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1 Reconciling tectonic shortening, sedimentation and spatial patterns of erosion from  $^{10}\text{Be}$   
2 paleo-erosion rates in the Argentine Precordillera

3

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20 **Highlights**

21 - Paleo-erosion rates were extracted from the 7.7 – 1.8 Ma sedimentary rocks

22 - Spatial control of erosion reveals upstream migration of an erosion wave

23 - Peak catchment-wide erosion lags shortening pulse by ~2 Ma

24

25 **Abstract**

26 The temporal evolution of erosion over million-year timescales is key to understand the  
27 development of mountain ranges and adjacent fold-and-thrust belts. While models of  
28 orogenic wedge dynamics predict an instantaneous response of erosion to pulses of rock  
29 uplift, stream-power based models predict that catchment-wide erosion maxima  
30 significantly lag behind a pulse of rock uplift. Here, we explore the relationships between  
31 rock uplift, erosion, and sediment deposition in the Argentine Precordillera fold-and-  
32 thrust belt at 30°S. Using a combination of  $^{10}\text{Be}$ -derived paleo-erosion rates, constraints  
33 on re-exposure using  $^{26}\text{Al}/^{10}\text{Be}$  ratios, geomorphic observations and detrital zircon  
34 provenance, we demonstrate that the attainment of maximum upland erosion rates lags  
35 the maximum rate of deformation over million-year timescales. The magnitudes and  
36 causes of the erosional delays shed new light on the catchment erosional response to  
37 tectonic deformation and rock uplift in orogenic wedges.

38

39 **Keywords**

40  $^{10}\text{Be}$  paleo-erosion rates, rock uplift, compressional mountain range, time-lag, detrital  
41 zircon provenance; Andes

42

43 **1. Introduction**

44 Temporal variations of erosion rates should be diagnostic of whether tectonic or climatic  
45 processes govern mass fluxes in and out of orogens (e.g. Whipple, 2009). Sediment  
46 accumulation rates derived from magnetostratigraphy provide important clues about  
47 those mass fluxes, but are often insufficient for completing mass balances (i.e. linking  
48 erosion to deposition). For example, magnetostratigraphy only allows calculation of  
49 upland erosion rates if a basin is underfilled. Otherwise, the best inference is that  
50 sediment supply keeps pace with or exceeds the rate of creation of accommodation in the  
51 basin. In the latter case, sediment is bypassed beyond the depocenter, thereby hindering  
52 mass balance approximations. Therefore, despite routinely providing rates of flexural  
53 foreland basin development related to the history of upland deformation (Jordan et al.,  
54 1993), sediment accumulation rates are only half of the story. The missing half consists  
55 of linking the histories of deformation and sediment accumulation to upland erosion.

56

57 The combination of spatial and temporal patterns of wedge-widening or narrowing, near-  
58 surface rock uplift rates, erosion and sedimentation rates can reveal the temporal  
59 evolution of an orogenic wedge (Whipple, 2009). Numerical and analytical modeling  
60 consistently predict abrupt or gradual changes in erosion coeval with rock uplift as  
61 diagnostic signs of tectonic or climatic perturbations, respectively (Fig. 1a, b; Stolar et al.,  
62 2006; Whipple and Meade, 2006). In these models, the coupling between erosion and  
63 rock-uplift stems from isostatic compensation as well as the assumed constancy of critical  
64 taper angles, which requires commensurate changes in mass fluxes (Fig. 1a, b; Stolar et

65 al., 2006; Whipple and Meade, 2006; Whipple, 2009). This set of diagnostic behaviors  
66 can be investigated in foreland basins (Whipple, 2009).

67

68 The contrast between landscape response time and the duration of perturbation is also  
69 highly important when determining climate *versus* tectonics (Braun et al., 2015; Whipple,  
70 2009). The predicted post-perturbation landscape response times of orogenic wedges are  
71 primarily modulated by erosional efficiency, which is in turn governed by bedrock  
72 erodibility and precipitation (Stolar et al., 2006; Whipple and Meade, 2006). If response  
73 times are fast enough, changes in deformation and erosion in orogenic systems like the  
74 Central Andes can be simultaneous (Whipple, 2009). However, the magnitude and  
75 timescale of the tectonic perturbation can also govern fluvial response times at the  
76 catchment scale due to a primary control of channel gradient on knickpoint propagation  
77 (e.g. Kooi and Beaumont, 1996; Whittaker and Boulton, 2012). This relationship is  
78 further supported by numerical modeling of catchment-wide  $^{10}\text{Be}$ -erosion rates that  
79 suggests a time-lag between a change in tectonically driven rock uplift and a  
80 corresponding change in catchment-wide erosion rate (Fig. 1c; Willenbring et al., 2013).  
81 Thus, for the tectonic end-member, not only may the erosional response be gradual, but  
82 also lagged with respect to the tectonic perturbation (Fig. 1b, c).

83

84 Time-lags between upland exhumation, erosion and deposition are often observed in low-  
85 temperature detrital thermochronology (e.g. Garver et al., 1999). Since detrital  
86 thermochronology typically integrates over the upper kilometers of the crust and  $10^6$  to  
87  $10^7$  yr timescales, the technique lacks the sensitivity needed to resolve changes in erosion

88 on sub-Ma timescales (Rahl et al., 2007). Alternatively,  $^{10}\text{Be}$ -derived erosion rates are  
89 much more sensitive to temporal variations in erosion since they average over the  
90 uppermost meters of the Earth's surface and integrate erosion over  $10^3$ - $10^5$  yr. This  
91 sensitivity is sufficient to detect changes in erosion rates on timescales shorter than  
92 typical time-lags present in detrital thermochronology. In alluvium, measured  $^{10}\text{Be}$   
93 concentrations constitute an average of the upstream area, and therefore reflect a mean  
94 catchment erosion rate (e.g. Granger et al., 1996). The same concept can be applied to  
95 buried sediments of known age and allows estimation of paleo-erosion rates using  $^{10}\text{Be}$   
96 alone. For example, when extracted from the stratigraphic record,  $^{10}\text{Be}$  concentrations can  
97 reveal the history of erosion rates of the adjacent uplands (e.g. Charreau et al., 2011).  
98 This can be complemented with a record of the spatial patterns in upland erosion through  
99 sediment provenance indicators of unique igneous sources, such as detrital zircon  
100 crystallization ages (e.g. Gehrels, 2011). Applied together in connected intermontane and  
101 foreland basins, these methods constitute a powerful means for constructing a complete  
102 view of the dynamics of tectonics, erosion and deposition in orogenic systems that is  
103 sensitive at sub-Ma timescales.

104

105 In this study, we combine  $^{10}\text{Be}$  paleo-erosion rates and detrital zircon U-Pb  
106 geochronologic age distributions from three stratigraphic sections to construct a spatial  
107 and temporal view of erosion in the Andean foreland at  $30^\circ$  S from 8-2 Ma (Fig. 2). The  
108 period investigated here is marked by increased shortening and rock-uplift rates in the  
109 Precordillera fold-and-thrust belt as well as a relatively steady semi-arid climate (Fig. 3,  
110 4). This framework allows the evaluation of the tectonic end-member of an orogenic-

111 wedge's evolution (Fig. 1b). Thus, based on the modeling predictions of the interplay  
112 between erosion and tectonics (Fig. 1b, c), we expect to find a gradual and potentially  
113 lagged erosional signal with respect to deformation in the Precordillera.

114

## 115 **2. History of deformation and sediment accumulation**

116 At 30°S the main tectonic provinces in the Andes are, from west to east, the Principal and  
117 Frontal Cordilleras, the Precordillera fold-and-thrust belt and the Sierras Pampeanas (Fig.  
118 2a). Deformation in the Precordillera is well-constrained through the chronostratigraphy  
119 of thrust-top synorogenic sediments (Jordan et al., 1993), balanced cross-sections  
120 (Allmendinger and Judge, 2014), and low-temperature thermochronology (Fig 3a;  
121 Fosdick et al., 2015). Major thrust activity began ca. 19 Ma in the Western Precordillera  
122 (Fig. 3a; Jordan et al., 1993). Shortening rates and exhumation peaked ca. 12-9 Ma  
123 during simultaneous movement on the Blanco, Blanquitos, and San Roque thrusts (Fig.  
124 3a, b and Fig. 4; Allmendinger and Judge, 2014; Levina et al., 2014). From 8 to 1 Ma,  
125 deformation continued on the San Roque thrust and migrated east to the Niquivil thrust  
126 (Fig. 3a, 4). Immediately to the east, deformation and exhumation of the Bermejo Basin  
127 deposits initiated ca. 3-2 Ma through west-verging blind thrusts forming the Eastern  
128 Precordillera (Fig. 3a, 4; Fosdick et al., 2015; Jordan et al., 2001; Zapata and  
129 Allmendinger, 1996). Here, we targeted deeply exhumed stratigraphic sections along  
130 cores of anticlines that contain, to a first order, the history of erosional unloading of the  
131 adjacent tectonic provinces (Fig. 4).

132

133 The development of the Precordillera fold-and-thrust belt segmented the foreland into  
134 sub-basins (wedge-top and foredeep) early in its history (Beer et al., 1990; Jordan et al.,  
135 1993). Out-of-sequence exhumation occurred in the edges of the wedge-top basin ca. 5  
136 Ma and marks the latest detectable unroofing of the Frontal Cordillera and of the Cerro  
137 Negro and Tranca thrusts (Fig. 3a, 4; Fosdick et al., 2015). Deformation of the wedge-top  
138 deposits continued to occur until at least 300 ka along the El Tigre fault system (Fig. 4b;  
139 Siame et al., 1997). Here, we sampled the wedge-top section to obtain a view of mass-  
140 fluxes upstream of the wedge in addition to the foreland-N and foreland-S (Fig. 4).

141

142 Sediment accumulation rates in the flexural foredeep reflect the pace of tectonic  
143 subsidence and/or sediment supply during upland crustal thickening and tectonic  
144 shortening in the Precordillera (Fig. 3b, c; Johnson et al., 1986; Jordan et al., 1993).  
145 Paleomagnetic data from the section foreland-N indicate two accumulation rate maxima  
146 at 8 Ma and 4.5 Ma since ca. 15 Ma (Fig. 3c; Jordan et al., 1993, 2001). At the section  
147 foreland-S, accumulation rates peaked ca. 6 Ma and declined until 1.8 Ma (Fig. 3c;  
148 Milana et al., 2003). At foreland-S, a major change in sedimentary facies from small,  
149 ephemeral drainage system to one with high transport capacity occurred ca. 4 Ma (Milana  
150 et al., 2003), coincident with the abandonment of the Río Jáchal of the foreland-N area  
151 (Jordan et al., 2001). Meanwhile, in the wedge-top section, the history of sediment  
152 accumulation appears uncorrelated with Precordillera thrust activity (Fig. 3c; Ruskin and  
153 Jordan, 2007).

154

155 Eastward propagation of deformation in the Precordillera is often compared to the critical  
156 Coulomb wedge model (e.g. Allmendinger et al., 1990) that balances tectonic and  
157 erosional mass fluxes (Dahlen, 1990). In a critical taper that is not at steady state, an  
158 increase in crustal shortening may cause the mass flux into the wedge to temporarily  
159 outpace erosional mass flux out and steepen the taper angle. However, the steeper taper  
160 will tend to increase erosion rates and return the wedge to its critical state. Alternatively,  
161 climatically driven erosion will remove mass from the wedge and decrease its taper. This,  
162 in turn, triggers renewed, out-of-sequence deformation coeval with a retreat of the  
163 deformation front (i.e. wedge narrowing) to reestablish its critical taper. If the  
164 Precordillera behaves similarly, we expect erosion highs to accompany wedge-narrowing  
165 if they are climate-driven. Conversely, we expect wedge-widening to precede erosional  
166 signals if they are tectonically driven.

167

### 168 **3. Methods**

#### 169 3.1 Cosmogenic Nuclides

##### 170 3.1.1 Sampling strategy

171 We targeted stratigraphic sections in key locations within the orogenic wedge, which  
172 allows monitoring of the dynamics of paleo-erosion upstream and downstream of the  
173 active Precordillera fold-and-thrust belt (Fig. 4, 5). Location and depositional age data  
174 were obtained from previously published paleomagnetic studies in the wedge-top,  
175 foreland-N and foreland-S sections (Fig. 3-5). The depositional ages were corrected to the  
176 2012 geologic time scale (Hilgen et al., 2012). We minimized post-burial re-exposure by  
177 sampling from the base of overhanging ledges with at least 3 m of rock above the sample

178 wherever possible (Fig. A.1); otherwise, we targeted vertical cliffs with obvious signs of  
179 outcrop instability, such as rock fall. Nonetheless, we measured  $^{26}\text{Al}$  and  $^{10}\text{Be}$   
180 concentrations in three replicate samples (one from each section) to quantify the modern  
181 cosmogenic nuclide dose.

182

### 183 3.1.2 Sample preparation

184 Medium to coarse-grained sandstones were collected from each section. Samples were  
185 crushed and sieved to 250-500  $\mu\text{m}$  (n=16), or 150-500  $\mu\text{m}$  (n=2) for samples with low  
186 quartz yield, and processed to obtain pure quartz at Syracuse University. Cosmogenic  
187 beryllium ( $^{10}\text{Be}$ ) was extracted from quartz separates spiked with a  $^9\text{Be}$  carrier (372.5  
188 ppm) at the GeoForschungsZentrum (GFZ) following standard methods (von  
189 Blanckenburg et al., 1996).  $^{10}\text{Be}/^9\text{Be}$  ratios were measured with accelerator mass  
190 spectrometry (AMS) at the University of Cologne relative to standards KN01-6-2  
191 ( $^{10}\text{Be}/^9\text{Be}$  of  $5.35 \times 10^{-13}$ ) and KN01-5-3 ( $^{10}\text{Be}/^9\text{Be}$  of  $6.32 \times 10^{-12}$ ; Nishiizumi et al.,  
192 2007). Procedural blanks yielded a  $^{10}\text{Be}/^9\text{Be}$  ratio of  $(6.6 \pm 5.4) \times 10^{-16}$ . Blank-corrected  
193 sample ratios range from  $9.9 \times 10^{-15}$  to  $1 \times 10^{-13}$  (supplementary Table A.1).

194

195  $^{26}\text{Al}$  and  $^{10}\text{Be}$  were measured from separate, newly weighted aliquots of three replicate  
196 samples.  $^{26}\text{Al}$  was extracted at GFZ following standard methods (Goethals et al., 2009).  
197 Total stable  $^{27}\text{Al}$  in each sample was measured using inductively coupled optical  
198 emission spectrometer (ICP-OES) at GFZ using the 396.152 nm wavelength. No  $^{27}\text{Al}$   
199 spike was added to samples. For the procedural blank, we added 2.029 g of 1,000 ppm  
200 Merck-Carrier  $^{27}\text{Al}$ .  $^{26}\text{Al}/^{27}\text{Al}$  were measured at the Purdue Rare Isotope Measurement

201 Laboratory (PRIME) relative to standards in Nishiizumi et al. (2004). Raw Al AMS and  
202 ICP-OES data are provided in supplementary material (Table A.2). The measured  
203 procedural Al blank  $^{26}\text{Al}/^{27}\text{Al}$  ratio is  $0.8 \pm 0.4 \times 10^{-15}$ , which is well below the measured  
204 sample ratios. The replicate  $^{10}\text{Be}/^9\text{Be}$  was measured at the University of Cologne and its  
205 procedural  $^{10}\text{Be}/^9\text{Be}$  blank was  $2.23 \times 10^{-15} \pm 29\%$ , causing a 15%-30% difference  
206 between this and the original measurements. Due to the better blank measurements from  
207 the original analysis, we use the  $^{10}\text{Be}$  data associated with the lower procedural blank.

208

### 209 3.1.3 $^{10}\text{Be}$ and $^{26}\text{Al}$ production rates

210 Spallation and muon production rates are scaled using the time-dependent LSD scaling  
211 routine implemented in CRONUScalc (Lifton et al., 2014; Marrero et al., 2016). Scaling  
212 factors are relative to a sea-level high latitude (SLHL)  $^{10}\text{Be}$  and  $^{26}\text{Al}$  productions of  $4.09$   
213  $\pm 0.35 \text{ at.g}^{-1}.\text{a}^{-1}$  and  $28.6 \pm 3.3 \text{ at.g}^{-1}.\text{a}^{-1}$ , respectively (Phillips et al., 2015). For  
214 production during aggradation and burial ( $P_{burial}$ ), the elevations of the paleo-depocenters  
215 were assumed to be similar to modern: 800 m at foreland-N, 700 m at foreland-S, and  
216 1,500 m in the wedge-top, all at  $\sim 30^\circ\text{S}$ . For modern production during recent exhumation  
217 ( $P_{outcrop}$ ), we used sample elevations, latitudes and outcrop geometry (Gosse and Phillips,  
218 2001).

219

220 The production rate on the hillslope during erosion ( $P_{hillslope}$ ) is calculated based on the  
221 modern elevation of both the Frontal Cordillera and Precordillera ranges (Fig. 2c; see  
222 Section 4.1 about paleo-elevations). This is necessary because the physiography of the  
223 area is complex and forms multi-modal distributions of elevations that precludes the

224 conventional pixel-by-pixel arithmetic averaging (Fig. 2c). Thus, we fit a Gaussian curve  
225 to the hypsometry of each physiographic province to obtain a mean elevation that is  
226 independent of changes in the configuration of the catchment while assuming that the  
227 main source areas are the mountains (Fig. 2c). Excluding quartz-poor lithologies does not  
228 significantly change the hypsometries (Fig. A.2).

229

230 Based on detrital zircon provenance data from the literature (see Section 5.4, Fosdick et  
231 al., 2015), we assume that the sampled foreland-N deposits are sourced from the  
232 Precordillera range, so its elevation was used to determine  $P_{hillslope}$  for all foreland-N  
233 samples. We assume this is the same for foreland-S samples >2 Ma, but not for the two  
234 youngest samples, which contain Frontal Cordillera-derived zircons (see Section 5.4). In  
235 this case, we used the Jáchal catchment average elevation (Fig. 2a, 3,410 m). As all  
236 wedge-top samples are sourced from the Frontal Cordillera range, we use its 4,000 m  
237 average elevation to determine  $P_{hillslope}$  for those. Topographic shielding in the study area  
238 is <5% (supplementary material), which is within uncertainty of final  $^{10}\text{Be}$  production  
239 rates and therefore not applied in  $P_{hillslope}$ .

240

241 Lastly, paleo-production rates varied in the past due to secular variations in the strength  
242 of the geomagnetic field (Lifton et al., 2014). Currently available calculators of paleo-  
243 production rates (i.e. CRONUScalc, Marrero et al., 2016) show up to 50% variability of  
244 production rates in the last 2 Ma. Beyond 2 Ma, they are not solvable due to poor  
245 resolution of paleo-intensity records (see Balco et al., 2008). Since most of our samples  
246 are older than 2 Ma, our calculations do not account for, or include errors associated with,

247 these variations in the geomagnetic field strength because they are unknown.

248 Nonetheless, we point out that variations in paleo-production rates are not capable of  
249 generating variability in paleo-erosion rates of the same magnitude as those observed in  
250 our data when averaged over >50 ka (supplementary material).

251

### 252 3.1.4 Calculating paleo-erosion rates from $^{10}\text{Be}$ measurements

253 For studies of modern landscapes, the  $^{10}\text{Be}$  concentrations of quartz in alluvium are used  
254 to determine the average erosion rate over the quartz-contributing area upstream of the  
255 sample point (e.g. Granger et al., 1996). This is possible because the  $^{10}\text{Be}$  concentration  
256 in active stream sediments is inferred to reflect the residence time of particles in the  
257 upper-most meters of Earth's surface (i.e. the erosion rate on the hillslopes; Lal (1991)).

258 We denote this concentration resulting from erosion on the hillslopes as  $C_{\text{hillslope}}$ . Since  
259 we measure  $^{10}\text{Be}$  concentrations in sandstones in the foreland basin deposits, the  
260 measured  $^{10}\text{Be}$  concentrations ( $C_M$ ) contain additional  $^{10}\text{Be}$  that was produced during  
261 progressive burial of the alluvium ( $C_{\text{burial}}$ ) in the basins (e.g. Clapp et al., 2001), recent  
262 exhumation during basin inversion ( $C_{\text{exhumation}}$ ), and a period of exposure in outcrop  
263 ( $C_{\text{exposure}}$ ; see Section 5.1). Thus, the measured  $^{10}\text{Be}$  concentration, in  $\text{at.g}^{-1}$ , is:

264

$$265 \quad C_M = (C_{\text{hillslope}} + C_{\text{burial}}) \cdot e^{-\lambda t} + C_{\text{exhumation}} + C_{\text{exposure}} \quad (1)$$

266

267 where  $t$  is time (yr) since deposition (i.e. the age of the deposit from

268 magnetostratigraphy), and  $\lambda$  the decay constant for  $^{10}\text{Be}$  ( $5.1 \times 10^{-7} \text{ a}^{-1}$ ; Chmeleff et al.,

269 2010; Korschinek et al., 2010) or  $^{26}\text{Al}$  ( $9.8 \times 10^{-7} \text{ a}^{-1}$ ; Nishiizumi, 2004). Equation 1 is

270 then rewritten in terms of  $C_{\text{hillslope}}$ :

271

$$272 \quad C_{\text{hillslope}} = (C_M - C_{\text{exhumation}} - C_{\text{exposure}}) \cdot e^{\lambda t} - C_{\text{burial}} \quad (2)$$

273

274 which yields the concentration due to erosion on the hillslopes at the time of deposition.

275 Once  $C_{\text{hillslope}}$  is obtained, it can be used to calculate a paleo-erosion rate ( $E_{\text{hillslope}}$  in cm/a)

276 from:

277

$$278 \quad C_{\text{hillslope}} = \sum_{i,j,k} \left[ \frac{P_{\text{hillslope}_{i,j,k}}}{\lambda + (\rho_{\text{hillslope}} \cdot E_{\text{hillslope}}) / \Lambda_{i,j,k}} \right] \quad (3)$$

279

280 where  $P_{\text{hillslope}_{(i,j,k)}}$  is the average production rate of the source area by each type of  $^{10}\text{Be}$

281 production (spallation ( $i$ ), stopped muons ( $j$ ), fast muons ( $k$ ), in  $\text{at.g}^{-1}.\text{a}^{-1}$ ),  $\rho_{\text{hillslope}}$  the

282 density of the source rocks, assumed to be  $2.5 \pm 0.3 \text{ g.cm}^{-3}$ ,  $\Lambda_{i,j,k}$  ( $\text{g.cm}^{-2}$ ) are attenuation

283 lengths for each type of production mechanism, adopted as  $160 \text{ g.cm}^{-2}$  for spallation,

284  $1,500 \pm 100 \text{ g.cm}^{-2}$  for stopped muons,  $5,300 \pm 950 \text{ g.cm}^{-2}$  for fast muons (Braucher et

285 al., 2003). Equation 3 describes the integral of  $^{10}\text{Be}$  production over a particle's path from

286 great depths to the surface. For deep and continuous burial, the equation is similar, thus

287  $C_{\text{burial}}$  is:

288

$$289 \quad C_{\text{burial}} = \sum_{i,j,k} \left[ \frac{P_{\text{burial}_{i,j,k}}}{\lambda + (\rho_{\text{burial}} \cdot A_r) / \Lambda_{i,j,k}} \right] \quad (4)$$

290

291 where  $P_{burial(i,j,k)}$  is the production rate at the depositional surface,  $\rho_{burial}$  the density of the  
 292 overburden, assumed to be of uncompacted sand ( $1.7 \pm 0.1 \text{ g.cm}^{-3}$ ), and  $A_r$  the  
 293 accumulation rate from magnetostratigraphy (cm/a). In this case,  $A_r$  is an average from  
 294 the time of deposition to the time of burial beyond  $^{10}\text{Be}$  production by muons.

295

296  $C_{exhumation}$  requires sample depth information because we collected samples underneath  
 297 overburden. Its equation is also similar to equations 3 and 4:

298

$$299 \quad C_{exhumation} = \sum_{i,j,k} \left[ \frac{P_{outcrop_{i,j,k}} \cdot e^{\frac{-\rho_{ss} \cdot z_{sample}}{\Lambda_{i,j,k}}}}{\lambda + (\rho_{ss} \cdot E_{exhum}) / \Lambda_{i,j,k}} \right] \quad (5)$$

300

301 where  $P_{outcrop(i,j,k)}$  is the production rate at the outcrop,  $\rho_{ss}$  the density of the overlying  
 302 sandstones (assumed  $2.5 \pm 0.3 \text{ g.cm}^{-3}$ ),  $z_{sample}$  is the depth of the sample to the top of the  
 303 outcrop,  $E_{exhum}$  the exhumation rate of the basin deposits from thermochronology (cm/a;  
 304 (Fosdick et al., 2015)).  $E_{exhum}$  at foreland-N and the wedge-top was adopted as the mean  
 305 long-term erosion rates of the Las Salinas anticline (3.1 mm/a, Fig. 4) and of the Frontal  
 306 Cordillera (1.5 mm/a) respectively (Fosdick et al., 2015). Since there is no  
 307 thermochronological data from the foreland-S section we calculate the long-term  
 308 exhumation differently. Here, the deepest stratigraphic units (2,400 m) were exhumed  
 309 after 1.8 Ma (Milana et al., 2003), which yields a 1.33 mm/a long-term erosion rate at the  
 310 core of the Mogna anticline (see Fig. 4a). We use the replicate  $^{26}\text{Al}$  and  $^{10}\text{Be}$   
 311 measurements to confirm this estimate (see Section 5.1). At foreland-N and foreland-S,

312 we assume that  $E_{exhum}$  is inversely proportional to distance from anticline axis and  
313 decrease linearly to zero where rock exposure ends (Table 1).

314

315 Lastly, the concentrations accumulated during recent exposure in outcrop can be inferred  
316 as:

317

$$318 \quad C_{exposure} = t_{outcrop} \cdot S \cdot P_{sample_i} \quad (6)$$

319

320 where  $t_{outcrop}$  is the exposure time in the outcrop in years and  $S$  is a shielding factor  
321 ranging from 0 to 100% depending on outcrop geometry (Gosse and Phillips, 2001). In  
322 general, the longer  $t_{outcrop}$  is, the higher and less accurate the  $^{10}\text{Be}$  paleo-erosion rates  
323 become. For this reason, we deem outcrop geometries providing less than 50% shielding  
324 as too vulnerable to even short exposure times and discard one of our samples from  
325 further interpretation (sample HU-11). Constraints on  $t_{outcrop}$  are derived from the  
326  $^{26}\text{Al}/^{10}\text{Be}$  ratio and a-priori constrains on local exhumation rates.

327

### 328 3.1.5 Error propagation

329 We performed the calculations described above using a Monte Carlo method for error  
330 propagation. We assume that each parameter containing a mean and a standard deviation  
331 forms a normal distribution (i.e.,  $P_{source}$ ,  $P_{burial}$ ,  $P_{exhumation}$ ,  $\lambda$ , sediment density, attenuation  
332 lengths, and sample age). We then draw 10,000 random values from each parameter to  
333 compute 10,000 possible outcomes of equations 2-6. Finally, a mean  $C_{hillslope} \pm 1\sigma$  is

334 obtained from the simulations and used to calculate the paleo-erosion rate. This approach  
335 yields uncertainties that are comparable to the standard propagation of errors.

336

### 337 3.2 Detrital zircon U-Pb geochronology

338 Detrital zircons were extracted from the 60-250  $\mu\text{m}$  grain-size fraction from samples  
339 MOG-07 (foreland-S) and HC-07 (foreland-N). Uranium and lead isotopic analysis of  
340 zircons was conducted by Laser Ablation Inductively Coupled Plasma Mass  
341 Spectrometry. Only analyses with <20% discordance between  $^{206}\text{Pb}/^{238}\text{U}$  and  $^{207}\text{Pb}/^{235}\text{U}$   
342 ages were used for provenance interpretations (n=105 for MOG-07, n=79 for HC-07).  
343 Histograms and probability density functions were generated to discretize different  
344 crystallization ages (see Section 5.4). Information about analytical methods and data is  
345 provided in the supplementary material (Tables A.3-5 and Fig. A.3-6).

346

## 347 **4. Assumptions and sensitivity tests**

### 348 4.1 Paleo-elevations of the sediment source area

349 By using modern elevations to calculate paleo-production rates we are assuming that the  
350 source area did not undergo significant amounts of surface uplift. While there is  
351 paleoaltimetry-based evidence that this is true for the wedge-top area (i.e. Hoke et al.,  
352 2014), elevations must have changed in the Precordillera after motion of the San Roque  
353 and Niquivil thrusts since 8 Ma (Fig. 3a, 4). However, these two thrusts represent ~20%  
354 of the areal extent of the Precordillera and largely expose the Cambrian-Ordovician  
355 limestones (Fig. 4b). Thus, they cannot have drastically changed the resulting  
356 Precordillera hypsometry relevant to cosmogenic nuclides after they uplifted. We

357 demonstrate that this assumption is acceptable by modeling  $^{10}\text{Be}$  concentrations for a  
358 fixed paleo-erosion rate in an uplifting source-area and inverting the signal back to paleo-  
359 erosion rates while assuming constant elevation (Fig. A.7b). This exercise reveals that  
360 small temporal changes in mean elevation (i.e. 20%) cause  $\leq 30\%$  difference between  
361 inferred and true rates (Fig. A.7b). Such difference is well within the range of uncertainty  
362 of individual paleo-erosion rates inferred in this study.

363

#### 364 4.2 Continuous deposition and exhumation

365 We assume that the time between sediment entrainment and deposition is unimportant in  
366 terms of  $^{10}\text{Be}$  production since the distance between the Precordillera and the foreland is  
367 typically  $<40$  km. While it is impossible to determine if this assumption is violated, we  
368 would expect the pre-depositional exposure time to be different for every sample, which  
369 would lead to random temporal patterns of  $^{10}\text{Be}$ , which is not what we observe.

370

371 We must further assume that burial was quasi-continuous and sediments were not  
372 exposed for long periods of time between depositional events given that the temporal  
373 resolution of the magnetostratigraphy is  $> 50$  ka. This assumption is acceptable given the  
374 high accumulation rates observed in the foreland (Beer, 1990). Similarly, we assume that  
375 the recent exhumation phase was also continuous, consistent with sustained high  
376 exhumation rates since 3 Ma (Fosdick et al., 2015) and climate change in the Quaternary  
377 would only have favored erosion instead of inhibiting it.

378

379 Discontinuous deposition cannot be entirely dismissed in the wedge-top deposits. There,  
380 the low accumulation rates are the result of hiatuses of uncertain duration either due to  
381 erosion or non-deposition (Beer et al., 1990; Ruskin and Jordan, 2007). To minimize  
382 these uncertainties, we collected samples at the base of thick sedimentary beds whenever  
383 possible and avoided clear evidence of hiatuses in outcrop. After deposition of the  
384 sedimentary sequences, the wedge-top valley may have been largely filled by the deposits  
385 forming the perched remnant Médano and San Guillermo surfaces (Fig. 2b), which are  
386 necessarily younger than 5.2 Ma (youngest unit in wedge-top). This wedge-top fill  
387 potentially served as a  $\geq 100$ -meter-thick cap-layer that shielded the deposits in the valley  
388 from cosmic rays until their recent excavation. Nonetheless, we show the effects on final  
389 paleo-erosion rates for the case where samples were exposed for 100 ka at the base of a 1  
390 m thick layer of sediment prior to being completely buried (supplementary material).

391

## 392 **5. Results**

### 393 5.1 Exposure in outcrop

394 The foreland-N area has the lowest measured concentrations and is the most sensitive to  
395 re-exposure (Table 1). There, the magnetostratigraphic section is broken in two sub-  
396 sections that are  $\sim 15$  km apart (I and II, Fig. 2b). We collected duplicate samples of the  
397 same age ( $7.5 \pm 0.1$  Ma) along strike: HU-07 (section I) and HU-14 (section II; Table 1).  
398 Sample HU-07 was collected in a well-shielded outcrop where cosmic rays are blocked in  
399 all directions ( $S=0$ ) except from muons through the vertical column of rock, which we  
400 already corrected for. Its  $^{10}\text{Be}$  dose is  $1,280 \pm 183$  at.g $^{-1}$ , while sample HU-14, a much  
401 less shielded sample by comparison ( $S=0.4$ ), has a  $^{10}\text{Be}$  dose of  $1,734 \pm 167$  at.g $^{-1}$ . Their

402 concentrations become indistinguishable after correcting for  $C_{burial}$ ,  $C_{exhumation}$ , and 135  
403 years of exposure in outcrop ( $C_{outcrop}$ ) using HU-14's shielding geometry and local  
404 production rate of  $\sim 5.2 \text{ at.g}^{-1}.\text{a}^{-1}$ . The fact that these samples are  $\sim 15 \text{ km}$  apart with  
405 different shielding geometries and their  $^{10}\text{Be}$  budget is matched after minor corrections is  
406 strong evidence that these outcrop exposures are short. Lastly, relief within a 120 m  
407 radius at all of our sample sites at foreland-N is low and near the noise level of the SRTM  
408 1-arc second data (Fig. A.8). The areas surrounding each outcrop have similar  
409 topography and relief, suggesting that outcrop ages must be similar in neighboring areas.  
410  
411 We further constrain exposure times at other locations using the replicate measurements.  
412 Overall, the  $^{26}\text{Al}/^{10}\text{Be}$  ratios in the replicate samples are slightly elevated compared to the  
413 expected ratios based solely on  $C_{burial}$  and indicate some modern dose (Table 1). This  
414 modern dose can be partitioned into exhumation and outcrop exposure components  
415 (equations 5 and 6). At foreland-N, 55% of the dose is the result of muogenic production  
416 during exhumation of the sections based on the HU-16 replicate (Table 1). This  
417 corroborates the rapid exhumation rates inferred from apatite (U-Th-Sm)/He reported by  
418 Fosdick et al. (2015). Based on outcrop geometry, the remaining 45% of the modern dose  
419 requires  $1,000 \pm 300$  years of exposure, which is higher than the along-strike pair (HU-07  
420 and HU-14), but those are also closer to the Las Salinas anticline. Together, these  
421 observations support the assumptions of short exposure time. For foreland-S, our inferred  
422 1.33 mm/a exhumation rate of the Mogna anticline yields a  $C_{exhumation}$  higher than that  
423 accounted for by the  $^{26}\text{Al}/^{10}\text{Be}$  measured in MOG-05. Thus, we re-scaled the anticline  
424 core exhumation to 2 mm/a, after which  $C_{exhumation}$  fully accounts for the modern doses

425 measured in the MOG-05 replicate. This re-scaled exhumation rate is then used as  $E_{exhum}$   
426 in the exhumation corrections for the other samples in the Mogna section (Table 1). In the  
427 wedge-top,  $C_{exhumation}$  accounts for a small portion of the inferred modern dose, thus  
428 requiring  $\sim 1,700$  years of additional exposure in outcrop. The  $t_{outcrop}$  inferred above were  
429 conservatively applied in all other samples in their respective vicinities (Table 1). A  
430 detailed description of the cosmogenic nuclide budget,  $t_{outcrop}$  calculation, and further  
431 evidence of short outcrop exposure is provided as supplementary material.

432

### 433 5.2 Post-depositional corrections

434  $\beta$  ratios  $(C_{exhumation} + C_{burial} + C_{exposure}) / (C_{hillslope} \cdot e^{-\lambda t})$  are useful for checking what proportion  
435 of the measured concentration is made of post-depositional dose; 0.8 has been reported as  
436 an acceptable maximum (e.g. Schaller and Ehlers, 2006).  $\beta$  ratios for wedge-top samples  
437 are consistently  $< 0.5$  for all but the 5.2 Ma sample due to low accumulation rates (Table  
438 1). Four out of nine foreland-N samples yield  $\beta$  ratios  $\leq 1$  while those at foreland-S are all  
439  $> 1$  except for the oldest one (Table 1). In cases where post-depositional doses dominate  
440 the measured  $^{10}\text{Be}$  concentration, i.e.  $\beta > 5$ , only a minimum possible paleo-erosion rate  
441 was determined from the  $C_{hillslope} \cdot e^{-\lambda t} + 1\sigma$  (Table 1). In case all of  $C_M$  is removed by  
442 post-depositional corrections, minimum rates are calculated from the maximum possible  
443 concentration ( $1\sigma_M \cdot e^{-\lambda t}$ ; Table 1). Decay corrections are large and range from a factor of  
444 two (2 Ma) to  $\sim 46$  (7.7 Ma). A supplementary Table A.6 is provided for input in  
445 CRONUS online calculator.

446

### 447 5.3 Paleo-erosion rates

448 Paleo-erosion rates upstream of the wedge-top were  $\geq 0.15$  mm/a before 7 Ma and  
449 decreased below 0.1 mm/a until 5.2 Ma, when they increased back to 0.25 mm/a (+1.2, -  
450 0.1;  $1\sigma$ ; Fig. 6a; Table 1). For comparison, we calculated a range of possible paleo-  
451 erosion rates using the preserved sediment volumes from seismic imaging of wedge-top  
452 deposits (Fig. 6a; Fernández, 1996). We divided the total sediment volumes by the  
453 current drainage area of the wedge-top basin (after removing the area of the depocenter)  
454 and estimated a total vertical amount of sediment eroded in the catchment for each  
455 sequence assuming none of it was recycled from within the basin. This total erosion was  
456 then divided by the timespan of each sequence to obtain a long-term paleo-erosion rate  
457 (Fig. 6a; Table 2). The range of literature-derived paleo-erosion rates was inferred for  
458 hypothetical scenarios in which the deposit volumes were up to 3x greater than those  
459 reported by Fernández (1996) (i.e. >100% error on reported values). This conservatively  
460 accounts for possible large volumes of erosion during hiatuses. These paleo-erosion rates  
461 agree well with the ones obtained using  $^{10}\text{Be}$  between 7 and 5.2 Ma.

462

463 At foreland-N, the paleo-erosion rates are overall higher than in the wedge-top (Fig. 6b).  
464 Here, paleo-erosion rates peak ca. 6.8 Ma at a minimum 1.7 mm/a, which is  $\sim 2$  Ma after  
465 the end of a major pulse of shortening (Fig. 6b). By ca. 4.8 Ma, they decrease to 0.48  
466 mm/a (+0.19, -0.1;  $1\sigma$ ), increase again to 0.75 (+0.44, - 0.2;  $1\sigma$ ) and drop dramatically to  
467  $0.03 \pm 0.01$  mm/a ca. 3.2 Ma (Fig. 6b). It is interesting to note that the foreland-N paleo-  
468 erosion rates increase when rates in the wedge-top decrease coeval with higher wedge-top  
469 accumulation rates (i.e. 7-5.2 Ma; Fig. 6a, b). Meanwhile, paleo-erosion rates at foreland-  
470 S ranged from  $\sim 0.06$  mm/a (+0.02, -0.01;  $1\sigma$ ) ca. 7.5 Ma to  $\sim 1.14$  mm/a (+2.57, -0.47;

471 1 $\sigma$ ) ca. 4.8 Ma, higher than foreland-N at that time (Fig. 6c). Rates then reached 2.4  
472 mm/a (+5.7, -1; 1 $\sigma$ ) ca. 2 Ma and remained high or potentially increased ca. 1.8 Ma,  
473 however, a meaningful paleo-erosion rate cannot be recovered due to a very low <sup>10</sup>Be  
474 dose (Table 1; Fig. 6c).

**Table 1.**  $^{10}\text{Be}$  and  $^{26}\text{Al}$  data and paleo-erosion rates

$A_r$  - accumulation rates;  $E_m$  -  $E_{\text{exhum}}$  (modern strat. section erosion rate);  $C_M$  - measured concentration;  $C_b$  -  $C_{\text{burial}}$  ( $^{10}\text{Be}$ );  $C_{\text{exh}}$  -  $C_{\text{exhumation}}$  ( $^{10}\text{Be}$ );  $C_{\text{hill}}$  -  $C_{\text{hillslope}}$  ( $^{10}\text{Be}$ );  $z$  - depth below top of outcrop;  $S$  - scaling coefficient;  $\beta$  -  $(C_{\text{ex}}+C_{\text{expo}}+C_b)/C_{\text{hill}}\cdot e^{-\lambda t}$ ;  $t_{\text{oc}}$  -  $t_{\text{outcrop}}$ ;  $E$  - paleo-erosion rate. Concentrations are for  $^{10}\text{Be}$  unless otherwise indicated.

Sample	Age <sup>a</sup> (Ma)	$A_r$ (mm/a)	$E_m^b$ (mm/a)	$C_M \times 10^3$ (at.g <sup>-1</sup> )		$^{26}\text{Al}/^{10}\text{Be}$		$C_b \cdot e^{-\lambda t}$ $\times 10^3$ (at.g <sup>-1</sup> )		$C_{\text{exh}}$ $\times 10^3$ (at.g <sup>-1</sup> )		$C_{\text{expo}}$ $\times 10^3$ (at.g <sup>-1</sup> )		$C_{\text{hill}}$ $\times 10^3$ (at.g <sup>-1</sup> )		$z$ cm	S	$\beta$	$t_{\text{oc}}$ yr	$E$ mm/a				
				$^{10}\text{Be}$	$^{26}\text{Al}$	$\mu$	$1\sigma$	$\mu$	$1\sigma$	$\mu$	$1\sigma$	$\mu$	$1\sigma$	$\mu$	$1\sigma$					$\mu$	$+1\sigma$	$-1\sigma$		
<b>foreland-N</b>																								
HU-11 <sup>c,d</sup>	8.8	0.88	2.9	4.06	0.2	-	-	-	-	-	-	-	-	-	-	100	0.7	-	-	-	-	-		
HU-13	7.8	0.94	2.6	1.88	0.15	-	-	-	0.17	0.04	0.50	0.14	-	-	65.16	19.3	600	0	0.50	-	0.24	0.10	0.05	
HU-07	7.5	0.92	2.7	1.73	0.17	-	-	-	0.19	0.04	0.49	0.14	-	-	28.98	13.9	600	0	1.12	-	0.53	0.49	0.17	
HU-14	7.5	0.89	2.5	1.28	0.18	-	-	-	0.21	0.05	0.65	0.16	0.28	0.02	27.78	13.8	250	0.4	1.90	135	0.56	0.55	0.18	
HU-15	6.9	0.74	2.5	1.25	0.17	-	-	-	0.34	0.07	0.66	0.17	0.11	0.01	5.12	9.21	250	0.2	8.10	135	1.4 <sup>e</sup>	-	-	
HU-02	6.2	0.32	2.3	2.64	0.35	-	-	-	1.11	0.22	0.64	0.17	-	-	22.27	13.3	410	0	1.94	-	0.69	1.04	0.26	
HU-16	4.8	0.47	2.3	5.52	0.53	11.4	2.66	2.1	0.5	1.55	0.26	0.60	0.16	0.53	0.05	32.06	8.91	500	0.1	0.70	1000	0.48	0.19	0.10
HU-05	4	0.46	2.0	8.39	0.64	-	-	-	2.38	0.37	0.77	0.20	2.63	0.23	20.54	7.60	330	0.5	2.20	1000	0.75	0.44	0.20	
HU-17	3.2	0.35	1.7	95.6	3.3	-	-	-	4.68	0.66	0.80	0.22	2.63	0.23	445.27	49.0	500	0.5	0.09	1000	0.03	0.01	0.01	
<b>foreland-S</b>																								
MOG-02	7.5	0.51	1.2	6.34	0.34	-	-	-	0.35	0.08	0.56	0.23	2.35	0.20	248.3	58.4	2000	0.4	0.16	-	0.06	0.02	0.01	
MOG-05 <sup>c</sup>	4.9	0.28	0.8	5.99	0.41	14.6	3.53	2.5	0.6	3.24	0.56	1.67	0.44	-	-	13.61	9.41	489	0.3	1.18	-	1.14	2.57	0.47
MOG-09	1.9	0.21	0.4	22.3	2.17	-	-	-	14.7	1.86	2.88	0.74	2.94	0.25	12.56	8.80	350	0.5	3.75	-	2.44	5.72	1.01	
MOG-07	1.7	0.21	0.4	16.9	0.82	-	-	-	16.2	2.04	3.16	0.85	-	-	-	1.95	500	0.5	-	-	-	-	-	
<b>wedge-top</b>																								
ANG-04	7.6	0.15	1.5	13.1	0.55	-	-	-	2.02	0.47	1.27	0.33	7.79	0.66	104.12	64.5	400	0.5	5.55	1700	0.38	0.62	0.15	
R1-02	7.1	0.22	1.5	13.5	0.68	-	-	-	1.71	0.38	1.29	0.34	3.89	0.33	259.96	71.9	360	0.5	1.04	1700	0.15	0.06	0.03	
MD-02 <sup>f</sup>	6.8	0.28	1.5	36.9	1.44	-	-	-	1.64	0.42	0.02	0.03	-	-	1170.7	296.6	1000	0	0.04	1700	0.03	0.01	0.01	
R1-01	6.8	0.22	1.5	25.6	1.05	66.3	7.9	2.6	0.3	2.05	0.43	1.27	0.33	7.44	0.64	472.09	105.2	400	0.5	0.47	1700	0.08	0.02	0.02
ANG-07	6.3	0.3	1.5	28.7	1.32	-	-	-	1.91	0.38	1.26	0.33	7.79	0.67	457.90	96.1	400	0.5	0.61	1700	0.09	0.02	0.02	
ANG-02	5.2	0.03	1.5	59.6	2.2	-	-	-	40.4	7.59	1.23	0.32	7.79	0.66	160.34	132.3	450	0.5	4.88	1700	0.25	1.18	0.11	

a. Magnetostratigraphic age (Johnson et al., 1986; Jordan et al., 2001, 1993; Milana et al., 2003; Ruskin and Jordan, 2007). Each age was propagated through with a 0.1 Ma error.

b. Long-term erosion rates obtained from Fosdick et al. (2015).

c. Grain size fraction: 100-500  $\mu\text{m}$

d. Sample is poorly shielded and not used for interpretation.

e. Minimum possible rate based on  $C_{\text{hillslope}}+1\sigma$ .

f. This sample is located at the base of the fill forming the Médano paleosurface (Fig. 2b). Age was obtained through U-Pb analysis on zircons from a tuffaceous sandstone sample collected directly below. See Table A.5 for results from U-Pb analysis.

475

**Table 2:** Preserved sequence volumes in the wedge-top reported by Fernández (1996) and inferred erosion rates (this study)

Stratigraphic sequence	Age <sup>a</sup> (Ma)	Timespan (Ma)	Compacted vol. (km <sup>3</sup> )	Vol./Area <sup>b</sup> (km)	Erosion rate (mm/a or km/Ma)		
					1x Vol.	2x Vol.	3x Vol.
	8.69						
3	7.65	1.04	109.62	0.02	0.01	0.02	0.03
4	6.93	0.72	187.31	0.04	0.03	0.05	0.08
5	6.56	0.37	163.06	0.03	0.04	0.09	0.13
6	5.2	1.4	123.9	0.02	0.01	0.02	0.03

a. Age of top of sequence

b. Wedge-top catchment area is 10,000 km<sup>2</sup>

476

#### 477 5.4 Detrital zircons

478 The two detrital zircon samples for U-Pb crystallization ages help evaluate changes in  
479 source areas. At 1.8 Ma in foreland-S, detrital zircon ages are prominently centered ca.  
480 231-258 Ma (Fig. 7a). At 7.1 Ma in foreland-N, detrital zircon ages yield a pronounced  
481 age population of 290-320 Ma (Fig. 7b). Both samples yield less prominent early  
482 Paleozoic and Proterozoic zircons. Following the interpretation of sources by Fosdick et  
483 al. (2015), we attribute the pronounced 290-320 Ma age-peaks present in the 7.1 Ma  
484 sandstone to recycling of Lower Permian strata eroded from the Precordillera.  
485 Furthermore, we attribute the considerably younger 231-258 Ma age-peaks to erosion of  
486 the Permo-Triassic Choiyoi Group that dominates the Frontal Cordillera, upstream of the  
487 wedge-top (Fig. 4a).

488

## 489 6. Discussion

### 490 6.1 Evidence in favor of a tectonic control of mass fluxes

491 The first distinction between climate and tectonic drivers we can make is based on the  
492 erosional signature of foreland-N (Fig. 6b). If climate changes were governing mass

493 fluxes, we would expect high erosion rates in the wedge-top followed by a retreat of the  
494 deformation front (i.e. narrowing of the wedge). The opposite behavior is observed in the  
495 Precordillera (Fig. 3a, 6), where deformation continues to propagate into the Bermejo  
496 basin today (Fosdick et al., 2015). Higher river discharge related to climate change could  
497 theoretically lead to increased erosion everywhere (Braun et al., 2015). There is little  
498 evidence of long-lasting climatic shifts in and around the study area (Latorre et al., 1997;  
499 Ruskin and Jordan, 2007; Hynek et al., 2012; Cotton et al., 2014; Hoke et al., 2014) that  
500 could sustain the high erosion rates from 7.5 to 4 Ma observed in foreland-N (Fig. 3d,  
501 6b). More importantly, erosion rates would vary similarly in all of the studied sections if  
502 that were the case. Nonetheless, increased seasonality in the Pliocene (Latorre et al.,  
503 1997) may have contributed to changes in erosion rates; however the periodicity and  
504 magnitude of such changes (Godard et al., 2013) would be irresolvable in our data due to  
505 the temporal resolution of the magnetostratigraphy and uncertainties in determining  
506 paleo-erosion rates this old (Schaller and Ehlers, 2006). Lastly, most of the evidence of  
507 early glaciers in Argentina is from the Late Miocene - Early Pliocene and hundreds of  
508 kilometers south of the study area (Fig. 3e; Mercer, 1983). In the latitudes of the study  
509 area, glaciers are localized, inefficient erosion agents (Bissig et al., 2002), and thus an  
510 unlikely cause for our observed low  $^{10}\text{Be}$  concentrations, especially at foreland-S ca. 2  
511 Ma. For these reasons, we discard the hypothesis that erosion rates are primarily  
512 controlled by shifts in climate.

513

514 6.2 Evidence in favor of a tectonic control of base-level

515 The observed factor of five difference in erosion rates between foreland-N and the  
516 wedge-top until 5.2 Ma indicates that sediment sources were independent and that  
517 erosion was focused in the actively deforming Precordillera. These findings are consistent  
518 with the shift in provenance records showing dominance of Precordillera-derived  
519 recycled Carboniferous zircons and sharp decrease in Frontal Cordillera-derived Choiyoi  
520 zircons starting ca. 10 Ma and lasting at least until ca. 7.1 Ma during and after the major  
521 pulse of crustal shortening (Fig. 6, 7; Fosdick et al., 2015).

522

523 Since erosion was focused in the uplifted hanging walls of the Precordillera thrusts, we  
524 suggest that thrust activity caused fluvial disconnection between the wedge-top and the  
525 foredeep as proposed by Beer et al. (1990) (Fig. 8). If this is the case, then the wedge-top  
526 must have been internally-drained, suggesting that the perched paleo-surfaces now being  
527 consumed by knickpoints are geomorphic remnants of a former wedge-top fill (Fig. 8b).  
528 This interpretation is consistent with a tectonic control of local base-level in which thrust-  
529 motion favors the formation and filling of internally drained wedge-top basins as  
530 proposed for another intermontane basin to the NE (Hilley and Strecker, 2005). Although  
531 our data is not sufficient to confirm hydrologic isolation, it still indicates that the wedge-  
532 top's base-level fluctuated. For example, during periods of low accumulation rates (or  
533 hiatuses), the paleo-erosion rates inferred in this study are higher than those derived using  
534 Fernández's (1996) reported sequence volumes (Fig. 6a). This suggests that the formation  
535 of hiatuses could be coeval with higher erosion in the wedge-top rather than non-  
536 deposition and, thus, could be caused by base-level fall. It is unknown, however, if the  
537 wedge-top remained isolated or if it overflowed into the neighboring intermontane valley

538 to the south. The available  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$  in local lake deposits within the wedge-top do  
539 not allow to test for hydrologic isolation (Ruskin and Jordan, 2007). Nonetheless, the  
540 presence of evaporitic deposits alone suggests some degree of basin isolation (Fig. 5c;  
541 Ruskin and Jordan, 2007).

542

543 If the wedge-top became fluvially isolated from the foreland during deformation in the  
544 Precordillera, at some point it reconnected to form the modern landscape (Fig. 8c). We  
545 suggest that this reconnection is evident in our data based on the migration of rapid  
546 erosion from the Precordillera to the Frontal Cordillera (in the upstream direction)  
547 between 5.2 and 2 Ma. The reported change in sedimentary facies ca. 4 Ma at foreland-S  
548 (Milana et al., 2003) to a permanent axial river is consistent with the sharp decline in  
549 paleo-erosion rates observed ca. 3.2 Ma at foreland-N followed by an increase ca. 2 Ma at  
550 foreland-S. Combined with the reappearance of abundant Choiyoi detrital zircons from  
551 the Frontal Cordillera ca. 2 Ma, the coeval change in sedimentary facies and trade-off  
552 between paleo-erosion rates in the two foreland sections suggests the southward  
553 deflection of the Río Jáchal and the reestablishment of its fluvial connectivity with the  
554 wedge-top (Fig. 8b, c). Here, we stress that most of the older foreland basin strata contain  
555 low proportions of Choiyoi zircons. Thus, old foreland strata are only secondary sources  
556 for the substantial resurgence of this age population in 1.8 Ma deposits at foreland-S  
557 (Levina et al., 2014; Fosdick et al., 2015).

558

559 If the observations outlined above do imply reconnection with the wedge-top after 5 Ma,  
560 then the post-2 Ma  $^{10}\text{Be}$  signal at foreland-S must be an admixture of recycled wedge-top

561 sediments and sediments sourced from the actively eroding Precordillera and Frontal  
562 Cordillera. Given the age range of the wedge-top strata, at least two half-lives of  $^{10}\text{Be}$   
563 decay occurred prior to being remobilized and deposited at foreland-S. Combined with  
564 the higher long-term erosion rates from the Precordillera and Frontal Cordillera (Fosdick  
565 et al., 2015), the low  $^{10}\text{Be}$  concentration at foreland-S is plausible.

566

### 567 6.3 Lagged erosional response to tectonic shortening

568 In the stream power construct, upstream migration of higher erosion rates coincides with  
569 the propagation of knickpoints through a catchment. Whether generated by rock uplift or  
570 base-level fall, knickpoints migrate upstream as a wave that separates the fast and slowly  
571 eroding parts of the landscape (Whipple and Tucker, 1999). Since  $^{10}\text{Be}$  concentrations in  
572 alluvium average over the entire upstream area, the mean  $^{10}\text{Be}$  concentration reflects the  
573 proportion of sediments derived from those fast and slowly eroding parts of the  
574 landscape. After the erosional wave covers half the catchment,  $^{10}\text{Be}$  erosion rates reach a  
575 maximum, lagging behind its trigger mechanism (Willenbring et al., 2013). We observe a  
576  $> 2$  Ma lag between erosion and shortening maxima at foreland-N (Fig. 6b). Combined  
577 with the later shift in provenance to a Frontal Cordillera source, this leads us to the  
578 interpret our data to reflect the time necessary for an erosion wave, which initiated at the  
579 thrust-front (i.e. knickzone), to migrate through the Precordillera and remove most of the  
580  $^{10}\text{Be}$ -rich, relict upper meter of rock (Willenbring et al., 2013). In this case, the  
581 occurrence of resistant limestones at the toes of the distal thrusts (Fig. 4), as well as  
582 continued thrusting, likely slowed these knickzones down and contributed to the observed  
583 lag.

584

585 Ultimately, it is the erosional efficiency, particularly modulated by precipitation, that  
586 rules an orogen's response time after an external perturbation (Stolar et al., 2006;  
587 Whipple and Meade, 2006). Perhaps it is not surprising that the Precordillera took at least  
588 2 Ma to reach its peak catchment-wide erosion rate under a semi-arid climate. Despite a  
589 long response time, the observed high erosion rates sustained from 7.5 to 4 Ma (Fig. 6b)  
590 are consistent with an erosional response to structural widening of critical tapers. This  
591 suggests that modeling efforts, such as those by Stolar et al. (2006) and Whipple and  
592 Meade (2006), accurately capture the overall nature of this system. However, the  
593 existence of a time-lag as predicted by Willenbring et al. (2013) and observed here may  
594 represent a way in which the transients (i.e. erosion and rock-uplift) are decoupled, thus  
595 deviating from Whipple and Meade's (2006) and Stolar et al.'s (2006) predictions (Fig.  
596 1b). This in turn suggests that the assumption of self-similar growth at constant taper  
597 angles may not always be appropriate, at least for arid orogens.

598

## 599 **7. Conclusion**

600 We extracted cosmogenic  $^{10}\text{Be}$  paleo-erosion rates from exhumed Upper Miocene-  
601 Pliocene foreland basin strata. The range of paleo-erosion rates observed in this study is  
602 typical of catchment-wide erosion rates in tectonically active areas both globally and  
603 locally (Pepin et al., 2013; Portenga and Bierman, 2011; Walcek and Hoke, 2012). The  
604 paleo-erosion rates reveal an  $\sim 2$  Ma lagged catchment-wide maximum in erosion relative  
605 to the end of a major period of tectonic forcing as predicted by Willenbring et al. (2013).  
606 Moreover, this lag is followed by sustained, high rates of erosion as envisaged by

607 Whipple and Meade (2006) and Stolar et al. (2006). Based on the magnitude of the  
608 observed lag, we conclude that semi-arid transient landscapes, such as the Precordillera,  
609 preserve geomorphic signals of base-level perturbation over Ma timescales. This implies  
610 that average catchment  $^{10}\text{Be}$  measurements may reflect perturbations that occurred  
611 millions of years earlier. Lastly, given the temporal and spatial dynamics of erosion  
612 observed in this study, we conclude that, in areas of similarly dry climate, structural  
613 widening of orogenic wedges is followed by erosional lags and sustained high erosion  
614 rates. During this period, headward waves of erosion may outpace wedge-widening and  
615 migrate from the active thrust-front, across the wedge-top basin, and ultimately into the  
616 core of the orogen.

617

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628

629 **Appendix A. Supplementary material**

630 Supplementary data to this article can be found online.

631 **FIGURE CAPTIONS**

632 **Fig. 1.** Contrasting models of erosional response to external perturbations in an orogenic  
633 wedge (a-b) and in a catchment (c). W&M' 06 denotes Whipple and Meade (2006) and  
634 W'13 denotes Willenbring et al. (2013). The gray rectangles mark the period of  
635 perturbation in each model. Colored lines denote the erosional response and black lines  
636 denote rock uplift rates. Magnitudes of changes in erosion and uplift rates were not kept  
637 to scale; only temporal variations matter. Note that the perturbation in c corresponds to a  
638 wave of uplift entering the catchment in the headwaters and exiting at its outlet. (For  
639 interpretation of the references to color in this figure legend, the reader is referred to the  
640 web version of this article.)

641  
642 **Fig. 2.** (a) Location of the study area showing tectonic provinces (colored polygons),  
643 outline of the Jáchal catchment (black line), outline of the assumed Precordillera paleo-  
644 catchment, and depths to subducting Nazca plate (solid white lines; Cahill and Isacks,  
645 1992). The black rectangle shows the location of Fig. 2b; (b) Topographic map from a 90  
646 m digital elevation model (<http://srtm.csi.cgiar.org>). Note that the foreland-N section is  
647 composed of two sub-sections. The wedge-top sub-basins are indicated by R (Rodeo) and  
648 Ang (Angualasto). The Médano (M) and San Guillermo (SG) paleosurfaces are outlined  
649 (solid white line). (c) Hypsometry of the Jáchal catchment and of the two major tectonic  
650 provinces as outlined in Fig. 2a. Note that hypsometry of the Frontal Cordillera contains a  
651 large area of intermontane, wedge-top valleys. (For interpretation of the references to  
652 color in this figure legend, the reader is referred to the web version of this article.)

653  
654 **Fig. 3:** Summary of tectonic (a, b), sedimentary (c), climatic (d), and other important  
655 events in South America (SA; e), such as the timing of flat-slab subduction and glaciation  
656 and inferred formation of the perched paleo-surfaces outlined in Fig. 2b. Abbreviations in  
657 (a) are: FC – Frontal Cordillera; CN – Cerro Negro; Tr – Tranca thrust; C – Caracol

658 thrust; B – Blanco thrust; Bq – Blanquitos thrust; SR – San Roque thrust; N – Niquivil  
659 thrust; NA – Niquivil anticline; SA – Las Salinas anticline. Thrusts are outlined in Fig.  
660 4a, except FC (not mapped at the surface).

661

662 **Fig. 4.** (a) Simplified bedrock geology based on published maps (Cardó and Díaz, 1999;  
663 Furque et al., 1998). Labeled, solid black lines are traces of major thrust faults. Sampled  
664 stratigraphic sections are marked by dashed red lines. (b) Modified structural cross-  
665 section of the area based on literature (line shown in Fig. 4a; Allmendinger and Judge,  
666 2014; Cardó and Díaz, 1999; Fosdick et al., 2015; Zapata and Allmendinger, 1996). (For  
667 interpretation of the references to color in this figure legend, the reader is referred to the  
668 web version of this article.)

669

670 **Fig. 5.** Stratigraphic positions for each sample. a) foreland-N (Johnson et al., 1986;  
671 Jordan et al., 2001, 1993); b) foreland-S (Milana et al., 2003); c) simplified wedge-top  
672 (Ruskin and Jordan, 2007). Samples R1-01 and R1-02 belong to the Rodeo sub-basin in  
673 the wedge-top (section not shown; see open circles for accumulation; locations shown in  
674 Fig. 2b) and are positioned in the profile based on age. Sample MD-02, also positioned  
675 by age, was collected from the base of the Médano surface (Fig. 2b, 3). Note difference in  
676 scales for section thicknesses. (For interpretation of the references to color in this figure  
677 legend, the reader is referred to the web version of this article.)

678

679 **Fig. 6.** Observed patterns of sediment accumulation and shortening history from literature  
680 (Allmendinger and Judge, 2014; Jordan et al., 2001; Milana et al., 2003; Ré et al., 2003;  
681 Ruskin and Jordan, 2007), and inferred paleo-erosion rates determined in this study  
682 (squares) with  $1\sigma$  uncertainty envelopes. (a) wedge-top paleo-erosion rates show opposite  
683 pattern to the foreland-N curve (b). Green band shows range of possible erosion rates  
684 using the volumes of deposits in the wedge-top (see Section 5.3 for details; Fernández,  
685 1996); (b) foreland-N results. Locus of erosion: Precordillera (see text). (c) foreland-S  
686 results show a pronounced increase in erosion rates ca. 2 Ma. Locus of erosion: wedge-  
687 top, Precordillera, and Frontal Cordillera (see Section 5.4). Numbers in parenthesis

688 denote  $1\sigma$  uncertainty where they fall out of the plot. (For interpretation of the references  
689 to color in this figure legend, the reader is referred to the web version of this article.)

690

691 **Fig. 7.** Probability density plot of detrital zircons (black line) from this study at foreland-  
692 S (Mogna Fm.) ca. 1.8 Ma (a) and at foreland-N (Río Jáchal Fm.) ca. 7.1 Ma (b) showing  
693 the relative contributions of each grain age. Colored bands highlight the age range for  
694 different sources (color-coded as in Fig. 4). 0-3.5 Ga plots are provided in Fig. A.3-6.  
695 Data from Fosdick et al. (2015) are shown for comparison (c-e).

696

697 **Fig. 8.** Conceptual model for landscape evolution of the study area as interpreted in this  
698 study overlain on the present day's topography. Blue arrows show hypothesized main  
699 fluvial flow directions. The San Roque and Niquivil thrusts are the only structures  
700 outlined but other Precordillera thrusts were active during the periods highlighted (Fig.  
701 3a). Faint contours are in 1 km (black) and 0.5 km (gray) intervals. Wedge-top fill  
702 depicted in (b) was obtained by interpolation of low-slope alluvial surfaces in the wedge-  
703 top. Paleo-surfaces: Médano (M) and San Guillermo (SG). (For interpretation of the  
704 references to color in this figure legend, the reader is referred to the web version of this  
705 article.)

706

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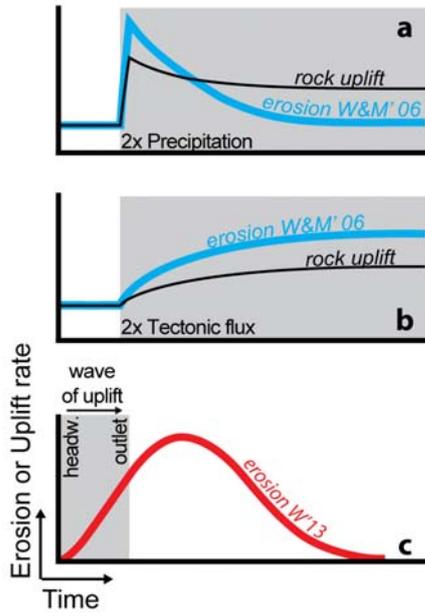
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903 **FIGURES**

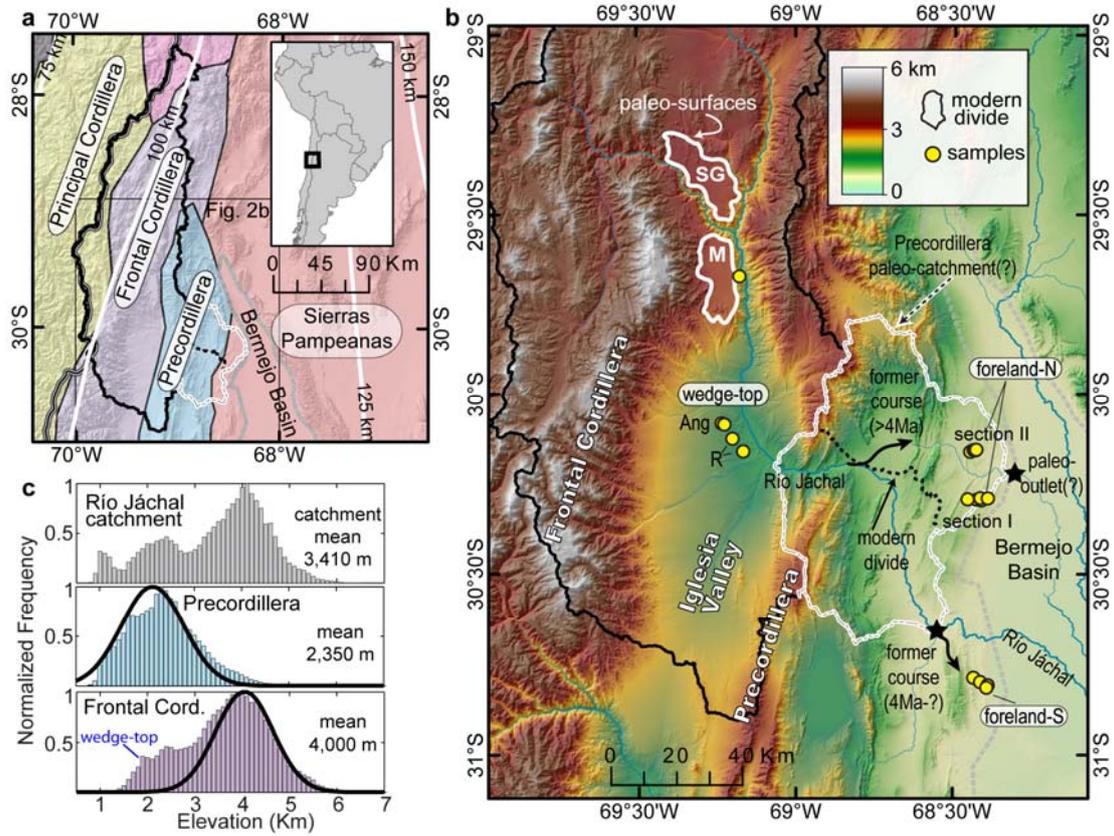
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Figure 1 (single column)



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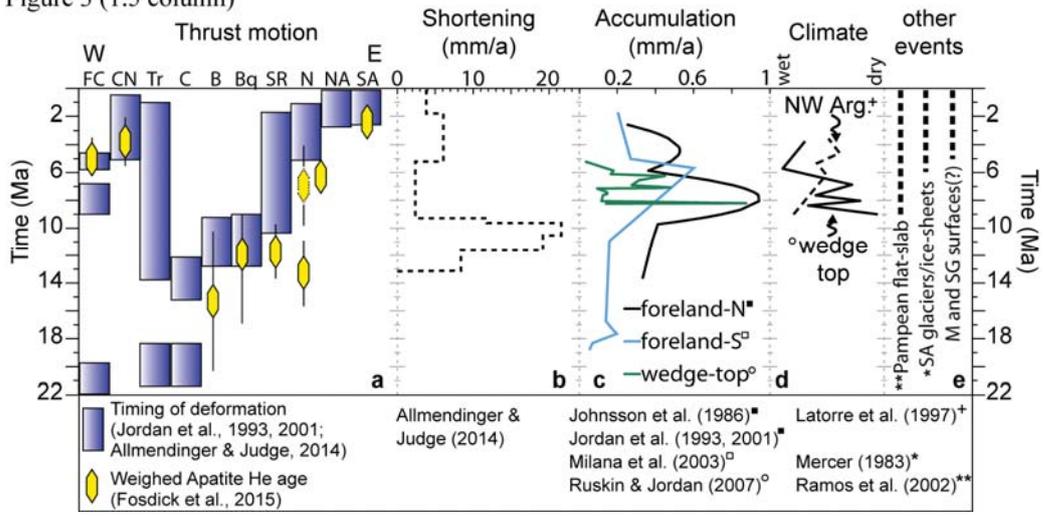
Figure 2 (double column)



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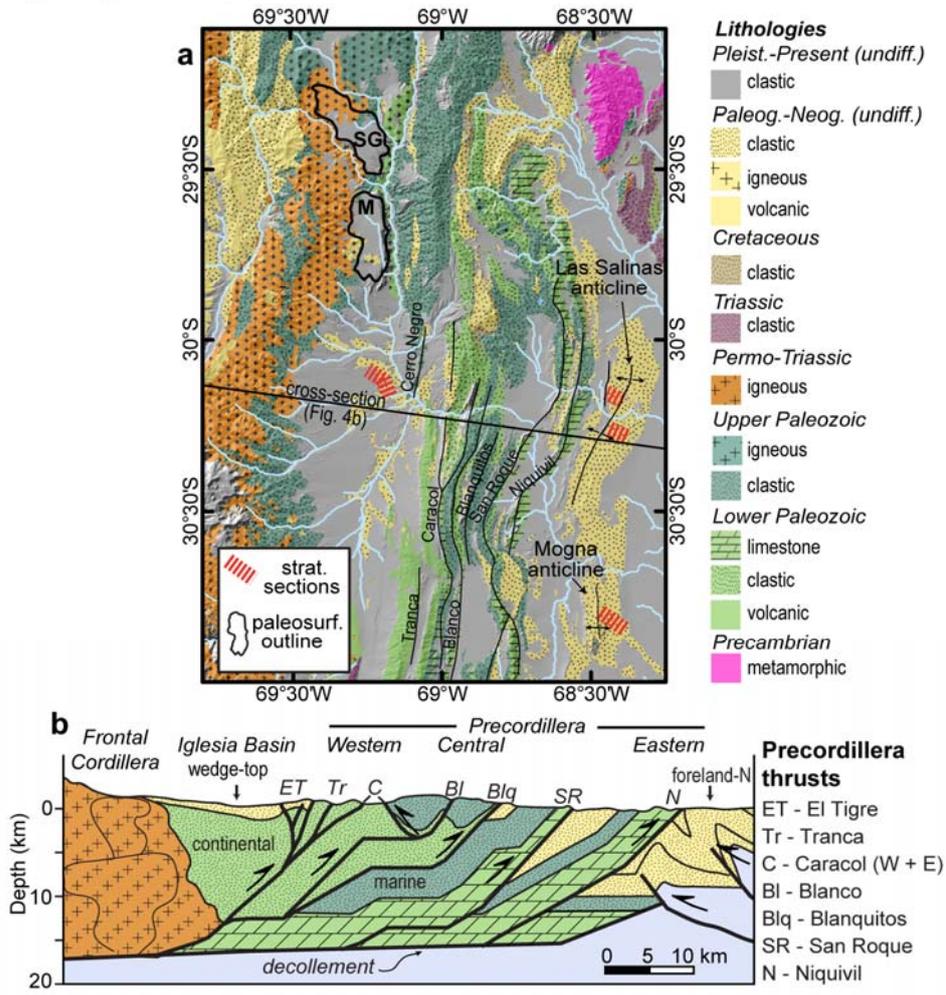
Figure 3 (1.5 column)



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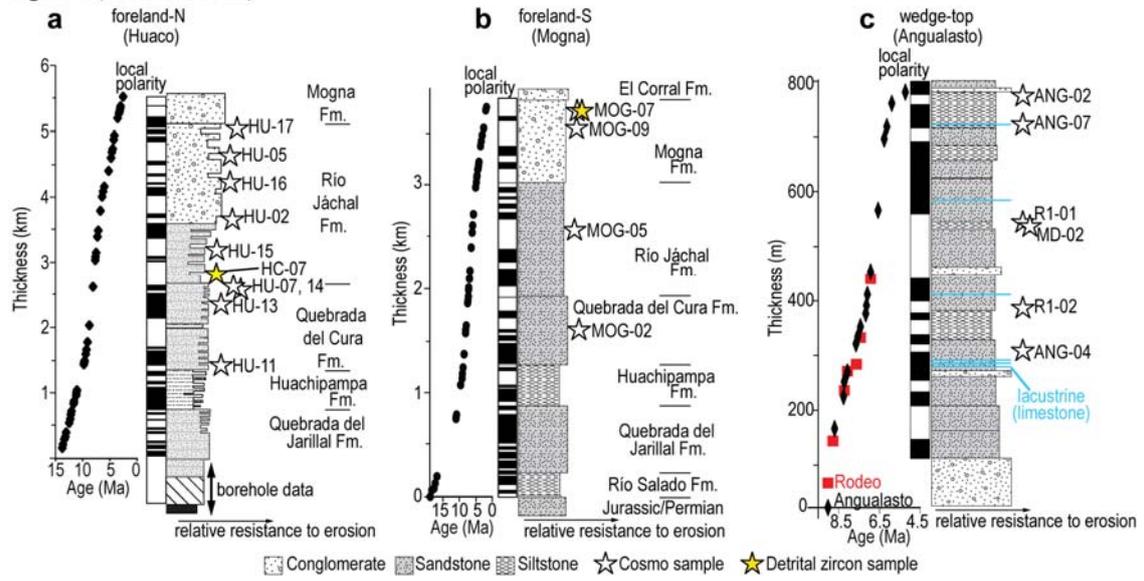
Figure 4 (1.5 column)



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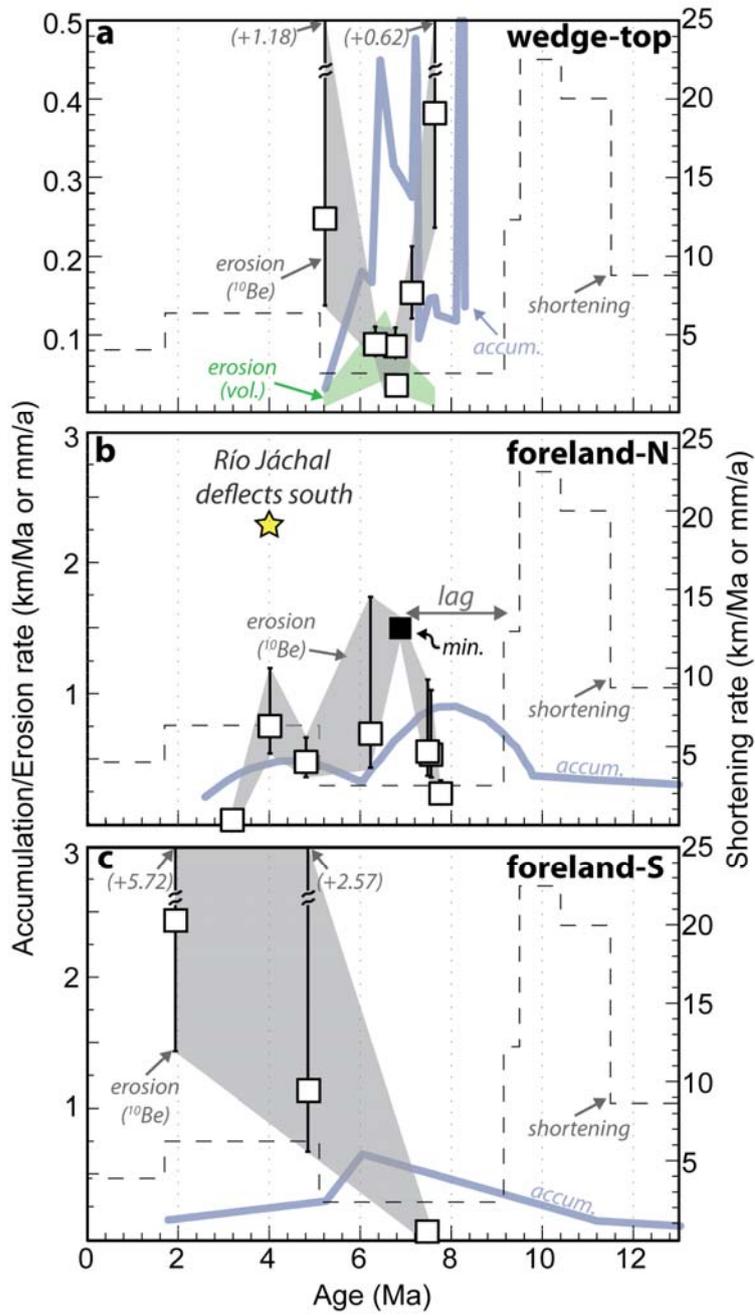
Figure 5 (double column)



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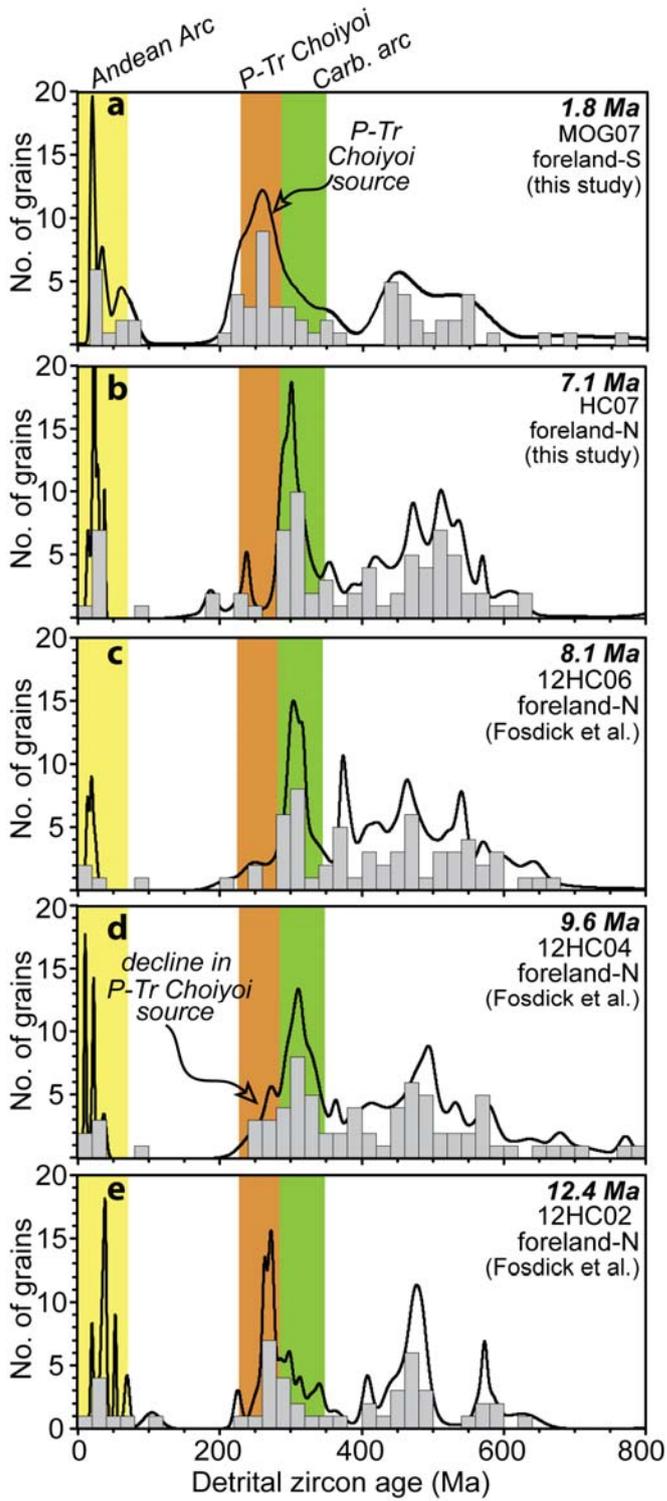
Figure 6 (single column)



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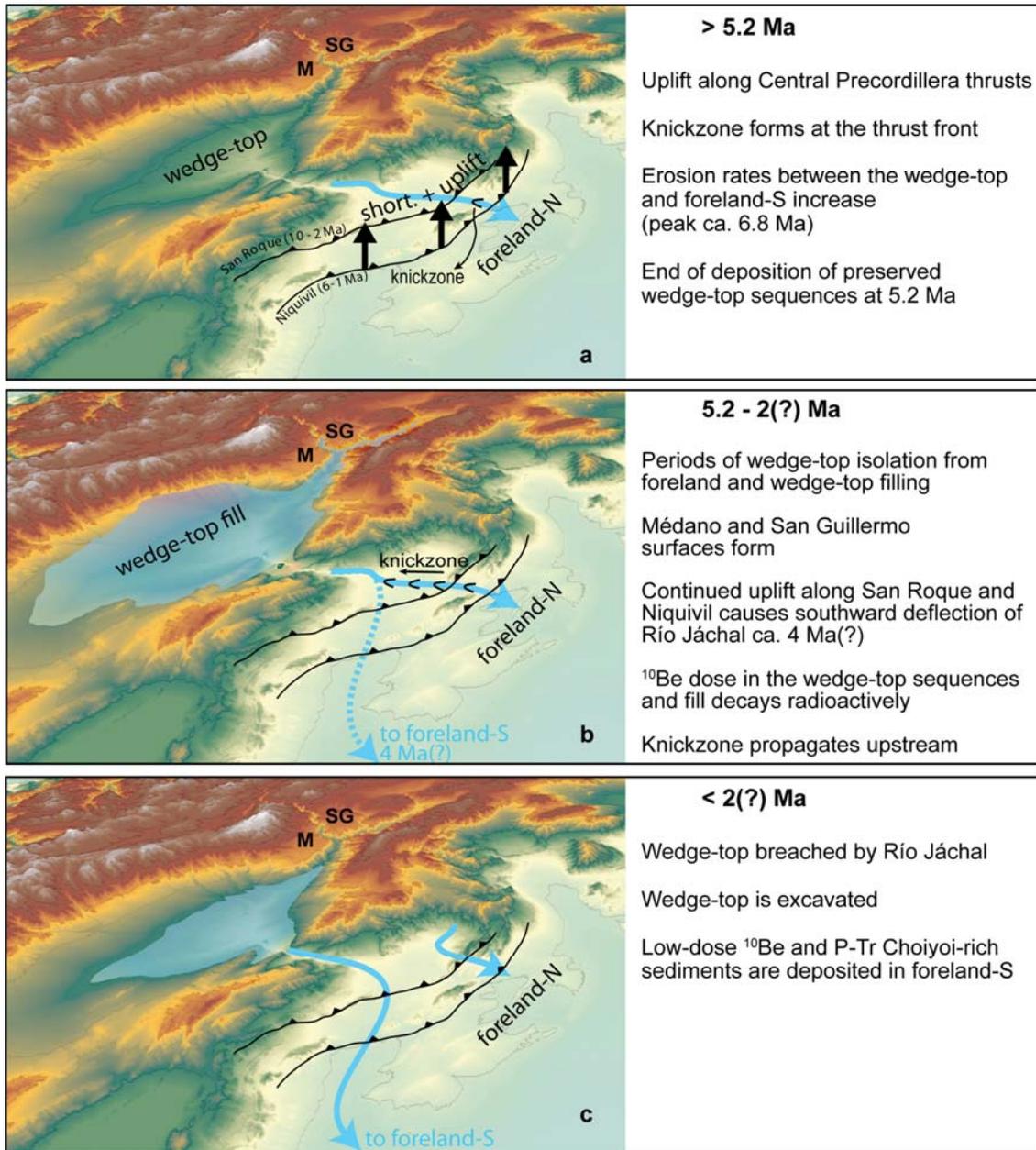
Figure 7 (single column)



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Figure 8 (double column)



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920 **Reconciling tectonic shortening, sedimentation and spatial patterns of erosion**  
921 **from  $^{10}\text{Be}$  paleo-erosion rates in the Argentine Precordillera**

922 Pedro Val, Gregory D. Hoke, Julie C. Fosdick, Hella Wittmann

923 **Supplementary Material**

924

925 **Detrital zircon U-Pb geochronology**

926 Detrital zircons were extracted from ~5 kg medium-grained sandstone hand-  
927 samples using standard mineral separation techniques, including crushing and  
928 grinding, fractionation of magnetic minerals with a Frantz isodynamic magnetic  
929 separator, and settling through heavy liquids to exclude phases with densities less  
930 than 3.3 g/cm<sup>3</sup>. Final zircon separates were mounted in epoxy resin and polished to  
931 expose interiors of grains. 127 grains from MOG-07 were selected at random for U-  
932 Pb isotopic analysis through Laser-Ablation Multicollector Inductively Coupled  
933 Plasma Mass Spectrometry (LA-ICPMS) at Boise State University relative to  
934 standards Plešovice zircon (Sláma et al., 2008). For sample HC-07, 110 grains were  
935 selected randomly for U-Pb isotopic analysis through LA-ICPMS at the University of  
936 Arizona relative to the Sri Lanka zircon standard following standard methods  
937 (Gehrels et al., 2006). Calculated U-Pb ages use the  $^{206}\text{Pb}/^{238}\text{U}$  ratio for >900 Ma  
938 grains and the  $^{206}\text{Pb}/^{207}\text{Pb}$  ratio for <1.0 Ga grains. Detrital zircon U-Pb analytical  
939 data for samples MOG-07 and HC-07 are summarized in Table A.3 and A.4,  
940 respectively.

941 At BSU, zircons were analyzed using a ThermoElectron X-Series II  
942 quadrupole ICPMS and New Wave Research UP0213 Nd:YAG UV (213 nm) laser  
943 ablation system (Rivera et al., 2013). For U-Th-Pb age analysis, instrumental  
944 fractionation of the background-subtracted  $^{206}\text{Pb}/^{238}\text{U}$ ,  $^{207}\text{Pb}/^{206}\text{Pb}$ , and  $^{208}\text{Pb}/^{232}\text{Th}$   
945 ratios is corrected, and ages calibrated with respect to interspersed measurements  
946 of the Plešovice zircon standard (Sláma et al., 2008). Signals at mass 204 are  
947 indistinguishable from zero following subtraction of mercury backgrounds  
948 measured during the gas blank (< 1000 cps  $^{202}\text{Hg}$ ), thus ages are reported without

949 common Pb correction; however, age error propagation includes an uncertainty  
950 contribution due to common Pb using the absolute value of the measured  
951  $^{206}\text{Pb}/^{204}\text{Pb}$  ratio. Secondary standards measured as unknown yielded ages within  
952 error of accepted values (Arizona Sri Lanka—563.5 Ma; R33—419.2 Ma; Zirconia—  
953 327.2 Ma; Fish Canyon Tuff—28.4 Ma).

954 At the University of Arizona, U-Pb detrital zircon geochronology was  
955 conducted through LA-ICPMS following the methods of Gehrels et al. (2006).  
956 Detrital zircons were randomly analyzed from a linear swath of grains across the  
957 sample mount to minimize sampling bias in characterizing all detrital populations.  
958 Zircons were ablated using a New Wave DUV193 Excimer laser (operating at a  
959 wavelength of 193 nm) using a spot diameter of 50  $\mu\text{m}$ . Each analysis lasted for 20 s,  
960 during which a pit  $\sim 20$   $\mu\text{m}$  in depth is excavated. U and Pb isotopes from the  
961 ablated material was measured simultaneously using a Micromass Isoprobe in static  
962 mode, using Faraday detectors for  $^{238}\text{U}$ ,  $^{232}\text{Th}$ , and  $^{208}\text{Pb}$ – $^{206}\text{Pb}$  and an ion-counting  
963 channel for  $^{204}\text{Pb}$ . Common Pb corrections are made by using the measured  $^{204}\text{Pb}$   
964 and assuming initial Pb composition from (Stacey and Kramers, 1975). In-run  
965 analysis of zircon grains of known isotopic and U-Pb composition (every fifth to  
966 sixth measurement) is used to correct for this fractionation. Concordia diagrams and  
967 relative-age-probability diagrams were constructed using the software of Ludwig  
968 (2008) (figures A.4, A.6).

969

#### 970 **Effects of ignoring time-dependent production rates on inferred paleo-erosion** 971 **rates**

972 Using CRONUScalc (Marrero et al., 2016), we compute a temporal signal of  $^{10}\text{Be}$  for  
973 two different paleo-erosion rate histories (0.2 mm/a and 0.3 mm/a) and a constant  
974 burial at 0.5 mm/a using time-dependent production rates. We then infer those  
975 paleo-erosion rates while assuming a constant production rate to assess the effects  
976 of potential variations in production rates past 2 Ma (Fig. A.7a). The results show  
977 that, averaged over  $>50$  ka, time-dependent production rates can generate variable  
978 structure in the paleo-erosion rates if assumed time-independent, but these

979 variations are generally  $\leq 20\%$ , which is smaller than the uncertainties of the paleo-  
980 erosion rates inferred in this study (Table 1).

981

982

983

984 **Effects of assuming constant catchment elevation on inferred paleo-erosion**  
985 **rates**

986 We generated a  $^{10}\text{Be}$  signal for a source area eroding at  $0.2 \text{ mm/a}$  and uplifting from  
987  $1800 \text{ m ca. } 8 \text{ Ma}$  to  $2350 \text{ m ca. } 3 \text{ Ma}$  using Equation 3 in the main text. We also used  
988 an arbitrary burial rate of  $0.5 \text{ mm/a}$  at  $\sim 800 \text{ m}$ , which is similar the accumulation  
989 rates observed in foreland-N. The signals are then inverted back to paleo-erosion  
990 rates through the same steps described in Section 3.1.4 in the main text while  
991 assuming the modern elevation ( $2350 \text{ m}$ ) remained unchanged through time. Two  
992 important predictions arise if this assumption is violated: 1) paleo-erosion rates  
993 deviate from the true (input) paleo-erosion rate in a systematic manner similar to  
994 the variation in source-area elevation; 2) for a  $20\%$  change in mean elevation,  
995 inferred paleo-erosion rates only marginally deviate from true (input) paleo-erosion  
996 rates (Fig. A.7b).

997

998 **Further evidence of short outcrop exposure times in the wedge-top**

999 The crucial assumption that our samples have only been marginally affected by  
1000 recent exposure of the stratigraphic sections combined with  $^{10}\text{Be}$  concentrations on  
1001 the order of  $10^3 \text{ at.g}^{-1}$  warrants discussion. While we are able to constrain what  
1002 outcrop exposure may look like based on  $^{26}\text{Al}$  and  $^{10}\text{Be}$  data from one replicate  
1003 sample from each section, we provide further evidence in support of minimal recent  
1004 exposure.

1005

1006 Similar to the foreland area, we expect recent exposure to be short in the wedge-top  
1007 due to ongoing excavation of the valley. For example, there is  $\sim 1 \text{ km}$  of fluvial  
1008 incision now separating the perched Médano and San Guillermo paleo-surfaces (Fig.  
1009 2b in main text). Moreover, parts of the stratigraphic sections were uplifted and

1010 exposed along blind-thrusts of the El Tigre fault system, which was active ca.  $\leq 300$   
1011 ka (Siame et al., 1997), making them particularly vulnerable to erosion. Lastly,  
1012 outcrops in the wedge-top usually contain rock-fall, indicating that they are  
1013 unstable. Taken together with the  $\sim 1,700$  years of outcrop exposure, these reasons  
1014 point to low recent exposure in the wedge-top.

1015

#### 1016 **Using $^{26}\text{Al}/^{10}\text{Be}$ to infer $t_{\text{outcrop}}$ and $E_{\text{exhum}}$**

1017 In this section, we describe how we use the replicate samples to infer  $t_{\text{outcrop}}$ . Of the  
1018 three replicate samples, the youngest is 4.8 Ma and the oldest, 7.7 Ma. Thus, we  
1019 expect low  $^{26}\text{Al}/^{10}\text{Be}$  ratios for all three. Sample HU-16 (4.8 Ma) is well shielded in  
1020 outcrop and yielded a ratio of  $2.1 \pm 0.5$  ( $1\sigma$ ), which is higher than the expected ratio  
1021 of  $0.8 \pm 0.2$  ( $1\sigma$ ) we infer for a sample of this age that was gradually buried at 0.47  
1022 mm/a. This observation indicates that this sample contains  $\sim 1,100$  at.g $^{-1}$  and  
1023  $\sim 7,430$  at.g $^{-1}$  of modern  $^{10}\text{Be}$  and  $^{26}\text{Al}$ , respectively. The production of modern  $^{10}\text{Be}$   
1024 due to exhumation accounts for  $\sim 600$  at.g $^{-1}$ . Since this sample is 90% shielded in  
1025 outcrop and the outcrop production rate is 5.2 at.g $^{-1}$ .yr $^{-1}$ , it takes  $\sim 1,000$  years to  
1026 produce 500 at.g $^{-1}$ . We use this exposure time as a possible maximum age for nearby  
1027 outcrops in the foreland-N section and conservatively apply this correction to  
1028 samples HU-05, HU-16, and HU-17.

1029

1030 Sample MOG-05 (4.8 Ma) is relatively well shielded in outcrop and yielded a  
1031  $^{26}\text{Al}/^{10}\text{Be}$  ratio of  $2.5 \pm 0.6$  ( $1\sigma$ ), which is also higher than the expected ratio of  $0.8 \pm$   
1032  $0.2$  ( $1\sigma$ ). The excess  $^{10}\text{Be}$  and  $^{26}\text{Al}$  concentrations responsible for this deviation is  
1033 1,600 at.g $^{-1}$  and 11,070 at.g $^{-1}$ , respectively. The inferred  $C_{\text{exhumation}}$  for the 1.33 mm/a  
1034 exhumation rate assumed for the Mogna anticline was  $2,550 \pm 690$  at.g $^{-1}$ , which is  
1035 higher than the inferred modern. Thus, to reach 1,600 at.g $^{-1}$  the local exhumation  
1036 rate ( $E_{\text{exhum}}$ ) must be 0.8 mm/a, not 0.5 mm/a. Based on our assumption of the  
1037 geometry of exhumation for the anticline cores, the exhumation rate at the core of  
1038 the Mogna anticline must be 2 mm/a, not 1.33 mm/a as inferred initially. The re-  
1039 scaled exhumation rates for all samples were used to calculate  $C_{\text{exhumation}}$  in this

1040 section and are shown in Table 1 in the main text. We do not attempt to determine  
1041  $t_{outcrop}$  here since it would imply even higher exhumation rates and at this point our  
1042 inferences would be somewhat arbitrary. Thus,  $C_{exposure}$  is not calculated for the  
1043 Mogna section.

1044

1045 Sample R1-01 (6.7 Ma) yielded a  $^{26}\text{Al}/^{10}\text{Be}$  ratio of  $2.59 \pm 0.33$  ( $1\sigma$ ), which is much  
1046 higher than an expected ratio of  $0.34 \pm 0.12$  ( $1\sigma$ ). This implies some modern  
1047 exposure of the sample, which was collected from an outcrop wall with 50%  
1048 shielding. To lower the  $^{26}\text{Al}/^{10}\text{Be}$  from  $\sim 2.6$  to 0.34, one needs to subtract  $\sim 9,000$   
1049  $\text{at.g}^{-1}$  of  $^{10}\text{Be}$  and  $\sim 60,750$   $\text{at.g}^{-1}$  of  $^{26}\text{Al}$  from the measured concentrations after the  
1050 burial correction, assuming a modern ratio of 6.75. The  $^{10}\text{Be}$  in  $C_{exhumation}$  accounts  
1051 for  $\sim 1,400 \pm 445$   $\text{at.g}^{-1}$  of  $^{10}\text{Be}$ , which leaves  $\sim 7,750$   $\text{at.g}^{-1}$  unaccounted for. With  
1052 50% outcrop shielding and an outcrop production rate of  $8.75$   $\text{at.g}^{-1}\text{.yr}^{-1}$ , this excess  
1053 dose requires  $\sim 1,770$  years of exposure in outcrop. We also take this exposure time  
1054 as a maximum for the wedge-top area and incorporate the correction when  
1055 calculating  $C_{exposure}$  for all wedge-top samples.

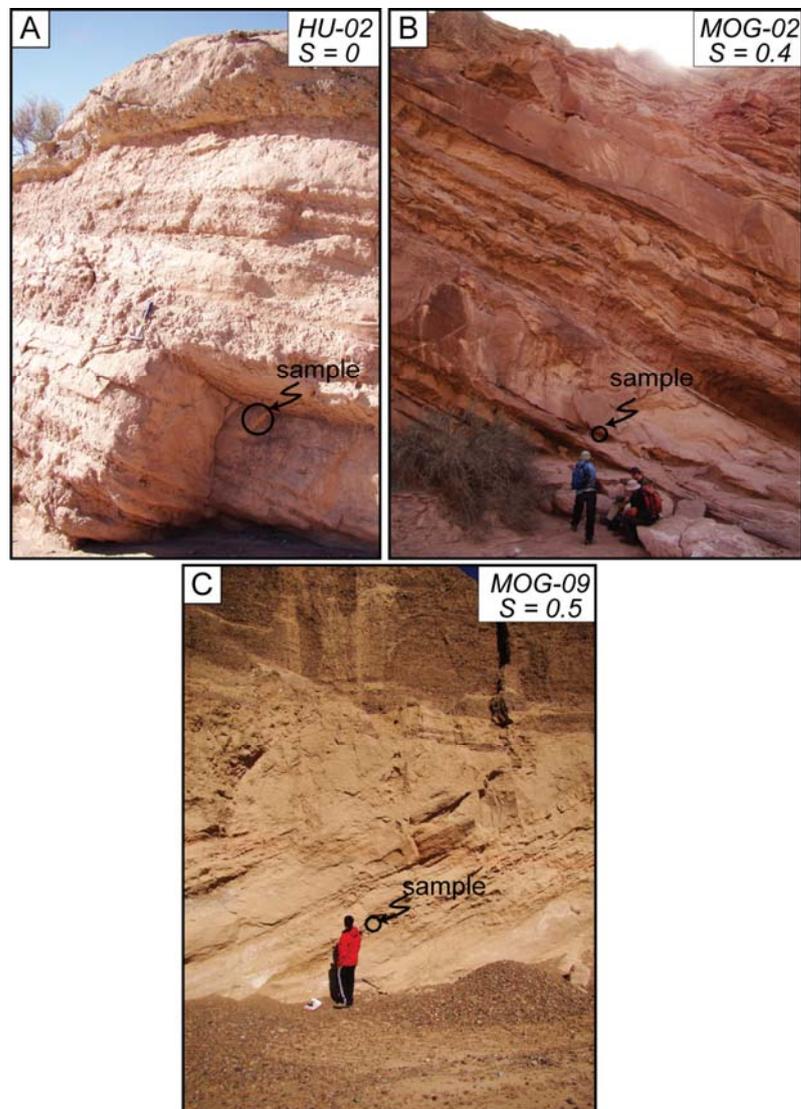
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1080 821X(75)90088-6
- 1081 **FIGURES**
- 1082 **A.1** Photographs of sampled outcrops showing outcrop geometry.
- 1083 **A.2** Outline of low-quartz lithologies
- 1084 **A.3** 0-2.5 Ga results for MOG-07
- 1085 **A.4** Concordia diagram from sample MOG-07
- 1086 **A.5** 0-2.5 Ga results for HC-07
- 1087 **A.6** Concordia diagram from sample HC-07
- 1088 **A.7** Testing the sensitivity of the assumption constant production and elevation  
1089 through time.
- 1090 **A.8** Relief of the foreland-N area.
- 1091 **A.9** Paleo-erosion rate results for hypothetical hiatus in the wedge-top.
- 1092 **A.10:** Histogram of scaling factors obtained pixel by pixel using a 90 m resolution  
1093 SRTM DEM. Note that the majority of scaling factors is higher than 98%.
- 1094 **TABLES**

- 1095 **A.1**  $^{10}\text{Be}$  AMS results
- 1096 **A.2**  $^{26}\text{Al}$  AMS and ICP-OES results
- 1097 **A.3** Detrital zircon U-Pb analytical data and geochronology, sample MOG-07
- 1098 **A.4** Detrital zircon U-Pb analytical data and geochronology, sample HC-07
- 1099 **A.5** Detrital zircon U-Pb analytical data and geochronology, sample MD-01
- 1100 **A.6** Table for input in CRONUS online calculator.  $^{10}\text{Be}$  concentrations are the estimated
- 1101 concentrations relative to paleo-erosion rates ( $C_{\text{hillslope}}$ ).

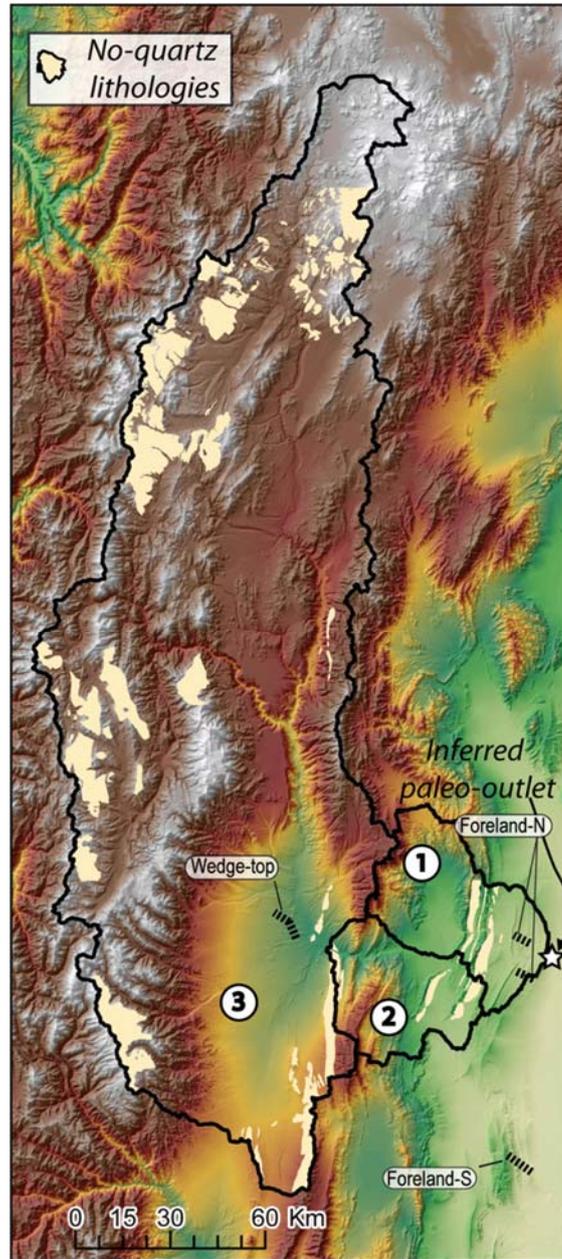


1102

1103 **Fig. A.1** Photographs of sampled outcrops showing shielding geometries. Sample names and  
1104 adopted  $S$  values for equation S2 is shown in the upper right corner. A) Ideal case in which  
1105 the sample is completely shielded by the outcrop geometry (see hammer for scale). B)  
1106 Example in a 30 m deep canyon. The vertical wall provides half the shielding, and the  
1107 surrounding topography constitutes about an overall  $40^\circ$  angle to the horizon, providing  
1108 another 10% of shielding. C) Example of an exposed wall without shielding by surrounding  
1109 topography.  
1110

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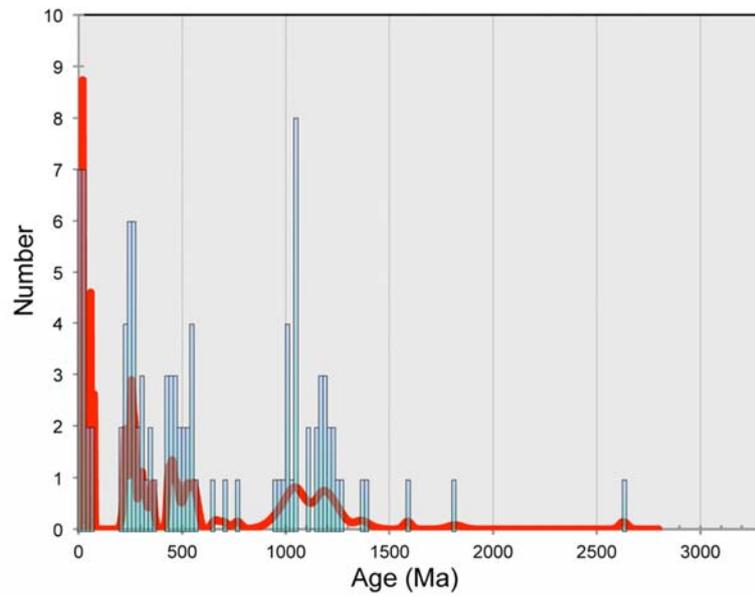
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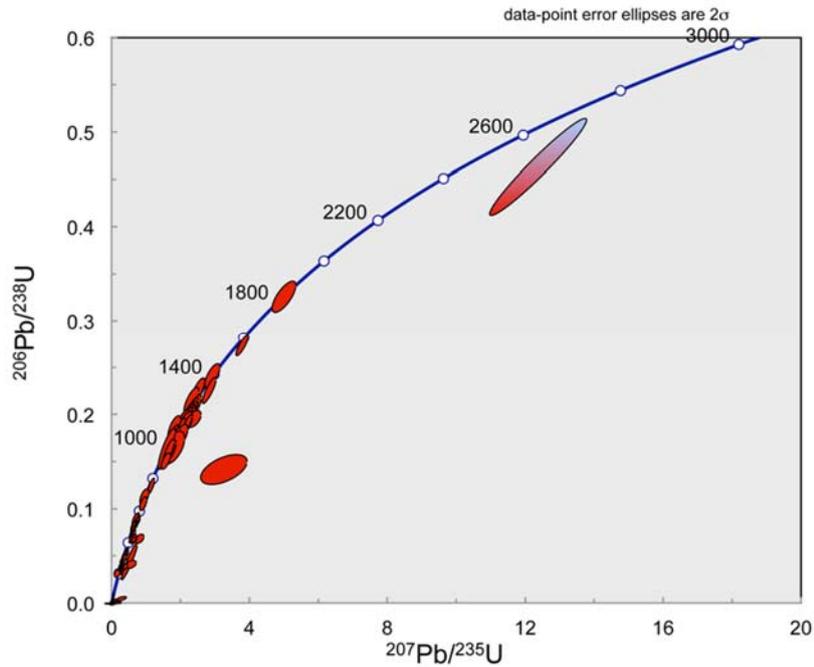
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1114 **Fig. A.2** Idealized Río Jáchal catchment with outlet near Huaco. Catchment #3 was used in  
1115 production rate calculations for wedge-top samples. Catchment #2 is the downstream  
1116 extension of catchment #3, i.e. the Río Jáchal catchment. Catchment #1 is the Huaco

1117 catchment. The entire area outlined in black yields the catchment wide scaling factors and  
1118 production rates depicted in figure S2.  
1119  
1120  
1121



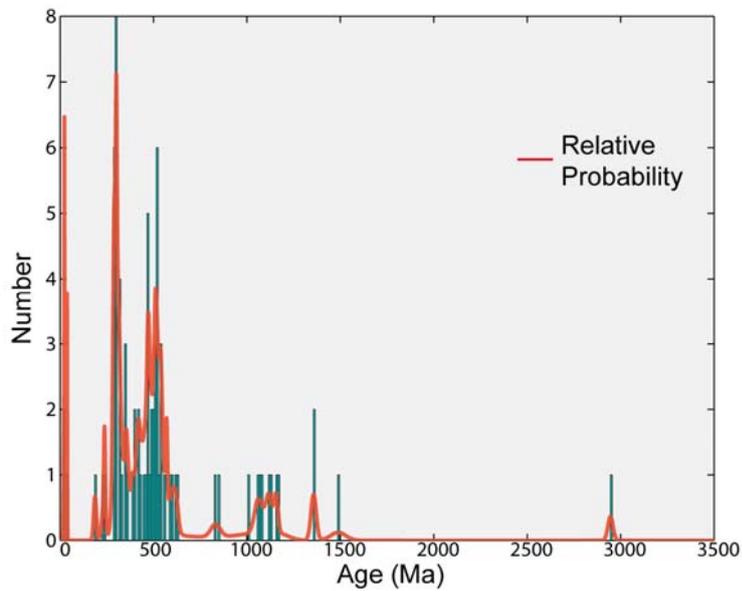
1122  
1123 **Fig. A.3** 0-2.5 Ga results for MOG-07 (1.8 Ma) detrital zircon sample from the Mogna Fm.,  
1124 Foreland-S section.



1125

1126 **Fig. A.4** Concordia diagram showing  $2\sigma$  ellipses of measured  $^{206}\text{Pb}/^{238}\text{U}$  and  $^{207}\text{Pb}/^{235}\text{U}$  ratios

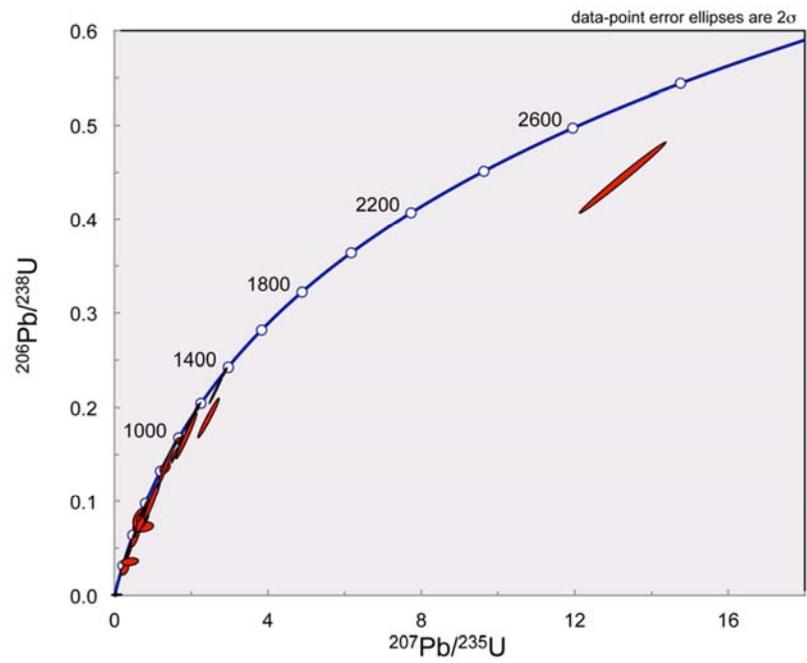
1127 measured in detrital zircons from sample MOG-07.



1128

1129 **Fig. A.5** 0-2.5 Ga results for HC-07 (7.1 Ma) detrital zircon sample from the Río Jáchal Fm.,

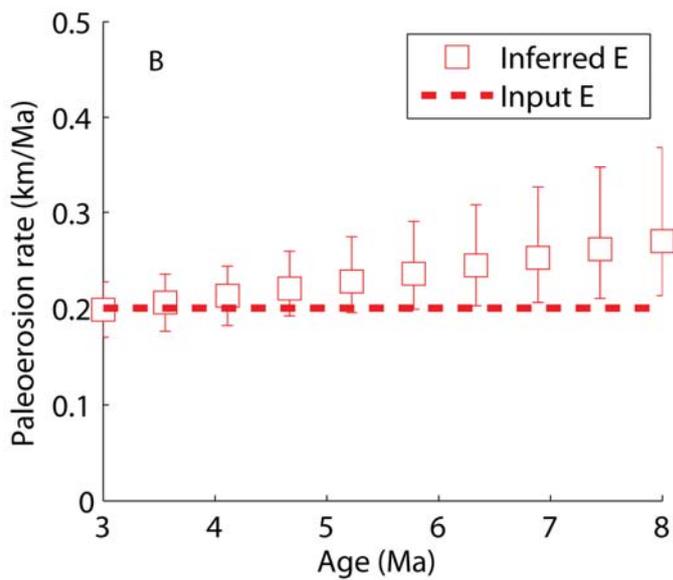
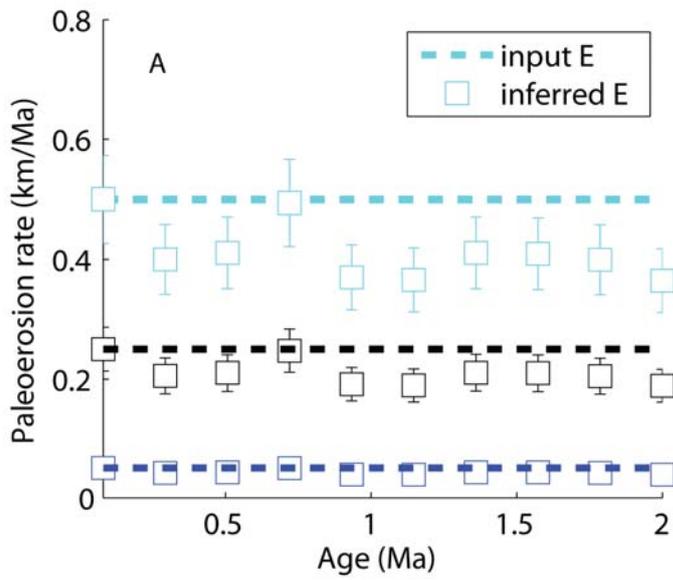
1130 Foreland-N section.



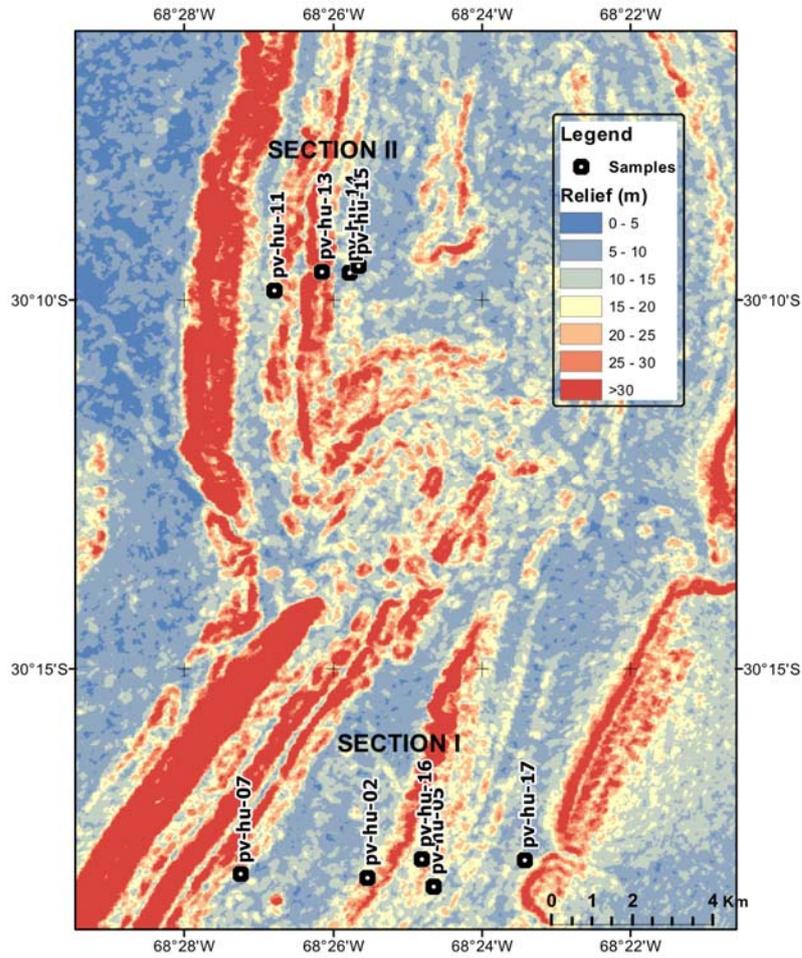
1131

1132 **Fig. A.6** Concordia diagram showing  $2\sigma$  ellipses of measured  $^{206}\text{Pb}/^{238}\text{U}$  and  $^{207}\text{Pb}/^{238}\text{U}$  ratios

1133 measured in detrital zircons from sample HC-07



1134  
 1135 **Fig. A.7** (A) Effects of ignoring changes in production rates due to secular variations in  
 1136 geomagnetic paleo-intensities for three different input paleo-erosion rates; (B) Effects of using  
 1137 the modern source-area elevation to calculate  $^{10}\text{Be}$  paleo-erosion rates. Input E denotes the input  
 1138 paleo-erosion rate and inferred E denotes the inverted paleo-erosion rate signal.



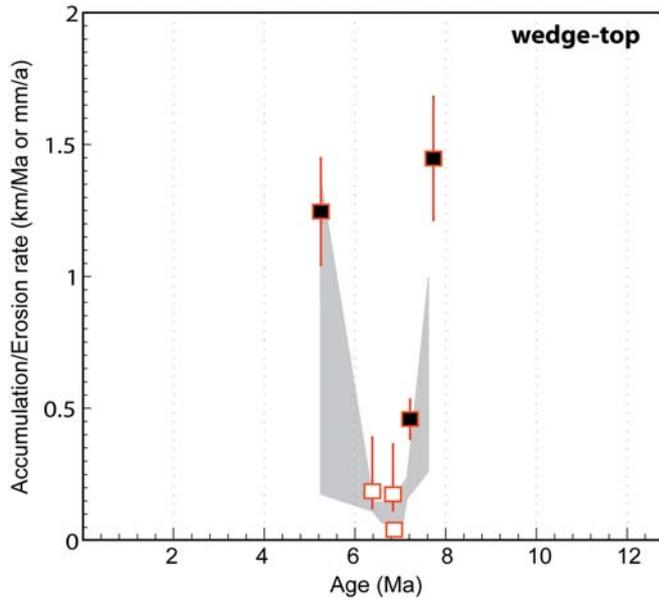
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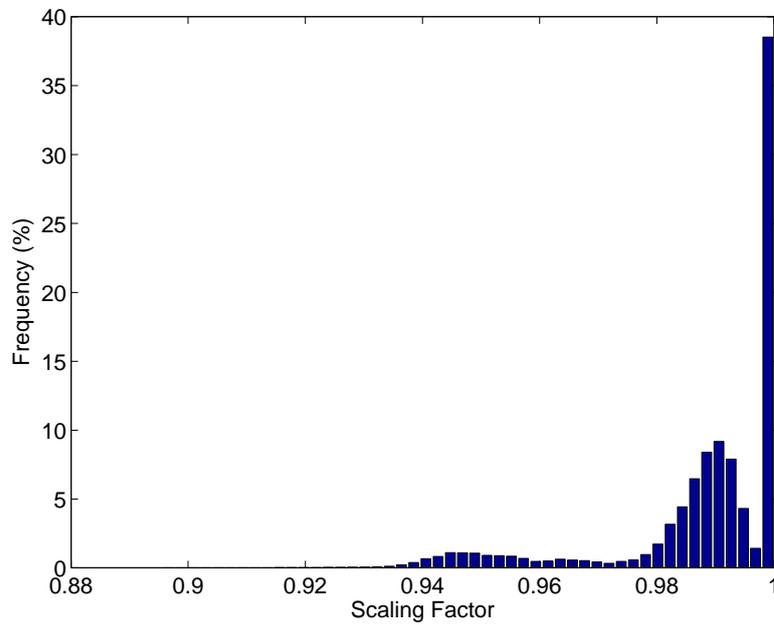
**Fig. A.8:** Relief map of the Foreland-N study area showing locations of samples. The relief was calculated within a 120 m radius using a 30 m resolution digital elevation model (1 arc-second SRTM Global Dataset - <http://earthexplorer.usgs.gov>).



1143

1144 **Fig. A.9:** Paleo-erosion rate results for wedge-top samples if each sample remained 1 m  
 1145 beneath the surface of the deposit 100 ka prior to burial. Black-filled squares denote  
 1146 minimum possible paleo-erosion rate. Gray envelope shows no-hiatus results from the main  
 1147 text.

1148



1149

1150 **Fig. A.10:** Histogram of scaling factors obtained pixel by pixel using a 90 m resolution SRTM  
 1151 DEM. Note that the majority of scaling factors is higher than 98%.