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Glacial isostasy and plate motion

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Abstract

The influence of glacial-isostatic adjustment (GIA) on the motion of tectonic plates is usually neglected. Employing a recently devel-12 oped numerical approach, we examine the effect of glacial loading 13 on the motion of the Earth's tectonic plates where we consider an elastic lithosphere of laterally variable strength and the plates 15 losely connected by low viscous zones. The aim of this paper is 16 to elucidate the physical processes which control the GIA induced 17 horizontal motion and to assess the impact of finite plate-boundary zones. We show that the present-day motion of tectonic plates induced by GIA is at, or above, the order of accuracy of the plate 20 motions determined by very precise GPS observations. Therefore, its contribution should be considered when interpreting the mechanism controlling plate motion.

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$_{4}$ 1 Introduction

Glacial isostatic adjustment (GIA) describes the ongoing readjustment of the Earth arising from the quasi-periodic mass redistribution between continental glaciers, ice sheets and the ocean 27 during glacial cycles, of about 100 kyr in periodicity (Petit et al., 1999). The fluctuation in global ice coverage during the last glaciation, extending over wide areas of North America, northern Europe and Antarctica, involved changes in ice-thicknesses of up 31 to 4km (Denton & Hughes, 1981; Tarasov & Peltier, 2004), with sea level in many areas being over 100 m below the present level 33 (Fleming et al., 1998), causing large deformations that remain active and detectable by modern geodetic techniques (Lambert et al., 35 2001; Pagiatakis & Salib, 2003). Therefore, the ongoing GIA is usually taken into account when interpreting, for example, presentday secular trends in the Earth's gravity field and rotation axis, and long-term variations in sea level. GIA-induced horizontal mo-39 tion inferred by space geodetic techniques has primarily been used in studies dealing with formerly glaciated regions such as Scandi-41 navia or North America (James & Morgan, 1990; Milne et al., 2001; Sella et al., 2007). However, the present-day accuracy of such tech-43 niques and improved modelling of GIA enables us to study the induced horizontal motion globally and to assess its effect on the motion of all of the Earth's major tectonic plates. The motion of the plates is predominantly driven by convective 47 processes in the Earth's mantle with characteristic timescales of

1–100 Myr (Knopoff, 1964; Richter, 1973; Mckenzie et al., 1974).

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There are two principal approaches for quantifying plate velocities: Geologically-based models like NUVEL-1A (DeMets et al., 1994), 51 that rely on the analysis of geomagnetic reversals and other geo-52 logical and seismic data, represent plate velocities averaged over 53 the last million years. In contrast, geodetically-based models like REVEL (Sella et al., 2002) rely on the analysis of Global Positioning System (GPS) observations. The latter represent plate velocities averaged over the last decade and a half (A.D. 1993–2008). The GIA process, with a period of 0.1 Myr, therefore, represents an undetectable transient perturbation of the geologically-inferred model, but a detectable component in the geodetically-inferred secular trends used to construct the model. The usual way to elim-61 inate the GIA-induced horizontal motion from GPS observations is to exclude those GPS observations that are recorded in areas 63 near formerly glaciated regions (Sella et al., 2002). The argument 64 for such a procedure is that the GIA-induced horizontal displacements are about one order of magnitude smaller than the associated vertical displacements and are concentrated to the formerly 67 glaciated regions (Peltier et al., 1986; Lambert et al., 2001). This approach does not, however, remove the effect of GIA from GPS 69 observations completely, since GIA-induced horizontal motion is 70 distributed over the entire globe with amplitudes of the order of one mm/yr (James & Morgan, 1990; Wang & Wu, 2006a). Since 72 the plate velocities are of the order of a few cm/yr, the GIA-induced horizontal motion generally contributes less than 10% of the total observed signal (Sella et al., 2002). Despite such a small amplitude, this signal is present in the extremely precise continuous tracking

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GIA induced horizontal motions are much more sensitive to

mechanical properties that lie above the mantle, and, as a

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data of GPS (Calais et al., 2006).

consequence, to lateral variations in lithosphere strength, than 80 are the GIA-induced vertical motions. Due to computational rigor inherent to the modelling of viscoelastic lateral hetero-82 geneities, various approximations, such as introducing some sym-83 metries like a 2-D structure for a half-space model (Sabadini et al., 1986; Kaufmann et al., 1997) or axial symmetry for a spherical earth model (e.g. Wu & van der Wal, 2003; Martinec & Wolf, 86 2005; Klemann et al., 2007) have been employed in past work. 87 Half-space models that consider structural features in 3-D are 88 constrained to regional GIA due to the neglection of sphericity: Europe (Marotta & Sabadini, 2004; Kaufmann & Wu, 1998, 90 2002; Steffen et al., 2006) Laurentia (Wu, 2005) and Antarctica 91 (Kaufmann et al., 2005). More recently, global approaches use spherical finite element 93 models with a grid strategy that incorporates lateral variations in 94 mantle viscosity and lithospheric strengths (Wang & Wu, 2006a,b) 95 or plate boundaries (Latychev et al., 2005; Whitehouse et al., 2006). Predictions from such models form a more realistic basis for comparison to the observed global horizontal motions. In this study, 98 we specifically demonstrate the importance of plate boundaries on 99 the GIA-induced horizontal motion and, therefore, better decipher 100 how GIA places an imprint on observed plate-tectonic motion.

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2 Modelling of GIA-induced motion

To examine how the GIA process effects plate motion, we approxi-103 mate the Earth by a self-gravitating sphere with a linear viscoelastic 104 rheology. The upper and lower mantles have uniform viscosities of 105 $5 \times 10^{20} \,\mathrm{Pa}\,\mathrm{s}$ and $10^{22} \,\mathrm{Pa}\,\mathrm{s}$, respectively; the model accommodates a fluid core. The elastic part of viscoelastic rheology is considered 107 incompressible with the elastic shear modulus and mass-density de-108 fined by the Preliminary Reference Earth Model. The spectral finite 109 element method developed by Martinec (2000) is applied which allows strong lateral viscosity contrasts to be considered, including 111 lateral variations in lithosphere thickness. The resolution in hori-112 zontal directions is represented by spherical harmonics up to degree 113 170 which corresponds to a spatial resolution of about 120 km, and vertically by finite elements of 5 km in the lithosphere and upper mantle. 116 We define the base of the elastic lithosphere to be at the depth of 117 the 1100 °C isotherm (Figure 1), assuming to be the characteristic temperature below which the Earth's material is dominated by elas-119

the 1100 °C isotherm (Figure 1), assuming to be the characteristic temperature below which the Earth's material is dominated by elasticity for strains having time scales shorter than 0.1 Myr. The thickness of continental lithosphere is directly derived from a global thermal model, where we used the temperature profiles on a 5° × 5° grid provided by Artemieva (2006). The thermal structure of the oceanic lithosphere is derived from its age (Müller et al., 1997) by applying a simple plate-cooling model. The 1100 °C isotherm is calculated by assuming heat conduction inside the oceanic plate between the mantle of 1300 °C and surface of 0 °C (e.g. Turcotte & Schubert,

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in the decoupling of shear stresses between the plates, the main 120 plate boundaries taken from Bird (2003) are approximated as 200-130 km narrow zones of viscoelastic material, and further approximated 131 assuming the same viscosity as the upper mantle (Figure 1). The surface loading applied considers the main areas of glacia-133 tion (Laurentide, Greenland, Fennoscandia and Antarctica) over 134 the last glacial cycle. The spatio-temporal evolution of the Lauren-135 tide and Greenland ice sheets are described by the standard ICE3G model (Tushingham & Peltier, 1991) for Laurentia, the Fennoscan-137 dian ice sheet by the SCAN model (Lambeck et al., 1998), and 138 the Antarctic ice sheet by the ICE3G model. This model gives a 139 global sea-level fall of -105 m at the Last Glacial Maximum (LGM) (Hagedoorn et al., 2006). The mass conservation principle is ap-141 plied for water exchange between ocean and ice sheets during the 142 Pleistocene glaciation where, for simplicity, we uniformally reduced the sea-level in accord with continental ice volume change assuming fixed coast lines. 145

2002, p. 161). In order to consider the role of plate boundaries

3 Influence of tectonic plates on GIAinduced motion

The spatio-temporal changes in surface-mass load during the Pleis-148 tocene glaciation induce vertical mantle-material flow resulting in 149 present-day surface-uplift rates of up to 20 mm/yr. This mate-150 rial flow is accompanied by horizontal material transport inside the 151

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mantle (James & Morgan, 1990). Since the viscosity of the upper mantle is significantly lower than that of the lower mantle, the horizontal flow is concentrated in the layer between a highly viscous lower mantle (below 670 km depth) and the elastic lithosphere.

Figure 2 shows the horizontal and vertical present-day veloci-156 ties induced by the assumed late-Pleistocene glaciation cycle. The 157 present-day GIA-induced horizontal flow displays a hemispheric 158 pattern and, generally, the geometry of this flow means that it 159 orients toward the areas of present-day uplift. The flow of the northern hemisphere is induced by the formerly glaciated regions 161 of Fennoscandia, Greenland and Laurentia and is directed north-162 ward, while the flow of the southern hemisphere, induced by the 163 glaciation of Antarctica is directed southward (e.g. Wang & Wu, 2006a). Furthermore, abrupt changes in the horizontal velocities 165 appear at several plate boundaries, for example at the boundary 166 between the Pacific and North American plates, the Indian and 167 Antarctic plates and between the Australian and Antarctic plates (Figure 2, right). These features do not appear if a model with a 169 uniform lithosphere thickness of 100 km is assumed (Figure 2, left). 170

More striking is to consider the spatial gradient of the displacement rate, $\nabla \dot{u}$. Regarding the strain rate, $\dot{\epsilon}$, we show in Figure 3 the divergence of the surface components and its second invariant:

$$\operatorname{div}_{h} \dot{\boldsymbol{u}} = \dot{\epsilon}_{\theta\theta} + \dot{\epsilon}_{\phi\phi} , \qquad (1)$$

$$\dot{\epsilon}_{h,II} = \sqrt{\dot{\epsilon}_{\theta\theta}^2 + \dot{\epsilon}_{\phi\phi}^2 + 2\dot{\epsilon}_{\theta\phi}^2} . \tag{2}$$

Regarding the vorticity, ω , it is important to monitor its radial

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component and its absolute value:

$$\omega_r = \boldsymbol{e}_r \cdot (\boldsymbol{\nabla} \times \dot{\boldsymbol{u}}) \,, \tag{3}$$

$$\omega = |\nabla \times \dot{\boldsymbol{u}}|. \tag{4}$$

Apart from the last quantity, wherein radial gradients are also in-

volved, the quantities (Eqs. 1-3) are defined by the surface com-

ponents of the displacement field. First, it should be mentioned that positive divergence means dilation, whereas negative diver-174 gence means compression of the surface plane. For the radial vor-175 ticity, a negative sign indicates a dextral rotation of an idealized 176 surface element whereas a positive sign indicates a sinistral rotation. The other two components are positive by definition. The diver-178 gence shows the known features of dilation of the formerly glaciated 179 regions which are surrounded by areas of compression as direct con-180 sequence of the uplift pattern. In addition, the plate boundaries 181 around Antarctica mainly show dilation whereas the Mid-Atlantic 182 ridge and the Arctic Lomonosov ridge show compression. 183 The second invariant of the surface strain rate shows a sim-184 ilar pattern like the divergence, only the plate boundaries are 185 more pronounced in this plot. In comparison to the values in-186 ferred for plate tectonic motion exceeding thousands of $10^{-9}/\text{yr}$ 187 Kreemer et al. (2003), the GIA contribution represents a small per-188 turbation. The radial vorticity directly resembles the toroidal mo-189 tion, which does not appear for a 1-D earth model. Such a model 190 if loaded by a surface pressure will only experience spheroidal mo-191 tions and, therefore, the radial vorticity is zero. Here, the largest 192 amplitudes appear at the plate boundaries while vorticity in the

regions of large uplift are negligible. The total vorticity is almost a factor of 10 larger than the radial component alone, as it collects 195 information pertaining to all components of the deformation-rate 196 tensor including the tilting of the surface, which tends to follow 197 the migrating peripheral bulge (James & Bent, 1994). Therefore, the largest amplitudes surround the former glacial maxima, where 190 the strongest uplift gradients appear (compare Figure 2). This fea-200 ture is generally common to all GIA models, appearing in the 1-D 201 earth model as well. The plate boundaries showing gradients in horizontal motion are diminished. 203

The degree variance spectra shown in Figure 4 provide insight into the partitioning of surface motion between the spheroidal and toroidal component (e.g. Forte & Peltier, 1987). Shown are the degree variances

$$\sigma_U^2(l) = \sum_{m=0}^l U_{lm} U_{lm}^* \tag{5}$$

$$\sigma_V^2(l) = \sum_{m=0}^l V_{lm} V_{lm}^* \tag{6}$$

$$\sigma_W^2(l) = \sum_{m=0}^l W_{lm} W_{lm}^* \tag{7}$$

of the vertical and horizontal spheroidal component, U_{lm} , V_{lm} , and the toroidal component, W_{lm} , of the velocity field at the surface as function of Legendre degree, l, and order m (see Eq. 8). Asterix, , denotes the conjugate complex. Proportional to these quantities are the surface divergence and radial vorticity which are scaled by l(l+1)/R with respect to V_{lm} and W_{lm} (e.g. Čadek & Ricard, 1992). The variances show that the toroidal motion reaches the amplitude of the spheroidal V component for $l \geq 5$. An equipartitioning of

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kinetic energy also appears in the motion of the tectonic plates (e.g. 212 Hager & O'Connell, 1978; Cadek & Ricard, 1992; Bercovici, 1995) 213 where, due to the existence of lithospheric plates the efficiency of 214 convection as a poloidal motion is enhanced by reducing the overall 215 dissipation in the system (Bercovici, 2003). Here, the equiparti-216 tioning gains strength at degrees larger than 3, whereas for longer 217 wavelengths the spheroidal motion dominates. An exposition on 218 this topic is beyond the scope of this paper, but the controlling me-219 chanics are likely to have strong analogies to the convective systems 220 described by Bercovici (2003). The vertical motion, dotted lines, 221 is dominated by buoyant forces that cause uplift in the previously 222 glaciated areas and does not noticably change if we consider lateral 223 lithospheric variations. This is also shown in Figure 2 where no 224 broad scale variations in the uplift process are predicted. 225 In order to better understand the dynamics appearing in these 226 227

numerical solutions, the material flow pattern along a cross-section passing through the main areas of former/current glaciation in the 228 northern hemisphere is shown in Figure 5. The lateral flow shown 220 in the upper panels is mainly confined to the upper mantle and 230 flow in the lower mantle is only visible below Laurentia, a feature which confirms the fact that the regional GIA is addition-232 ally affected by the lower mantle (Wolf et al., 2006) whereas for 233 Fennoscandia motion is, relatively, confined to upper mantle pro-234 cesses (Wieczerkowski et al., 1999; Martinec & Wolf, 2005). The 235 lateral mantle-material flow manifests a shear traction on the over-236 lying plates. The traction, consequently, pulls the tectonic plates 237 toward the former glaciation centres of Laurentia, Greenland and 238

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Fennoscandia. The role of the mid-Atlantic ridge as a soft plate 239 boundary is complex because ongoing GIA is occurring on both 240 sides of the cross-section. Figure 6 more closely examines the upper 241 mantle for the same profile. In the right panel of Figure 6 the flow 242 pattern for a 1-D earth model is shown for comparison. Although the flow shown in the upper panels of Figure 6 are similar (which 244 corresponds to the similar flow around the previously glaciated re-245 gions at the surface, see Figure 2), it is evident that the lateral 246 flow pattern tracks the viscous portion of the upper mantle and is strikingly discontinuous at the lithospheric base. The discontinu-248 ity in the motion becomes more evident in the vorticity (Figure 6, 249 lower panels), as the intensity of the parameter makes a step-like 250 change across the rheological boundaries. So, the vector component 251 normal to the plane (off-plane) describes the rotation of idealized 252 rigid particles in the plane, sinistral in red and dextral in blue. The 253 vertical component of the vector shown here describes the radial 254 vorticity as in Figure 2 and the horizontal in-plane component ex-255 hibits a tilting of the particle out of the plane. For the 1-D earth 256 model it is evident that within the earth radial vorticity is absent. 257 Again, recall that this component is solely described by a toroidal field. The pattern of the horizontal components for the two models 259 are similar in this portion of the spherical earth, especially in the 260 lithosphere. The amplitude of vorticity is largest near the bound-261 aries of the upper mantle, whereas they are almost constant inside 262 the elastic lithosphere, a consequence of near-unidirectional motion 263 in the lithosphere. Due to the direction of the profile normal to the 264 load margins, the in-plane components of vorticity are quite small 265

in comparison to the normal component. Inside the lithosphere, the

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normal component of rotation resembles the tilting of the surface 267 according to the uplift gradient with distance from the former cen-268 tres of glaciation. The rotation is sinistral to the right and dextral 269 to the left of the respective uplift centres. The horizontal in-plane component is one order of magnitude smaller than the off-plane 271 component. The vertical component, ω_r , becomes evident in the 272 upper mantle where the lithosphere base shows a significant slope, 273 e.g. at 35° on the profile. Here, ω_r is also comparable to the hori-274 zontal in-plane components and the discontinuity in the off-plane 275 component across the lithospheric base is also evident. 276 In order to resolve the effect of a low-viscosity plate boundary on 277 the GIA-induced flow, we choose a second cross-section (Figure 7) 278 that passes through the Southeast-Indian ridge between the Aus-270 tralian and Antarctic plates. For this cross section it is striking 280 that the GIA-induced horizontal velocities are relatively subdued 281 north of the Southeast-Indian ridge and abruptly promoted south 282 of the ridge. In the vicinity of the ridge, the sub-surface horizontal 283 velocities increase toward the surface, approximating the behaviour 284 of material flow for the case of a free-slip boundary condition. 285 The amplitude of the horizontal-velocity contrast is 2 mm/yr (Fig-286 ure 8). Although this is only a small part of the observed 70 mm/yr 287 spreading rate of this ridge, the presence of the ridge significantly 288 changes the material flow pattern in comparison to the case with-289 out a plate boundary (Figure 8, bottom panel). Traversing across the Southeast-Indian ridge, the abrupt change in horizontal veloc-291 ity is predicted in the 3-D earth model, while the change is much 292

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smoother in the prediction of the corresponding laterally homogeneous model. The top panel shows that the uplift rate is relatively little influenced by the presence of the ridge.

²⁹⁶ 4 Influence of GIA on plate motion

The motion of tectonic plates is usually represented by a rigid 297 motion around a rotation pole. In order to compare the GIA-298 induced motion with the geodetically inferred plate motion in 299 Altamimi et al. (2007), we calculate the incremental rotation of in-300 dividual tectonic plates from the GIA induced horizontal velocities at the sites provided in their Table 7 applying a least squares fit 302 (Appendix, Eq. 10). Table 1 shows the rotation poles and veloci-303 ties given in Altamimi et al. (2007, Table 8) with those determined 304 for the GIA-induced motion. For the chosen models, the largest 305 rotation is exhibited in North America, as is shown in Figure 9. 306 The vectors west of the North American continental divide show 307 the considerable influence of the modeled plate boundaries (Fig-308 ure 2). The latter facilitates an enhanced north-northwesterly and, consequently, a stronger rotation of the continent. This numerical 310 experiment suggests that models of the geophysical forces that con-311 trol North American plate motion might have to further consider 312 GIA induced motions before advancing dynamical models that ex-313 plain crustal motion at the 1–2 mm/yr level. 314 The model reference system is defined by fixing the center of 315 mass and that no net rotation of the surface is allowed (Appendix, 316 Eq. 9). In order to compare the GIA induced poles with those of

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the observed plate motions, we subtract the respective GIA induced rotation vectors from the ITRF rotation vectors, $\Omega_{ITRF} - \Omega_{GIA}$ and 310 consider the resulting changes in longitude, latitude and rotation 320 velocity in the 4th to 7th column of Table 1. The rotational velocity 321 of Antarctica and Europe are slightly reduced, whereas those of the 322 Pacific and South American plate are uneffected by GIA. The other 323 plates listed are enhanced by the effect of GIA. The shifts in the 324 location of the rotation poles are at the level of about 1 degree. 325 From the intercomparison of the perturbations considered here, we conclude that there is a ubiquitous influence of rhelogically soft 327 plate boundaries on plate rotations, intraplate horizontal velocities 328 and on the generation of finite global toroidal deformation field. 329 Each of these features may be detectable by modern space geodetic 330 techniques. 331

5 Summary

This study shows that the GIA-induced horizontal velocities are 333 globally of the order of 1 mm/yr, and are strongly influenced by lateral variations in lithosphere thickness and the presence of tectonic 335 plate boundaries. The current level of accuracy in the detection of 336 present-day horizontal motions approaches 1 mm/yr using space-337 geodetic techniques such as GPS. Therefore, GIA signals must be considered for modeling the mantle-wide dynamical causes of plate 339 motion and of intraplate deformations. The present study essen-340 tially considers only one element of lateral variability, lithosphere 341 thickness. Other viscosity variations in the upper mantle have to

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be considered as shown by Wang & Wu (2006a) and Paulson et al. (2007). It should be noted that this study probes but one subset of 344 the parameter space for predicting GIA-induced surface horizontal 345 motion. Our results indicate the global importance of ongoing GIA in the interpretation of GPS time series for sites far away from the formerly glaciated areas, as these are non-negligible with respect 348 to state-of-the-art kinematic and dynamic models of plate motion. 349 The global consequences of GIA are amplified by the existence of 350 soft plate boundaries. Further amplification could occur in model predictions that incorporate lateral variations of lithosphere thick-352 ness and variability in the mantle viscosity. 353

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A Appendix

The spectral representation of the displacement field follows the no-

tations in Martinec (2000), and splits the motion into a spheroidal

part represented by U_{jm} and V_{jm} and a toroidal part W_{jm} :

$$\boldsymbol{u}(r,\Omega) = \sum_{j=0}^{j_{\text{max}}} \sum_{m=0}^{j} \left[U_{jm}(r) \, \boldsymbol{e}_r \, Y_{jm}(\Omega) + V_{jm}(r) \, \boldsymbol{\nabla}_{\Omega} \, Y_{jm}(\Omega) + W_{jm}(r) \, \boldsymbol{L}_{\Omega} \, Y_{jm}(\Omega) \right]$$
(8)

where r is the radial distance and $\Omega=(\theta,\phi)$ are the colatitude and longitude. The summations extend in Legendre degrees, j, from 0 to the maximum degree, j_{max} , considered and in order m from 0 to j. $Y_{jm}(\Omega)$ are the normalized scalar spherical harmonics (e.g Varshalovich $et\ al.$, 1988). ∇_{Ω} denotes the angular part of the gradient operator and $\boldsymbol{L}_{\Omega}=\boldsymbol{e}_r\times\nabla_{\Omega}$ stands for the angular part of the angular momentum operator.

The conditions of center of mass invariance and no net rotation are considered by assuming that for each epoch

$$\int_{V_0} \mathbf{u} \, \rho \, dV = 0$$

$$\int_{\Omega_0} \mathbf{u} \times \mathbf{e}_r \, d\Omega = 0$$
(9)

 $_{\text{544}}$ holds, where V_0 is the earth's volume, Ω_0 is the earth's surface,

ho is the material density, $m{e}_r$ is the radial unit vector and $d\Omega =$

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 $\sin \theta \, d\theta \, d\phi$.

In order to determine the rigid-plate rotation, we solve

$$\frac{1}{R} \begin{pmatrix} u_{\theta}(\theta_i, \phi_i) \\ u_{\phi}(\theta_i, \phi_i) \end{pmatrix} = \begin{pmatrix} -\sin \phi_i & \cos \phi_i & 0 \\ -\cos \theta_i & \cos \phi_i & -\cos \theta_i & \sin \phi_i & \sin \theta_i \end{pmatrix} \begin{pmatrix} \omega_x \\ \omega_y \\ \omega_z \end{pmatrix}$$
(10)

where R is the earth's radius. The displacement, u_{θ} , u_{ϕ} , at colatitude θ_i and longitude ϕ_i is given in polar coordinates originating from an incremental rotation vector $\boldsymbol{\omega} = \omega_x \, \boldsymbol{e}_x + \omega_y \, \boldsymbol{e}_y + \omega_z \, \boldsymbol{e}_z$ in Cartesian coordinates. The orientation of the Cartesian system with respect to the polar coordinates, (r, θ, ϕ) , is $\boldsymbol{e}_x = (1, \pi/2, 0)$, $\boldsymbol{e}_y = (1, \pi/2, \pi/2)$ and $\boldsymbol{e}_z = (1, 0, 0)$. The rotation vector is determined from the given displacement rates of the respective plate by applying a least-squares fit. The rotation pole then follows:

$$\Omega = \sqrt{\omega_x^2 + \omega_y^2 + \omega_z^2},$$

$$\Theta = \arccos \frac{\omega_z}{\Omega},$$

$$\Phi = \arctan \frac{\omega_y}{\omega_x}.$$
(11)

Its geographical coordinates are Φ = longitude and 90° - Θ = latitude. Figure 9 shows as an example of Table 1 the predicted contribution to the plate motion of North America inferred from modelled velocities at the ITRF stations.

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	ITRF/GIA			error/corr.		
Plate	Lon. (°E)	Lat. (°N)	vel. (°/Myr)	Lon. (°E)	Lat. (°N)	vel. (°/Myr)
Antarctica	-125.3	59.8	0.223	± 1.7	±0.9	± 0.007
	135.5	34.0	0.005	2.2	-0.6	-0.002
	150.3	28.8	0.008	3.6	-0.4	-0.004
Australia	37.4	32.4	0.628	± 0.4	± 0.3	± 0.003
	-118.3	-3.9	0.009	0.4	-0.3	0.007
	-105.8	1.9	0.006	0.4	-0.2	0.004
Eurasia	-96.0	56.3	0.261	±1.0	± 0.5	± 0.003
	-52.4	-16.3	0.009	-2.4	1.4	-0.001
	-56.7	-13.0	0.006	-1.6	1.0	-0.001
N. America	-87.4	-4.3	0.192	± 0.6	± 0.9	± 0.002
	153.7	3.5	0.004	1.1	-0.0	0.002
	136.5	-39.8	0.015	2.3	2.9	0.008
Pacific	112.9	-62.6	0.682	± 0.7	± 0.2	± 0.004
	52.0	36.3	0.002	0.3	-0.1	0.001
	9.5	11.4	0.003	0.5	-0.0	0.001
S. America	-129.6	-16.8	0.121	± 2.1	±1.6	± 0.003
	-19.9	-18.4	0.008	-3.6	1.6	0.002
	-9.9	-27.0	0.006	-2.1	1.5	0.002

Table 1: Absolute rotation poles for different plates from ITRF2005 (1st line of respective plate), GIA-induced component for 1-D earth model (2nd line) and for 3-D earth model (3rd line). The 4th to 7th collumn presents the respective errors in the determination of the ITRF poles and below the corrections to the ITRF poles if they are reduced by the GIA induced components. The table is designed to show the influence of a plate structure, not to give a realistic prediction

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Figure 1: Mosaic of elastic lithosphere thickness defined by the

depth of the 1100 °C isotherm inferred from the thermal struc-

ture of continental lithosphere (Artemieva, 2006) and age structure

of oceanic lithosphere (Müller et al., 1997). Plate boundaries (pur-

ple lines) are approximated by 200 km wide zones with the same

viscosity as the upper-mantle.

Figure 2: Prediction of present-day surface velocities induced by the

last glacial cycle with a laterally homogeneous lithosphere structure

(left) and the considered 3-D earth model (right). Purple areas

denote the considered plate boundaries.

Figure 3: Spatial surface fields related to the deformation rate ac-

cording to (1)-(4): divergence in surface plane, second invariant of

surface strain rate, radial component of rotation rate and absolute

value of rotation rate.

Figure 4: Degree variance spectra of spheroidal and toroidal surface motion (left) and the derived quantities of surface divergence and radial vorticity (right). The spheroidal velocities are splitted into its horizontal contribution (solid lines) and its radial contribution (dotted lines).

Figure 5: Present-day GIA-induced velocities along a cross-section passing through areas that contain the Laurentide (LIS), Greenland (GIS) and Fennoscandian (FIS) ice sheets, respectively. The upper panel shows the ice-sheet extent (blue line) and thickness at the last glacial maximum relative to present day. The profile of the cross section is marked in red. The plate boundaries are plotted in purple. The lower panel shows the present-day GIA-induced velocities in the plane of the cross section (vectors) and perpendicular to the plane (blue indicates a flow directed into the plane). The inverted triangle denotes the position of the mid-Atlantic ridge, dashed red lines denote the base of the effective elastic lithosphere and the boundary between upper and lower mantle.

Figure 6: Present day velocity (upper panels) and vorticity (lower panels) on profile I over a limited mantle/lithosphere range. Fields for a 1-D earth model with 100 km lithosphere thickness are presented on the right. The red lines indicate the top and base of the viscoelastic upper mantle. At the top, the inverted triangle shows the position of the mid-Atlantic ridge and the rectangulars the extension of the glacial loads at LGM.

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Figure 7: Present-day GIA-induced velocities along a cross-section

beginning in Australia, passing through the Southeast-Indian ridge

and ending in the Antarctic ice sheet (AIS). For details see Figure

5.

Figure 8: Horizontal and vertical present-day GIA-induced surface

velocities along the cross-section shown in Figure 7. Velocities

for the cases where lateral variations in lithosphere thickness and

the presence of the Southeast-Indian ridge are considered are plot-

ted with solid lines (3-D) and where they are not considered with

dashed lines (1-D). Thick and thin lines in the bottom panel de-

note the horizontal motion in- and perpendicular to the plane of

the cross section, respectively. The inverted triangle denotes the

position of the Southeast-Indian ridge.

Figure 9: Horizontal present-day GIA-induced surface velocities for

rigid plate sites of North America considered in the ITRF2005

(Altamimi et al., 2007) for the 1-D earth model (left) and 3-D earth

model (right) of teh analysis presented in the paper. Black arrows

denote the induced surface velocities and red arrows the velocities

due to a rigid rotation with parameters given at the top of each

plot.























