^{26}\text{Al}/^{10}\text{Be}$   
431 ratio. Overall, our results suggest that the age of the top of the Sedom Formation ranges  
432 between  $5.0\pm 0.5$  Ma and  $6.2_{-2.1}^{+inf}$  Ma, depending on the different assumptions made in the  
433 burial age calculations. These ages lie at the upper limit of the age range of the cosmogenic  
434 burial dating method. The absolute age of the Miocene-Pliocene boundary is somewhat  
435 debated and ranges between 4.6 and 5.3 Ma (e.g. Hilgen and Langeries, 1993). Our age for the  
436 Sedom Formation is taken from 250 meters below the top of the formation and ranges  
437 between  $6.2_{-2.1}^{+inf}$  and  $5.0\pm 0.5$  Ma. Therefore, the age of the very top of the Sedom Formation  
438 maybe very early Pliocene. Of course, this constrains a Miocene age for most of the Sedom  
439 formation sediments.

440

#### 441 ***Cave ages***

442

443 Sediment burial ages have been used to estimate the time when active drainage systems  
444 operated at the level of cave openings (e.g. Granger et al. 1997, 2001; Stock et al., 2004;  
445 Haeuselmann et al., 2007; Wagner et al., 2010). In this study, we date the oldest sediments  
446 relating to active drainage or shoreline activity are preserved within caves close to the top of  
447 the escarpment (Frumkin, 2001; Lisker et al., 2010). Sediment burial ages in the Masada cave  
448 (MZ2) would be indicative of the time the surface of the Dead Sea basin was at the level of the  
449 cave. Similarly, sediment burial ages in the Cave of the Letters would indicate the time the floor  
450 of Wadi Hever canyon was at the level of the cave while incising into the desert plateau in  
451 response to subsidence of the Dead Sea basin. Therefore, as both caves are relatively close to  
452 the top of the escarpment, burial ages of sediment trapped in them would represent the initial  
453 stages of rapid subsidence of the Dead Sea basin and formation of significant relief in the region.



454 Apart from sample COL3, collected from the front face of the sedimentary cave fill, which was  
455 exposed by canyon side-cliff retreat, all other Cave of the Letters samples were collected from  
456 deep within the cave (>10 m from the entrance) and are well shielded. Nevertheless, Cave of  
457 the Letters samples do not converge to a common burial age. We suspect that biogenic activity,  
458 namely the presence of wasp nests in the cave, is responsible for this age scatter. Wasps collect  
459 wet silt and fine sand from potholes in the canyons and use it when building their nests in the  
460 cave. Wasps have been nesting in the soft sediments of the Cave of the Letters since the  
461 passages in the cave were excavated ~2000 yr ago by humans. Therefore, the samples we  
462 collected, although not showing evidence of present wasp activity, may have been  
463 contaminated with silt and fine sand brought in by wasps during the past 2000 yr. Therefore,  
464 the resultant  $^{26}\text{Al}/^{10}\text{Be}$  ratio and calculated burial age depend on the ratio of sand deposited in  
465 the cave when the cave was at the level of the active drainage system and wasp-delivered  
466 sediment. Any addition of silt which postdates cave abandonment and thus characterized by  
467 higher  $^{26}\text{Al}/^{10}\text{Be}$  ratios, to the sample would result in an underestimation of the burial age. The  
468 observation of possible mixing of ancient cave sediment with modern sediment is supported by  
469 the location of all the COL samples on the Al-Be burial plot (Fig. 7). All COL samples fall on a  
470 mixing curve between modern sediment and undisturbed cave sediment. We therefore,  
471 consider the oldest age as representing the minimum burial age of the sediment in the cave.  
472 This age of  $3.4 \pm 0.2$  Ma is given by sample COL1, which was collected 21 m from the entrance of  
473 the cave. Since the cave opening is located ~70 m below the upper desert surface, incision must  
474 have started some time prior to this age.

475 The burial age given by sample COL1 enables to calculate an average bedrock incision rate  
476 ( $\sim 0.07$  mm yr<sup>-1</sup>) of Wadi Hever since the deposition of the Cave of the Letters sediments. This  
477 rate is slightly higher than those calculated in the southern Sierra Nevada during times of  
478 tectonic quiescence but an order of magnitude lower during major uplift, 3-1.5 Ma (Stock et al.,  
479 2005). Similarly, incision rate in Wadi Hever is similar to that calculated in the eastern Alps for  
480 the past 2.5 Ma but an order of magnitude slower than the eastern Alps in the Pliocene  
481 (Wagner et al., 2010). The more active Swiss (western) Alps have been incised an order to two  
482 orders of magnitude faster (Haeuselmann et al., 2007). Overall, the incision rate calculated for  
483 the DSF canyon is similar to those calculated in other regions of moderate tectonic activity. In

484 contrast, it is somewhat faster than those calculated in regions of tectonic quiescence, such as  
485 in the Appalachian Mountains (Granger et al., 1997, 2001). It must be kept in mind, however,  
486 that in Wadi Hever, and in all other canyons that drain into the Dead Sea, incision slowed down  
487 dramatically during periods of lake high stand, such as the late Pleistocene Lake Lisan high  
488 stand between 20-30 ka (Bartov et al., 2002; Lisker et al., 2009). This implies, of course, that  
489 incision must have been more rapid during other times.

490 It is interesting to consider the burial age calculated for sample COL3 that was collected at the  
491 front of the cave, where the sedimentary section is exposed by the horizontal retreat of the  
492 canyon cliff. The age given by this sample ( $3.0 \pm 0.1$  Ma) is only 0.4 Ma younger than the oldest  
493 age given by sample COL1. This decrease in the calculated burial age is most likely caused by  
494 the production of  $^{26}\text{Al}$  and  $^{10}\text{Be}$  in the sediments as they were exhumed at the front of the cave  
495 by the retreating canyon wall. If actually caused by recent exposure, the increase in  
496 the  $^{26}\text{Al}/^{10}\text{Be}$  ratio from 1.11 (COL1) to 1.45 (COL3), would take only  $\sim 15$  ka if the face was  
497 exposed instantaneously (using a site  $^{10}\text{Be}$  production rate of  $\sim 2.5$  atoms  $\text{g}^{-1} \text{yr}^{-1}$  and a  $^{26}\text{Al}$   
498 production rate of  $\sim 16.9$  atoms  $\text{g}^{-1} \text{yr}^{-1}$ ). However, this is a theoretical end-member case. The  
499 retreat rate of the major knick point of Wadi Hever was estimated to range between 800 and  
500 1600  $\text{mm ka}^{-1}$  (Haviv et al., 2006) and cliff retreat rates that range between 100 and 850  $\text{mm ka}^{-1}$   
501 have been published for various arid regions in the world (e.g. Schumm and Chorley, 1966;  
502 Yair and Gerson, 1974; Mayer, 1979; Cole and Mayer, 1982; Matmon et al., 2005). Applying cliff  
503 retreat rates in the range of 100 to 800  $\text{mm ka}^{-1}$ , implies that the time required to shift  
504 the  $^{26}\text{Al}/^{10}\text{Be}$  ratio from 1.11 to 1.45 would be longer but that the sediments presently exposed  
505 at the front of the cave were shielded until recently.

506  $^{21}\text{Ne}-^{10}\text{Be}$  and  $^{21}\text{Ne}-^{26}\text{Al}$  ratios of sample COL2 yield apparent burial ages of 2.4 Ma and 1.9 Ma,  
507 respectively. These ages are older than the  $^{10}\text{Be}-^{26}\text{Al}$  burial age of this sample (0.9 Ma). It is  
508 reasonable that some of the quartz grains transported into the Cave of the Letters by Wadi  
509 Hever originated from Miocene Hazeva sand outcrops that existed on the Judea Desert Plateau  
510 during the Pliocene. Therefore, as with the AMZ samples, the excess in the cosmogenic  $^{21}\text{Ne}$   
511 concentrations can be explained by its accumulation during previous cycles of exposure and its  
512 retention in the quartz crystals.

513 In contrast to sample COL2,  $^{21}\text{Ne}-^{10}\text{Be}$  and  $^{21}\text{Ne}-^{26}\text{Al}$  ratios of sample COL1 yield apparent burial

514 ages of 0.5 Ma and 2.3 Ma, respectively. These ages are younger than the  $^{10}\text{Be}$ - $^{26}\text{Al}$  burial age of  
515 this sample (3.4 Ma). It is difficult to explain such a result as typically  $^{21}\text{Ne}$  based ages generally  
516 equal or exceed those determined using  $^{26}\text{Al}/^{10}\text{Be}$  ratios. Other than instrumental or analytical  
517 errors, field related complications could result in diffusive loss of  $^{21}\text{Ne}$ , most likely through  
518 heating. During Roman times, 2000 yr ago, the Cave of the Letters was occupied for over a  
519 century (this is the cave in which many of the Dead Sea ancient documents were found and the  
520 reason for its name). We conjecture that frequent use of small camp fires would be sufficient to  
521 heat the sand in the immediate surrounding of the hearth to several hundred degrees causing  
522 diffusive loss of Ne from quartz grains but not affecting the  $^{10}\text{Be}$  and  $^{26}\text{Al}$  concentrations.  
523 Niedermann (2002) showed that a high percentage (up to 70%) of cosmogenic  $^{21}\text{Ne}$  is loosed  
524 from quartz grains at temperatures  $<600^\circ\text{C}$ . This would result in underestimated  $^{21}\text{Ne}$ - $^{10}\text{Be}$   
525 and  $^{21}\text{Ne}$ - $^{26}\text{Al}$  burial ages. Such a process has been previously suggested to explain the loss of  
526 other noble gases from rock samples (Gillespie et al., 1985). Although speculative, we offer this  
527 scenario as a possible explanation for the relatively low Ne concentration.

528 The  $^{26}\text{Al}/^{10}\text{Be}$  ratio in the Masada cave (sample MZ2) is  $1.09\pm 0.42$ . This ratio corresponds to a  
529 simple burial age of  $3.6\pm 0.4$  Ma. At this site we did not observe any wasp activity. However,  
530 shielding of the sediment is not as thick as in the Cave of the Letters. MZ2 burial age is identical  
531 within error to the age of sample COL1 ( $3.4\pm 0.2$  Ma) and suggests that post-burial production is  
532 not significant. This points to the possibility that shielding must have been greater than  
533 observed today during most of the time that the sediment of sample MZ2 was buried. Based on  
534 the burial ages of the cave sediments, initial incision and relief formation along the Dead Sea  
535 escarpment was well expressed by  $\sim 3.5$  Ma.

536

### 537 ***Implications for the morphotectonic evolution of the Dead Sea basin and its western margin***

538

539 Stein et al. (2000) observed discordant dolomitic bodies along the western margin of the Dead  
540 Sea basin that attest to intense dolomitization, which occurred within the Upper Cretaceous  
541 limestone by interaction with highly evaporated sea water that filled the basin. Ryb et al.  
542 (2009a,b) studied iron mineralization of veins included in these dolomitic bodies and  
543 emphasized the involvement of evaporated sea water in the process. Using U-Pb, Ryb et al.

544 (2009a) dated the mineralized veins to  $11.8 \pm 2.0$  Ma. The geomorphic implication of these  
545 observations and results is that the structure, and more importantly the topography of the  
546 western margin of the Dead Sea fault enabled, at that time, the penetration of Mediterranean  
547 water into the transform valley and the formation of saline marine lagoons. Although this  
548 process has been proposed and described earlier (Katz and Starinsky, 2009), the work of Ryb et  
549 al. (2009a) was the first to provide robust age constraints. Passage between the evolving  
550 transform valley and the Mediterranean was possible as long as low topography persisted along  
551 the western margin of the Dead Sea fault. The new cosmogenic nuclide age determinations  
552 provided here for the top of the Sedom Formation, between  $5.0 \pm 0.5$  Ma and  $6.2_{-2.1}^{+inf}$  Ma, and  
553 for the base of the overlying Amora Formation,  $3.3_{-0.8}^{+0.9}$  Ma, constrain the time that this passage  
554 ceased and terminal lakes commenced. Based on these ages, the transition from water bodies  
555 connected to the Mediterranean to those not connected may have occurred roughly between 5  
556 and 3.5 Ma. This time range is essential in any further research of the Quaternary topographic  
557 evolution and the contemporary drainage pattern development of the western margin of the  
558 DSF.

559 Why did such a transition occur at that time? A change of only a few degrees in the direction of  
560 motion of the Arabian plate relative to the African plate would introduce an extension  
561 component to the Dead Sea fault and trigger significant subsidence (to form a rift) and  
562 significant uplift of the margins. Such a change in the direction of motion has been previously  
563 proposed to explain observations of major importance. To the south, Bosworth et al., (2005)  
564 estimate that the rate of extension increases at  $\sim 5$  Ma which is followed by the first appearance  
565 of MORB in the Red Sea. To the north, Westaway, (2004) observes that the East Anatolian fault  
566 (EAF) was initiated at  $\sim 4$  Ma following the change in stress field. Both these observations point  
567 to an increase in the E-W vector of the motion of the Arabian plate (Fig. 1). Such a change  
568 would introduce noticeable extension along the Dead Sea fault. We suggest that the timing of  
569 increased rift subsidence and major relief expression along the Dead Sea segment of the  
570 Arabian-African plate boundary, based on our new cosmogenic burial ages, is commensurate  
571 with the tectonic modifications observed in the Red Sea to the south and Anatolian Fault to the  
572 north.

573 Changes in the direction of plate motion over time scales of  $10^6$  yr are not a unique event (e.g.

574 Sharp and Clague, 2006). Nor is the reorganization of plate boundary directions. The direction  
575 of the Arabia-African plate boundary has changed over time, as did the direction of relative  
576 motion between the two plates. For example, at 14–12 Ma a major structural event occurred  
577 along the northern part of the Arabia-African plate boundary. Tectonic activity decreased  
578 dramatically along the Gulf of Suez, which is the northern and direct extension of the Red Sea,  
579 the Red Sea motion switched from rift-normal movement to highly oblique extension and strike  
580 slip motion commenced along the DSF (Bosworth et al., 2005). Steckler and ten Brink (1986)  
581 suggested that the cause for this switch in plate boundary configuration was the inability of the  
582 Red Sea rift to propagate northward through the stronger lithosphere of the Mediterranean  
583 continental margin. Later research (e.g. Bosworth and McClay, 2001) established that, although  
584 the greater strength of the Mediterranean continental margin lithosphere did play a role in  
585 determining the Red Sea plate boundary geometry, it was not in as complete a fashion as  
586 previously thought.

587 A change of smaller magnitude occurred along the Arabia-African plate boundary in the  
588 Pliocene. Garfunkel (1981) analyzed the kinematics of the DSF motion and concluded that  
589 Miocene motion along the plate boundary had to be of a more N-S direction than that of the  
590 Pliocene to prevent inconsistent overlap of lithospheric portions of the Arabian and African  
591 plates. Joffe and Garfunkel (1987) showed that the slip rates for the Miocene ( $\sim 6.5 \text{ km Ma}^{-1}$ )  
592 were slower than those calculated for the Pliocene-Pleistocene ( $\sim 9 \text{ km Ma}^{-1}$ ). These rates also  
593 agree with initial observations by Quennell (1951, 1958). Late Quaternary slip rates as well as  
594 present ones, based on GPS measurements, are slower and are on average  $5 \text{ km Ma}^{-1}$  (LeBeon  
595 et al., 2010, 2012; Daëron et al., 2004, 2007; McClusky et al., 2003).

596 The association between rift subsidence and margin uplift has been described for several rifts  
597 (Rosendahl, 1987; Wdowinski and Zilberman, 1996, 1997; ten-Brink et al., 1990). Uplifted rift  
598 shoulders are recognized in rifts of various ages (e.g. Bloom, 1998; Hutchinson et al., 1992;  
599 Bohannon et al., 1989). Furthermore, the memory of uplifted rift margins from a time of  
600 significant tectonic activity still persists along most continental passive margins (van der Beek  
601 and Braun, 1999; Gilchrist and Summerfield, 1994; Beaumont et al., 2000). We argue that the  
602 transition from a subdued topography of the Sedom Lagoon during the late Miocene to the  
603 more contemporary disconnected water bodies such as documented over the more recent

604 Quaternary is the consequence of the initiation of significant subsidence along the DSF  
605 accompanied by significant margin uplift (Wdowinski and Zilberman, 1996, 1997) about 5-3 Ma.  
606 The subsequent construction of a continuous water divide west of the rift then allowed major  
607 fluvial incision along the evolving escarpment. These processes dominated the slip along the  
608 DSF long enough to modify the drainage system pattern and significantly affect sediment  
609 transport and deposition capabilities throughout the Quaternary. Consequently, in contrast to  
610 the tectonic evolution of the Dead Sea fault that started at about 18 – 14 Ma but did not affect  
611 topography in a significant way for at least 10 Ma, the style of deformation since 5 Ma and into  
612 the Quaternary, which included rifting, also controlled the evolution of the current drainage  
613 systems and topographical framework of the water bodies in the DSF valley.

614 The oldest burial age of alluvial sediments derived from the Cave of the Letters ( $3.4\pm 0.2$  Ma)  
615 and the burial age derived from the sediment in the Masada cave ( $3.6\pm 0.4$  Ma) indicate that  
616 incision had commenced earlier than  $\sim 3.5$  Ma and that significant relief had also been present.  
617 The cave sediment burial ages obtained in this study are similar to other age constraints of  
618 initial canyon incision related to Dead Sea subsidence. On the eastern side of the Dead Sea, in  
619 the Zarka Ma'in canyon region (Fig. 3), a flood basalt located on the plateau and dated to  $\sim 6$  Ma  
620 (Steinitz and Bartov, 1991) indicates that incision initiated after it flowed. On the other hand,  
621 about 100 m below the Jordanian desert plateau and 200 m above the Zarka Ma'in canyon floor,  
622 a younger basalt that flowed towards the rift through the entrenched canyon was dated using  
623 K-Ar to 3.4 Ma (Steinitz and Bartov, 1991). This age is identical to the burial ages given by the  
624 cave sediments on the western DSF margin. Furthermore, the basalt is positioned in the canyon  
625 below the desert plateau at a similar elevation as the Cave of the Letters, suggesting that along  
626 both margins flanking the Dead Sea basin, incision began at roughly similar times and that initial  
627 canyons were already formed. The beginning of relief formation along the DSF western margin  
628 is implied by other observations, such as the water-table fall in the northern Negev dated by U-  
629 Pb of the earliest vadose cave deposits to  $\sim 3.1$  Ma (Vaks et al., 2013) and by exhumation,  
630 exposure to air oxygen, and combustion of bituminous rocks that occurred between 3.8 and 2.6  
631 Ma along the western margin of the Dead Sea (Gur et al., 1995).

632 Similar age constraints are obtained from the northern part of the Dead Sea fault. In the  
633 northern Jordan Valley, the Yarmuch canyon incises between the Jordanian Plateau and the

634 Golan Heights into Pliocene flood basalts (the Cover Basalt Formation) which flowed prior to  
635 incision and are dated to ~4.5 Ma (Mor and Steinitz, 1985; Shaliv, 1991). Thus, incision of this  
636 major canyon began after 4.5 Ma as a response to major DSF subsidence. Davis et al. (2011)  
637 dated the first fresh water body in the Jordan Valley (the Erk El Ahmar Lake) to ~4.5 Ma. The  
638 sediments of this lake overlie the Cover Basalt Formation (Rotstein et al., 1992; Inbar et al.,  
639 2010), and although of the same age, their stratigraphic order indicates that the lake sediments  
640 were deposited immediately after the basalt covered the region. The presence of a fresh water  
641 body indicates that disconnection from the Mediterranean Sea and formation of non-terminal  
642 fresh water bodies (similar to present-day Sea of Galilee) had occurred by 4.5 Ma. Our results  
643 from the region of the Dead Sea and the results presented above from the northern Jordan  
644 Valley, ~200 km north, suggest a similar time, 5-4 Ma, for the disconnection of the rift valley  
645 from Mediterranean Sea. This similarity suggests that significant subsidence accompanied by  
646 margin uplift and the establishment of the western water divide occurred simultaneously.  
647 These processes, therefore, set the scene for the Quaternary morphologic development and  
648 sedimentologic regime of the entire DSF western margin.

649  
650 **Conclusions**

651  
652 Direct radiometric dating of deeply buried Dead Sea intra-rift and rift shoulder sediments was  
653 possible for the first time using TCN measurements. Complexities due to post-burial production  
654 were avoided by sampling from a deep borehole (bottom of Amora Formation), from sediments  
655 which were rapidly uplifted from within the deep basin by a rising diapir (Sedom Formation),  
656 and sediments that were buried in caves. These approaches enabled the burial dating of late  
657 Miocene and Pliocene sediments – pushing the limit of the cosmogenic burial dating method.  
658 The absence of  $^{26}\text{Al}$  in the top of the Sedom Formation sediments indicates both the antiquity  
659 of the sediments and the very rapid exposure by the rising diapir. Using different constraining  
660 methods and assumptions these sediments are dated to 6-5 Ma. The oldest Amora Formation  
661 sample, collected from ~50 m above the base of the formation, yielded a TCN burial age of  
662  $3.3^{+0.9}_{-0.8}$  Ma.  $^{21}\text{Ne}$  ages of the Amora Formation and of cave sediments do not conform to  
663 the  $^{26}\text{Al}$ - $^{10}\text{Be}$  burial ages due to  $^{21}\text{Ne}$  retention in the quartz grains from previous erosional

664 episodes.  
665 Burial ages of cave sediments range between 3.6 and 3.4 Ma and suggest initial incision of  
666 canyons flowing into the Dead Sea shortly before this time. These ages agree well with other  
667 indications of topographic relief and incision from other locations along the fault. We show that  
668 the major difference between the Sedom Lagoon and the later water bodies is the consequence  
669 of the initiation of DSF significant subsidence accompanied by significant uplift of the its  
670 margins, and the establishment of a continuous water divide west of the DSF. The change in the  
671 style of faulting along the DSF coincides with the first MORB's in the Red Sea and the initiation  
672 of the East Anatolian fault. All of these processes resulted from a change in the direction of  
673 motion of the Arabian plate relative to the African plate. Along the Dead Sea rift, this change  
674 was followed by a dramatic reorganization of relief and drainage patterns in the region.

675

## 676 **Acknowledgments**

677

678 This study was supported by the ISF-Bikura grant 362/06 and Hebrew University Ring internal  
679 grant. We thank Roi Porat and Uri Davidovich who noted to us the potential significance of the  
680 Masada cave sediments to the rift margin evolution.

681

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981

982 Figure captions

983

984 Figure 1: Satellite image of the Middle East marking the location of the major plate boundaries  
 985 around the Arabian Plate. The Dead Sea fault (DSF) transforms motion between the  
 986 spreading zones in the south (marked with dash-point lines) and the convergence in the  
 987 north (marked with dashed lines). EAF - East Anatolian fault. NAF - North Anatolian fault.  
 988 Location of 5 Ma MORB in Red Sea is marked with red circle. Location of pole of rotation  
 989 between the African and Arabian plates (Westaway, 2004) is marked with star.

990 Figure 2: Composite stratigraphic column of the Dead Sea Group (Zak, 1967). Sampled sections  
 991 are marked with red open boxes. Notice that most of Sedom Formation, particularly  
 992 within its uppermost ~250 m, is composed of halite and not amenable for  
 993 cosmogenic  $^{26}\text{Al}/^{10}\text{Be}$  burial dating except for the Benot Lot member section (at ~1500  
 994 m). The Amora Formation is mostly detrital.

995 Figure 3: A) Shaded relief map with locations of study sites (marked with yellow dots) within the  
 996 Dead Sea rift valley (marked with a blue line). The speculated extent of the Sedom  
 997 Lagoon is marked with a light gray polygon. Sites marked with red dots indicate

998 locations mentioned in discussion (EEA is Erk el Ahmer) but not sampled for TCN in this  
999 study. B) Close-up shaded relief image with TCN sample locations (yellow dots): (1) SDM  
1000 samples at Mt. Sedom. (2) AMZ samples from the Amora Formation collected from the  
1001 Amiaz-1 borehole. (3) Surface sediment sample derived from the Yamin Plane (sample  
1002 MK22W from the Hazeva Formation; Fruchter et al., 2011).

1003 Figure 4: A) Schematic geological cross section through the Mt. Sedom diapir (Frumkin, 2009)  
1004 showing the location of the SDM samples within the diapir and the Amiaz-1 borehole  
1005 from which AMZ samples were collected. Subvertical lines are bedding planes. Black  
1006 dotted filled polygon capping Mt. Sedom represents late Pleistocene Lake Lisan  
1007 sediments that were uplifted by the rising diapir. These sediments can also be seen as  
1008 horizontal beds in the upper right side of the photo in (C). B) Photograph showing the  
1009 landscape presented in (A) with the Pleistocene Lake Lisan sediments in the foreground.  
1010 C) Photograph of sampling site of the SDM samples on the eastern edge of Mt. Sedom  
1011 (persons in photo for scale). Sediment beds of the Sedom formation are overturned  
1012 ( $\sim 28^\circ$  dip to the west) due to the uplift of the salt. Late Pleistocene Lake Lisan sediments  
1013 appearing at the top of the photo are horizontal.

1014 Figure 5: The Cave of the Letters. The main cave entrance is marked with a white polygon which  
1015 includes the 13.5 m thick sediment fill truncated by the cliff. Sample COL 3 was collected  
1016 from this truncated wall. The entrance to the passage in the soft sediment that was  
1017 excavated  $\sim 2000$  yr ago by humans can be seen at the top of the white polygon. Left  
1018 inset: Samples COL 1 and COL 2 were collected inside this passage 15-21 m from its  
1019 entrance. Samples COL 5 and COL 7 were collected from the middle passage, with its  
1020 entrance indicated by the arrow. Both these samples were collected  $>10$  m from the  
1021 entrance of the middle passage. The upper desert plateau is  $\sim 70$  m above the cave  
1022 opening. The right inset shows a close-up of sedimentary cross bedding in the soft  
1023 sediment and some wasp drillings in the upper part of the photo.

1024 Figure 6: The Masada Cave. (A) Photo, from south to north, of the Dead Sea escarpment  
1025 showing Masada, as part of the escapement, in the foreground. The general location of  
1026 Masada Cave shown in (B) is marked with open box. (B) A series of natural caves  
1027 exposed below the ruins of the Masada archeological site. The sampled cave is marked

1028 with a black arrow. (C) A measured cross section of the cave indicates the location of  
1029 sample MZ2 (full dot). Well-cemented sediment is exposed at three locations: a  
1030 conglomerate on the upper desert surface above the cave, slope debris at the entrance  
1031 of the cave, and fine shoreline sediments deep at the back of the cave where it is  
1032 cemented to the bedrock wall. The shoreline sandy sediments are bedded and include  
1033 mostly chert fragments.

1034 Figure 7: Burial ages for Amora Formation (AMZ) samples, Sedom Formation (SDM) samples,  
1035 and cave samples (COL, MZ) plotted on a  $^{10}\text{Be}$  vs.  $^{26}\text{Al}/^{10}\text{Be}$  exposure-burial diagram.  
1036 Also shown is a modern analog for possible pre-burial concentrations of Sedom samples  
1037 (sample MK22W in Fruchter et al., 2011). Notice that all COL samples fall on a mixing  
1038 curve (designated by an empty elongated rectangle) of ancient buried sediment and  
1039 modern exposed sediment.

1040 Figure 8: Ne stepwise heating extraction results: a - for samples AMZ-1 and AMZ-3. b - for  
1041 samples COL-1 and COL-2.  $2\sigma$  uncertainties are shown in both panels. Extraction  
1042 temperatures: white- 400°C, gray- 600°C, red- 800°C, black- 1200°C, green- crusher  
1043 extraction. Crusher extraction measurement is available only for sample AMZ-1 and  
1044 COL-2. mfl – mass fractionation line.

1045 Figure 9: Simulations of post-burial production of  $^{26}\text{Al}$  in SDM samples as a function of the initial  
1046 overburden (20-40 m) at the time the diapir breached the surface and the time when  
1047 this overburden started eroding (14-3 ka). Lines in A and B are equal-concentration lines.  
1048 (A) Simulations considering sample collection after removing 100 cm of surface  
1049 sediment cover. If erosion of the overburden (~40 m) and late incision, since 3.3 ka, are  
1050 assumed as suggested by dated cave passages (Frumkin, 1996), then post-burial  
1051 production of  $^{26}\text{Al}$  may amount to no more than 2000 atoms  $\text{g}^{-1}$  quartz. (B) Simulations  
1052 considering samples collection after removing 50 cm of cover. If erosion of the  
1053 overburden (~20 m) and late incision, since 3.3 ka, are assumed as suggested by dated  
1054 cave passages (Frumkin, 1996), then post-burial production of  $^{26}\text{Al}$  may amount to no  
1055 more than 3000 atoms  $\text{g}^{-1}$  quartz. (C) In both end member cases described in (A) and (B),  
1056 the TCN ages presented above may be underestimated by only 0.7-1.3 Ma. The initial  
1057 burial age that corresponds to zero post-burial production (i.e. 6.2 Ma) is derived from

1058 the ratio of  $^{26}\text{Al}/^{10}\text{Be}$  based on  $0.5 \times 10^4$  atoms of  $^{26}\text{Al} \text{ g}^{-1}$  quartz measured in sample  
1059 SDM-2 and  $1.56 \times 10^4$  atoms of  $^{10}\text{Be} \text{ g}^{-1}$  quartz measured in sample SDM-3. X axis  
1060 indicates  $^{26}\text{Al}$  atoms produced only during exhumation in the last 14 ka that are  
1061 subtracted from the initial  $^{26}\text{Al}$  concentration ( $0.5 \times 10^4$  atoms  $\text{g}^{-1}$  quartz) to estimate the  
1062 true burial age. As the fraction of post-burial  $^{26}\text{Al}$  increases it lowers the true  $^{26}\text{Al}/^{10}\text{Be}$   
1063 ratio and hence correlates to a longer burial age. Parameters used in the calculation:  
1064 altitude: -300 m, overburden density:  $2 \text{ g cm}^{-3}$ , average dipole moment:  $8 \times 10^{22} \text{ Am}^2$ ,  
1065 scaling: Dunai (2001), muon production: Braucher et al. (2011), spallation production:  
1066 Balco et al. (2008).  
1067  
1068  
1069

Table 1: Cosmogenic nuclide data

Sample name	Quartz (g)	Be carrier (mg)	Al carrier (mg)	Al (ICP-OES) ( $10^{18}$ atoms $g^{-1}$ )	$^{26}Al/^{27}Al$ ( $\times 10^{-15}$ )	$^{26}Al$ ( $\times 10^4$ atoms $g^{-1}$ )	Be spike ( $\times 10^{17}$ atoms $g^{-1}$ )	$^{10}Be/^{9}Be$ ( $\times 10^{-15}$ )	$^{10}Be$ ( $\times 10^4$ atoms $g^{-1}$ )	$^{26}Al/^{10}Be$	$^{26}Al/^{10}Be$ Corrected <sup>d</sup>
SDM-2 <sup>a</sup>	40.037	0.286	0	2.79±0.11	1.9±1.9 <sup>c</sup>	0.5±0.5	4.77	N/A	N/A	N/A	N/A
SDM-3 <sup>b</sup>	35.002	0.300	0.49	2.99±0.12	<b>&lt; 9</b>	< 2.7	5.73	27.2±5.8	1.56±0.34	< 1.7	0.35
SDM-4 <sup>b</sup>	35.007	0.299	0.49	3.20±0.13	<b>&lt; 21</b>	< 6.7	5.71	29.1±9.0	1.66±0.52	< 4.1	0.33
SDM-5 <sup>b</sup>	35.010	0.298	0.49	2.89±0.12	<b>&lt; 12</b>	< 3.5	5.69	18.4±3.8	1.05±0.22	< 3.3	0.52
AMZ-1 <sup>a</sup>	35.009	0.307	0	3.15±0.13	8.8±2.2	2.77±0.71	5.86	24.6±1.8	1.44±0.12	1.92±0.51	
AMZ-2 <sup>a</sup>	35.482	0.303	0	3.05±0.12	6.6±2.4	2.02±0.74	5.71	16.9±1.6	0.96±0.10	2.09±0.80	
AMZ-3 <sup>a</sup>	35.003	0.306	0	3.23±0.13	7.7±2.2	2.13±0.62	5.84	25.4±2.9	1.48±0.18	1.44±0.45	
COL1 <sup>a</sup>	35.061	0.294	0.491	5.74±0.23	111±8	63.5±5.2	5.60	988±19	57.5±1.6	1.11±0.10	
COL2 <sup>a</sup>	35.074	0.295	0.496	4.44±0.18	175±6	59.6±3.2	5.62	237±7	13.8±0.5	4.32±0.28	
COL3 <sup>a</sup>	26.971	0.300	0.293	7.07±0.28	147±7	69.4±4.2	7.43	620±15	47.8±1.5	1.45±0.10	
COL5 <sup>a</sup>	35.024	0.222	0.160	5.38±0.22	92±13	49.5±7.2	4.24	402±9	12.0±0.4	4.11±0.61	
COL7 <sup>a</sup>	35.005	0.235	0.134	7.34±0.29	89±9	65.2±7.4	4.48	858±20	27.2±0.8	2.40±0.28	
MZ2 <sup>a</sup>	34.171	0.199	0.141	1.02±0.04	19.4±7.4	19.7±7.5	3.90	657±14	18.1±0.5	1.09±0.42	
Sample name	$^{21}Ne$ ( $10^6$ atoms $g^{-1}$ )										
AMZ-1	6.7±0.7										
AMZ-3	5.3 +1.5/-1.2										
COL-1	4.3±0.6										
COL-2	4.05±0.44										

SDM samples are from the Sedom Formation. AMZ samples are from the base of the Amora Formation collected from the Amiaz-1 borehole. COL samples are from the Cave of the Letters. Sample MZ2 is from the Masada cave.

AMZ-1 collected from 1180-1212 m borehole depth, AMZ-2 collected from 1585-1640 m borehole depth, AMZ-3 collected from 1686-1723 m borehole depth.

$^{26}Al/^{27}Al$  ratios in bold indicate no  $^{26}Al$  counts were measured. The ratio presented here was calculated assuming one count.

<sup>a</sup> Measured at LLNL.  $^{10}Be/^{9}Be$  were normalized to 07KNSTD3110 =  $2.85 \times 10^{-12}$ .  $^{26}Al/^{27}Al$  were normalized to KNSTD 10650 =  $1.065 \times 10^{-11}$ .

<sup>10</sup>Be/<sup>9</sup>Be procedural blank was  $21.2 \pm 1.1 \times 10^{-15}$ . Standard reference materials at both ANSTO and LLNL are self-consistent.

<sup>b</sup> Measured at ANSTO. <sup>10</sup>Be/<sup>9</sup>Be were normalized to NIST SRM-4325 =  $2.790 \times 10^{-11}$ . <sup>26</sup>Al/<sup>27</sup>Al were normalized to Vogt SRM Z93-0221 =  $1.680 \times 10^{-11}$ .

<sup>10</sup>Be/<sup>9</sup>Be average of procedural blanks was  $20.1 \pm 2.0 \times 10^{-15}$ .

<sup>c</sup> Sample SDM-2 had a relatively high <sup>26</sup>Al/<sup>27</sup>Al background ratio ( $2.4 \times 10^{-15}$ ).

<sup>d</sup> Ratio calculated with <sup>26</sup>Al concentration of sample SDM-2 and <sup>10</sup>Be concentrations of the respective samples.

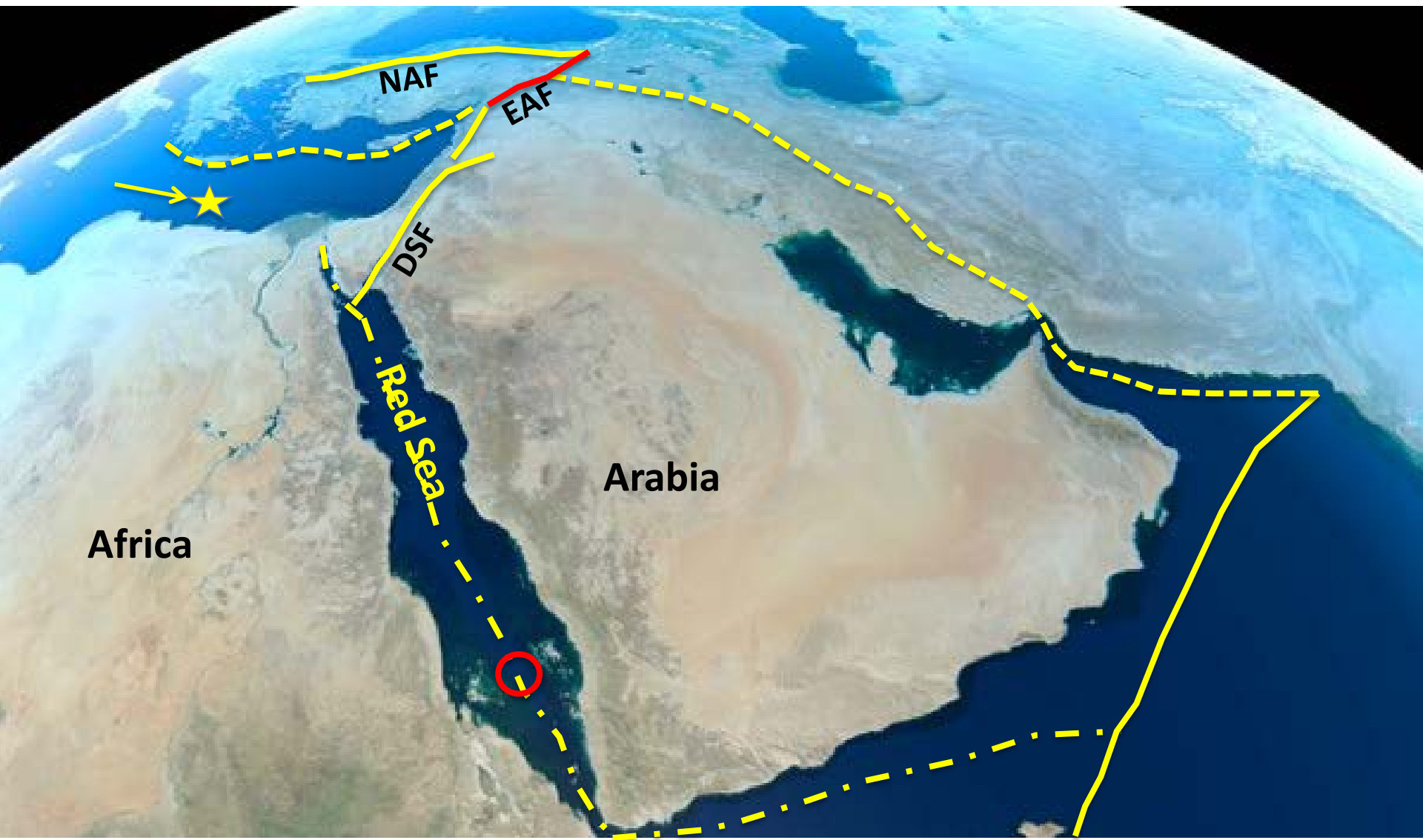
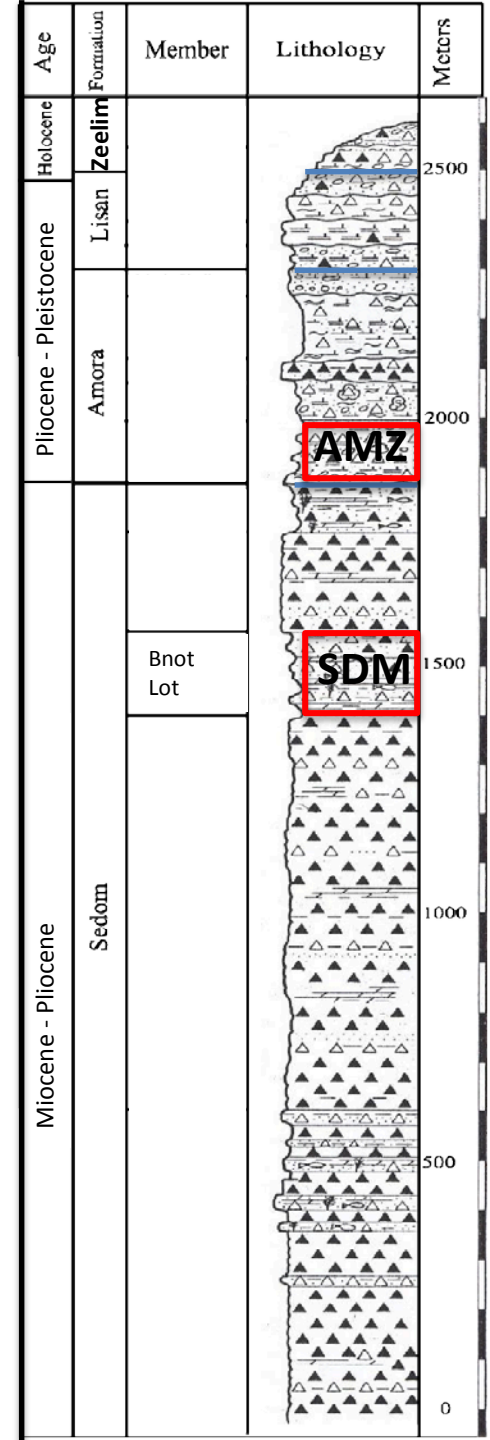
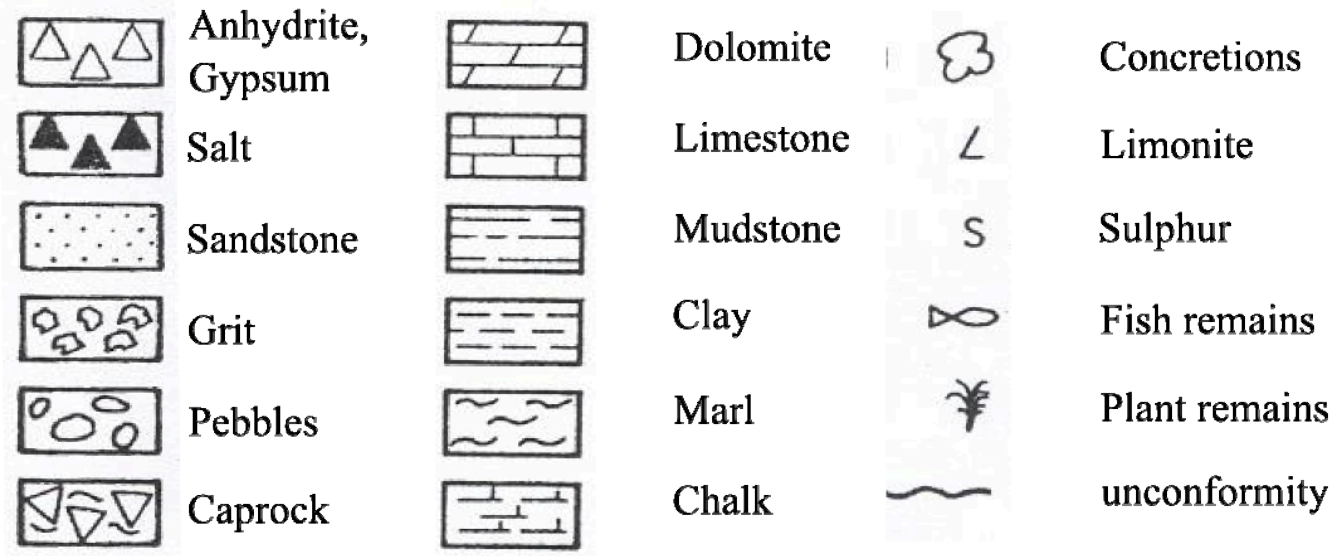
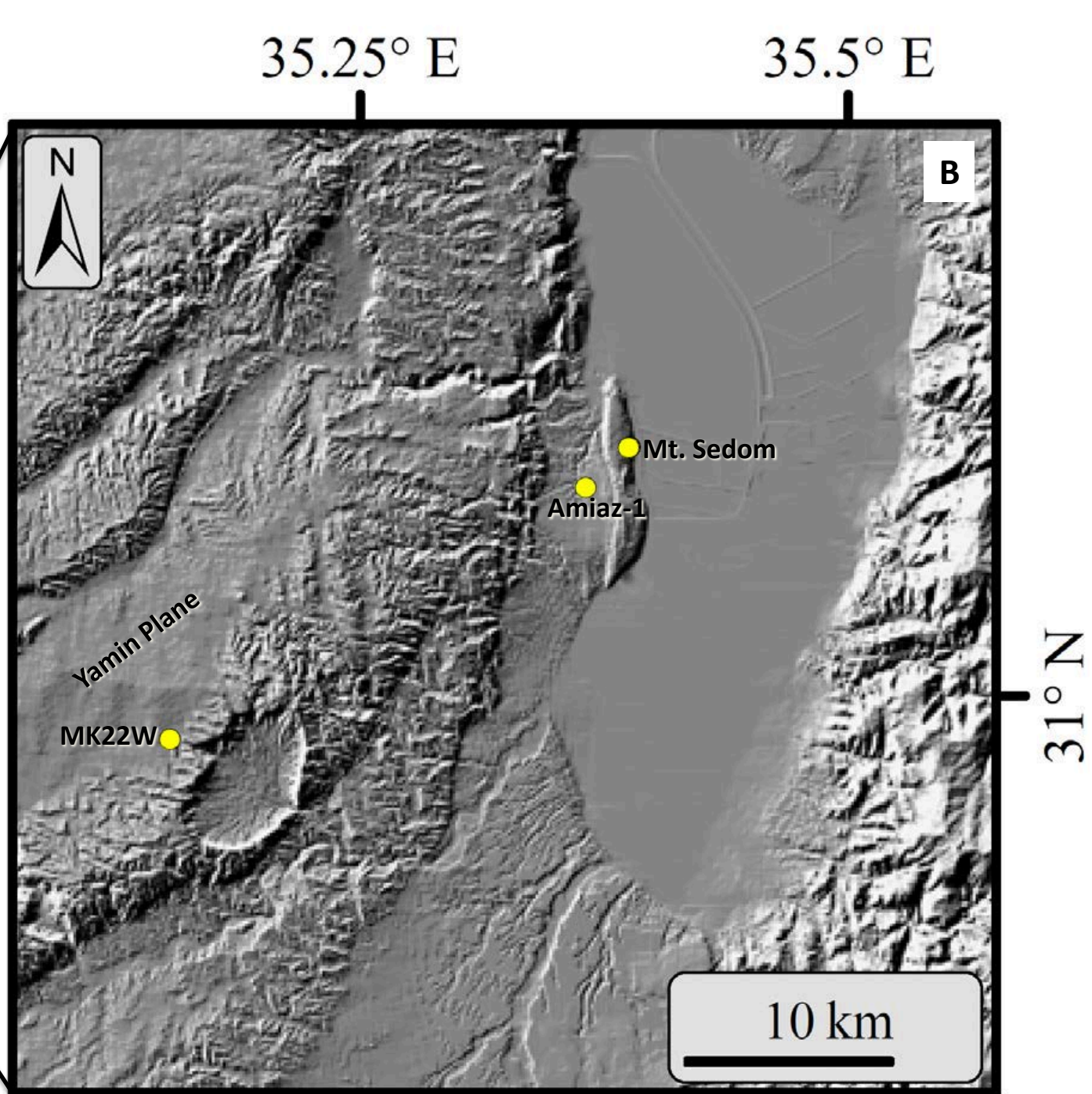
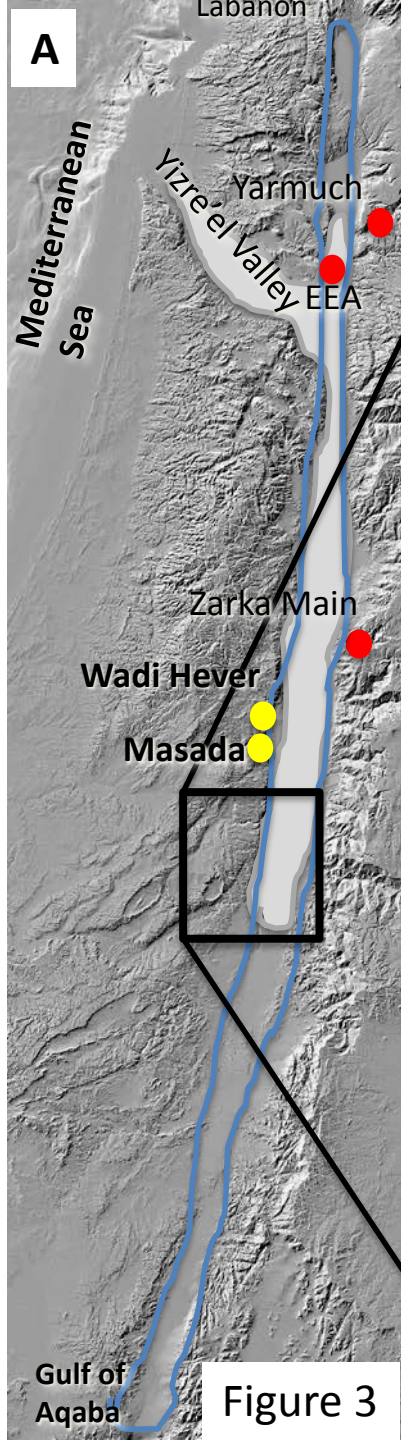


Figure 1



Figure 2





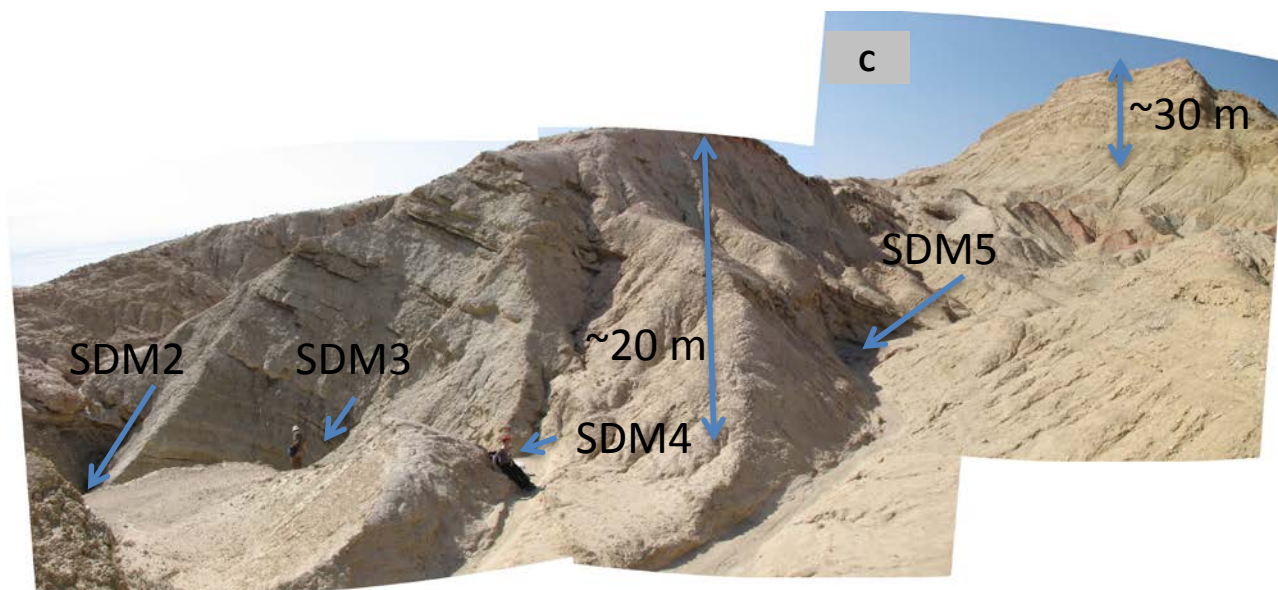
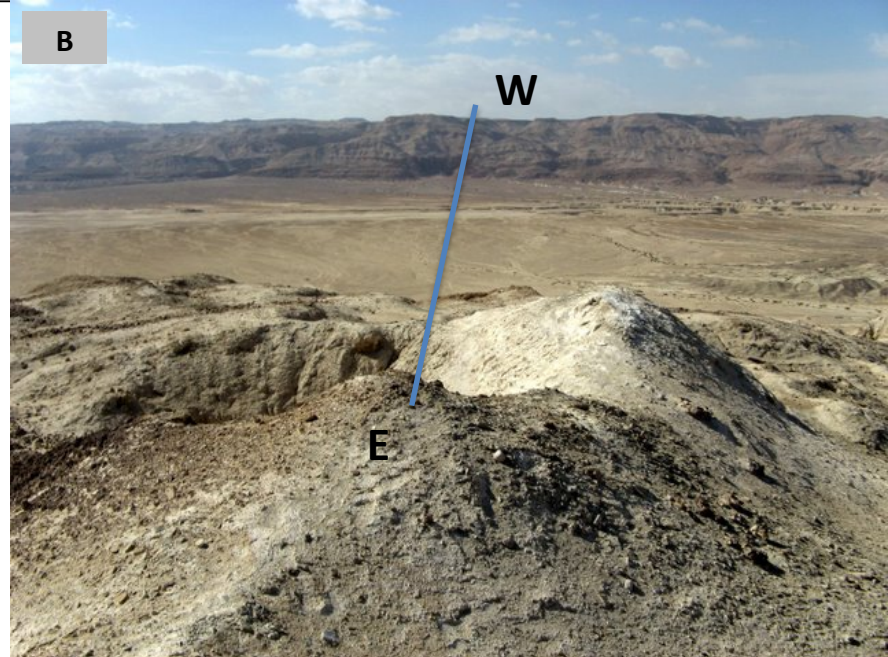
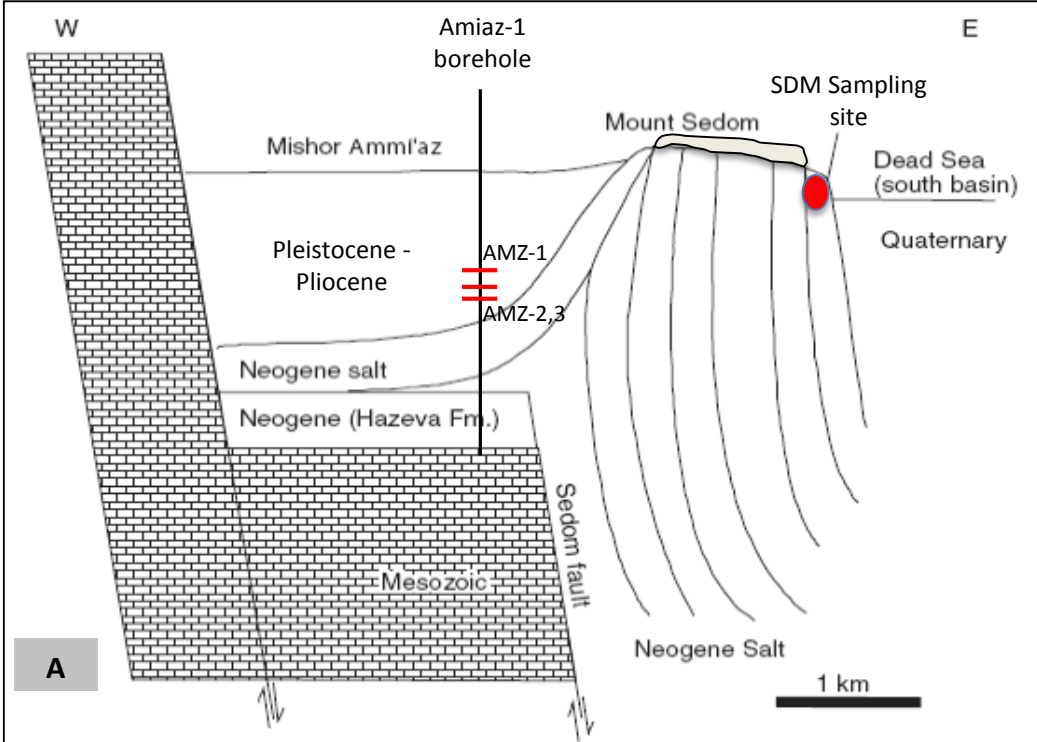


Figure 4

Figure 5

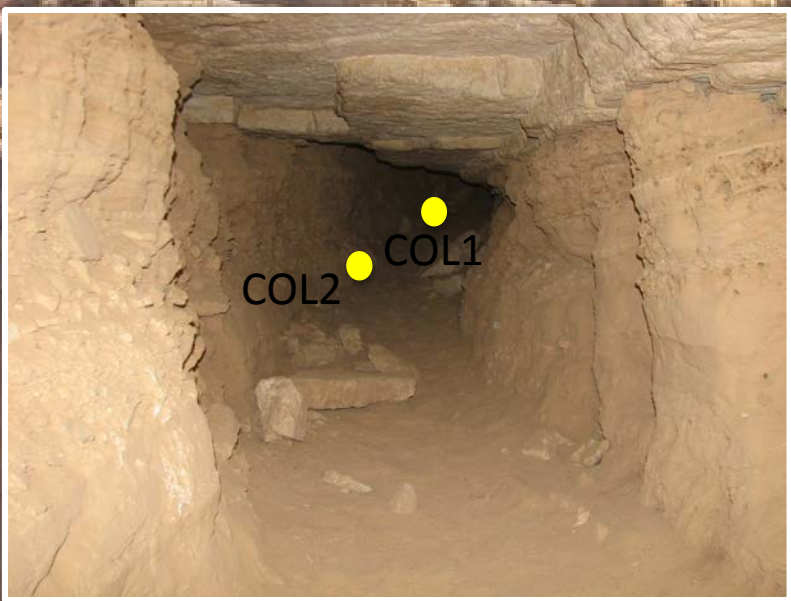
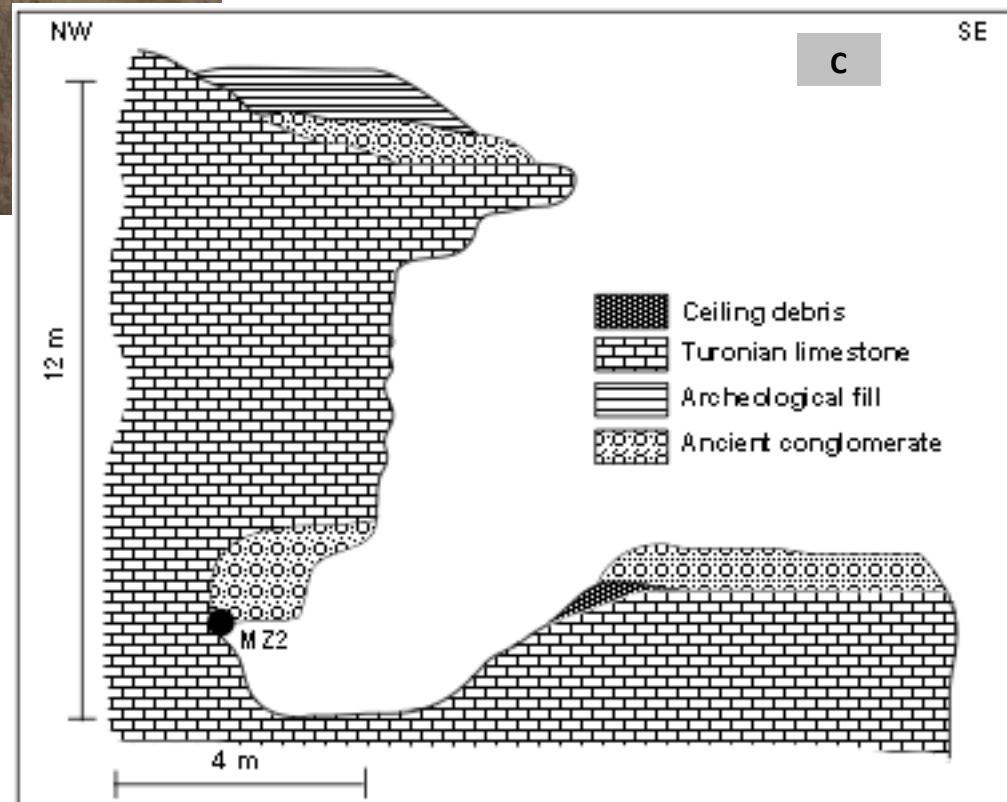


Figure 6



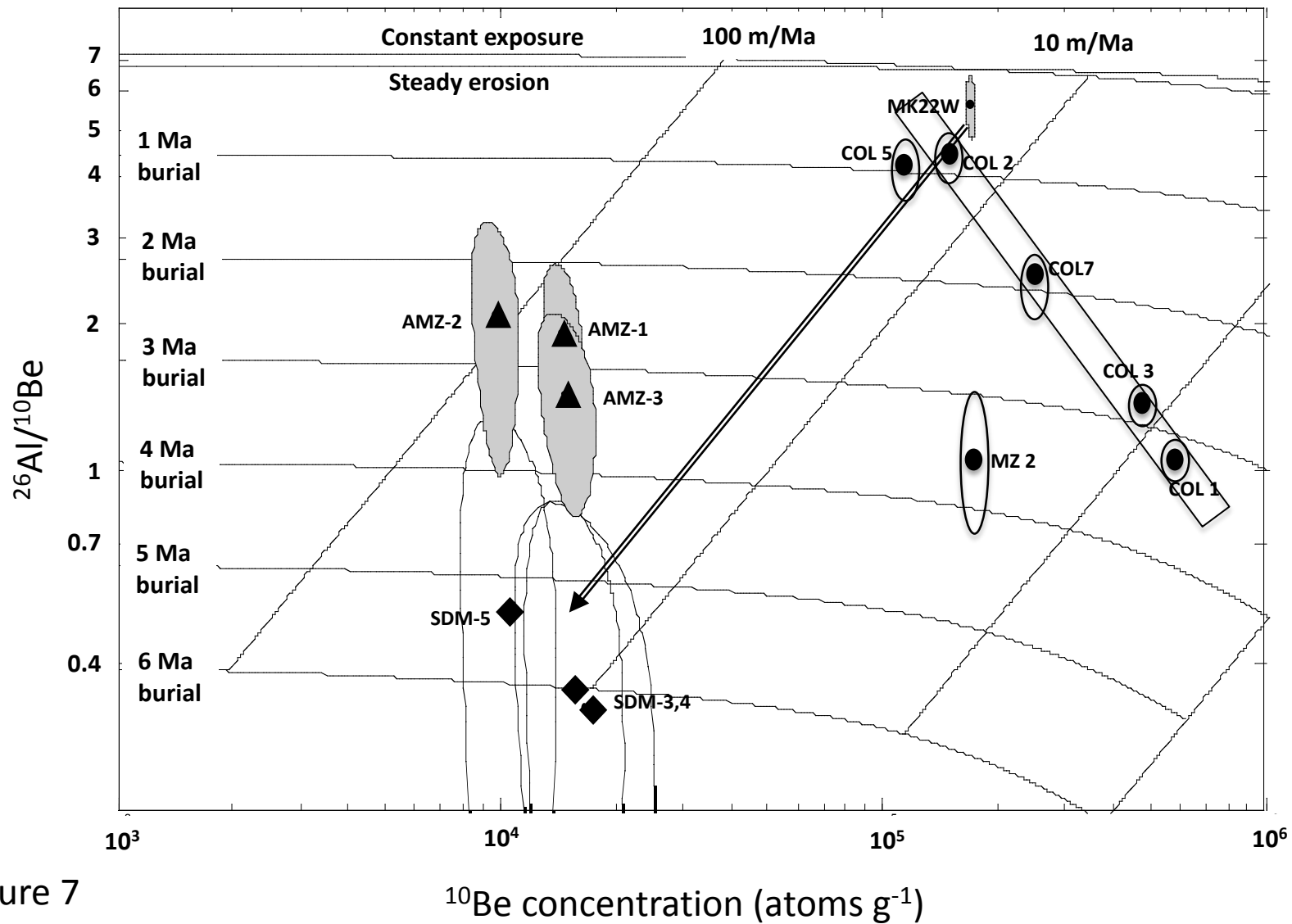


Figure 7

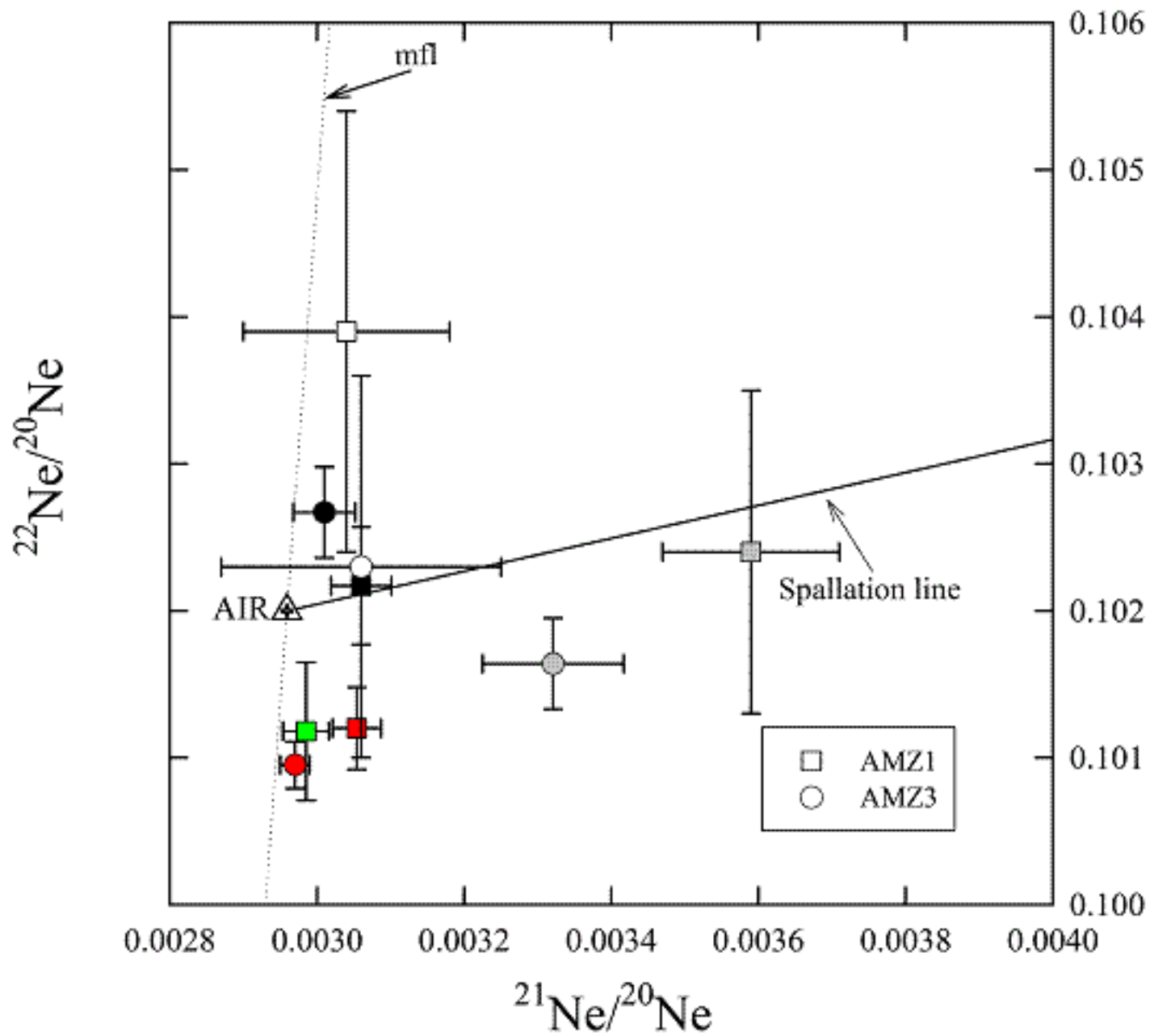


Figure 8a

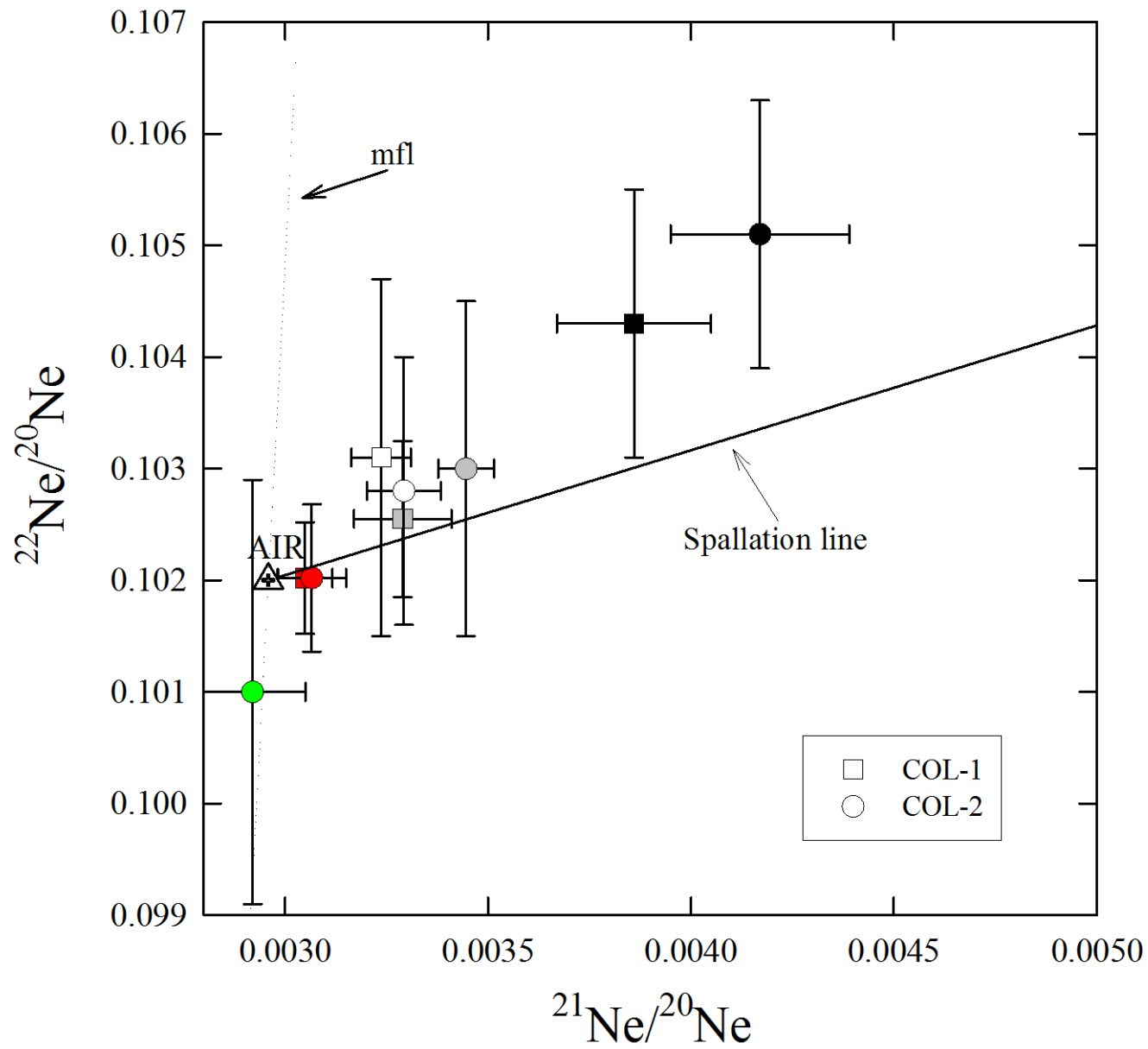


Figure 8b



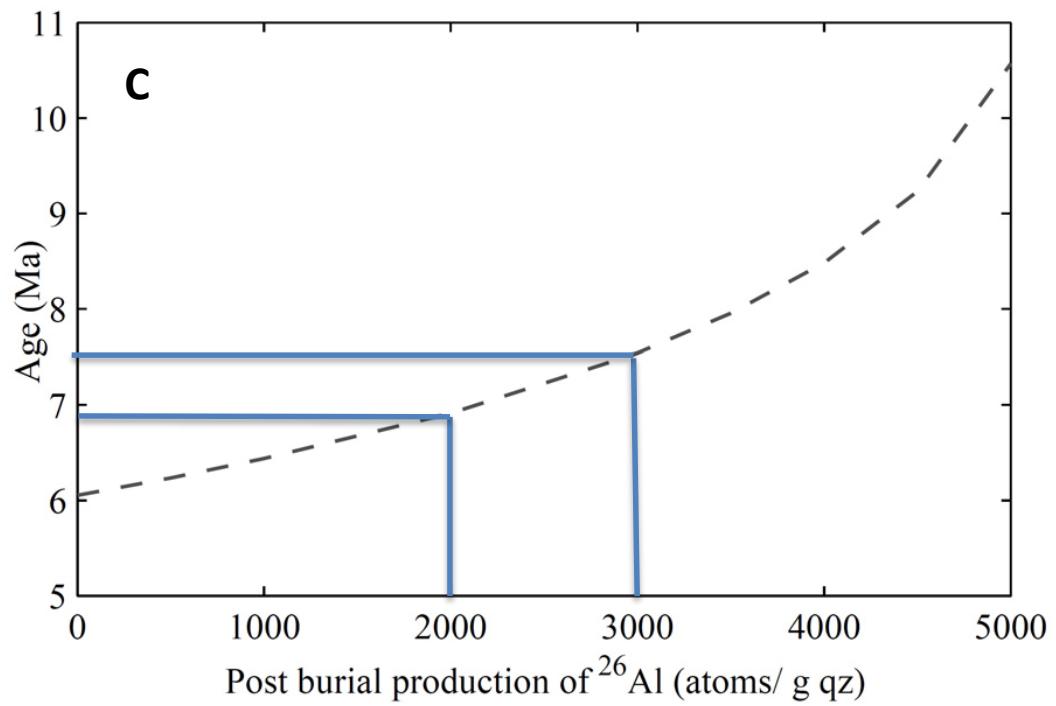
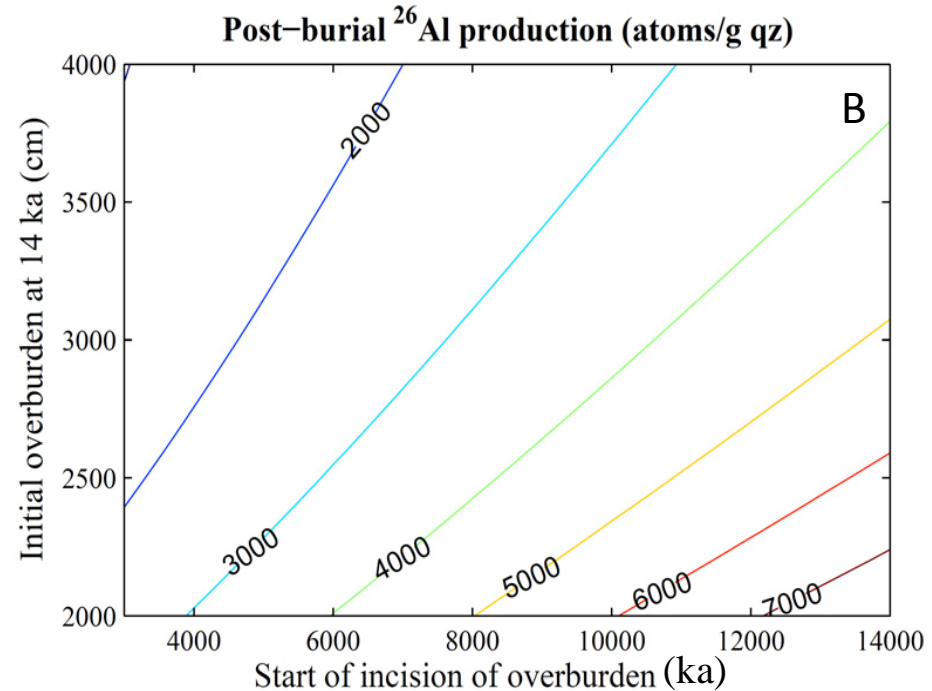
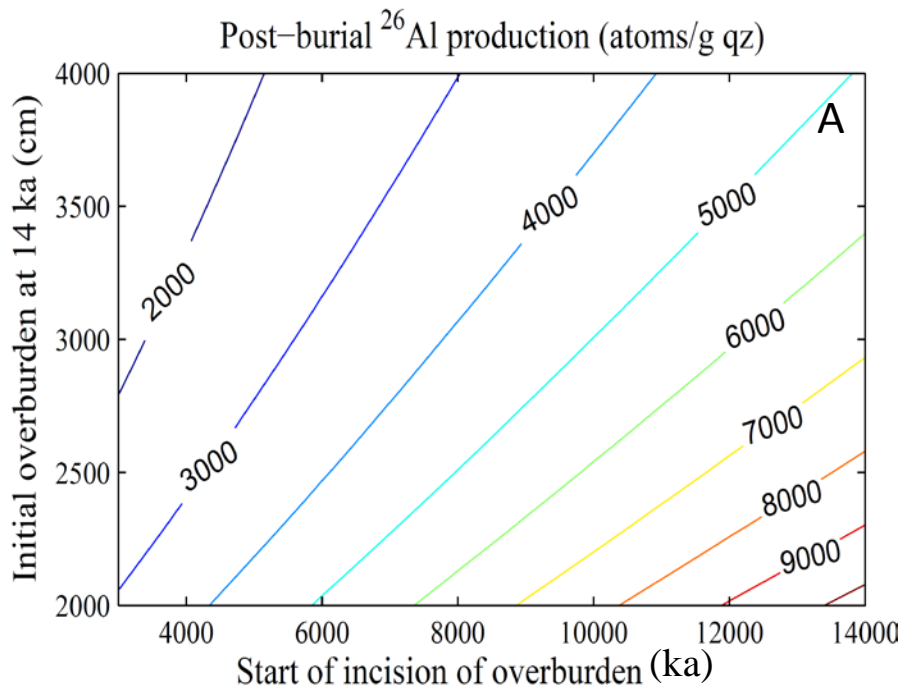


Figure 9