



Originally published as:

Uzbekov, B., Shprits, Y., Orlova, K. (2016): Scattering of Relativistic and Ultra-relativistic Electrons by Obliquely Propagating Electromagnetic Ion Cyclotron Waves. - *Journal of Atmospheric and Solar-Terrestrial Physics*, 148, pp. 22–31.  
DOI: <http://doi.org/10.1016/j.jastp.2016.08.004>

# 1 Scattering of Relativistic and Ultra-relativistic Electrons by Obliquely Propagating 2 Electromagnetic Ion Cyclotron Waves

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10

## 11 **Key points**

- 12 1. The method presented can be applied to EMIC waves or any other wave modes;
- 13 2. First order resonance dominates the scattering of relativistic electrons up to multi-MeV  
14 energies;
- 15 3. Ion composition is a crucial parameter for quantifying scattering rates.

16

17 Electromagnetic Ion Cyclotron (EMIC) waves are transverse plasma waves that are generated in  
18 the Earth magnetosphere by ring current protons with temperature anisotropy in three different  
19 bands: below the  $H^+$ ,  $He^+$  and  $O^+$  ion gyrofrequencies. EMIC events are enhanced during the  
20 main phase of a geomagnetic storm when intensifications in the electric field result in enhanced  
21 injections of ions and are usually confined to high-density regions just inside the plasmopause or  
22 within drainage plumes. EMIC waves are capable of scattering radiation belt electrons and thus  
23 provide an important link between the intensification of the electric field, ion populations, and  
24 radiation belt electrons. Bounce-averaged diffusion coefficients computed with the assumption  
25 of parallel wave propagation are compared to the results of the code that uses the full cold  
26 plasma dispersion relation taking into account oblique propagation of waves and higher-order  
27 resonances. We study the sensitivity of the scattering rates to a number of included higher-order  
28 resonances, wave spectral distribution parameters, wave normal angle distribution parameters,  
29 ambient plasma density, and ion composition. Inaccuracies associated with the neglect of higher-  
30 order resonances and oblique propagation of waves are compared to potential errors introduced  
31 by uncertainties in the model input parameters.

32

## 33 1. Introduction

### 34 1.1. Scattering by EMIC waves

35 Wave-particle interaction is an important aspect of radiation belt physics, and is responsible for  
36 the variability of particle fluxes observed. One type of plasma wave that is supported by the  
37 Earth's magnetosphere, and is able to effectively scatter radiation belt particles is the  
38 electromagnetic ion cyclotron (EMIC) wave [*Thorne and Kennel, 1971; Lyons and Thorne,*  
39 *1972*]. Recent studies have suggested that EMIC waves are most efficient at scattering ultra-  
40 relativistic electrons [e.g., *Usanova et al., 2014; Shprits et al., 2014; Drozdov et al., 2015*].  
41 EMIC waves are able to cause relativistic and ultra-relativistic electron loss in a matter of hours  
42 [e.g. *Ni et al., 2015; He et al., 2016; Summers et al., 2007b; Usanova et al., 2014; Su et al.,*  
43 *2011*], and thus quantification of the effect of EMIC waves on radiation belt electrons is crucial  
44 for predicting the evolution of the radiation belt electron fluxes.

45 While the calculations of diffusion coefficients due to very low frequency (VLF) whistlers are  
46 done for an oblique distribution of waves [e.g., *Glauert and Horne, 2005; Albert, 2005; Shprits*  
47 *and Ni, 2009; Ni et al., 2008; Gu et al., 2012; Orlova and Shprits, 2014; Orlova et al., 2014*],  
48 calculations of the scattering rates for EMIC waves usually assume parallel propagating waves  
49 [e.g., *Summers and Thorne, 2003; Summers et al., 2007a; Ukhorskiy et al., 2010; Usanova et al.,*  
50 *2014; Shprits et al., 2014*]. The likely reason is a cumbersome plasma dispersion relation and  
51 related computational difficulties required to quantify effects of oblique EMIC waves in the  
52 presence of heavy ions on the radiation belt particles. In this study, we use a semi-analytical  
53 approach described in the text below. It allows us to obtain accurate and reliable solutions for  
54 diffusion coefficients, which is challenging to obtain using a purely analytical or purely  
55 numerical approach. The method and the code developed can be used to study other plasma  
56 wave modes obliquely propagating in the presence of heavy ions and their effects on radiation  
57 belt particles with high accuracy, opening opportunities for future research. The presented  
58 method guarantees that all intersections of dispersion relation and cyclotron resonance condition  
59 curves are obtained.

60 *Albert* [2003] considered all three bands of oblique EMIC waves and investigated the sensitivity  
61 of the diffusion coefficients for EMIC waves to wave normal angle distribution and resonance  
62 numbers. He also proposed a computationally efficient technique to avoid integration over  
63 ranges of latitudes, wave normal angles and resonant numbers that are not consistent with the  
64 resonance in a prescribed wave population. Yet *Albert* [2003] made slight approximations to the  
65 exact cold plasma theory, namely a high-density approximation, which is reasonable in the case  
66 of EMIC waves, and allows for an analytical solution as a result of simpler equations. *Glauert*

67 *and Horne* [2005] have taken the results of the PADIE code calculated numerically the diffusion  
68 coefficients for the case of oblique EMIC waves, and compared them to the results of [*Summers*  
69 *and Thorne*, 2003], where the assumption of parallel wave propagation is made, and outlined the  
70 importance of EMIC wave obliquity. The study conducted by *Glauert and Horne* [2005],  
71 however, fixes ambient plasma density, ion composition and parameters of the wave spectral  
72 distribution. *Ni et al.* [2015] provided a useful general 14th-order polynomial equation of  
73 resonant frequencies concerning oblique EMIC wave interactions with charged particles in cold,  
74 magnetized, multi-ion plasma. The study considered multi-band EMIC waves propagating at  
75 different L-shells and having various wave normal angles. Present study fixes L-shell at  $L = 4$   
76 and only considers He band waves, but complements work of *Ni et al.* [2015] by conducting  
77 extensive sensitivity analysis of the diffusion coefficients for oblique EMIC to other model input  
78 parameters, including number of cyclotron resonances, wave spectral distribution, ambient  
79 plasma density and ion composition.

80 While *Albert* [2003], *Glauert and Horne* [2005] and *Summers et al.* [2007b] considered an  
81 arbitrary ratio of plasma to gyrofrequency, in the current study we use statistical models  
82 [*Carpenter and Anderson*, 1992] to calculate the ratio of plasma frequency to gyrofrequency.

83 In this study we focus on the He band left-hand polarized EMIC waves (L-mode), due to its  
84 typically higher amplitude and subsequent comparison to H and O bands [*Saikin et al.*, 2015]. At  
85 the same time, right-hand polarized (R-mode) EMIC waves are rarely observed experimentally  
86 [e.g. *Meredith et al.*, 2003, *Allen et al.*, 2015]. In addition, the R-mode waves are not as efficient  
87 at scattering ultra-relativistic electrons as the L-mode waves.

## 88 **1.2. Code Description**

89 We have performed calculations of the bounce-averaged diffusion coefficients for oblique EMIC  
90 waves using Full Diffusion Code (FDC) [e.g. *Ni et al.*, 2008; *Shprits and Ni*, 2009; *Gu et al.*,  
91 2012], following the theoretical approach presented in *Glauert and Horne* [2005]. The details of  
92 the calculations are presented in the supporting information. The code is capable of computing  
93 resonant scattering rates, including first-order, Landau, and higher-order resonances and uses the  
94 methodology of *Orlova and Shprits* [2011] to integrate bounce period and momentum diffusion  
95 coefficients, avoiding uncertainty at the mirror point.

96 Substituting resonance conditions for wave-particle interactions into the cold plasma dispersion  
97 relation yields a fourteenth order polynomial with respect to the wave frequency  $\omega$ . For the  
98 version of the code developed for this study, the coefficients are derived semi-analytically using  
99 the Matlab Symbolic Toolbox and then incorporated into the code explicitly. Coefficients of the

100 polynomial are not presented in this work, as they are very cumbersome. The roots of the  
 101 polynomial, satisfying both cold plasma dispersion relation and resonance conditions, are  
 102 computed using a standard polynomial root finder. Only roots that lie between the lower and  
 103 upper cut-off frequencies of the wave are considered. The roots are then checked on their R-  
 104 mode or L-mode origin, and filtered according to the mode under consideration (e.g. R-mode  
 105 roots are not considered, when studying the effects of the L-mode waves only). The dispersion  
 106 relation derivative,  $(\partial\omega/\partial k)|_{\mathcal{D}=0}$ , which is required to compute local diffusion coefficients, is  
 107 also calculated semi-analytically with the help of the Matlab Symbolic Toolbox. The resulting  
 108 derivative formula is very cumbersome and therefore not presented here.

109 From the mathematical standpoint, the dispersion relation derivative  $(\partial\omega/\partial k)|_{\mathcal{D}=0}$ , as well as  
 110 coefficients of the polynomial derived substituting resonance conditions into the dispersion  
 111 relation  $\mathcal{D}(k, \omega, X) = 0$ , are functions of local magnetic field value, ambient plasma density  $n_e$ ,  
 112 plasma ion composition (parameters  $\zeta_{H^+} = n_{H^+}/n_e$ ,  $\zeta_{He^+} = n_{He^+}/n_e$ , and  $\zeta_{O^+} = n_{O^+}/n_e$ ),  
 113 particle energy, local pitch angle  $\alpha$ , wave normal angle  $\theta$  and resonant number  $n$ . The choice of  
 114 these variables makes the code relatively easy to adapt to studies of other plasma wave modes  
 115 and their effects on the trapped radiation at various magnetospheres.

116 We assume distributions of wave normal angle  $g(\theta)$  and wave spectral density  $B^2(\omega)$  to be  
 117 Gaussian:

$$118 \quad B^2(\omega) = \begin{cases} A^2 \exp\left(-\left(\frac{\omega - \omega_m}{\delta\omega}\right)^2\right), & \omega_{lc} \leq \omega \leq \omega_{uc} \\ 0, & otherwise \end{cases}$$

119 where  $\omega_m$  and  $\delta\omega$  are distribution maximum and bandwidth respectively,  $\omega_{lc}$  and  $\omega_{uc}$  are  
 120 lower and upper cut-off frequencies, and  $A$  is a normalisation factor. And for the wave angular  
 121 spread:

$$122 \quad g(X) = \begin{cases} \exp\left(-\left(\frac{X - X_m}{X_w}\right)^2\right), & X_{min} \leq X \leq X_{max} \\ 0, & otherwise \end{cases}$$

123 where  $X = \tan(\theta)$ ,  $X_w$  is the angular distribution bandwidth,  $X_m$  is the peak value, and  $X_{min}$   
 124 and  $X_{max}$  are the cut-off values.

125 We set MLT presence of the waves to 100%. In order to obtain more realistic scattering rates,  
 126 our diffusion coefficients need to be scaled by the width of the MLT sector where waves are  
 127 present. The scaling is performed by simple multiplication of the coefficients by realistic MLT  
 128 presence. Maximum latitude of the wave propagation is set to  $20^\circ$  following statistical study of  
 129 *Meredith et al.* [2003].

130 We have validated the code by comparing our results with those from *Glauert and Horne* [2005].

## 131 **2. Sensitivity Simulations**

132 In this section, we first study contribution of the Landau, first-order and higher-order resonances  
133 to the bounce-averaged pitch angle diffusion coefficients for oblique EMIC waves. This analysis  
134 allows us to estimate the number of resonances that needs to be included into calculations to give  
135 sufficiently accurate results. We then move to comparison of the scattering rates calculated for  
136 EMIC waves propagating parallel to the magnetic field and for the case of oblique propagation  
137 of waves. Finally, we study sensitivity of the diffusion coefficients for oblique EMIC waves to  
138 the key model input parameters – wave spectral distribution, wave normal angle distribution,  
139 ambient plasma density, and ion composition.

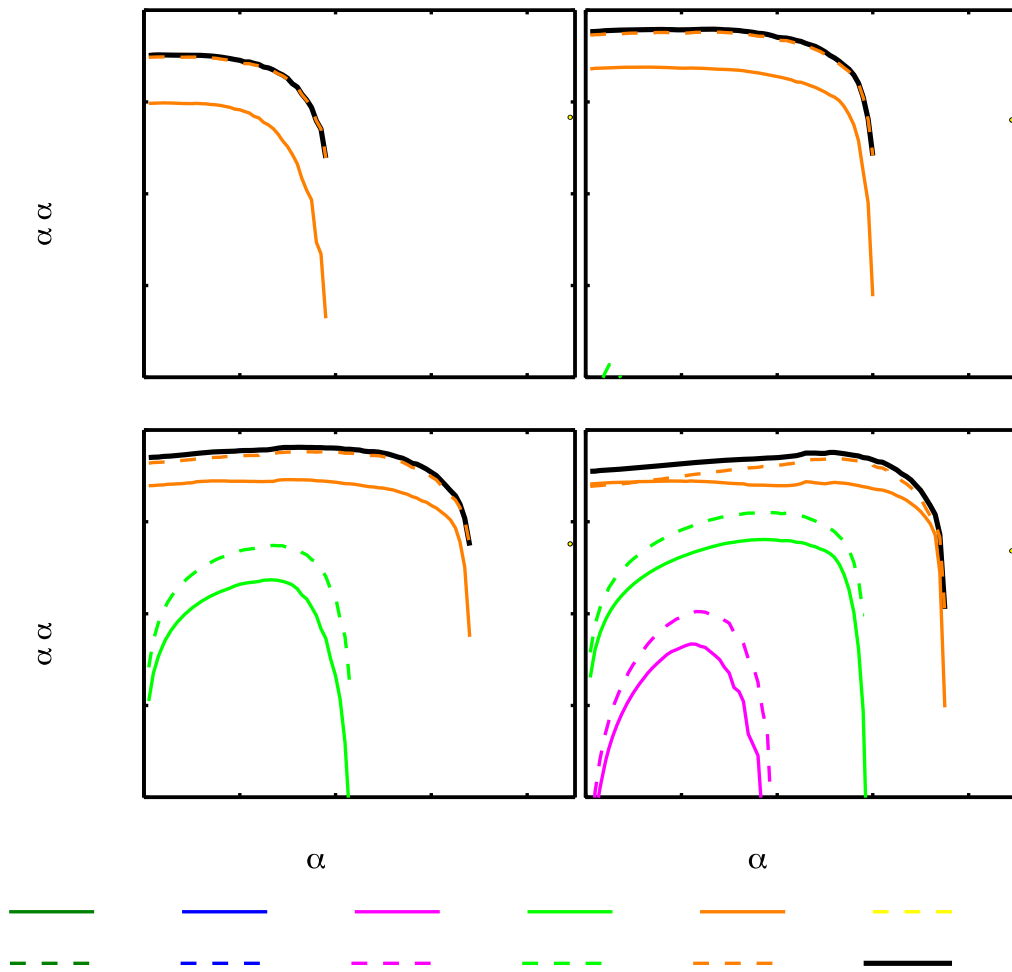
140 **2.1 Contribution of Various Resonances**

141 Figure 1 shows the contribution of various resonances to the bounce-averaged pitch angle  
 142 diffusion coefficients for electron energies of 3, 5, 7, and 10 MeV. The scattering rates were  
 143 computed at  $L = 4$  with the input parameters summarized in Table 1.

144 *Table 1. Model input parameters used to obtain the results shown in Figure 1 and Figure 2*

| <i>Parameters</i>              | <i>Values</i>   |
|--------------------------------|---|
| Wave normal angle distribution | $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 30^\circ, \delta\theta = 30^\circ$   |
| Wave spectral distribution     | $\omega_{lc} = 2.5\Omega_{O^+}, \omega_{uc} = 3.5\Omega_{O^+}, \omega_m = 3\Omega_{O^+}, \delta\omega = 0.5\Omega_{O^+}$<br>[Summers et al., 2007b] |
| Wave amplitude                 | $dB = 1 \text{ nT}$   |
| Plasma density model           | <i>Carpenter and Anderson</i> [1992]  |
| Ion composition                | $\zeta_{H^+} = 0.7n_e, \zeta_{He^+} = 0.2n_e, \zeta_{O^+} = 0.1n_e$ [Meredith et al.,<br>2003]  |
| Resonant numbers               | $n = -5, \dots, 0, \dots, 5$  |

145



146

147 *Figure 1. Contribution of various resonances into the bounce-averaged pitch angle diffusion coefficients for*  
 148 *oblique EMIC waves at  $L=4$  for energies of a) 3 MeV, b) 5 MeV, c) 7 MeV and d) 10 MeV, respectively.*

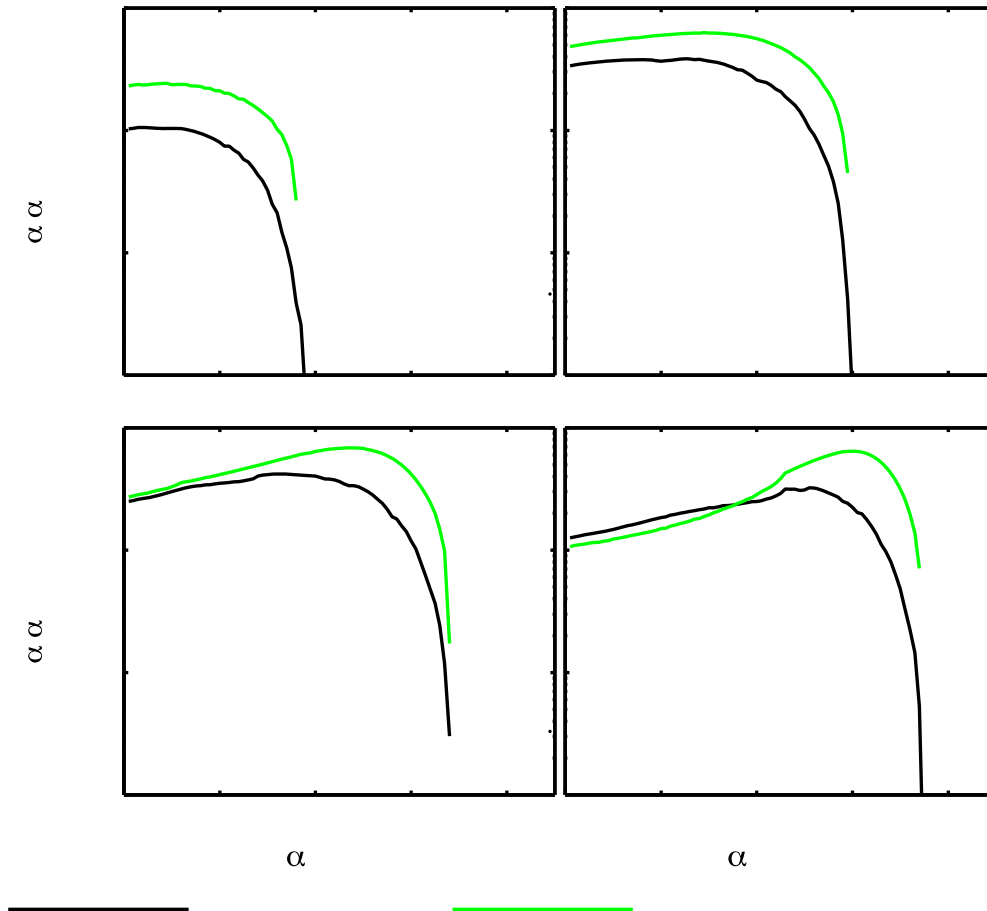
149 From Figure 1 we see that the principal resonance ( $n = 1$ ) can give a fairly good estimate of the  
 150 diffusion coefficients at energies of up to 7 MeV, while at  $E = 10$  MeV, the resonance  
 151 represented by  $n = -1$  number overtakes the principal resonance at pitch angles below  $20^\circ$ ,  
 152 which should be taken into account. It is interesting to note resonances appearing in pairs  
 153 corresponding to  $|n|$ , and how the range of electrons' pitch angles affected by diffusion is  
 154 decreasing with growing  $|n|$ .

155 Another interesting observation is the absence of  $n = 0$  (Landau) resonance in Figure 1. Zeroth  
 156 resonance has appeared, however, in the case of four-fold increase in the plasma density value  
 157 (not shown in this paper). Nevertheless, the contribution of Landau resonance was at least  
 158 several orders of magnitude smaller compared to contributions from cyclotron resonances in all  
 159 of our simulations. Finally we notice that more resonances contribute to the scattering rates with  
 160 energy growing.



## 161 2.2 Comparison with the Case of Parallel Wave Propagation

162 We continue by comparing the diffusion coefficients calculated for oblique EMIC waves and for  
 163 the case of field-aligned propagation. We illustrate the results of the simulations in Figure 2, and  
 164 present model input parameters that were used in Table 1. We have treated the case of field-  
 165 aligned waves using the theoretical approach from *Summers et al.* [2007a], taking into account  
 166 factor of two discrepancy reported by *Albert* [2007].



167

168 *Figure 2. Comparison of bounce-averaged pitch angle diffusion coefficients calculated for oblique EMIC waves*  
 169 *and for the case of field-aligned wave propagation at  $L=4$ . Panels correspond to energies of a) 3 MeV, b) 5 MeV,*  
 170 *c) 7 MeV and d) 10 MeV.*

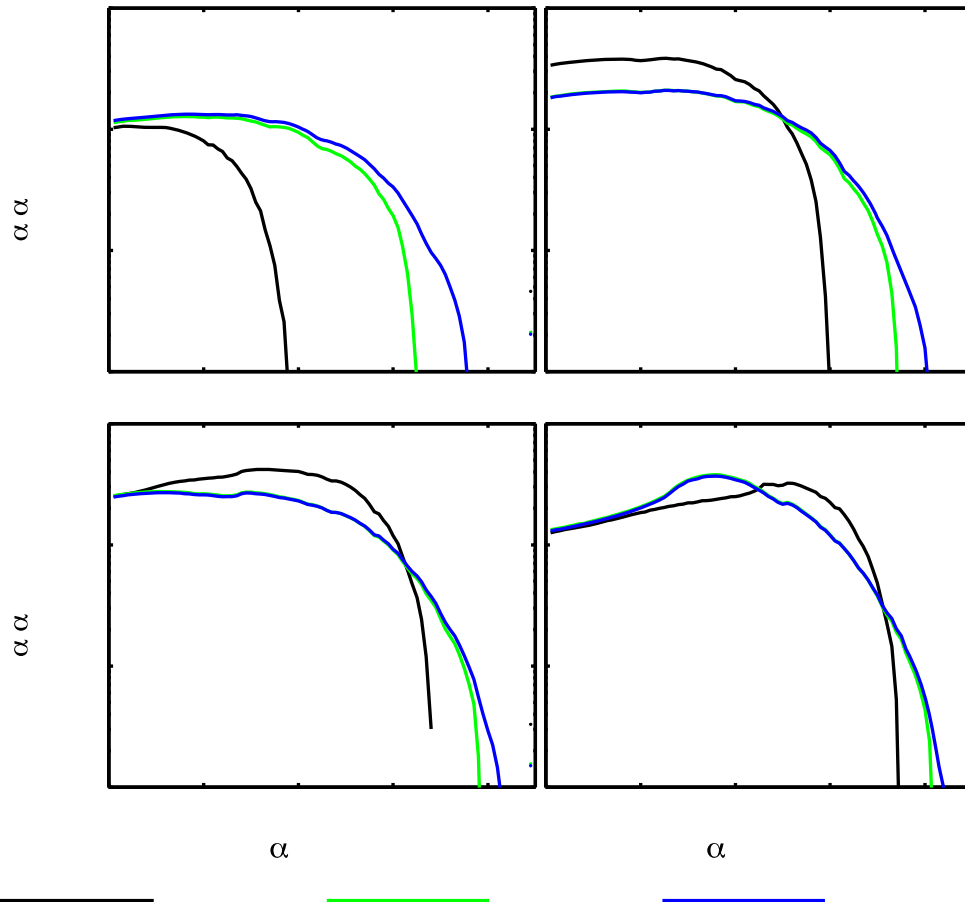
171 We find that obliquity of EMIC waves leads to an overall decrease of their ability to scatter  
 172 ultra-relativistic outer radiation belt electrons, when compared to field-aligned waves.  
 173 Importantly, we can also see that obliquity of EMIC waves does not change the range of  
 174 equatorial pitch angles that can be affected by electron scattering. Both implications are further  
 175 confirmed below, when looking at a sensitivity of electron scattering rates by EMIC waves to the  
 176 wave normal angle distribution.

### 177 2.3 Sensitivity of the Scattering Rates to Changes in the Wave and Plasma Properties

178 We start sensitivity studies of the diffusion coefficients for EMIC waves by varying the wave  
 179 spectral distribution parameters. The inputs used are summarized in Table 2, and the results are  
 180 comprised in Figure 3. We used three different sets of parameters of the wave spectral  
 181 distribution: the first is adopted from *Summers et al.* [2007b], the second set is from *Usanova et*  
 182 *al.* [2014], and finally, the third set of spectral distribution parameters is the same as the latter,  
 183 but with the higher upper cutoff frequency, which is the same as used in *Li et al.* [2007].

184 **Table 2. Model input parameters used to study sensitivity of the diffusion coefficients to the parameters of the**  
 185 **wave spectral distribution**

| Parameters                     | Values  |
|--------------------------------|---|
| Wave normal angle distribution | $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 30^\circ, \delta\theta = 30^\circ$   |
| Wave spectral distribution     | <p><u>Set 1</u> [<i>Summers et al.</i>, 2007b]: <math>\omega_{lc} = 2.5\Omega_{O^+}, \omega_{uc} = 3.5\Omega_{O^+}, \omega_m = 3\Omega_{O^+}, \delta\omega = 0.5\Omega_{O^+}</math></p> <p><u>Set 2</u> [<i>Usanova et al.</i>, 2014]: <math>\omega_{lc} = 2.16\Omega_{O^+}, \omega_{uc} = 3.84\Omega_{O^+}, \omega_m = 2.96\Omega_{O^+}, \delta\omega = 0.4\Omega_{O^+}</math></p> <p><u>Set 3</u>: <math>\omega_{lc} = 2.16\Omega_{O^+}, \omega_{uc} = 3.95\Omega_{O^+}, \omega_m = 2.96\Omega_{O^+}, \delta\omega = 0.4\Omega_{O^+}</math></p> |
| Wave amplitude                 | $dB = 1\text{ nT}$  |
| Plasma density model           | [ <i>Carpenter and Anderson</i> , 1992]   |
| Ion composition                | $\zeta_{H^+} = 0.7n_e, \zeta_{He^+} = 0.2n_e, \zeta_{O^+} = 0.1n_e$ [ <i>Meredith et al.</i> , 2003]  |
| Resonant numbers               | $n = -5, \dots, 0, \dots, 5$  |



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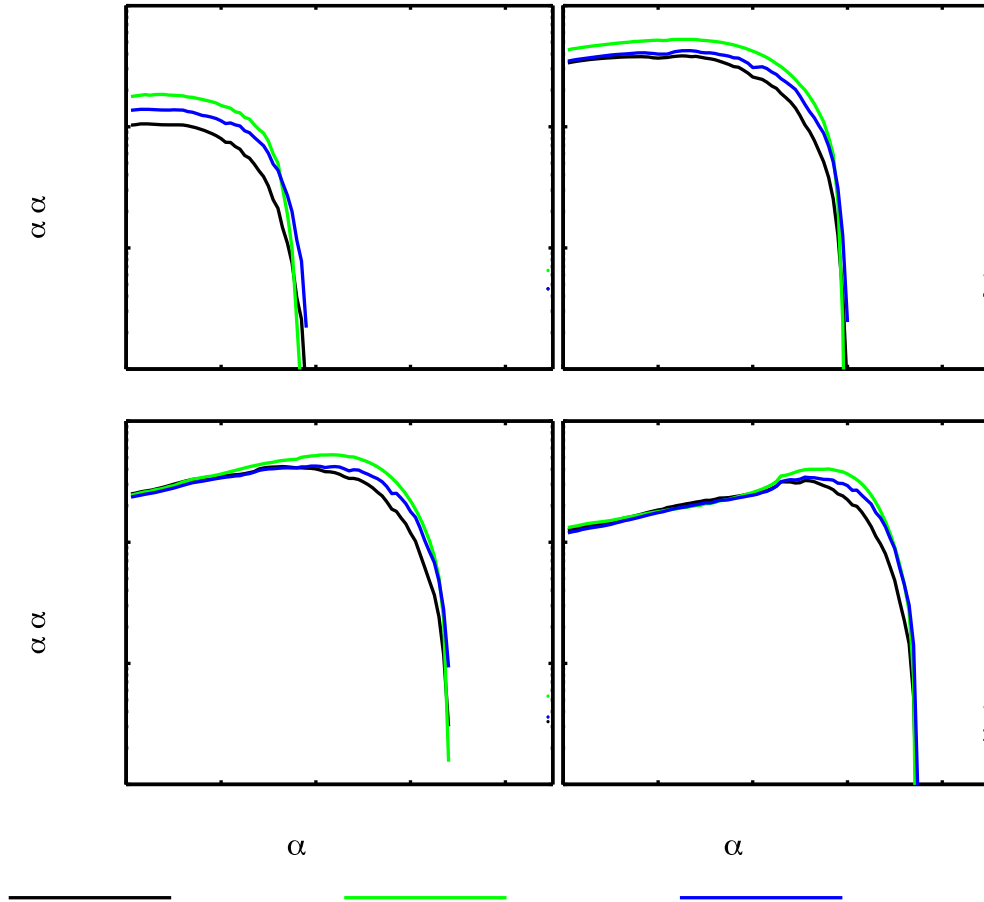
188 *Figure 3. Sensitivity of the bounce-averaged pitch angle diffusion coefficients for oblique EMIC waves at  $L=4$  to*  
 189 *the parameters of the wave spectral distribution. Black, green and blue lines denote results obtained using input*  
 190 *parameters of set 1, set 2, and set 3, respectively, from Table 2. Panels correspond to energies of a) 3 MeV, b) 5*  
 191 *MeV, c) 7 MeV and d) 10 MeV.*

192 Changes in the parameters of the wave spectral distributions lead to changes in the range of  
 193 electrons' equatorial pitch angles that can be affected by EMIC waves through diffusion. We  
 194 note that electron scattering rates appear to be more sensitive to the parameters of wave spectral  
 195 distribution at lower energies. For lower energy wave distribution changes the range of pitch  
 196 angles for which waves can resonate with electrons, while for higher energies for a broader range  
 197 of electron pitch angles, resonance conditions are satisfied and wave parameters play less of a  
 198 role. Indeed, the greatest sensitivity is observed at energy of 3 MeV, where we see a shift of the  
 199 maximum equatorial pitch angle of electrons to be scattered into the loss cone by about 30 and  
 200 40 degrees for the input parameters of set 2 and set 3 respectively. Finally, high sensitivity of  
 201 electron scattering rates due to interaction with EMIC waves to the upper cutoff frequency is to  
 202 be emphasized. We can see it by comparing scattering rate curves in Figure 3 corresponding to  
 203 sets 2 and 3. The effect was previously confirmed by studies of field-aligned EMIC waves, e.g.  
 204 *Li et al. [2007].*

205 We continue the sensitivity studies by varying parameters of the wave normal angle  
 206 distribution. We give model input parameters that were used in Table 3 and demonstrate  
 207 the results of simulations in Figure 4.

208 *Table 3. Model input parameters used to study sensitivity of the diffusion coefficients to the parameters of the*  
 209 *wave normal angle distribution (corresponding results are shown in Figure 4)*

| <i>Parameters</i>              | <i>Values</i>   |
|--------------------------------|---|
| Wave normal angle distribution | <u>Set 1:</u> $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 30^\circ, \delta\theta = 30^\circ$<br><u>Set 2:</u> $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 30^\circ, \delta\theta = 10^\circ$<br><u>Set 3:</u> $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 15^\circ, \delta\theta = 30^\circ$ |
| Wave spectral distribution     | $\omega_{lc} = 2.5\Omega_{O^+}, \omega_{uc} = 3.5\Omega_{O^+}, \omega_m = 3\Omega_{O^+}, \delta\omega = 0.5\Omega_{O^+}$<br>[Summers et al., 2007b]   |
| Wave amplitude                 | $dB = 1 \text{ nT}$   |
| Plasma density model           | <i>Carpenter and Anderson [1992]</i>  |
| Ion composition                | $\zeta_{H^+} = 0.7n_e, \zeta_{He^+} = 0.2n_e, \zeta_{O^+} = 0.1n_e$ [Meredith et al.,<br>2003]  |
| Resonant numbers               | $n = -5, \dots, 0, \dots, 5$  |



211

212 *Figure 4. Sensitivity of the bounce-averaged pitch angle diffusion coefficients for oblique EMIC waves at  $L=4$  to*  
 213 *the parameters of the wave normal angle distribution. Black, green and blue lines denote results obtained using*  
 214 *input parameters of set 1, set 2 and set 3 respectively from Table 3. Panels correspond to energies of a) 3 MeV, b)*  
 215 *5 MeV, c) 7 MeV and d) 10 MeV.*

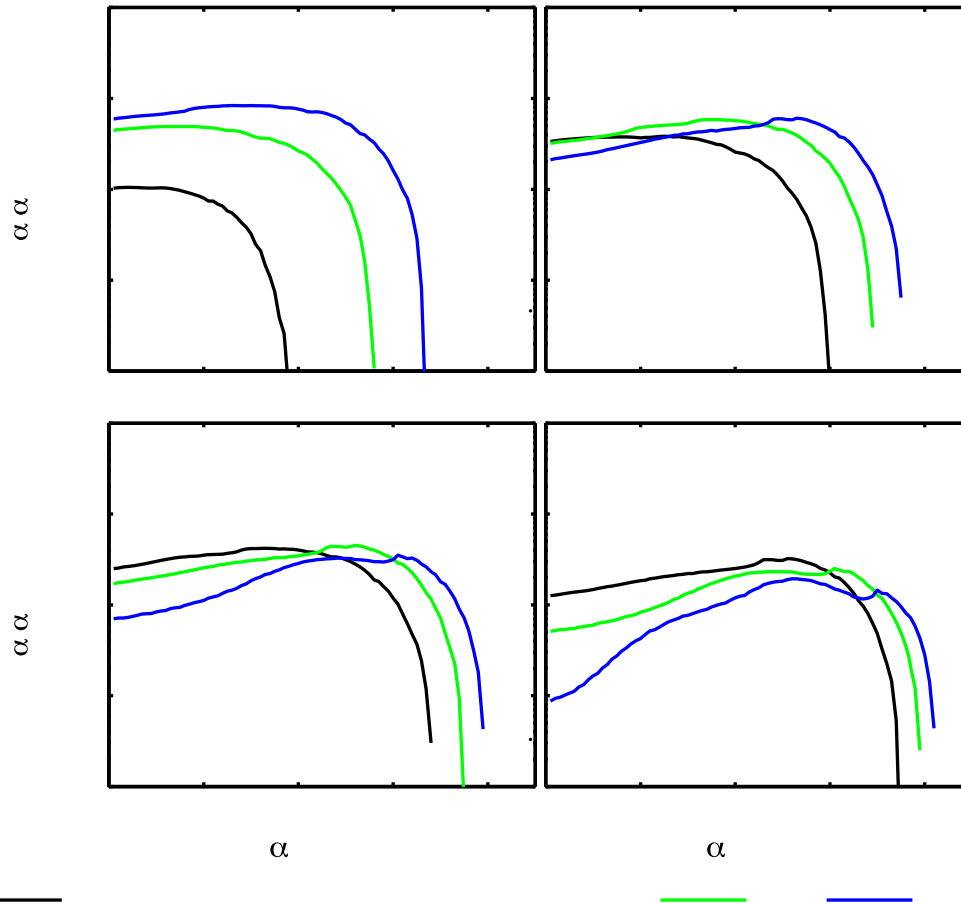
216 From Figure 4 we confirm a conclusion drawn from Figure 2 above, namely that changing  
 217 obliquity of EMIC waves does not lead to substantial changes in the range of equatorial pitch  
 218 angles to be affected by diffusion, unless waves turn out to be more oblique than assumed in this  
 219 study. At each panel of Figure 4, we see that all the plots corresponding to various wave normal  
 220 angle distributions terminate at almost exactly the same value of the equatorial pitch angle.

221 We move on to analysis of sensitivity of the diffusion coefficients for oblique EMIC waves to  
 222 ambient plasma density. Input parameters that were used in the study are depicted in  
 223 Table 4. The reference plasma density corresponds to statistical model of *Carpenter and*  
 224 *Anderson* [1992] at  $L = 4$ . Plasma density is then increased by a factor of two and four,  
 225 which approximately corresponds to three and six standard deviations, respectively, of the  
 226 statistical plasmasphere density model by *Sheeley et al.* [2001]. The results of the study are  
 227 shown in Figure 5.

228 **Table 4. Model input parameters used to study sensitivity of the diffusion coefficients to plasma density.**

| <i>Parameters</i>              | <i>Values</i>  |
|--------------------------------|--|
| Wave normal angle distribution | $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 30^\circ, \delta\theta = 30^\circ$  |
| Wave spectral distribution     | $\omega_{lc} = 2.5\Omega_{O^+}, \omega_{uc} = 3.5\Omega_{O^+}, \omega_m = 3\Omega_{O^+}, \delta\omega = 0.5\Omega_{O^+}$<br>[ <i>Summers et al.</i> , 2007b] |
| Wave amplitude                 | $dB = 1 \text{ nT}$  |
| Plasma density model           | <u>Case 1</u> : $n_e$ according to <i>Carpenter and Anderson</i> [1992]<br><br><u>Case 2</u> : $2n_e$<br><br><u>Case 3</u> : $4n_e$                          |
| Ion composition                | $\zeta_{H^+} = 0.7n_e, \zeta_{He^+} = 0.2n_e, \zeta_{O^+} = 0.1n_e$ [ <i>Meredith et al.</i> ,<br>2003]  |
| Resonant numbers               | $n = -5, \dots, 0, \dots, 5$   |

229



230

231 *Figure 5. Sensitivity of the bounce-averaged pitch angle diffusion coefficients for oblique EMIC waves at  $L=4$  to*  
 232 *plasma density. Black, green and blue lines represent the case of plasma density  $n_e$  corresponding to Carpenter*  
 233 *and Anderson [1992], and cases of two-time and four-time increase of  $n_e$  respectively. Panels correspond to*  
 234 *energies of a) 3 MeV, b) 5 MeV, c) 7 MeV and d) 10 MeV.*

235 From Figure 5 we note high sensitivity of the diffusion coefficients to plasma density – much  
 236 higher, when compared to sensitivity to the parameters of wave normal angle distribution. We  
 237 see that at all energies, growing plasma density leads to an increase of the range of equatorial  
 238 pitch angles of electrons to be scattered by EMIC waves.

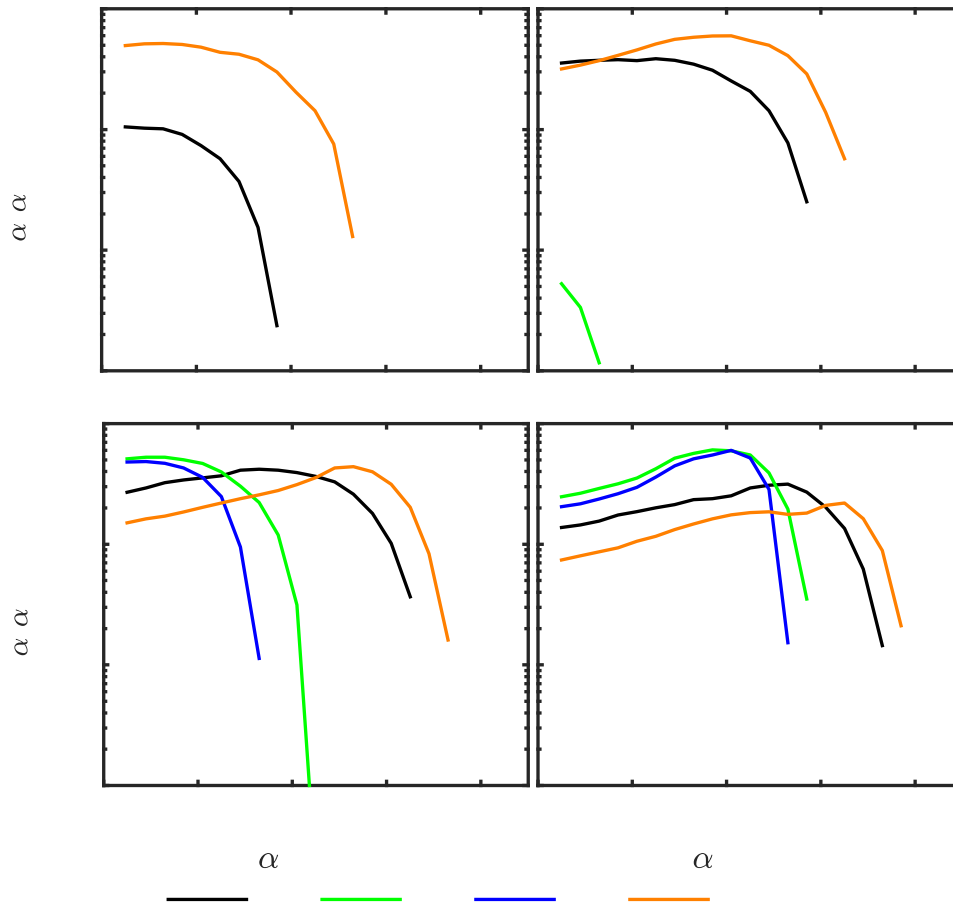
239 Significantly, ignoring the increasing range of electron equatorial pitch angles to be deflected by  
 240 EMIC waves, we observe two opposite effects of increasing plasma density at electron energies  
 241 of 3 MeV and 10 MeV. We note a dramatic enhancement in the ability of EMIC waves to scatter  
 242 3 MeV electrons with increasing plasma density, while at the same time a remarkable weakening  
 243 of their effect on 10 MeV electrons, with panels b) and c) of Figure 5 corresponding to a  
 244 transition between the two modes.

245 Finally, we conduct a sensitivity analysis of the diffusion coefficients for oblique EMIC waves  
 246 to ion composition. We use four different sets of ion number densities: given by *Meredith et al.*  
 247 [2003], *Horwitz et al.* [1984], and also simply  $\mathbf{H}^+$  plasma and plasma containing only 2% of  
 248 each  $\mathbf{He}^+$  and  $\mathbf{O}^+$  ions. All the input parameters used in the simulation are summarized in *Table*  
 249 *5*, and the results are presented in *Figure 6*.

250 *Table 5. Model input parameters used to study sensitivity of the diffusion coefficients to ion composition*

| <i>Parameters</i>              | <i>Values</i>  |
|--------------------------------|--|
| Wave normal angle distribution | $\theta_{lc} = 0^\circ, \theta_{uc} = 75^\circ, \theta_m = 30^\circ, \delta\theta = 30^\circ$  |
| Wave spectral distribution     | $\omega_{lc} = 2.5\Omega_{O^+}, \omega_{uc} = 3.5\Omega_{O^+}, \omega_m = 3\Omega_{O^+}, \delta\omega = 0.5\Omega_{O^+}$<br>[ <i>Summers et al.</i> , 2007b]   |
| Wave amplitude                 | $dB = 1 \text{ nT}$  |
| Plasma density model           | <i>Carpenter and Anderson</i> [1992]   |
| Ion composition                | <p><u>Set 1</u> [<i>Meredith et al.</i>, 2003]: <math>\zeta_{H^+} = 0.7n_e, \zeta_{He^+} = 0.2n_e, \zeta_{O^+} = 0.1n_e</math></p> <p><u>Set 2</u>: <math>\zeta_{H^+} = 0.96n_e, \zeta_{He^+} = 0.02n_e, \zeta_{O^+} = 0.02n_e</math></p> <p><u>Set 3</u>: <math>\zeta_{H^+} = n_e, \zeta_{He^+} = 0, \zeta_{O^+} = 0</math></p> <p><u>Set 4</u> [<i>Horwitz et al.</i>, 1984]: <math>\zeta_{H^+} = 0.6n_e, \zeta_{He^+} = 0.37n_e, \zeta_{O^+} = 0.03n_e</math></p> |
| Resonant numbers               | $n = -5, \dots, 0, \dots, 5$   |





252

253 *Figure 6. Sensitivity of the bounce-averaged pitch angle diffusion coefficient for oblique EMIC waves at L=4 to*  
 254 *ion composition. Black, green, blue and orange lines denote results obtained using input parameters of set 1, set*  
 255 *2, set 3 and set 4 respectively from Table 5.*

256 Figure 6 suggests that ion composition, which is very much unknown during storms [e.g.  
 257 *Moldwin, 1997; Kistler and Mouikis, 2015*], appears to be a key factor determining the lowest  
 258 energy that can be scattered by EMIC waves. Comparing the curves for sets 2 and 3, one can  
 259 also notice that the addition of even a small number of heavy ions leads to notable changes in  
 260 scattering by EMIC waves, as mentioned by *Albert [2003]*. Importantly, both scattering  
 261 timescales associated with EMIC waves and the range of electrons' pitch angles affected are  
 262 altered significantly by varying ion composition.

263 Given great variability of the ion composition in the plasmasphere during storm time [e.g. *Kistler*  
 264 *and Mouikis, 2015; Gallagher et al., 2000*] and high sensitivity (compared to wave spectral and  
 265 wave normal angle distributions and ambient plasma density) of electron scattering by EMIC on  
 266 the composition, we conclude that it is the key ingredient necessary for a good agreement of  
 267 theoretical models with in-situ observations of radiation belt electrons lifetimes.

## 268 3. Discussion

### 269 3.1 Contribution of various resonances

270 By studying the contribution of various cyclotron resonance numbers into an overall diffusion  
 271 coefficient associated with EMIC waves, we showed that for a wide range of energies (from 3  
 272 MeV to 10 MeV in our work) and given a realistic ambient plasma density, it is reasonable to  
 273 neglect higher order resonances, as well as Landau resonances, when studying oblique EMIC  
 274 waves. This does not mean obliquity should be ignored, because the wave normal angle is still a  
 275 part of the resonance condition and dispersion relation equations, and therefore affects scattering  
 276 rates (see Figure 2). We observed a growing contribution of higher-order resonances with  
 277 increasing electron energies. Scattering rates associated with those resonances are still orders of  
 278 magnitude smaller than the rates provided by  $|n| = 1$  resonances.

### 279 3.2 Sensitivity analysis

#### 280 *Wave spectral distribution*

281 We studied the sensitivity of the diffusion coefficients by adopting wave spectral distributions  
 282 from *Summers et al.* [2007b] and *Usanova et al.* [2014] and assuming the wave normal angle of  
 283  $\theta_m = 30^\circ$ . We showed that changing the spectral distribution leads to a dramatic alteration of the  
 284 range of electron equatorial pitch angles being affected by oblique EMIC waves. For the energy  
 285 of 3 MeV, the changes accounted for  $30^\circ - 40^\circ$ . We confirmed high sensitivity of the diffusion  
 286 coefficients for oblique EMIC to the upper cutoff frequency reported in *Li et al.* [2007] for the  
 287 case of field-aligned waves.

#### 288 *Wave normal angle distribution*

289 Even though obliquity of EMIC waves changes scattering rates of ultra-relativistic electrons in  
 290 comparison with field aligned waves, we have not noted significant changes in scattering rates  
 291 when varying the maximum of wave normal angle distribution from  $\theta_m = 30^\circ$  to  $\theta_m = 15^\circ$ , and  
 292 varying the distribution bandwidth from  $\delta\theta = 30^\circ$  to  $\delta\theta = 10^\circ$ . Neither the value of the  
 293 diffusion coefficients, nor the range of equatorial pitch angles of electrons scattered by the  
 294 prescribed wave population were changing appreciably. It is important to note that these  
 295 observations are valid unless waves turn out to be more oblique, say  $\theta_m = 60^\circ$  or  $\theta_m = 80^\circ$ ,  
 296 which has been observed for selected storms in the recent statistical studies [e.g. *Saikin et al.*,  
 297 2015; *Allen et al.*, 2015].

#### 298 *Plasma density*

299 Adopting the values of ambient plasma density prescribed by *Carpenter and Anderson* [1992] at  
 300  $L = 4$ , and increasing the density by three and by six standard deviations defined by the model  
 301 of *Sheeley et al.* [2001], we demonstrated high sensitivity of the scattering rates to this

302 parameter. For instance, one can notice one-two orders of magnitude changes in the value of the  
303 diffusion coefficients when changing the value of plasma density in Figure 5, as well as changes  
304 of affected pitch angles by  $20^\circ - 30^\circ$  at electron energies of 3 MeV and  $5^\circ - 10^\circ$  for the energy  
305 of 10 MeV. In our simulation, increasing ambient plasma density always leads to an increase in  
306 electron pitch angles that are able to resonantly interact with EMIC waves. As for the scattering  
307 rates, we observed its one-two orders of magnitude increase for electron energy of 3 MeV and its  
308 significant decrease for electron energies of 10 MeV, especially at the lower pitch angles.

### 309 *Ion composition*

310 In order to investigate sensitivity of electron scattering rates caused by EMIC waves, we used  
311 four different sets of parameters. Two were adopted from *Meredith et al.* [2003] and from  
312 *Horwitz et al.* [1984]. The other two represented pure hydrogen plasma and plasma mainly  
313 consisting of hydrogen ions, but for 2% additional amount of each helium and oxygen ions. It  
314 was shown that the addition of even a small number of heavy ions leads to notable changes in  
315 diffusion coefficients. Both scattering rates and the range of pitch angles are affected.  
316 Furthermore, given high variability of the ion composition during magnetic storms, we conclude  
317 that ion composition is a key factor determining the minimum electron energy to interact with  
318 EMIC waves, as well as scattering rates of ultra-relativistic electrons caused by EMIC waves.

### 319 **3.3 General remarks**

320 We should emphasize that conclusions of our work are based on the assumption that waves are  
321 quasi-parallel, e.g.  $\theta_m = 30^\circ$ . If waves turn out to be more oblique than in this study, higher  
322 order resonances may play a more important role. Future detailed statistical analysis of the wave  
323 measurements will be required to validate this assumption. Additionally, we focused on the case  
324 of left-hand polarized He band waves. The choice is based on the statistical study of *Saikin et al.*  
325 [2015], who showed that He band waves statistically have a higher overall occurrence rate, as  
326 well as higher amplitudes, when compared to H and O EMIC bands.

327 Notably, the results obtained in the study possess the accuracy of an analytical solution as a  
328 result of using the full cold plasma dispersion relation. Thus, presented research can be used as a  
329 benchmark for future validation of various numerical methods.

330 With regards to the assumption of cold plasma, *Silin et al.* [2011] and *Chen et al.* [2011], [2013]  
331 showed that assumption of cold plasma theory may lead to significant errors when quantifying  
332 the effects of EMIC waves on radiation belt relativistic and ultra-relativistic electrons. Non-  
333 resonant and non-linear interactions may also contribute to electron scattering. It was shown in  
334 the past that for EMIC waves with amplitude of several nT nonlinear effects such as phase  
335 bunching and phase trapping may play important role and cause significant deviations from the

336 results of quasi-linear model [e.g. *Albert and Bortnik, 2009; Su et al., 2012; 2013*]. Thus, the  
337 assumption of non-thermal motion of plasma particles, as well as quasi-linear approach, has to  
338 be challenged in our future studies.

339 It is interesting to notice that EMIC waves can violate the second adiabatic invariant of nearly  
340 equatorially mirroring electrons and provide pitch-angle scattering for a wide range of electron  
341 energies and L-shells inside and outside the plasmasphere [*Shprits, 2009*]. Another potential  
342 mechanism for scattering equatorially mirroring electrons is bounce resonance by fast  
343 magnetosonic waves [*Shprits, 2016*].

344 In the future we are planning to use the new diffusion coefficients for oblique EMIC waves with  
345 one of the radiation belt codes that computes particle phase space densities and fluxes – Versatile  
346 Electron Radiation Belt code [e.g. *Shprits et al., 2008*]. This study would help to further quantify  
347 the effect of oblique EMIC waves on the radiation belt particle fluxes, which is of utmost interest  
348 from the perspective of space weather applications.

#### 349 **4. Summary and Conclusions**

350 We have studied pitch angle scattering of ultra-relativistic electrons by obliquely propagating  
351 EMIC waves in the Earth's outer radiation belt. Note that the conclusion below corresponds to  
352 the most frequently observed He band and may or may not be valid for other bands. The  
353 investigation of other EMIC wave bands will be a subject of future studies. We used a full cold  
354 plasma dispersion relation that was derived semi-analytically using Matlab Symbolic Toolbox,  
355 and improved Full Diffusion Code (FDC) to calculate bounce-averaged pitch angle diffusion  
356 coefficients for oblique EMIC waves.

357 We compared bounce-averaged diffusion coefficients computed with the assumption of parallel  
358 wave propagation to the results of our code that uses the full cold plasma dispersion relation,  
359 taking into account the oblique propagation of waves and higher-order resonances. We have  
360 quantified the sensitivity of ultra-relativistic electron pitch angle scattering rates by EMIC waves  
361 to the model input parameters – wave spectral distribution, wave normal angle distribution,  
362 ambient plasma density and ion composition.

363 The main conclusions of the work are as follows:

364 1. From the results of studying contribution of various resonances, we demonstrate that in our  
365 simulations for energies up to approximately 7 MeV, the primary resonance ( $n = 1$ ) plays a  
366 dominant role. For higher energies, both  $n = \pm 1$  resonances should be taken into account. For  
367 the presented simulations, other resonances, including Landau resonance, can be neglected.

368 2. Sensitivity analysis conducted in this work leads us to the following conclusions:

- 369 • Obliquity of EMIC waves leads to an overall decrease of ability of EMIC waves to  
370 scatter ultra-relativistic electrons;
  - 371 • Obliquity of EMIC waves does not substantially change the range of equatorial pitch  
372 angles of electrons that are able to resonate with EMIC waves, unless the waves turn out  
373 to be much more oblique than assumed in this study;
  - 374 • Electrons of higher values of equatorial pitch angles can resonate with EMIC waves  
375 when the plasma density is higher and there is a substantial wave power close to the local  
376 ion gyrofrequency;
  - 377 • Increases in background plasma density cause enhancement of electron scattering by  
378 EMIC waves at energies of approximately (3-5 MeV), and weakening at higher energies  
379 (7-10 MeV);
  - 380 • Diffusion coefficients for oblique EMIC waves appear to be most sensitive to the  
381 assumed ion composition. The sensitivity to realistic variations of wave spectral  
382 distribution, wave normal angle distribution, and ambient plasma density show that these  
383 factors can also determine scattering rates but to a lesser degree;
  - 384 • The lowest electron energy to be scattered by EMIC waves is most sensitive to ion  
385 composition of ambient plasma;
  - 386 • Uncertainty in the model input parameters and neglect of oblique wave propagation lead  
387 to comparable errors in diffusion coefficients for EMIC waves (yet uncertainty in ion  
388 composition leads to greater errors).
- 389 3. The approach used in this work and the improved Full Diffusion Code (FDC) for oblique  
390 EMIC waves enables studies of other plasma waves affecting magnetically trapped radiation in  
391 magnetic fields of Earth and outer planets.
- 392 4. The presented method can be used to quantify scattering rates by other wave modes that  
393 depend on the ion composition. The results obtained with the presented method can serve as a  
394 benchmark for future validation of various approximate numerical methods.

**395 Acknowledgments**

396 We would like to thank Lynn Kistler, Matina Gkioulidou, Maria Usanova, Adam Kellerman and  
397 Alexander Drozdov for their contribution and advice. Support at UCLA was provided by NASA  
398 grant NNX10AK99G, NNX13AE34G, NSF grant 443869-YS-21686, and UC Lab Fee grant  
399 #116720, Horizon 2020 #637302. We would like to thank International Space Science Institute  
400 (Bern) for the support provided.

401 All data for this paper are properly cited and referred to in the reference list.

402 **References**

- 403 Albert, J. M. (2003) Evaluation of quasi-linear diffusion coefficients for EMIC waves in a  
 404 multispecies plasma, *J. Geophys. Res. Space Physics*, 108, A6, ISSN 2156-2202,  
 405 <http://dx.doi.org/10.1029/2002JA009792>.
- 406 Albert, J. M. (2007), Simple approximations of quasi-linear diffusion coefficients, *J.*  
 407 *Geophys. Res.*, 112, A12202, doi:10.1029/2007JA012551.
- 408 Albert, J. M., and J. Bortnik (2009a), Nonlinear interaction of radiation belt electrons with  
 409 electromagnetic ion cyclotron waves, *Geophys. Res. Lett.*, 36, L12110,  
 410 doi:10.1029/2009GL038904.
- 411 Allen, R. C., J.-C. Zhang, L. M. Kistler, H. E. Spence, R.-L. Lin, B. Klecker, M. W. Dunlop,  
 412 M. André, and V. K. Jordanova (2015), A statistical study of EMIC waves observed by Cluster:  
 413 1. Wave properties, *J. Geophys. Res. Space Physics*, 120, 5574–5592, doi:10.1002/  
 414 2015JA021333.
- 415 Carpenter, D. L., R.R. Anderson, (1992) An ISEE/Whistler model of equatorial electron  
 416 density in the magnetosphere, *J. Geophys. Res. Space Physics*, 97, A2, 1097-1108, ISSN 2156-  
 417 2202, <http://dx.doi.org/10.1029/91JA01548>.
- 418 Chen, L., R. M. Thorne, and J. Bortnik (2011), The controlling effect of ion temperature on  
 419 EMIC wave excitation and scattering, *Geophys. Res. Lett.*, 38, L16109,  
 420 doi:10.1029/2011GL048653.
- 421 Chen, L., R. M. Thorne, Y. Shprits, and B. Ni (2013), An improved dispersion relation for  
 422 parallel propagating electromagnetic waves in warm plasmas: Application to electron scattering,  
 423 *J. Geophys. Res. Space Physics*, 118, 2185–2195, doi:10.1002/jgra.50260.
- 424 Gallagher, D. L., Craven, P. D., Comfort, R. H. (2000), Global core plasma model, *J.*  
 425 *Geophys. Res. Space Physics*, 105, A8, 18819-18833 ISSN 2156-2202, doi:  
 426 10.1029/1999JA000241.
- 427 Glauert, S. A., and R. B. Horne (2005), Calculation of pitch angle and energy diffusion  
 428 coefficients with the PADIE code, *J. Geophys. Res.*, 110, A04206, doi:10.1029/2004JA010851.
- 429 Gu, X., Y. Y. Shprits, and B. Ni (2012), Parameterized lifetime of radiation belt electrons  
 430 interacting with lower-band and upper-band oblique chorus waves, *Geophys. Res. Lett.*, 39,  
 431 L15102, doi:10.1029/2012GL052519.
- 432 He, F., et al. (2016), Combined scattering loss of radiation belt relativistic electrons by  
 433 simultaneous three-band EMIC waves: A case study, *J. Geophys. Res. Space Physics*, 121, 2169-  
 434 9402, doi:10.1002/2016JA022483.
- 435 Kistler L.M. AND C.G. Mouiskis (2015), The Inner Magnetosphere Ion Composition and  
 436 Local Time Distribution over a Solar Cycle, submitted to *J. Geophys. Res.*
- 437 Li, W., Shprits, Y. Y., and Thorne, R. M. (2007), Dynamic evolution of energetic outer zone  
 438 electrons due to wave-particle interactions during storms, *J. Geophys. Res.*, 112, A10, 2156-  
 439 2202, 10.1029/2007JA012368.
- 440 Lyons, L. R., and R. M. Thorne (1972), Parasitic pitch angle diffusion of radiation belt  
 441 particles by ion cyclotron waves, *J. Geophys. Res.*, 77, 5608–5616,  
 442 doi:10.1029/JA077i028p05608.
- 443 Meredith, N. P., R. B. Horne, T. Kersten, B. J. Fraser, and R. S. Grew (2014), Global  
 444 morphology and spectral properties of EMIC waves derived from CRRES observations, *J.*  
 445 *Geophys. Res. Space Physics*, 119, 5328–5342, doi:10.1002/2014JA020064.

- 446 Meredith, N. P., R. M. Thorne, R. B. Horne, D. Summers, B. J. Fraser, and R. R.  
 447 Anderson (2003), Statistical analysis of relativistic electron energies for cyclotron resonance  
 448 with EMIC waves observed on CRRES, *J. Geophys. Res.*, 108, 1250,  
 449 doi:10.1029/2002JA009700.
- 450 Moldwin, M. B. (1997), Outer Plasmaspheric Plasma Properties: What We Know from  
 451 Satellite Data, *Space Science Reviews*, 80, ISSN 0038-6308, 181-198, doi:  
 452 10.1023/A:1004921903897.
- 453 Ni, B., R. M. Thorne, Y. Y. Shprits, and J. Bortnik (2008), Resonant scattering of plasma  
 454 sheet electrons by whistler-mode chorus: Contribution to diffuse auroral precipitation, *Geophys.*  
 455 *Res. Lett.*, 35, L11106, doi:10.1029/2008GL034032.
- 456 Ni, B., et al. (2015), Resonant scattering of outer zone relativistic electrons by multiband  
 457 EMIC waves and resultant electron loss time scales, *J. Geophys. Res. Space Physics*, 120, 9,  
 458 7357-7373, ISSN 2169-9402, doi:10.1002/2015JA021466
- 459 Orlova, K., and Y. Y. Shprits (2011), On the bounce-averaging of scattering rates and the  
 460 calculation of bounce period, *Phys. Plasmas* 18, 092904, doi:10.1063/1.3638137.
- 461 Saikin, A. A., J.-C. Zhang, R. C. Allen, C. W. Smith, L. M. Kistler, H. E. Spence, R. B.  
 462 Torbert, C. A. Kletzing, and V. K. Jordanova (2015), The occurrence and wave properties of H<sup>+</sup>-  
 463 , He<sup>+</sup>-, and O<sup>+</sup>-band EMIC waves observed by the Van Allen Probes, *J. Geophys. Res. Space*  
 464 *Physics*, 120, 7477–7492, doi:[10.1002/2015JA021358](https://doi.org/10.1002/2015JA021358).
- 465 Sheeley, B. W., Moldwin, M. B., Rassoul, H. K., Anderson, R. R. (2001), An empirical  
 466 plasmasphere and trough density model: CRRES observations, *J. Geophys. Res. Space Physics*,  
 467 106, A11, ISSN 2156-2202, <http://dx.doi.org/10.1029/2000JA000286PY>
- 468 Shprits Y. Y., Subbotin D. A., Meredith N. P., Elkington S. R. (2008), Review of modeling of  
 469 losses and sources of relativistic electrons in the outer radiation belt II: Local acceleration and  
 470 loss, *J. Atmos. Sol. Terr. Phys.*, 70, 14, 1694-1713, ISSN 1364-6826,  
 471 <http://dx.doi.org/10.1016/j.jastp.2008.06.014>.
- 472 Shprits, Y. Y. (2009), Potential waves for pitch-angle scattering of near-equatorially  
 473 mirroring energetic electrons due to the violation of the second adiabatic invariant, *Geophys.*  
 474 *Res. Lett.*, 36, L12106, doi:[10.1029/2009GL038322](https://doi.org/10.1029/2009GL038322).
- 475 Shprits, Y. Y., et al. (2015), Combined convective and radial diffusive transport with pitch and  
 476 energy scattering: 4D VERB results, in preparation for *Geophys. Res. Lett.*
- 477 Su, Z., F. Xiao, H. Zheng, and S. Wang (2011), CRRES observation and STEERB simulation  
 478 of the 9 October 1990 electron radiation belt dropout event, *Geophys. Res. Lett.*, 38, L06106,  
 479 doi:10.1029/2011GL046873.
- 480 Su, Z., H. Zhu, F. Xiao, H. Zheng, C. Shen, Y. Wang, and S. Wang (2012), Bounce-averaged  
 481 advection and diffusion coefficients for monochromatic electromagnetic ion cyclotron wave:  
 482 Comparison between test-particle and quasi-linear models, *J. Geophys. Res.*, 117, A09222,  
 483 doi:10.1029/2012JA017917.
- 484 Su, Z., H. Zhu, F. Xiao, H. Zheng, C. Shen, Y. Wang, and S. Wang (2013), Latitudinal  
 485 dependence of nonlinear interaction between electromagnetic ion cyclotron wave and radiation  
 486 belt relativistic electrons, *J. Geophys. Res. Space Physics*, 118, 3188-3202,  
 487 doi:10.1002/jgra.50289.
- 488 Summers, D., and R. M. Thorne (2003), Relativistic electron pitch-angle scattering by  
 489 electromagnetic ion cyclotron waves during geomagnetic storms, *J. Geophys. Res.*, 108(A4),  
 490 1143, doi:10.1029/2002JA009489.
- 491 Summers, D., B. Ni, and N. P. Meredith (2007a), Timescales for radiation belt electron



- 492 acceleration and loss due to resonant wave-particle interactions: 1. Theory, *J. Geophys. Res.*,  
493 112, A04206, doi:[10.1029/2006JA011801](https://doi.org/10.1029/2006JA011801).
- 494 Summers D., Ni B., Meredith N.P. (2007b), Timescales for radiation belt electron  
495 acceleration and loss due to resonant wave-particle interactions: 2. Evaluation for VLF chorus,  
496 ELF hiss, and electromagnetic ion cyclotron waves, *J. Geophys. Res. Space Physics*, 112, A4,  
497 ISSN 2156-2202, <http://dx.doi.org/10.1029/2006JA011993>
- 498 Thorne, R. M., and C. F. Kennel (1971), Relativistic electron precipitation during magnetic  
499 storm main phase, *J. Geophys. Res.*, 76(19), 4446–4453, doi:10.1029/JA076i019p04446.
- 500 Ukhorskiy, A. Y., Y.Y. Shprits, B. J. Anderson, K. Takahashi, and R. M. Thorne (2010),  
501 Rapid scattering of radiation belt electrons by storm time EMIC waves, *Geophys. Res. Lett.*, 37,  
502 L09101, doi:10.1029/2010GL042906.
- 503 Usanova, M.E., et al. (2014), Effect of EMIC waves on relativistic and ultrarelativistic  
504 electron populations: Ground-based and Van Allen Probes observations, *Geophys. Res. Lett.*, 41,  
505 doi:10.1002/2013GL059024.
- 506

Supporting Information for

## Scattering of Relativistic and Ultra-relativistic Electrons by Obliquely Propagating Electromagnetic Ion Cyclotron Waves

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### Introduction

Material below describes quasi-linear formulation that was used in our study to calculate diffusion coefficients for oblique EMIC waves propagating in multi-ion plasma.

### Full Cold Plasma Dispersion Relation

In our study we used theory of cold plasma which is defined as a collection of charged particles without any net charge that have no kinetic thermal motion on their own [Swanson, 2003]. Cold plasma particles can only be induced to move by self-consistent electric and magnetic fields of the wave. Cold plasma dispersion relation (CPDR) is derived in plasma physics textbooks, for example in Stix [1962]. CPDR in multicomponent plasma is given by the following set of expressions

$$A\mu^4 - B\mu^2 + C = 0, \quad (1)$$

where  $\mu = kc/\omega$  is refractive index, and

$$A = SX^2 + P, \quad (2)$$

$$B = RLX^2 + PS(2 + X^2), \quad (3)$$

$$C = PRL(1 + X^2), \quad (4)$$

where  $X = \tan(\theta)$ ,  $\theta$  is the angle between the wave normal and local geomagnetic field vector and  $P, R, L$  and  $S$  are the Stix parameters defined as

$$R = 1 - \sum_i \frac{\omega_{pi}^2}{\omega} \frac{1}{\omega + \Omega_i} \quad (5)$$

$$L = 1 - \sum_i \frac{\omega_{pi}^2}{\omega} \frac{1}{\omega - \Omega_i} \quad (6)$$

$$P = 1 - \sum_i \frac{\omega_{pi}^2}{\omega^2} \quad (7)$$

$$S = \frac{R + L}{2} \quad (8)$$

where the summations is over all the plasma species (electrons and ions),  $\omega$  is the wave frequency, and  $\omega_{pi}$  and  $\Omega_i$  are plasma and cyclotron frequencies respectively

$$\omega_{pi}^2 = \frac{n_i e^2}{m_i \epsilon_0} \quad (9)$$

$$\Omega_i = \frac{|q_i| B}{m_i} \quad (10)$$

Here  $n_i$  and  $m_i$  are number density and mass, respectively, of particles of species  $i$ , and  $B$  is local magnetic field value.

Substituting expressions for plasma and cyclotron frequencies (9)-(10) into expressions for the Stix parameters (5)-(8), then Stix parameters into the coefficients  $A$ ,  $B$ , and  $C$  given by (2)-(4), and finally substituting resulting expressions into (1) yields cold plasma dispersion relation of the form  $\mathcal{D}(k, X, \omega) = 0$ . In the case of plasma consisting of electrons, hydrogen, helium, and oxygen ions expression  $\mathcal{D}(k, X, \omega)$  is fourteenth order polynomial with respect to the wave frequency  $\omega$ .

### Computation of the Diffusion Coefficients

Bounce-averaged diffusion coefficients for oblique EMIC waves are calculated using formalism reported in [Glauert and Horne, 2005; Albert, 2005]. In this section we give summary of the equations implemented in our model.

Charged particles populating radiation belts can undergo resonant interactions with plasma waves that are present in the Earth's magnetosphere. Resonance condition is met when Doppler-shifted frequency (frequency of the wave in the electron frame of reference) of obliquely propagating circularly polarized wave is equal to multiples of electron gyrofrequency. The condition can be written as [Stix, 1962]

$$\omega - k_{\parallel} v_{\parallel} = \frac{n \Omega_e}{\gamma} \quad (11)$$

where  $k_{\parallel}$  and  $v_{\parallel}$  are parallel components (with respect to local geomagnetic field vector) of wave normal and electron velocity vectors,  $n = 0, \pm 1, \pm 2, \pm 3, \dots$  is called resonance harmonic number, and  $\gamma$  is relativistic Lorentz factor.

Thus, waves satisfying resonance condition (11) and plasma dispersion relation  $\mathcal{D}(k, X, \omega) = 0$  can resonantly interact with RB electrons causing diffusion in adiabatic invariants.

Pitch angle and momentum local diffusion coefficients are defined as

$$D_{\alpha\alpha} = \frac{1}{2} \left\langle \frac{(\Delta\alpha)^2}{\Delta t} \right\rangle \quad (12)$$

$$D_{\alpha p} = \frac{1}{2p} \left\langle \frac{\Delta\alpha\Delta p}{\Delta t} \right\rangle \quad (13)$$

$$D_{pp} = \frac{1}{2p^2} \left\langle \frac{(\Delta p)^2}{\Delta t} \right\rangle \quad (14)$$

where  $\langle (\Delta\alpha)^2 / \Delta t \rangle$  denotes the expected value of change of  $(\Delta\alpha)^2$  per unit time  $\Delta t$  of for a particle with pitch angle  $\alpha$ , and term  $\langle \Delta\alpha\Delta p / \Delta t \rangle$  has similar definition.

In order to calculate diffusion coefficients we assume that wave spectral power density  $B^2(\omega)$  is distributed normally following Lyons [1974]

$$B^2(\omega) = \begin{cases} A^2 \exp\left(-\left(\frac{\omega - \omega_m}{\delta\omega}\right)^2\right), & \omega_{lc} \leq \omega \leq \omega_{uc} \\ 0, & otherwise \end{cases} \quad (15)$$

where  $B^2$  is measured in  $T^2/Hz$ ,  $\omega_m$  and  $\delta\omega$  are distribution maximum and bandwidth respectively,  $\omega_{lc}$  and  $\omega_{uc}$  are lower and upper cut-off frequencies, and  $A$  is a normalization factor given by

$$A^2 = \frac{|B_w|^2}{\delta\omega} \frac{2}{\sqrt{\pi}} \left[ \operatorname{erf}\left(\frac{\omega_m - \omega_{lc}}{\delta\omega}\right) + \operatorname{erf}\left(\frac{\omega_{uc} - \omega_m}{\delta\omega}\right) \right]^{-1} \quad (16)$$

where  $B_w$  is the wave amplitude measured in Tesla.

The distribution of the tangent of wave normal angle  $\theta$  is also assumed to be Gaussian:

$$g(X) = \begin{cases} \exp\left(-\left(\frac{X - X_m}{X_w}\right)^2\right), & X_{min} \leq X \leq X_{max} \\ 0, & otherwise \end{cases} \quad (17)$$

where  $X_w$  is the angular distribution bandwidth,  $X_m$  is the peak, and  $X_{min}$  and  $X_{max}$  are the cut-off values.

Given these definitions, local diffusion coefficients can be written as

$$D_{\alpha\alpha} = \frac{1}{p^2} \sum_{n=n_l}^{n=n_h} \int_{X_{min}}^{X_{max}} X dX D_{\alpha\alpha}^{nX} \quad (18)$$

$$D_{\alpha p} = \frac{1}{p^2} \sum_{n=n_l}^{n=n_h} \int_{X_{min}}^{X_{max}} X dX D_{\alpha p}^{nX} \quad (19)$$

$$D_{pp} = \frac{1}{p^2} \sum_{n=n_l}^{n=n_h} \int_{X_{min}}^{X_{max}} X dX D_{pp}^{nX} \quad (20)$$

where the summation is over the resonant harmonics and

$$D_{\alpha\alpha}^{nX} = \sum_i \frac{q_e^2 \omega_i^2}{4\pi(1+X^2)N(\omega_i)} \left[ \frac{n\Omega_e/\gamma\omega_i - \sin^2\alpha}{\cos\alpha} \right]^2 B^2(\omega_i) g(X) \left. \frac{|\Phi_{n,k}|^2}{\left| \vartheta_{\parallel} - \frac{\partial\omega}{\partial k_{\parallel}} \right|_{\mathcal{D}=0}} \right|_{k_{\parallel i}} \quad (21)$$

$$D_{\alpha p}^{nX} = D_{\alpha\alpha}^{nX} \left[ \frac{\sin(\alpha)\cos(\alpha)}{n\Omega_e/\gamma\omega_i - \sin^2\alpha} \right] \quad (22)$$

$$D_{pp}^{nX} = D_{\alpha\alpha}^{nX} \left[ \frac{\sin(\alpha)\cos(\alpha)}{n\Omega_e/\gamma\omega_i - \sin^2\alpha} \right]^2 \quad (23)$$

and quantities  $D_{\alpha\alpha}^{nX}$ ,  $D_{\alpha p}^{nX}$  and  $D_{pp}^{nX}$  are evaluated at the resonant parallel wave number and resonant frequency,  $k_{\parallel i}$  and  $\omega_i$  respectively, that are found by solving a set of resonance condition (11) and of plasma dispersion relation equations:

$$\begin{cases} \omega - k_{\parallel}v_{\parallel} = \frac{n\Omega_e}{\gamma} \\ \mathcal{D}(k, \omega, X) = 0 \end{cases} \quad (24)$$

It is important to notice that derivative  $\partial\omega/\partial k_{\parallel}|_{\mathcal{D}=0}$  in (21) is calculated along the dispersion relation line  $\mathcal{D}(k, X, \omega) = 0$ .

Functions  $|\Phi_{n,k}|^2$  and  $N(\omega)$  are given by

$$\begin{aligned} |\Phi_{n,k}|^2 &= \\ &= \left[ \left( \frac{\mu^2 - L}{\mu^2 - S} \right) J_{n+1} + \left( \frac{\mu^2 - R}{\mu^2 - S} \right) J_{n-1} \right. \\ &\quad \left. + \tan^{-1}(\alpha) \sin(\psi) \cos(\psi) J_n \right] \left( \frac{\mu^2 \sin^2(\psi) - P}{2\mu^2} \right)^2 \left[ \left( \frac{R - L}{2(\mu^2 - S)} \right)^2 \left( \frac{P - \mu^2 \sin^2(\psi)}{\mu^2} \right. \right. \\ &\quad \left. \left. + \left( \frac{P \cos(\psi)}{\mu^2} \right)^2 \right)^{-1} \right] \end{aligned} \quad (25)$$

$$N(\omega) = \frac{1}{2\pi^2} \int_{X_{min}}^{X_{max}} \frac{g(X)k^2}{(1+X^2)^{1/2}} \frac{\partial\mathcal{D}}{\partial\omega} \left[ \frac{\partial\mathcal{D}}{\partial k} \right]^{-1} X dX \quad (26)$$

where  $R$ ,  $L$ ,  $P$  and  $S$  are the Stix parameters defined above,  $\mu$  is the refractive index, and  $J_n$  is the  $n^{th}$  order Bessel function with argument  $k_{\perp}\vartheta_{\perp}/m_e\Omega_e$  ( $n$  being resonant harmonic).

Finally, we average the local diffusion coefficients over the bounce motion. Bounce averaging is obtained by integrating between the mirror points along the bounce motion path and weighting each element of the path by the time that particle spends at that increment.

It can be shown that bounce averaged values for  $D_{\alpha\alpha}$ ,  $D_{\alpha p}$ , and  $D_{pp}$  are given by

$$\langle D_{\alpha_{eq}\alpha_{eq}} \rangle = \frac{1}{\tau_B} \int_0^{\tau_B} D_{\alpha\alpha} \left( \frac{\partial \alpha_{eq}}{\partial \alpha} \right)^2 dt \quad (27)$$

$$\langle D_{\alpha_{eq}p} \rangle = \frac{1}{\tau_B} \int_0^{\tau_B} D_{\alpha p} \left( \frac{\partial \alpha_{eq}}{\partial \alpha} \right) dt \quad (28)$$

$$\langle D_{pp} \rangle = \frac{1}{\tau_B} \int_0^{\tau_B} D_{pp} dt \quad (29)$$

where  $\tau_B$  is particle's bounce period and  $\alpha_{eq}$  is particle's equatorial pitch angle that can be obtained from the local pitch angle using conservation of the first adiabatic invariant (magnetic moment). Changing integration variable to magnetic latitude and using dipole geomagnetic field model yields

$$\langle D_{\alpha_{eq}\alpha_{eq}} \rangle = \frac{1}{T} \int_0^{\lambda_m} D_{\alpha\alpha} \frac{\cos \alpha}{\cos^2 \alpha_{eq}} \cos^7 \lambda d\lambda \quad (30)$$

$$\langle D_{\alpha_{eq}p} \rangle = \frac{1}{T} \int_0^{\lambda_m} D_{\alpha p} \frac{\cos^4 \lambda (1 + 3 \sin^2 \lambda)^{1/4} d\lambda}{\cos \alpha} \quad (31)$$

$$\langle D_{pp} \rangle = \frac{1}{T} \int_0^{\lambda_m} D_{pp} \frac{\cos \lambda (1 + 3 \sin^3 \lambda)^{1/2} d\lambda}{\cos \alpha} \quad (32)$$

where  $\lambda_m$  is the mirror latitude, and  $T(\alpha_{eq})$  gives variation of  $\tau_B$  with the equatorial pitch angle

$$T(\alpha_{eq}) = 1.38 - 0.32 \sin(\alpha_{eq}) \quad (33)$$

Mirror latitude is found by solving

$$x^6 + 3x \cdot \sin^4(\alpha_{eq}) - 4\sin^4(\alpha_{eq}) = 0 \quad (34)$$

where  $x = \cos^2(\lambda_m)$ .

## References

Albert, J. M. (2005), Evaluation of quasi-linear diffusion coefficients for whistler mode waves in a plasma with arbitrary density ratio, *J. Geophys. Res.*, **110**, A03218, doi:10.1029/2004JA010844.

Glauert, S. A., and R. B. Horne (2005), Calculation of pitch angle and energy diffusion coefficients with the PADIE code, *J. Geophys. Res.*, **110**, A04206, doi:10.1029/2004JA010851.

Lawrence R. Lyons (1974). Pitch angle and energy diffusion coefficients from resonant interactions with ion-cyclotron and whistler waves. *Journal of Plasma Physics*, **12**, pp 417-432 doi:10.1017/S002237780002537X

Stix, T. H., *The Theory of Plasma Waves*, McGraw-Hill, New York, 1962.

Swanson D. G., *Plasma Waves*, 2nd Edition, Taylor and Francis, 2003.

Highlights for

**Scattering of Relativistic and Ultra-relativistic Electrons by Obliquely Propagating Electromagnetic Ion Cyclotron Waves**

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1. Presented method can be applied to EMIC waves or any other wave modes;
2. First order resonance dominates the scattering of relativistic electrons up to multi-MeV energies;
3. Ion composition is a crucial parameter for quantifying scattering rates.

