

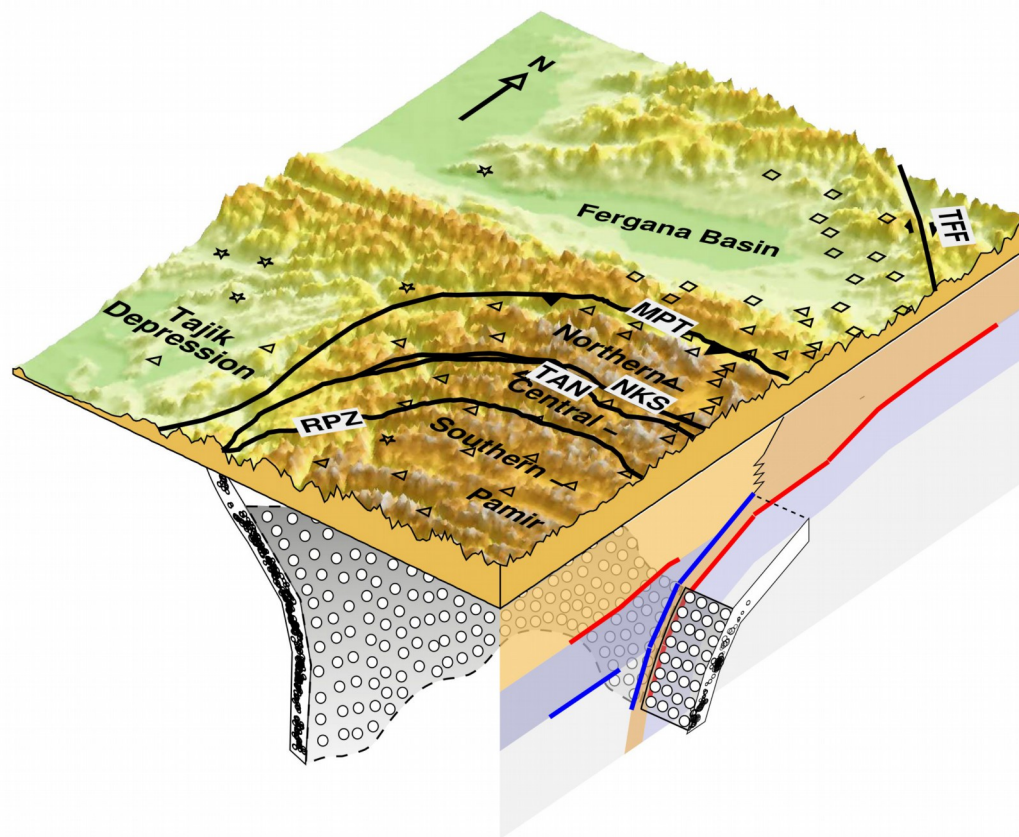


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## Graphical Abstract



### Highlights

We operated a modern seismic array in the Pamir, a key area for Indo-Asian collision.

We constructed receiver function cross sections traversing Tien Shan and Pamir.

We used a modified CCP stacking to image strongly dipping interfaces.

We observed subduction of Eurasian continental crust beneath the Pamir.

A south-dipping low-velocity zone coincides with the intermediate-depth seismicity.

1 **Seismic imaging of subducting continental lower crust beneath the Pamir**

2

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15

16 **ABSTRACT**

17 **Exhumation of ultra-high pressure metamorphic rocks testifies that the continental**  
18 **crust can subduct to significant depth into the mantle despite its buoyancy. However,**  
19 **direct observation of ongoing subduction of continental crust is rare. The Pamir is**  
20 **regarded as a possible place of active continental subduction because of the**  
21 **intermediate-depth seismicity, crustal xenoliths and estimates of crustal shortening**  
22 **versus convergence rates. Here we present for the first time receiver function images**  
23 **from a passive-source seismic array traversing the Tien Shan and the Pamir plateau**  
24 **showing southward subduction of Eurasian continental crust. In the eastern Pamir, we**

25 observe a southerly dipping 10-15 km thick low-velocity zone (LVZ) that extends from  
26 50 km depth near the base of the crust to more than 150 km depth with a dip angle  
27 increasing to subvertical. While the upper- and mid- crustal material seems to be  
28 shortened and incorporated into the Pamir, the lower Eurasian crust detaches and  
29 subducts. In its deeper part ( $> 80$  km) the LVZ envelopes the intermediate-depth  
30 earthquakes. Our observations imply that the complete arcuate intermediate depth  
31 seismic zone beneath the Pamir traces a slab of subducting Eurasian continental lower  
32 crust.

33

#### 34 1. Introduction

35 The Pamir, situated north of the western Himalayan syntaxis, is thought to consist of the same  
36 collage of continental terranes as Tibet, that were progressively accreted to Eurasia prior to  
37 the Indian-Eurasian collision around 50 million years ago (Schwab et al., 2004). To regard  
38 the Pamir as a place where subduction of continental crust may occur is motivated  
39 tectonically, since the correlation of the Pamir and Tibetan belts and sutures indicates that  
40 structures in the Pamir were translated northwards by around 300 km with respect to Tibet.  
41 Thereby the crust connecting the Tajik and Tarim basins disappeared (Burtman and Molnar,  
42 1993; Robinson et al., 2004), leading to the hypothesis that the Eurasian continental  
43 lithosphere is subducted along the Pamir's deformation front, which is today formed by the  
44 Main Pamir Thrust (MPT) (Burtman and Molnar, 1993; Hamburger et al., 1992; Sobel et al.,  
45 2013). Today, 10-15 mm/year of concentrated shortening occurs across the MPT (Zubovich et  
46 al., 2010), which is about one third of the total present-day Indian-Eurasian convergence rate.  
47 In the Pamir, the total amount of crustal shortening that occurred during the Cenozoic is even  
48 higher than in Tibet, since a similar amount of total convergence has been accommodated

49 over a much smaller distance. Considering the relation of present to pre-Cenozoic crustal  
50 thicknesses, these extremely high amounts of crustal shortening in the Pamir yield realistic  
51 scenarios that require crustal excess (Schmidt et al., 2011). Besides, an indication for  
52 subduction of continental crust in this region was given by Roecker (1982), who first  
53 observed a low-velocity zone in the upper mantle dipping north beneath the Hindu Kush by  
54 local earthquake tomography. The most striking observation indicating continental subduction  
55 in the Pamir-Hindu Kush region is the vigorous intermediate-depth seismicity occurring from  
56 70 to 250 km depth in this intra-continental setting. In map view, this seismicity forms a  
57 narrow S-shaped band of approximately 450 km length from the Hindu Kush in northeastern  
58 Afghanistan to the eastern Pamir (Fig. 1). The hypocentres of these mantle earthquakes form  
59 two separated Wadati-Benioff zones, one beneath the Pamir and one beneath the Hindu Kush  
60 (Burtman and Molnar, 1993; Fan et al., 1994; Negredo et al., 2007; Pegler, and Das, 1998).  
61 N-S oriented cross sections reveal opposite dips for both zones. There is a long standing  
62 debate whether the geometry of the seismic zone results from a single originally northward  
63 dipping subduction interface, which was contorted and overturned under the eastern Pamir  
64 (Billington et al., 1977; Pegler, and Das, 1998; Pavlis and Das, 2000), or if it is the result of  
65 two subduction zones, one dipping to the north beneath the Hindu Kush and one dipping to  
66 the south beneath the Pamir (Burtman and Molnar, 1993; Chatelain et al., 1980; Fan et al.,  
67 1994; Negredo et al., 2007). Recent high resolution earthquake locations from the local  
68 TIPAGE network (see next paragraph) show the seismicity in the Pamir as a curvilinear arc  
69 dipping to the south in the eastern Pamir and bending to an eastward dipping direction  
70 beneath the south-western Pamir, while the separated seismic zone beneath the Hindu Kush is  
71 a more complex structure striking east-west and dipping subvertically north to north-west  
72 (Sippl et al., 2013).

## 73 **2. TIPAGE seismic experiment**

74 The seismic data analyzed here were collected by a temporary experiment operated from 2008  
75 to 2010 (Fig. 1). The experiment is the seismological component of the multi-disciplinary  
76 **Tien Shan Pamir Geodynamic Program (TIPAGE)** (Mechie et al., 2012) and ran in two  
77 configurations. In the first year, 24 broadband stations were installed in the eastern Pamir  
78 forming a 350 km long north-south linear array from southern Kyrgyzstan to the Tajik-Afghan  
79 border with an average station spacing of 15 km. Additionally, 8 broadband and 8 short period  
80 stations were distributed in an areal network covering the whole Pamir. In the second year the  
81 network geometry was changed. 17 stations from the eastern Pamir linear array were  
82 relocated to densify the areal network in order to achieve an equidistant mesh with a station  
83 spacing of approximately 40 km. The areal array helped to accurately locate the local  
84 earthquakes (Sippl et al., 2013). We also included 5 permanent stations in western Tajikistan,  
85 operated by PMP International, Dushanbe. In the present work we concentrate on the receiver  
86 function analysis along the linear array. The profile was prolonged northwards with 10  
87 stations from the FERGANA seismic experiment, running from 2009 to 2010 in the Fergana  
88 Valley (Haberland et al., 2011).

89

## 90 **3. Receiver function processing**

91 The receiver function (RF) method is a widely used seismological tool to image crustal and  
92 upper mantle seismic discontinuities underneath seismic stations (Kind et al., 2012;  
93 Rondenay, 2009; Bostock, 2007). This technique identifies the discontinuities, e.g. the Moho,  
94 by detecting mode conversion from compressional (P) to shear (S) waves. The time delay  
95 between the converted wave (Ps) and the mother phase (P) determines the conversion depth.  
96 The amplitude and phase of the Ps conversions reflects the nature of the discontinuity, such as

107 the sign, strength and sharpness of the velocity contrast. The kernel of the receiver function  
108 algorithm is composed of two main steps, a coordinate rotation that isolates the Ps converted  
109 waves from the P wave; and a deconvolution that removes the source-side complications and  
110 the propagation effects. Further processing steps such as moveout correction, common-  
111 conversion-point (CCP) stack and migration enable the construction of station stack averages  
112 and receiver function profiles for dense seismic arrays (Bostock et al., 2001; Kind et al., 2002;  
113 Yuan et al., 1997). There are different flavors in RF processing in terms of coordinate rotation  
114 and deconvolution. Here we calculate receiver functions in the manner described by  
115 Yuan et al. (1997). We used high quality teleseismic earthquake records originating at  
116 epicentral distances between  $30^\circ$  and  $90^\circ$  (Fig. 1). All records were bandpass filtered (0.5-30  
117 s). For event selection records of all events with magnitudes bigger than 5.5 were inspected  
118 manually. If needed, additional high pass filters (20 s ,15 s ,10 s ,7.5 s or 5 s) were applied.  
119 We rotate the originally recorded Z-N-E components to the L-Q-T coordinate system using  
120 theoretical back azimuths and incidence angles and deconvolve the L component from the Q  
and T components by a time-domain Wiener filtering approach. The Q and T components  
correspond to the polarizations of the SV and SH component, respectively. The deconvolved  
Q and T components are called Q- and T- receiver functions (QRF and TRF). For horizontally  
layered isotropic media the Ps converted waves are polarized in the radial direction and can be  
analyzed by the QRF. In the case of dipping interfaces or anisotropy significant converted  
energy will fall in the transverse direction and can be detected by the TRF. Using the obtained  
RFs, we then create a receiver function depth section (Fig. 2) by a CCP stack approach (Kind  
et al., 2002). The approach divides the depth section into a fine grid and stacks at each grid  
cell those samples of the receiver functions that have originated by Ps conversion in the same  
grid cell. The conversion points are calculated by ray tracing through a reference velocity

121 model.

122

#### 123 **4. Depth sections along the N-S profile**

##### 124 **4.1 CCP stacking for horizontal converters**

125 Figure 2 shows the CCP cross section constructed along the north-south directed linear array.  
126 Beneath the Pamir we observe an ~70 km thick crust, whereas in the north towards the  
127 Fergana Valley crustal thickness decreases rapidly to ~45 km beneath the northern edge of the  
128 southern Tien Shan. A prominent southward dipping structure is evident in the uppermost  
129 mantle south of the MPT. The dipping structure consists of two parallel phases with a  
130 negative phase above and a positive phase below. The two phases define a low velocity zone  
131 (LVZ) that dips towards the south, slightly offset from the deep seismicity in this cross  
132 section. However, since the CCP section assumes horizontally layered discontinuities,  
133 strongly dipping structures are not properly reconstructed. They tend to shift to shallower  
134 positions.

135

##### 136 **4.2 CCP stacking for dipping layers**

137 The CCP stack using a horizontally layered (1D) reference model remains valid if the object  
138 to be reconstructed is horizontal or slightly inclined. If mode conversion occurs on strongly  
139 dipping interfaces, the corresponding ray paths differ significantly from ray paths that belong  
140 to mode conversions at horizontally oriented interfaces (Fig. 3). Thus the locations of the  
141 conversion points from inclined structures are not properly calculated if a horizontally layered  
142 earth is assumed. In order to reconstruct inclined structures we calculate ray paths that satisfy  
143 Snell's law at a converter with an assumed inclination angle. This method was already applied  
144 by Li et al. (2000) and Kawakatsu and Watada (2007).



#### 145 **4.2.1 Synthetic tests**

146 Strongly inclined interfaces can only be properly reconstructed by taking into account the  
147 corresponding inclination in the migration. We conducted synthetic experiments to test the  
148 CCP stack of receiver functions assuming horizontal and dipping converters. We used the  
149 RAYSUM (Frederiksen and Bostock, 2000) code to compute QRF and TRF for models  
150 involving a horizontal Moho and a dipping low-velocity zone (LVZ) with dip angles of 30°,  
151 45° and 65°, respectively (Fig. 4). The dipping LVZ produces a double phase with the  
152 negative one above and the positive one below, marked by black lines in each figure. The  
153 structures are reconstructed by CCP stacking of QRF and TRF assuming horizontal and  
154 dipping converters, respectively. In the QRF, both the Moho and the LVZ are visible, whereas  
155 in the TRF only the dipping LVZ is present. For both QRF and TRF, an aberration in the  
156 reconstruction of the dipping slab phase assuming horizontal converters already exists for a  
157 30° dip and increases with increasing dip angle.

158 Assuming horizontal converters, the dipping interface tends to be shifted to a shallower  
159 position. The dipping phases have been properly reconstructed by assuming dipping  
160 converters with corresponding dips. Interestingly, a strongly dipping interface changes its  
161 polarity in the QRF, which may reduce the coherence in the CCP stack section. On the  
162 transverse (T) component (TRF) the inclined interfaces produce significant or even clearer  
163 signals than on the QRF (Abe et al., 2011) (Fig. 4). Here, only events from azimuths between  
164 0° and 180° for a south-dipping LVZ are plotted. Events with azimuths between 180° and  
165 360° would show reversed polarity for the TRF.

166

#### 167 **4.2.2 QRF and TRF cross sections assuming inclined converters**

168 Fig. 5 shows migrated QRFs and TRFs assuming converting interfaces dipping 65° to the

169 south. The QRF section (Fig. 5a) shows that by assuming inclined converters the dipping  
170 structure is migrated to the proper position which coincides with the deep seismicity. The TRF  
171 section (Fig. 5b) is dominated by the dipping structure, since horizontal converters such as the  
172 Moho do not produce signals on the T-component. The double phase in the TRF section  
173 connects the intermediate depth seismicity to the base of the crust just below and south of the  
174 MPT. It provides clear evidence for a southward subduction configuration for the eastern  
175 Pamir. Tectonic models for the Pamir-Hindu-Kush seismic zone with a single originally north-  
176 dipping overturned slab of Indian provenance should thus be discarded.

177

#### 178 **4.2.3 Reconstruction of the dipping low-velocity zone**

179 The migration shown in Fig. 5 is strictly only valid for structures with an inclination angle of  
180 about  $65^\circ$ . Interfaces with other dips are shifted to wrong positions. Figure 6 shows the CCP  
181 stacks of QRF and TRF along the linear array assuming different inclination angles ranging  
182 from  $0^\circ$  (horizontally layered) to  $65^\circ$ . From each image the corresponding information is  
183 picked and displayed in the right column. The summation of all images is displayed in the  
184 bottom figure. The resulting image shows two parallel running phases, the upper one with  
185 negative and the lower one with positive amplitude dipping to the south. The inclination  
186 angles increase with depth from  $30^\circ$  around the Moho to  $65^\circ$  around 150 km depth. Circles  
187 mark the locations of the earthquake hypocenters occurring between  $73^\circ$  and  $75^\circ$ E that have  
188 been recorded and located by our network (Sippl et al., 2013). The dipping structure coincides  
189 with the location of the Wadati-Benioff zone indicated by the hypocenters. The signs of the  
190 imaged velocity contrasts (velocity decrease followed by velocity increase) indicate the  
191 existence of a low-velocity zone (LVZ) between the phases.

192

### 193 **4.3 Combined 2D - model from CCP depth sections**

194 Figure 7a shows the entire 2D model along the N-S profile obtained from the information  
195 displayed in Fig. 2 and 6. It summarizes the RF results from CCP stacking along the N-S  
196 profile. They show the crustal thicknesses as discussed in section 4.1 and the LVZ connected  
197 to the Eurasian lower crust south of the MPT. In this reconstruction, that takes into account  
198 the raypaths of converted waves from the inclined conversion layers of the LVZ, the  
199 intermediate depth earthquakes are located within its borders. Since CCP stacking results tend  
200 to overestimate the thickness of the LVZ due to lateral smoothing, we perform forward  
201 modeling of its phases to estimate its thickness and velocity contrast in the following section  
202 6. The red plane in Fig. 7b shows the position of the N-S cross section relative to the  
203 intermediate depth seismic zone at its eastern end. The green plane shows the position of a  
204 subsidiary depth section constructed from our data, that is discussed in the next section.

205

### 206 **5 Depth section in NW-SE direction**

207 Subsidiary to depth sections along the main TIPAGE profile, an additional cross section is  
208 constructed in the northwest to southeast direction using the areally distributed stations of the  
209 TIPAGE network (Fig. 8). The orientation is chosen perpendicular to the strike of the  
210 intermediate depth seismic zone (see green line in Fig. 1 and green plane in Fig. 7b).

211 Although the resolution of the additional cross section is, due to the sparser station  
212 distribution, not as good as that of the N-S cross section along the main profile, a very similar  
213 configuration as in the main section can be identified. The QRF section shows a 70 km thick  
214 crust in the southeast that shallows towards the northwest (Fig. 8a). A southeast dipping  
215 structure is evident starting at  $\sim 38^\circ$  N. Figure 8b shows the TRF section that is migrated  
216 assuming inclined converters (see section 4.2). As in the north-south cross section, the TRFs

217 show coherent negative and positive dipping phases coinciding with the earthquake  
218 hypocenters. It connects the intermediate depth seismicity to the lower crust in the north-west.  
219 It is likely that the observed structure is the same as in the north-south profile, but rotated  
220 anticlockwise by 45°.

221

## 222 **6 Forward modeling**

223 The basic characteristics of the double phase imaging the LVZ are determined by forward  
224 modeling. An example of the comparison between observed and synthetic receiver functions  
225 at station P18, located in the southern part of the linear array (P18), imaging the LVZ at  
226 around 100 km depth is given in Fig. 9. By modeling the double phase we obtained a  
227 thickness of the low velocity zone of 10.5 +/- 1 km at station P18. At five further stations in  
228 the vicinity of P18, the LVZ phases could be identified in the TRFs with back azimuths  
229 ranging from 30° to 110°. The LVZ thickness appears to vary laterally between 10 and 15 km  
230 (see Figs. S4, S5 and S6 in the Supplementary material for estimation of error and lateral  
231 variation). Amplitude modeling of the Moho Ps phase in the QRFs and the double phase in  
232 the TRFs, reveals the velocity contrast between the LVZ and the surrounding mantle at  
233 100 km depth to be only slightly smaller compared to the velocity contrast at the Moho  
234 ( $\Delta v_{LVZ} = 0.97 \Delta v_{Moho}$ ), implying that the velocity in the LVZ is very similar to the velocity of  
235 the lower crust (Fig. 10)

236

### 237 **6.1 Comparison of modeled and observed RFs for the entire azimuth range**

238 Figs. 9a and 9c show all the individual QRF- and TRF- traces of station P18, sorted by back  
239 azimuth. The summation traces are shown in the upper panels. In the QRF the Moho at ~8 s is  
240 the dominant phase. The double slab phase with negative-positive polarities is clearly seen at

241 10-15 s in the TRF, and can also be recognized in the individual QRF traces. We calculated  
242 the synthetic QRF (Fig. 9b) and TRF (Fig. 9d) with the RAYSUM code (Frederiksen and  
243 Bostock, 2000) using a simple three layered regional model, including a 10.5 km thick LVZ  
244 that dips to the south at an angle of  $50^\circ$  and a 65 km thick crust with a flat Moho (illustrated  
245 by the black lines in Figures 9e-9h). The values of slownesses and back azimuths for the  
246 synthetic RFs were chosen identical to those of the observed RFs. The majority of the RFs  
247 have back azimuths of  $0-180^\circ$ . Few events come from the west (see Fig. 1). Furthermore, the  
248 same post-processing as for the observed data (low pass filter and deconvolution) is applied to  
249 the synthetics.

250

251 Moho conversions are observed in the QRFs at 8s. The first multiple phase (PpPs) of the  
252 Moho (at  $\sim 29$  s) is very (to the point of being invisible) weak in the observed RFs due to  
253 noise, lateral heterogeneity or a crust-mantle transition gradient. Surprisingly, the second  
254 multiple (PpSs) is clearly visible, though at reduced amplitude.

255 The primary phases of the dipping layers are well matched by both the synthetic QRFs and  
256 TRFs between 10 s at  $0^\circ$  azimuth and 16 s at  $180^\circ$  azimuth. Although too few TRFs from  
257 western backazimuths ( $\sigma$ ) are available to verify the polarity flip of the inclined phase seen in  
258 the synthetics, there are sufficient RFs recorded that show both quadrants in the east to have  
259 the same polarity, ruling out anisotropy, which would show a  $2\sigma$  symmetry. Thus, the  
260 waveform modeling verifies the existence of a dipping LVZ. Fig. 9e-9h shows the comparison  
261 of the CCP stack of QRFs and TRFs assuming converters inclined at an angle of  $50^\circ$ . The  
262 amplitude of the QRFs changes its sign on the inclined interface, which may explain the  
263 reduced coherence of the reconstructed dipping phase in the QRFs compared to the TRFs.

264

## 265 **6.2 Waveform modeling for thickness and velocity contrast of the low velocity zone**

266 Modeling waveforms of the slab phases can provide information on the velocity of the LVZ.  
267 However, due to diverse sources of noise, lateral heterogeneity, anisotropy and data  
268 processing, amplitudes of receiver function summations often underestimate the true velocity  
269 contrast. These effects should be similar for the slab and the Moho phases. In Fig. 10 we  
270 compare modeled amplitudes of the slab phase with that of the Moho conversion for station  
271 P18. We consider the most coherent part of the receiver functions in the azimuthal range of  
272 100-110° in order to reduce noise and heterogeneity. We stack QRFs and TRFs and extract the  
273 Moho Ps phase from the QRFs and the slab Ps phase from the TRFs. Synthetic QRFs and  
274 TRFs are modeled for the same azimuth and slowness values and treated in the same way as  
275 the observed data. The summation traces of synthetic and observed seismograms are  
276 compared. First, the amplitudes of the data are normalized to fit the velocity contrast ( $\Delta v_m$ ) of  
277 Vs from 4.6 km/s to 4.15 km/s at the Moho (Mechie et al., 2012). The velocity contrast  
278 between the LVZ and the mantle is chosen relative to  $\Delta v_m$ . At station P18, the best fit between  
279 modeled and observed amplitudes of the slab Ps phase is achieved for a 10.5 km thick LVZ  
280 with a velocity contrast of  $0.97 \Delta v_m$ , implying that the velocity of the LVZ is very similar to  
281 that of the lower crust. Assuming an S-wave velocity (Vs) of 4.6 km/s in the upper mantle, we  
282 obtain a Vs of 4.16 km/s in the LVZ. Thus we conclude that the LVZ is composed of lower-  
283 crustal material.

284

## 285 **7 Crustal thickness in the Pre-Cenozoic**

286 The initial crustal thickness of the Pamir prior to the Cenozoic deformation is crucial to  
287 understand its evolution (Schmidt et al., 2011). The thickness of pre-deformed Pamir crust  
288 may be best estimated by examining the crustal thicknesses of the flat basin areas surrounding

289 the Pamir. Controlled-source seismic profiling revealed a crustal thickness of 45-50 km in the  
290 central Tarim basin (Li et al., 2006). A similar estimate of the crustal thickness beneath the  
291 eastern Fergana basin can be obtained by our receiver function profile (Fig. 2). Here, we  
292 analyze receiver functions of stations (SHAA, IGRN and FRK9) in the central part of the  
293 Tajik basin, that yield a crustal thickness of ~45 km.

294

### 295 **7.1 Crustal thickness in the Tajik basin**

296 The crustal thickness of the Tajik basin is determined using the depth -  $V_p/V_s$  slant stacking  
297 technique (Zhu and Kanamori, 2000) at the stations SHAA, IGRN and FRK9 (see Fig. 1 for  
298 their location). Receiver functions are often complicated for stations located in basins due to  
299 strong multiple reverberations in the sediments (Zheng et al., 2005). In order to suppress these  
300 inner crustal phases the data are low pass filtered with a corner frequency of 2 seconds. In  
301 Fig. 11, for each of these stations the slant stacking results show crustal thicknesses larger  
302 than 45 km (left column). The right column shows the corresponding receiver functions sorted  
303 by slowness (lower panels) and the summation traces moveout-corrected for the  $P_s$ ,  $P_pP_s$  and  
304  $P_pS_s$  phases (as labeled, upper panels). The theoretical arrival times of the  $P_s$ ,  $P_pP_s$  and  $P_pS_s$   
305 phases, calculated from the depth and  $V_p/V_s$  values of the slant stack result, are marked by  
306 black, yellow and blue lines respectively. Direct Moho conversions and their multiples can be  
307 distinguished due to the reversed moveout of the multiples. For all stations the Moho  $P_s$   
308 conversion (at around 8 s) is supported by the  $P_pP_s$  and  $P_pS_s$  phases at around 19 s and 25 s.  
309 The observed  $V_p/V_s$  values of 1.78, 1.81 and 1.86 at stations SHAA, FRK9 and IGRN are in  
310 a realistic range but higher than the global average, probably due to the very thick  
311 sedimentary cover in the Tajik Depression (Nikolaev, 2002).

312

## 313 **8 Discussion**

314 While low velocity layers hosting earthquakes at intermediate depths in subduction settings  
315 are a common feature in oceanic subduction zones (Bostock, 2012; Ferris et al., 2003), they  
316 are rare in intra-continental collisional settings. Moreover, we think that the material we  
317 observe as the LVZ is of continental origin, since the LVZ is clearly connected to the  
318 overlying Eurasian lower crust (Fig. 5) and subduction of oceanic material should have  
319 stopped with the final closure of the Tethys ocean no later than 40 million years ago (Yin and  
320 Harrison, 2000). The absence of Cenozoic volcanism (Schwab et al., 2004) supports this  
321 interpretation. Besides, findings of crustal xenoliths of clearly Eurasian provenance in the  
322 south-eastern Pamir testify that crustal rocks have resided at least at 90 km depth in this  
323 region (Hacker et al., 2005). Together with our observation of the south-dipping LVZ this  
324 provides strong evidence for ongoing subduction of Eurasian continental crust beneath the  
325 Pamir (Fig. 7a). The crustal thicknesses of the surrounding basins of the Pamir can be  
326 regarded as a proxy for the pre-Cenozoic crustal thickness of the Pamir, as the basins are  
327 thought to have been stable during the Cenozoic convergence. Since we observed crustal  
328 thicknesses of more than 45 km in the Fergana and Tajik Basin (Fig 2 and 11) and similar  
329 crustal thicknesses have been observed for the Tarim Basin (Li et al., 2006) it appears  
330 reasonable to assume a similar pre-Cenozoic crustal thickness of at least 40 km in the Pamir.  
331 Estimation of the budget of crustal shortening during the Cenozoic Indian-Eurasian  
332 convergence suggests that, based on this initial crustal thickness, a significant amount of  
333 continental crust must have been lost, most likely having been recycled into the mantle  
334 (Schmidt et al., 2011), which is consistent with our observations. In contrast to previous  
335 models of continental subduction in the Pamir that inferred the subduction of intact albeit  
336 thinned crust (Burtman and Molnar, 1993), our observation suggests that only the lower part



337 of the crust subducts (Fig. 7a). Upper and mid-crustal Eurasian material is underthrust and  
338 incorporated into the Pamir. While the upper crust undergoes brittle deformation reflected by  
339 crustal seismicity in the upper 30 km, the mechanically weakened mid-crustal material  
340 (dashed region in Fig. 7a) is ductilely shortened and decoupled from the lower crust.  
341 Although the buoyancy of the continental crust tends to prevent it from being subducted into  
342 the mantle, the outcrops of ultra-high pressure metamorphic rocks suggest that continental  
343 crust can exist deep in the mantle (Chopin, 1984; Hacker et al., 2006; Okay et al., 1989). Due  
344 to crustal thickening in continental collision zones lower crust can reach pressures where it  
345 transforms to denser eclogite, which reverses the buoyancy and hence may drive its  
346 subduction (Molnar and Gray, 1979; Krystopowicz and Currie, 2013). The seismic velocity  
347 contrast between crust and mantle should then be decreased through eclogitization (Gubbins  
348 et al., 1994), which is supported by our observation of the decreasing intensity of the velocity  
349 contrasts of the LVZ below 100 km depth in the migrated cross sections (Figs. 5 and 6). Since  
350 fully eclogitized crustal material would be seismically invisible our observation suggests that  
351 eclogitization has not been completed above 150 km depth.

352

353 The occurrence of the intermediate-depth earthquakes inside the LVZ implies that these  
354 earthquakes are caused by processes in the subducted crust, which may be attributed to  
355 dehydration embrittlement (Kirby et al., 1996) or shear instability (John et al., 2009; Kelemen  
356 and Hirth, 2007). The most recent study of the Pamir – Hindu Kush seismicity reveals an arc-  
357 like geometry of the seismic zone beneath the Pamir (Sippl et al., 2013) (Fig. 7b). It is  
358 resolved as a relatively smoothly connected layer with a dip direction changing from  
359 eastwards to southwards from its southwestern to its eastern end. Our observation that the  
360 earthquakes occur within the crust at the eastern end and assuming the same petrological

361 mechanisms enable seismicity within the whole arc, implies that the complete arc is of crustal  
362 material. In the subsidiary NW-SE cross section (Fig. 8), constructed using the stations of the  
363 areally distributed stations of the TIPAGE network and crossing the intermediate depth  
364 seismic zone perpendicularly about 130 km to the west, a dipping LVZ can also be recognized  
365 in the TRF section (Fig. 8b), connecting the intermediate depth seismic zone to the lower  
366 crust. This additional observation supports our conclusion, that the arc described by the  
367 earthquake hypocentres traces a slab of Eurasian continental lower crust that subducts beneath  
368 the Pamir.

369

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377 used Seismic Handler (Stammler, 1993) for receiver function processing and GMT (Wessel  
378 and Smith, 1991) for production of the most figures.

379

### 380 **Appendix A. Supplementary material**

381 Supplementary material associated with this article can be found in the online version.

382

383

384

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29, 23 PP.

386

### 387 **Figure Captions**

388 Figure 1. Map of the Pamir with locations of the seismic stations used in this study. Major  
389 tectonic features such as sutures and faults marked in the map are (from north to south): Talas  
390 Fergana fault (TFF), Main Pamir thrust (MPT), Northern Pamir / Kunlun Suture (NKS),  
391 Tanymas Suture (TAN), Rushan Pshart Zone (RPZ), and Shyok suture (SHY). Red and green  
392 highlighted stations were used for the construction of the N-S and the NW-SE depth sections,  
393 respectively. The blue highlighted stations are used to determine the crustal thickness in the  
394 Tajik depression (section 7.1). Insets show a map of teleseismic events used and a map of the  
395 Indian-Eurasian collision zone. The area of altitudes exceeding 3000m is highlighted in  
396 brown, showing the area of Tibet, Himalaya, Pamir and Tien Shan. The red and green lines  
397 denote the projection planes of the N-S (Figs. 2 and 5) and NW-SE (Fig. 8) cross sections,  
398 respectively.

399

400 Figure 2. Common conversion point (CCP) image of Q-component receiver functions (QRF)  
401 along the N-S main profile. Positive velocity contrasts such as the Moho are shown in red  
402 colors, negative velocity contrasts in blue colors. Dashed lines mark the Moho and the double  
403 phase of the subducting low velocity zone (LVZ). Black dots denote the intermediate-depth  
404 earthquakes located with our network (Sippl et al., 2013) within 45 km of the profile. After



405 CCP stacking a boxcar filter with 20 km diameter was applied. The data were amplified by a  
406 factor of 20. The inset shows the color scale of the RF- amplitudes. This figure is shown  
407 without interpretation lines in Fig. S1 in the Supplementary material. Key: see Fig. 1.

408

409 Figure 3. Ray geometry of P to S converted waves at horizontal vs. inclined discontinuities. In  
410 the case of a horizontal boundary (left) the conversion point is situated on the left side of the  
411 station. In contrast, the conversion point at the inclined structure (dip angle  $\delta=45^\circ$ ) is located  
412 on the right side of the station. Incidence and conversion angles in this figure are calculated  
413 for a P-wave velocity of 8 km/s below and an S-wave velocity of 4.15 km/s above the  
414 discontinuity, as may be realistic for a discontinuity between lower-crustal material and the  
415 mantle.

416

417 Figure 4. Synthetic migration test for a single station assuming an inclined geometry of the  
418 slab. The discontinuities of the input model are depicted as black lines. The slab dips to the  
419 south ( $180^\circ$ ). Synthetics are calculated for events with a slowness of 6.4 s/ $^\circ$  and backazimuths  
420 between  $0^\circ$  and  $180^\circ$  with an increment of  $10^\circ$ . Since the model is symmetric (2.5D), events  
421 from backazimuths between  $180^\circ$  to  $360^\circ$  and those from backazimuths between  $0^\circ$  and  $180^\circ$   
422 would be identical for the QRFs, and would sum to zero for the TRFs.

423

424 Figure 5. Migrated CCP stack receiver function cross sections assuming  $65^\circ$  inclined  
425 conversion interfaces. Q-component and T-component RF are shown in (A) and (B),  
426 respectively. The same processing on the CCP stacked data was applied as in Fig. 2. This  
427 figure is shown without interpretation lines in Fig. S2 in the Supplementary material. Key: see  
428 Fig. 1.

429

430 Figure 6: Migrated cross sections along the TIPAGE main seismic profile assuming different  
431 inclination angles of the converters. The assumed angles are depicted on the left. In the left  
432 and middle columns the amplitudes of the Q- and T-components are shown, respectively. For  
433 the T-component only events from the east (azimuthal angles  $0^{\circ}$ - $180^{\circ}$ ) are taken into account.  
434 In the right column the positions of the conversion boundaries corresponding to the relevant  
435 dip angle are highlighted in red (positive converters) and blue (negative converters). The  
436 bottom figure shows the amalgamation of all picked elements from the right columns.

437

438 Figure 7. A: Cartoon showing reconstruction of the crust and mantle lithosphere at the  
439 northern margin of the India (Pamir) – Eurasian collision zone based on the observations of  
440 this study. Positive velocity contrasts such as the Moho are shown in red colors, negative  
441 velocity contrasts in blue colors. Bold interfaces and events are taken from Figs. 2 and 6. The  
442 shaded region depicts a zone that is expected to be weak, which allows crustal material to be  
443 ductilely shortened and/or Eurasian upper and mid crust to penetrate into the Pamir.  
444 B: Schematic sketch showing the geometry of the Pamir intermediate depth seismic zone,  
445 figure modified from Sippl et al. (2013). Green and red planes show the locations of the main  
446 N-S and the subsidiary NW-SE RF profile, respectively.

447

448 Figure 8. Diagonal north-west to south-east cross section using the areal network. In a) the  
449 amplitude of the Q-component is migrated assuming horizontal converters. In b) the T-  
450 component of the RFs is migrated assuming inclined converters with an inclination angle of  
451  $45^\circ$ . RFs from events with azimuths between  $135^\circ$  and  $315^\circ$  are migrated with reversed  
452 amplitudes for the T-component. The data are lowpass filtered with a cut-off period of 3s.  
453 This figure is shown without interpretation lines in Fig. S3 in the Supplementary material.

454

455 Figure 9. (A-D) Comparison of observed receiver functions and synthetic tests for station P18.  
456 Traces are low pass filtered with a cut-off period of 3s. (E-H) Migrated observed QRFs (E)  
457 and TRFs (G) and synthetic QRFs (F) and TRFs (H) for station P18. The traces are migrated  
458 assuming  $50^\circ$  dipping conversion layers.

459

460 Figure 10. Comparison and modeling of waveforms of (A) the Moho (on the Q component)  
461 and (B) the slab (on the T component) phases. To match the amplitudes for the LVZ (slab), the  
462 velocity contrast at the base of the LVZ has to be chosen to be very similar (97%) to that at  
463 the Moho. Amplitudes of the data are normalized to a velocity drop from  $V_s=4.6$  km/s  
464 (mantle) to  $V_s=4.15$  km/s (lower crust) at the Moho (Mechie et al., 2012). QRF and TRF data  
465 traces used for the modeling are shown in (C) and (D), respectively.

466

467 Figure 11. Slant stack results (left column) and receiver functions sorted by slowness (right  
468 column) for stations SHAA, FRK9 and IRGN, located in the Tajik basin showing a crustal  
469 thickness of ~45 km. On top of each receiver function panel, moveout corrected summation  
470 traces for the direct (PS) and multiple phases (PPPS and PPSS) are shown. The theoretical  
471 arrival times determined from the slant stack results for the Ps, PpPs and PpSs phases are  
472 marked on top by black, yellow and blue lines, respectively.

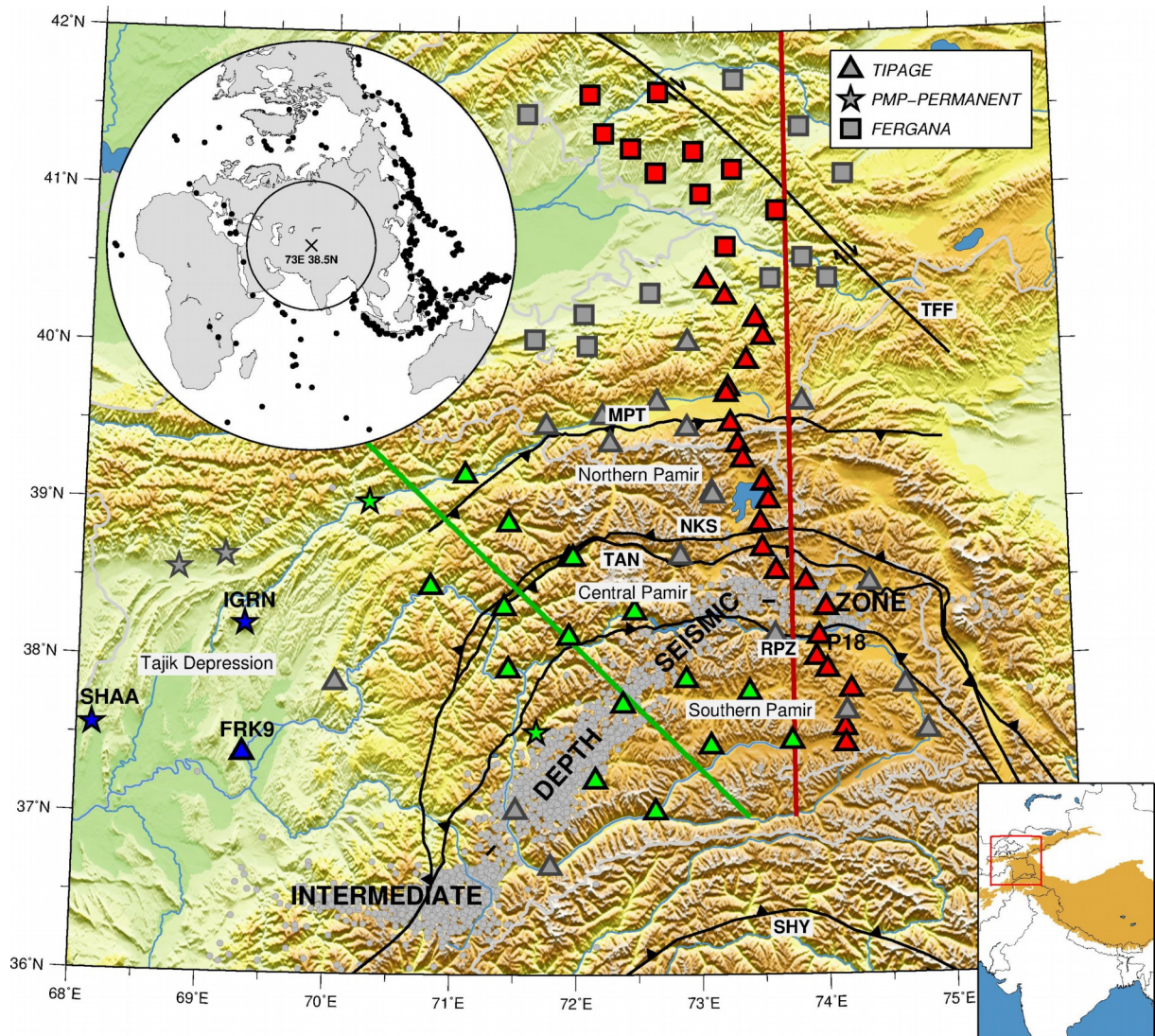


Figure 1

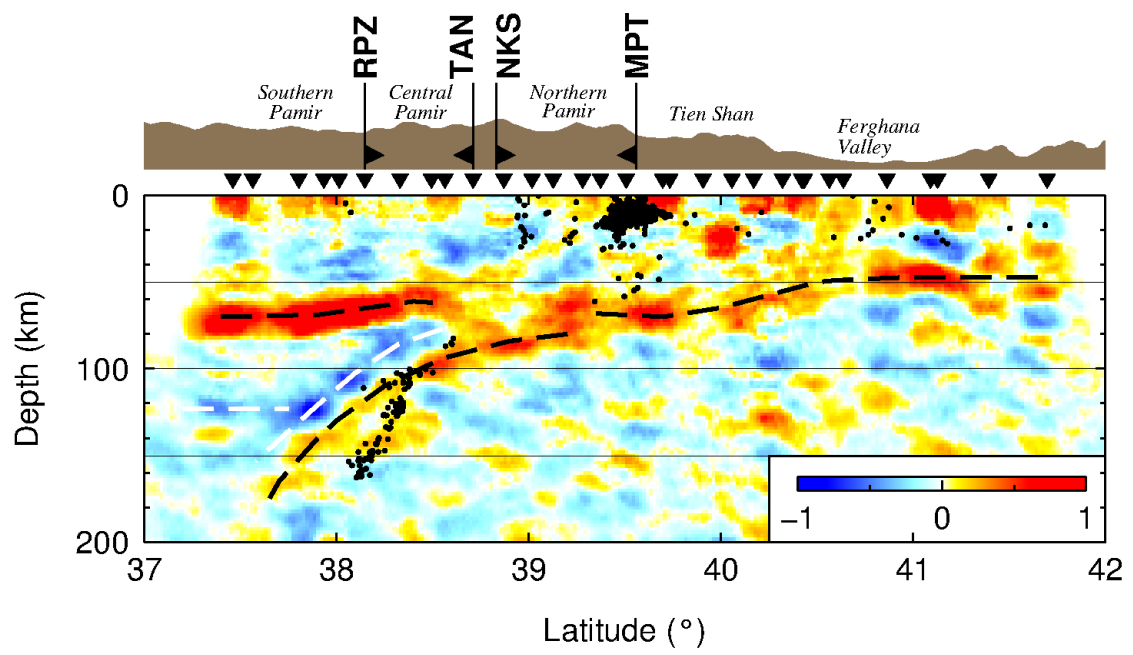


Figure 2

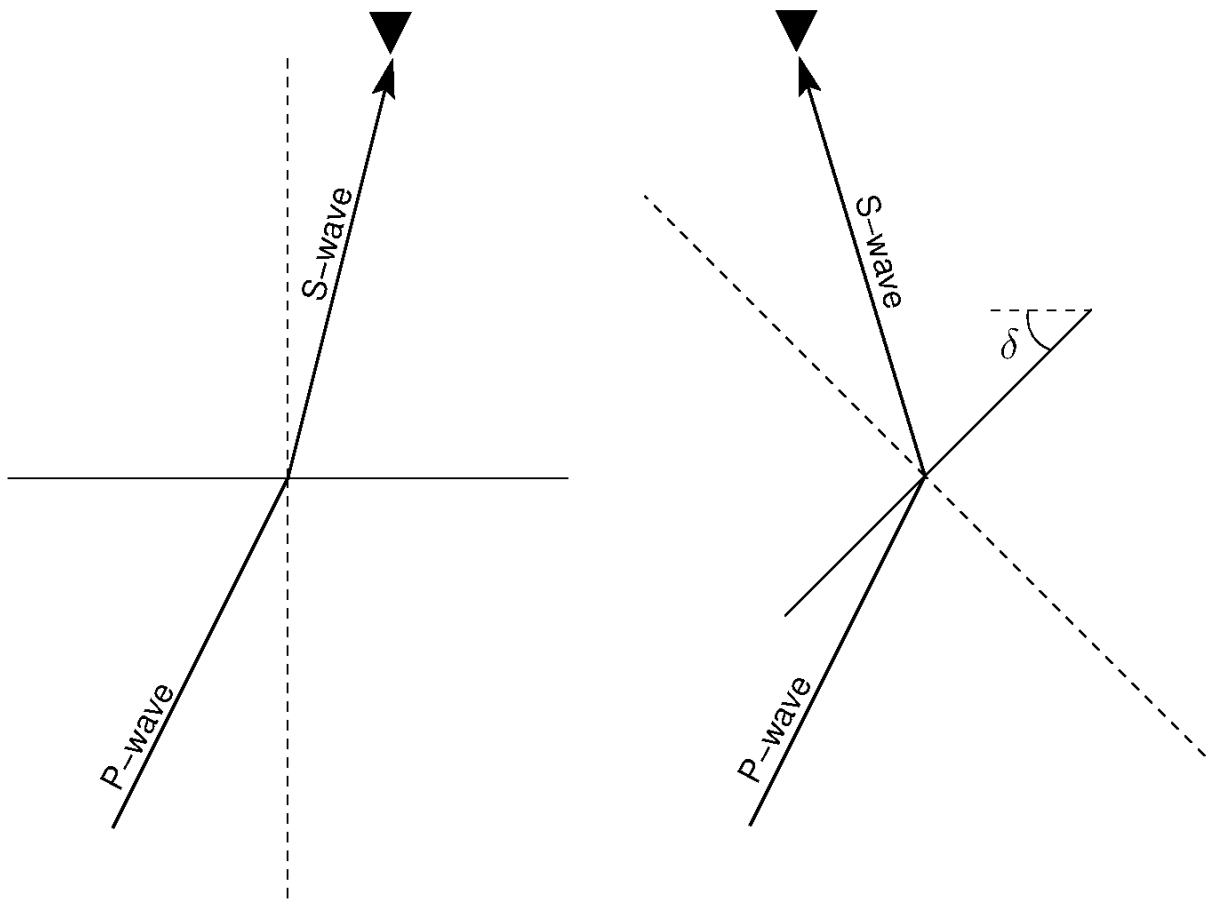


Figure 3

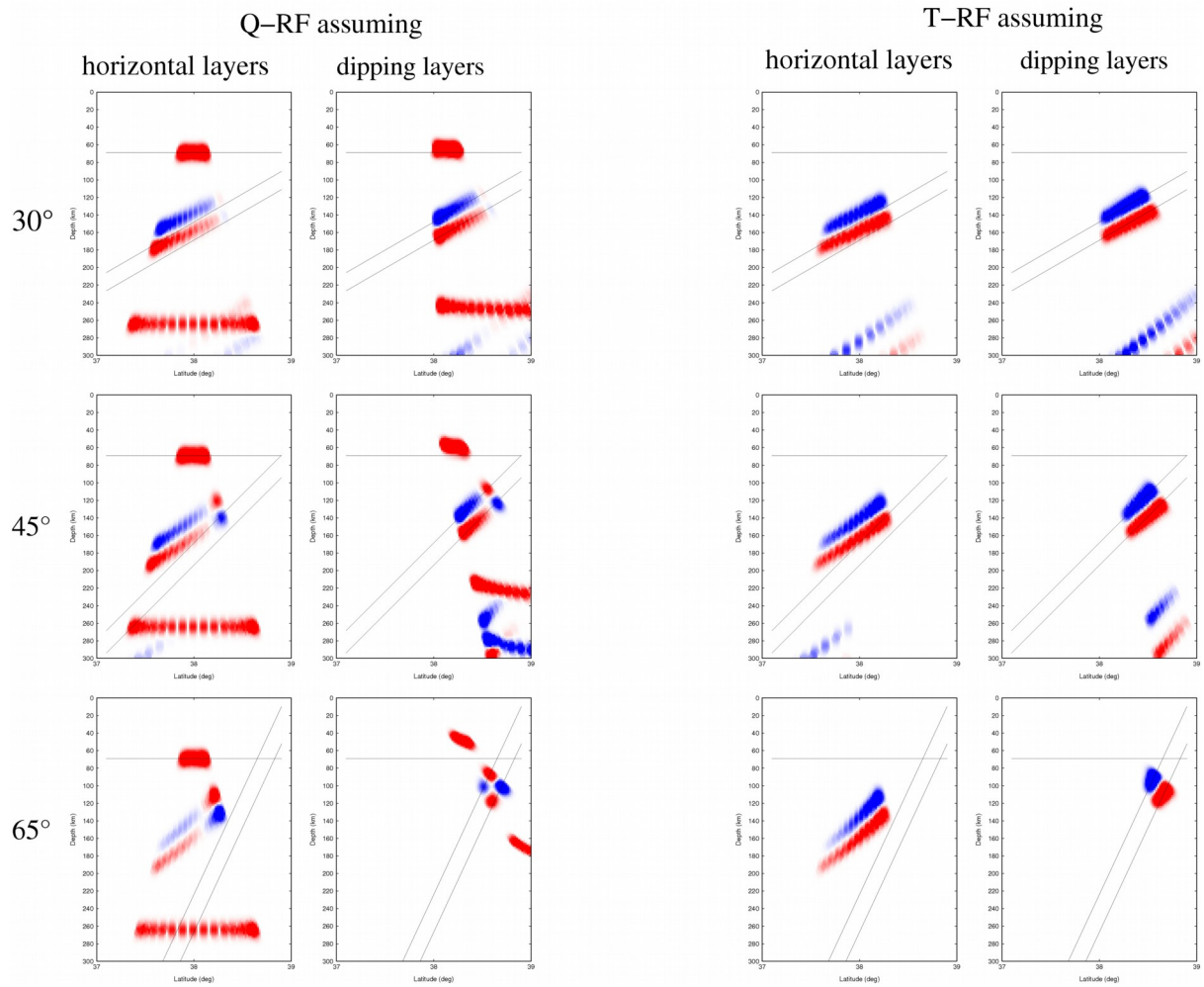


Figure 4



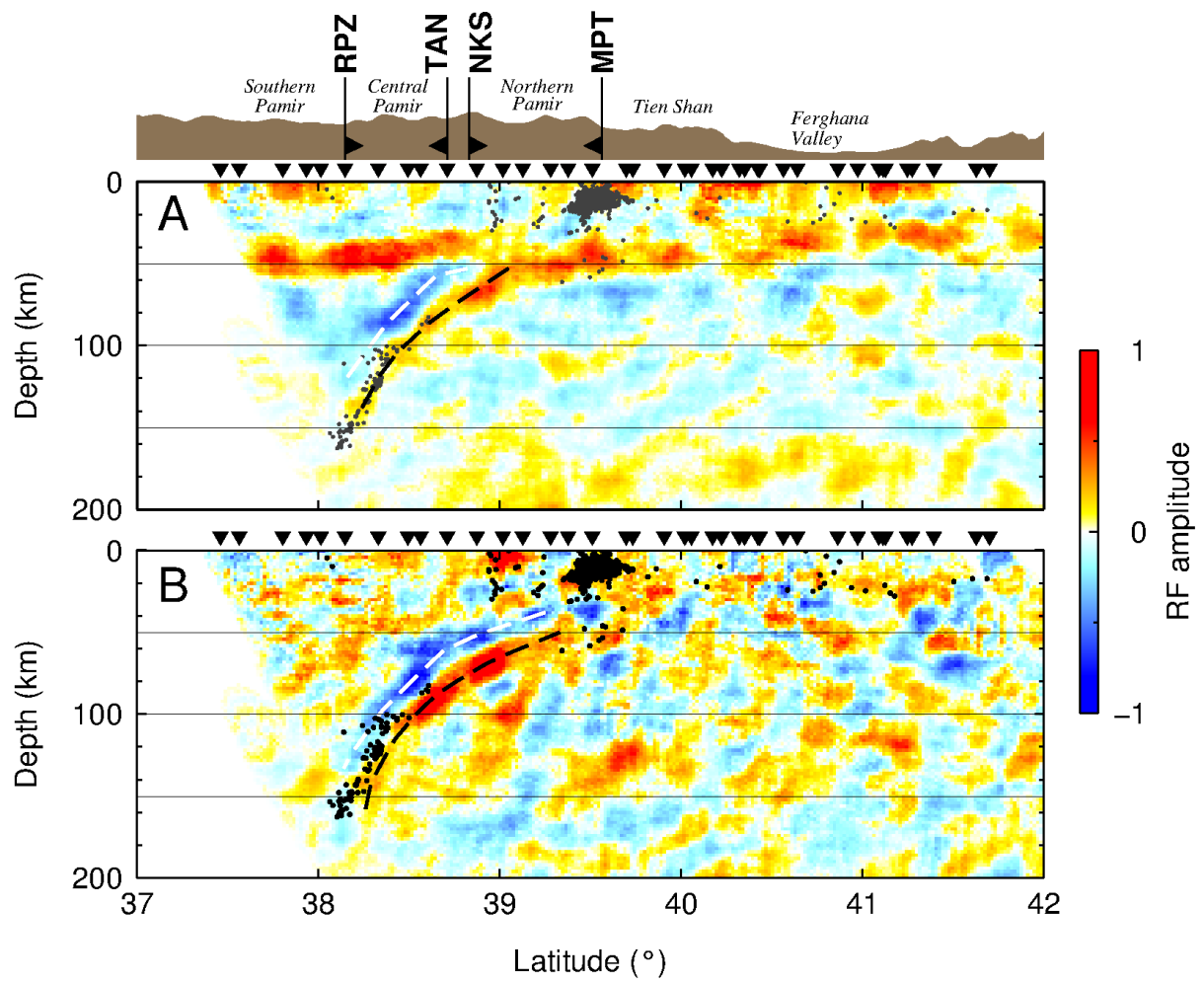


Figure 5

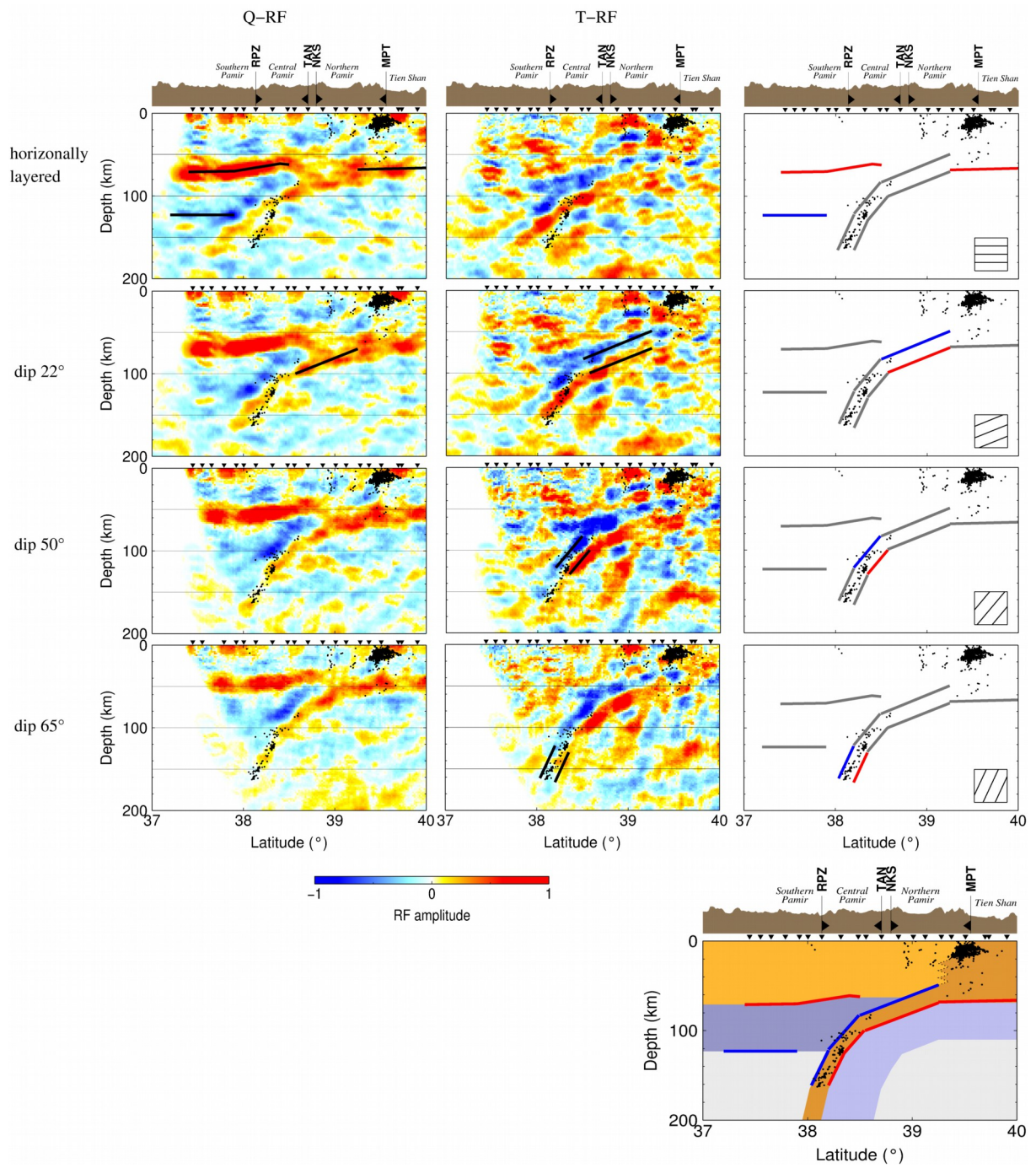


Figure 6

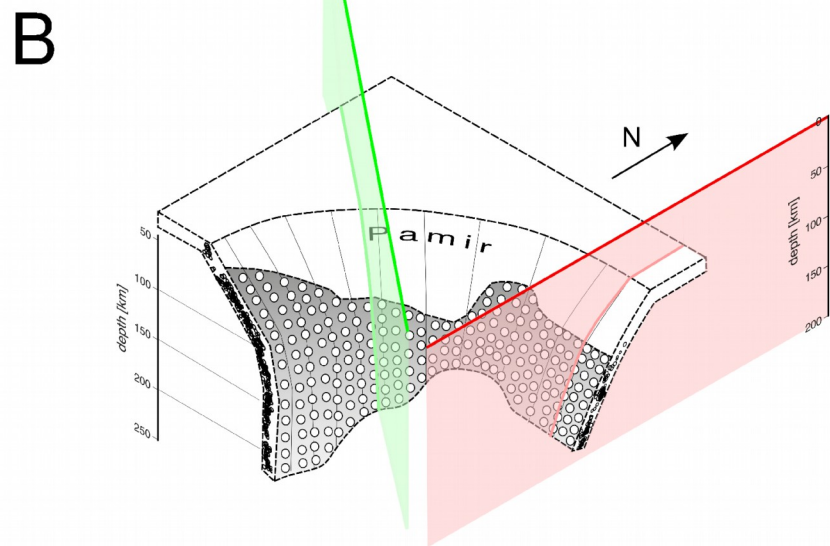
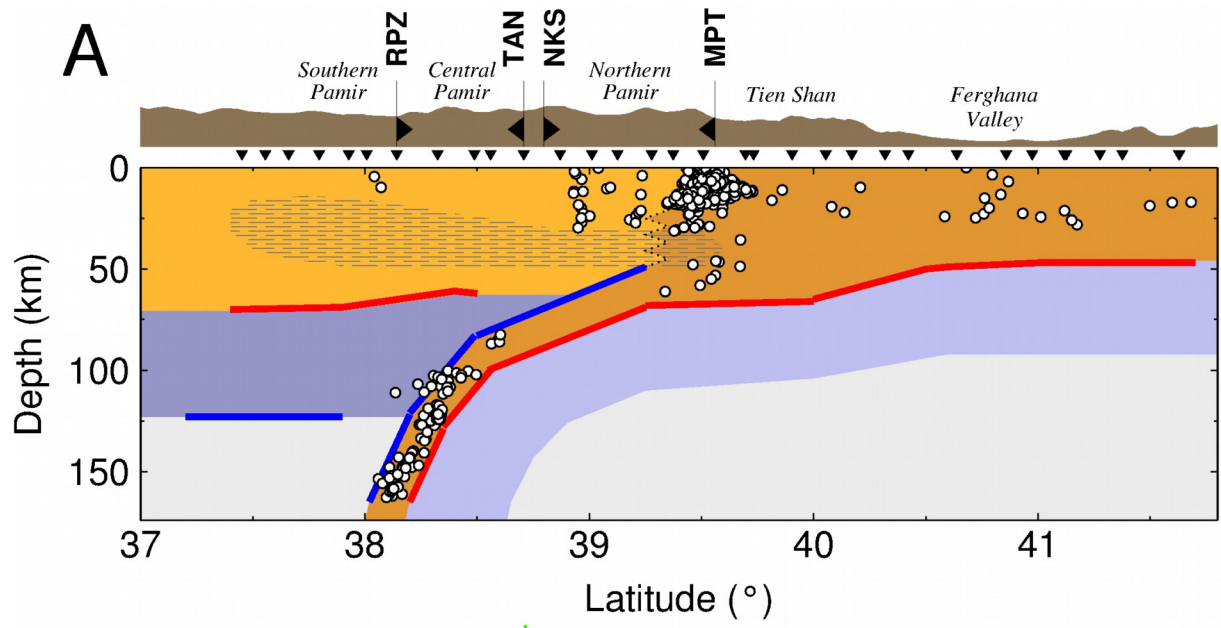


Figure 7

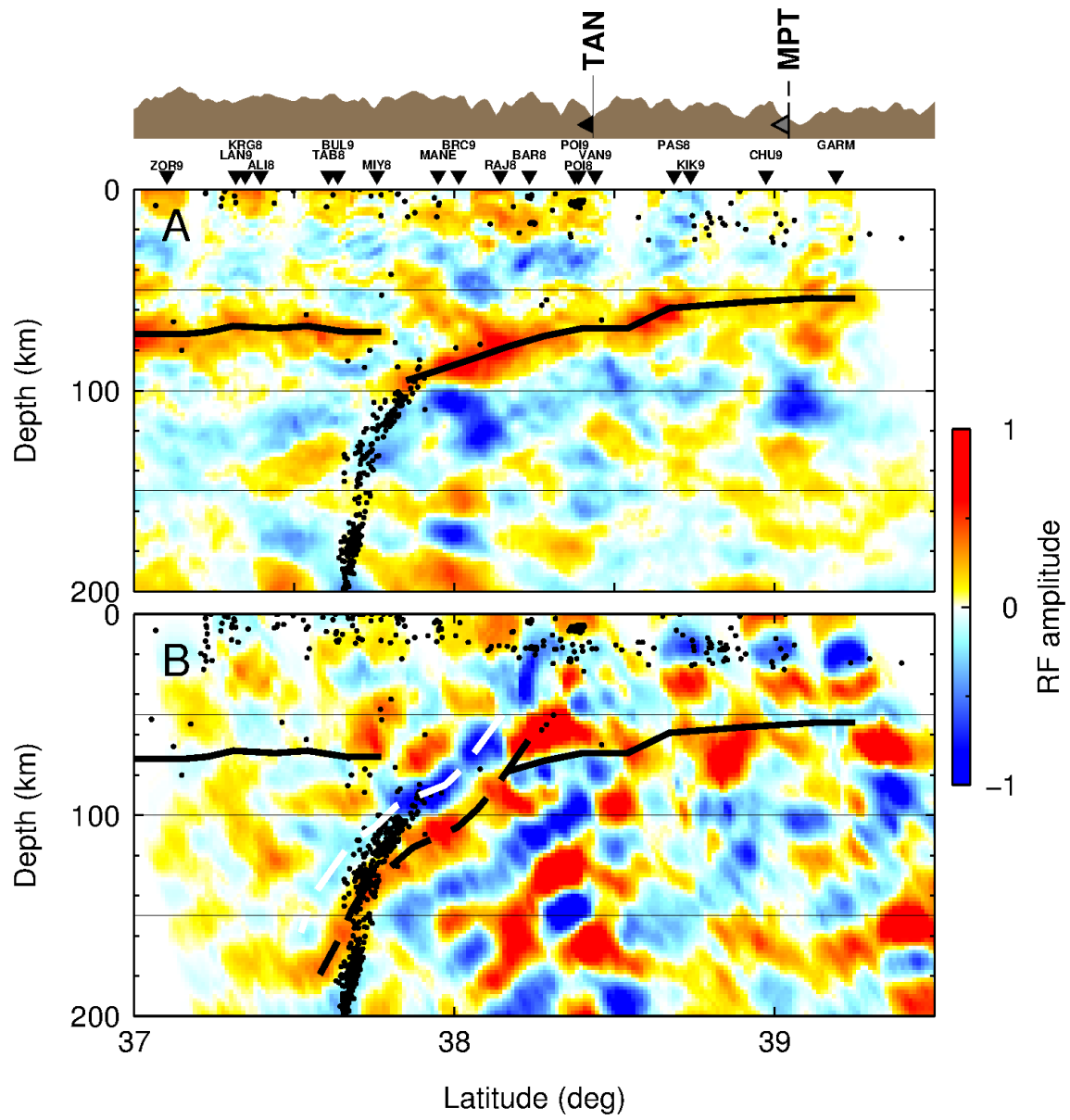


Figure 8

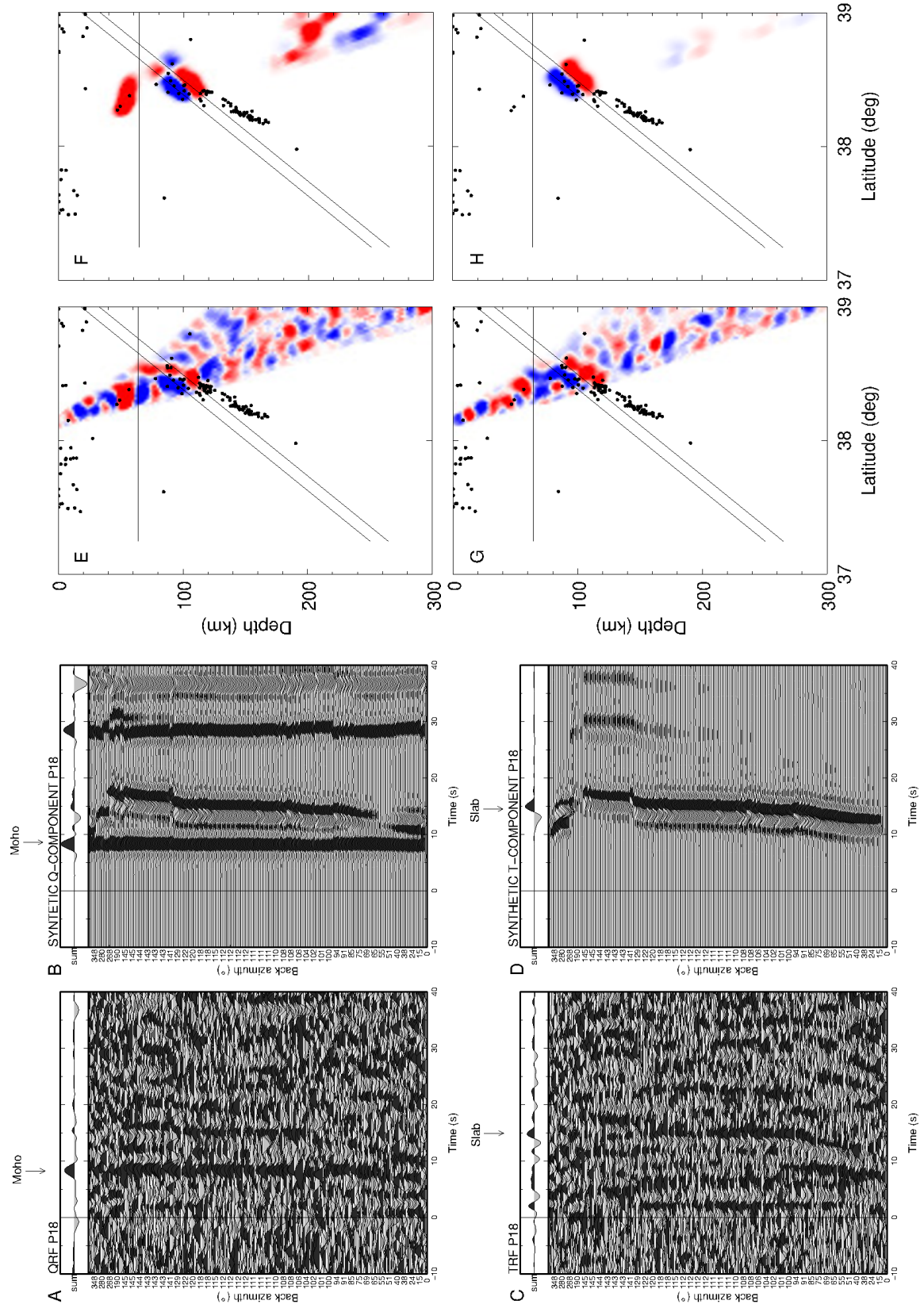


Figure 9

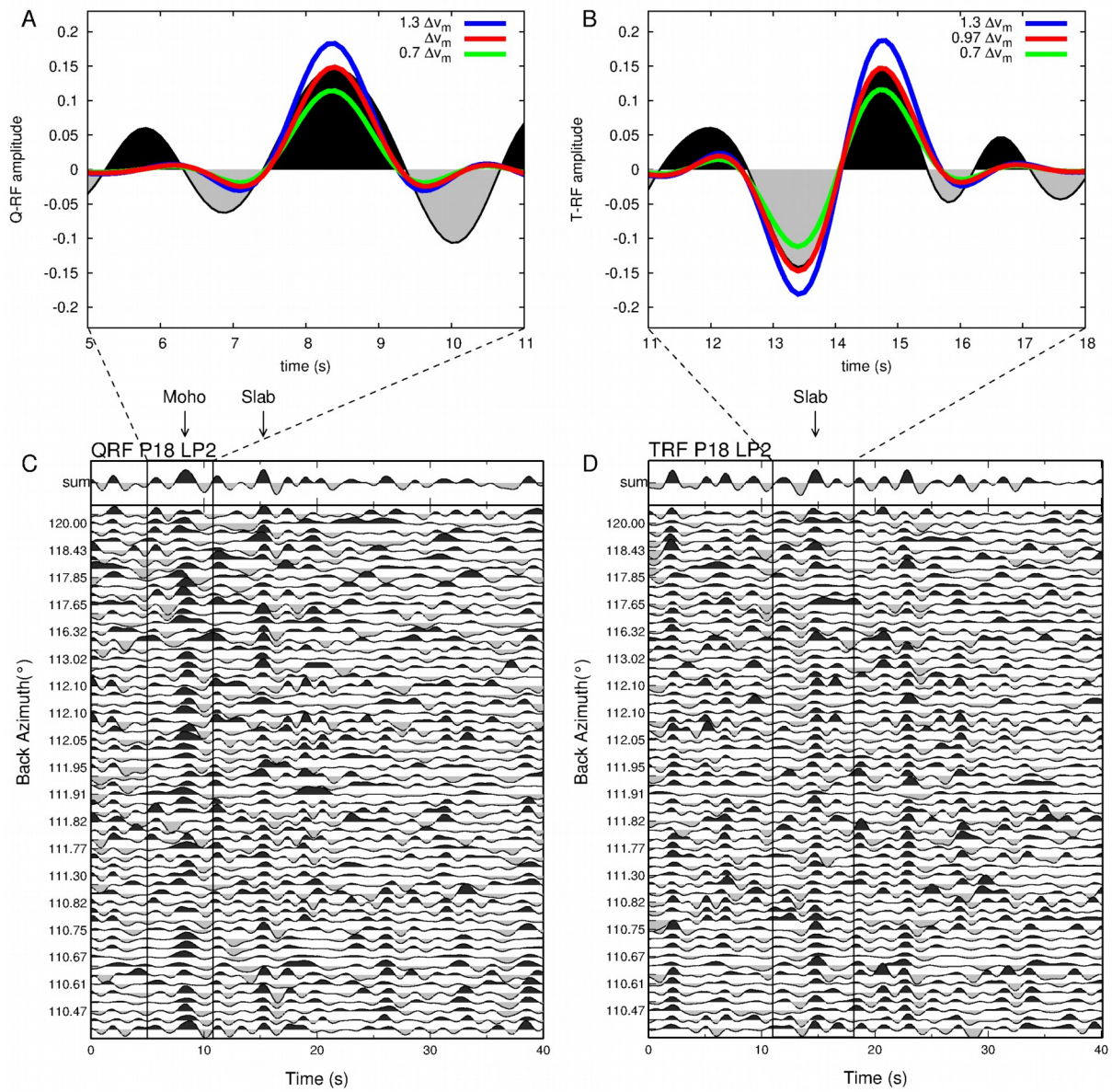


Figure 10

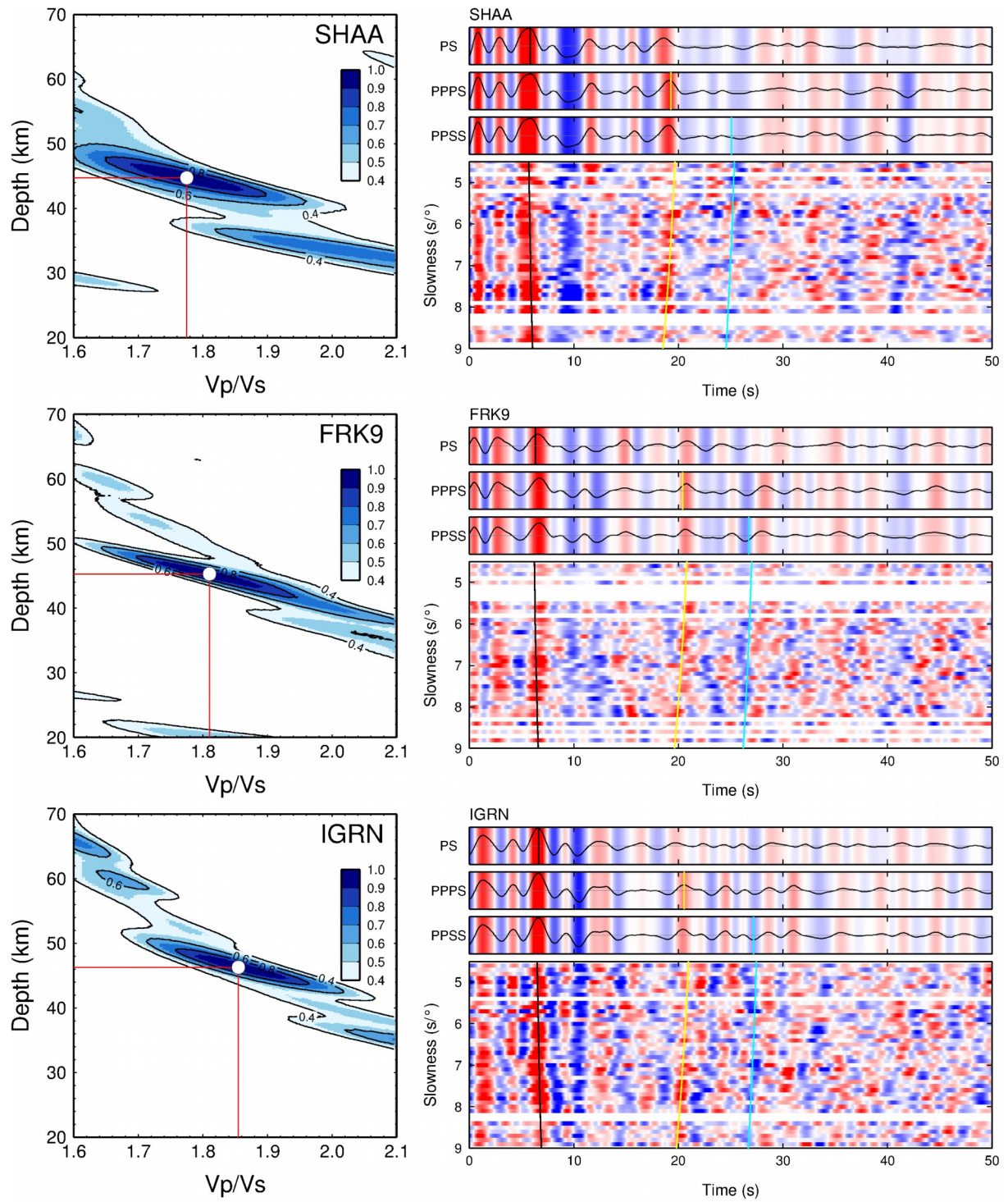
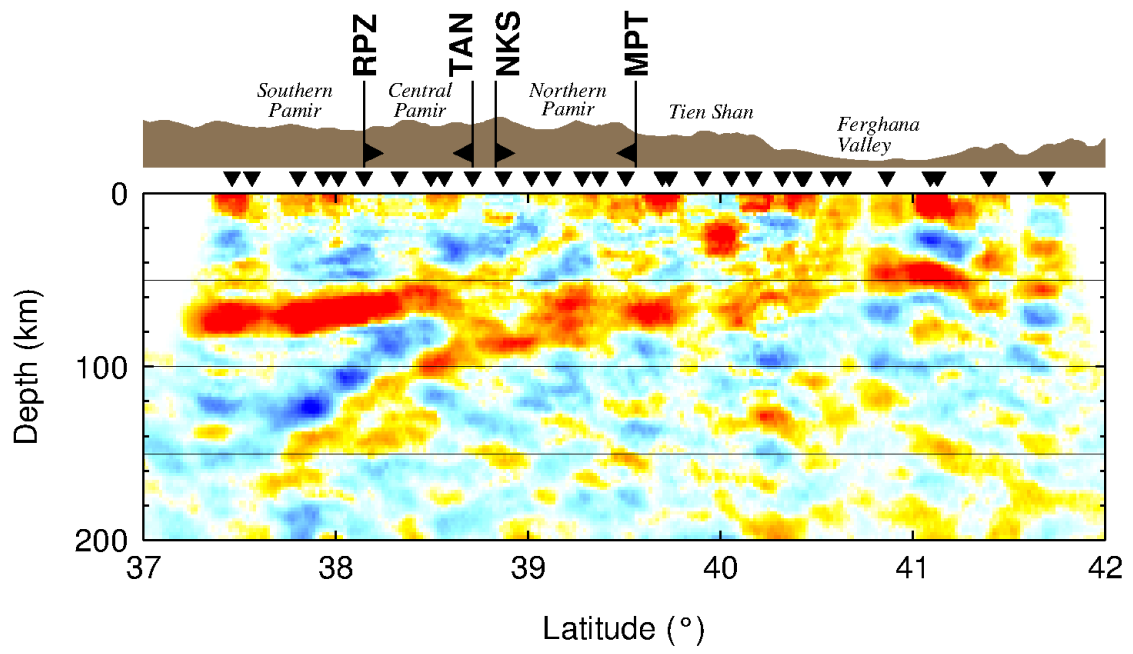
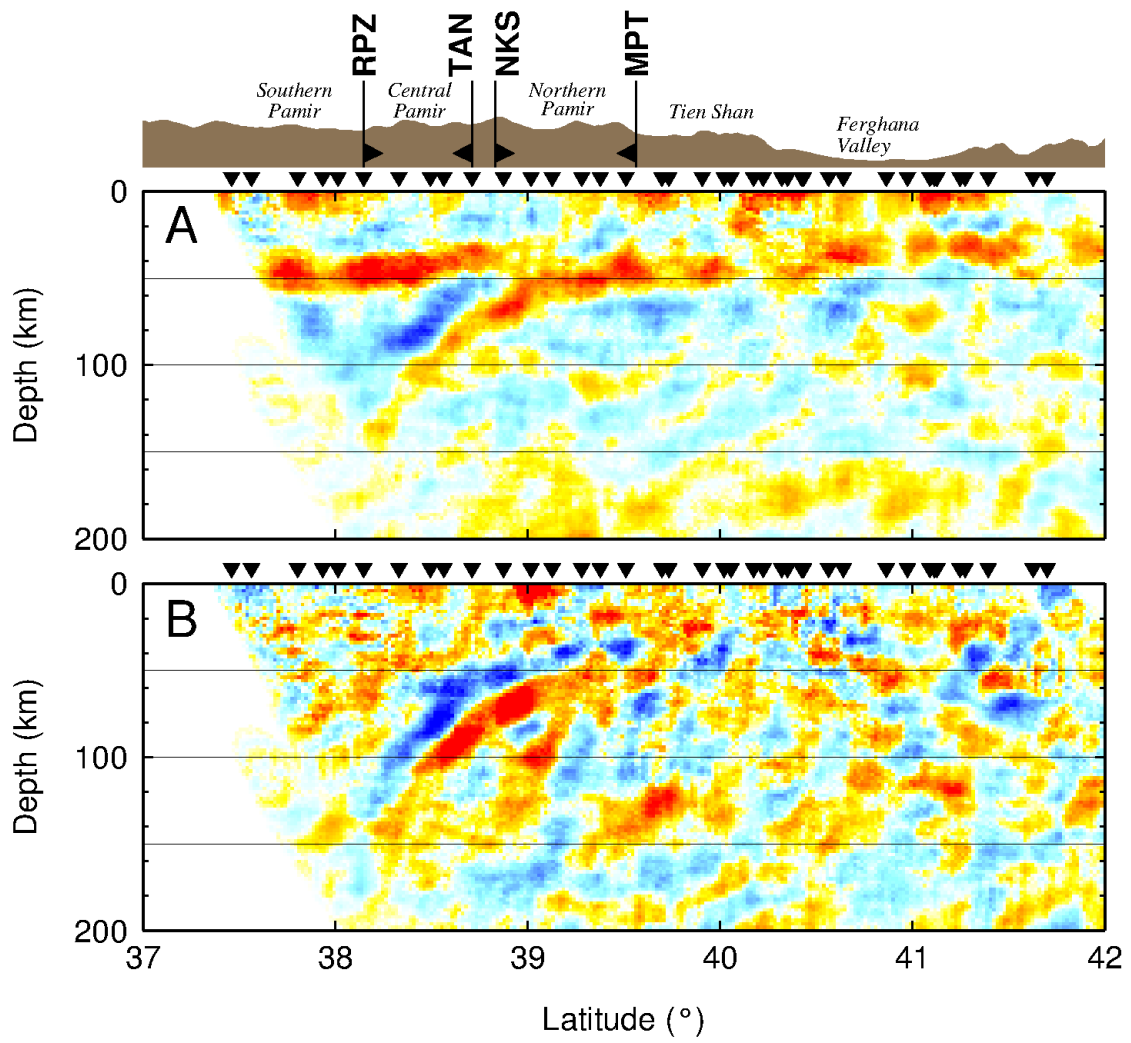


Figure 11

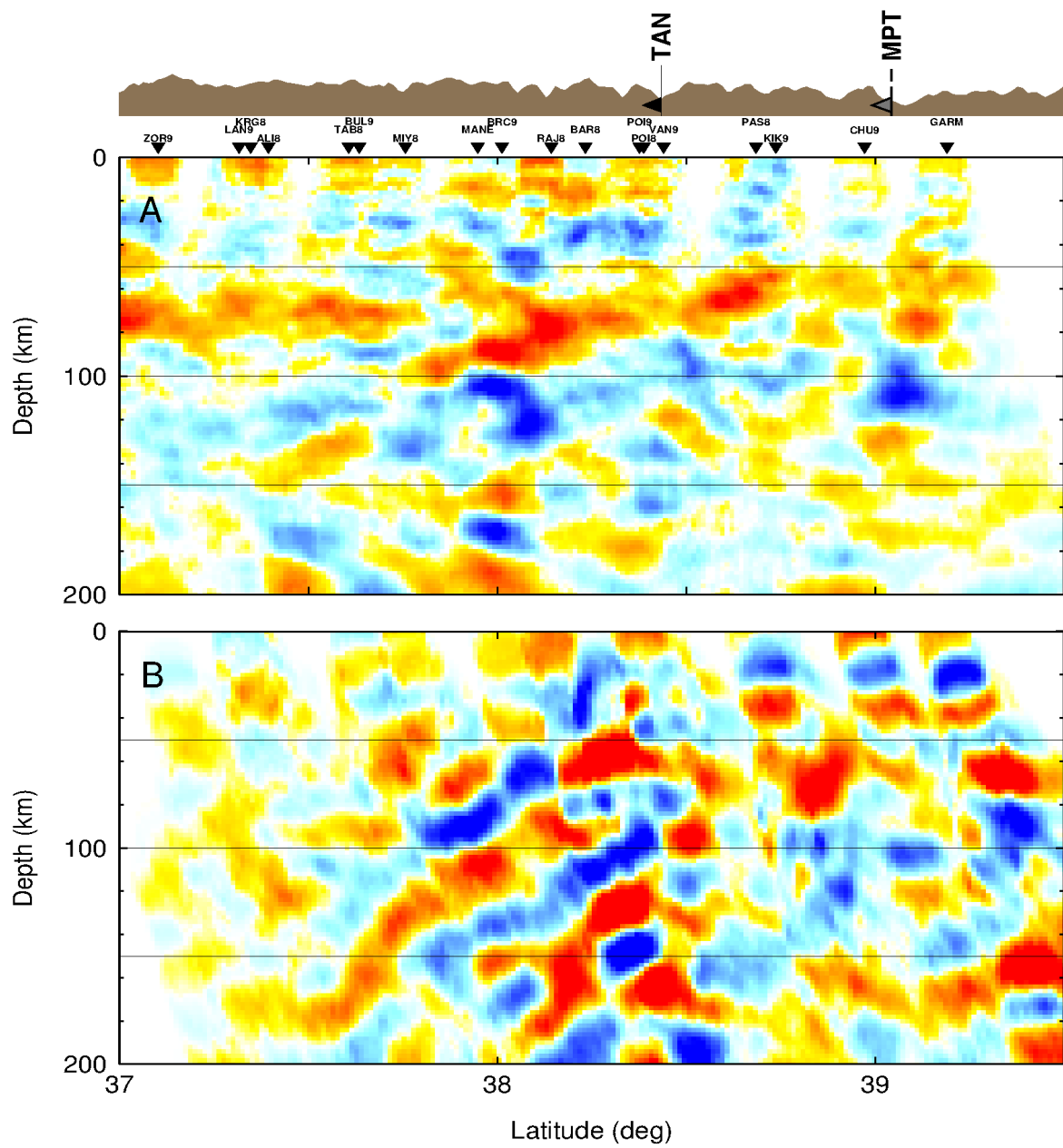


Supplementary Figure S1. Same as Fig. 2 without interpretation lines.

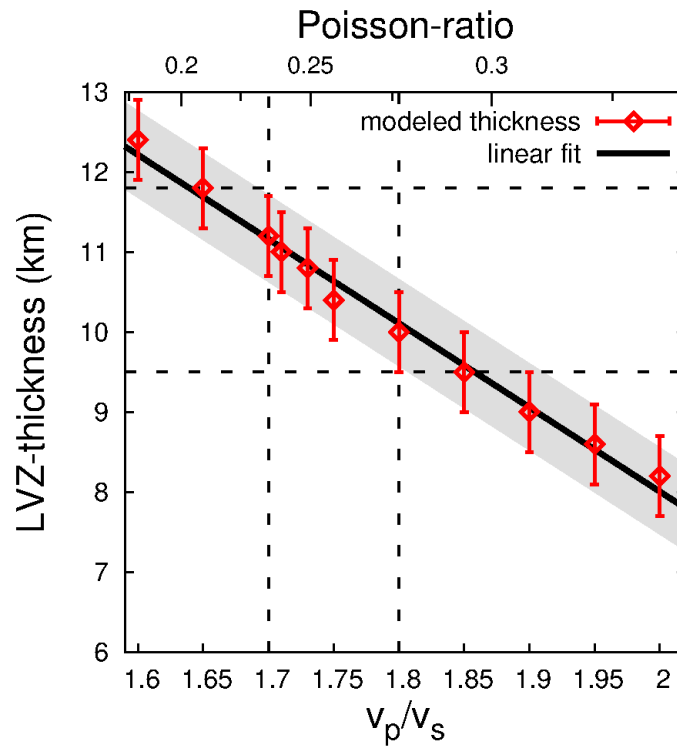




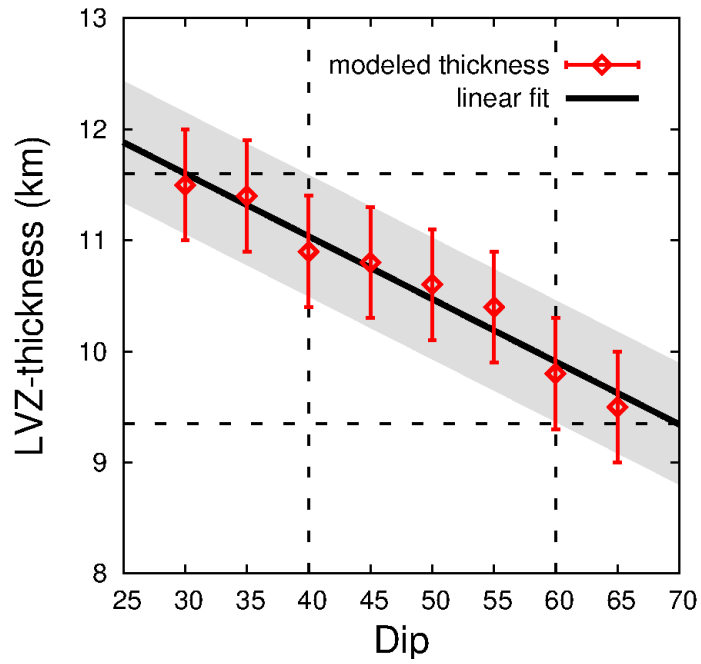
Supplementary Figure S2. Same as Fig. 5 without interpretation lines.



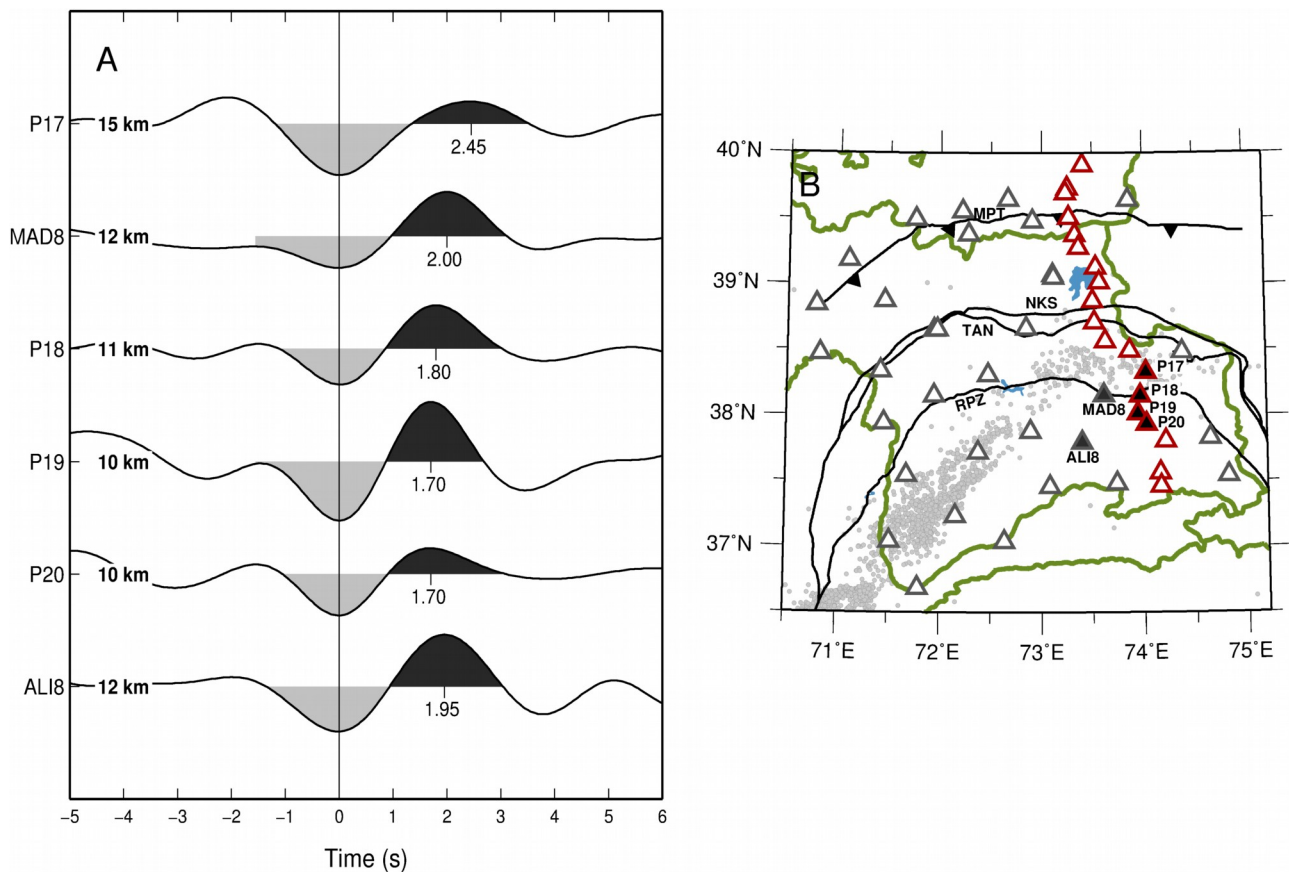
Supplementary Figure S3. Same as Fig. 8 without interpretation lines.



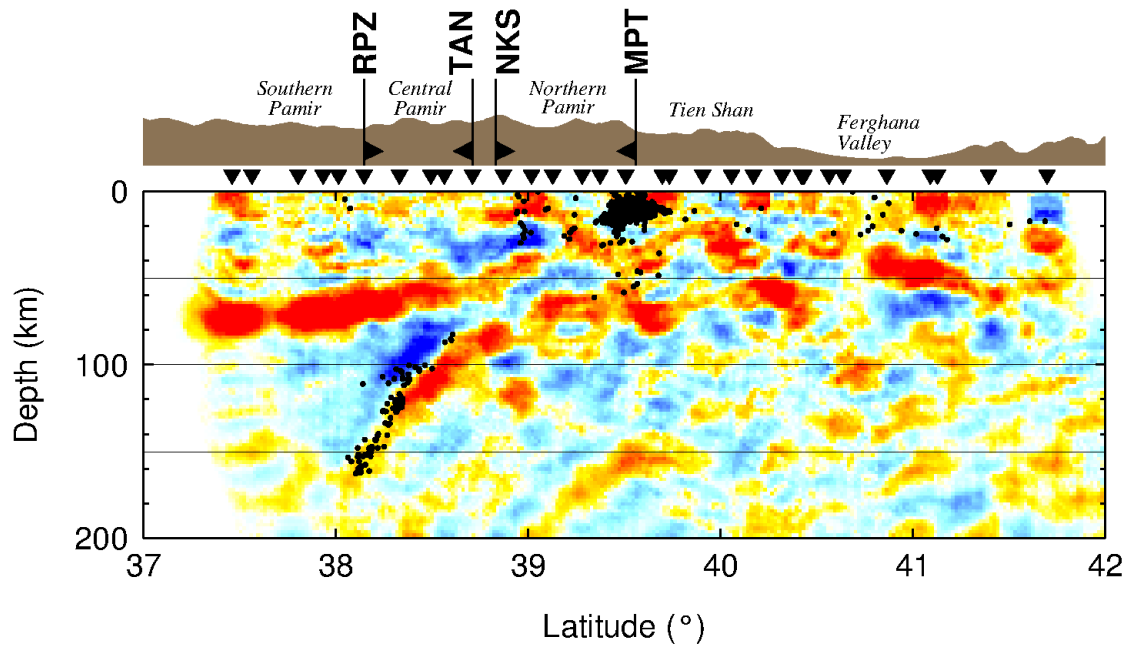
Supplementary Figure S4. Forward modeling of the LVZ phases has been performed for varying  $V_p/V_s$  inside the LVZ. A linear dependency between modeled thickness and  $V_p/V_s$  is observed. The minimum unit of thickness difference that can be distinguished visually by comparing the synthetic TRFs with the data is 0.5 km, which is thus given as the error for each model. In the  $V_p/V_s$  interval from 1.7 to 1.8 the LVZ thickness ranges between 9.5 and 11.8 km.



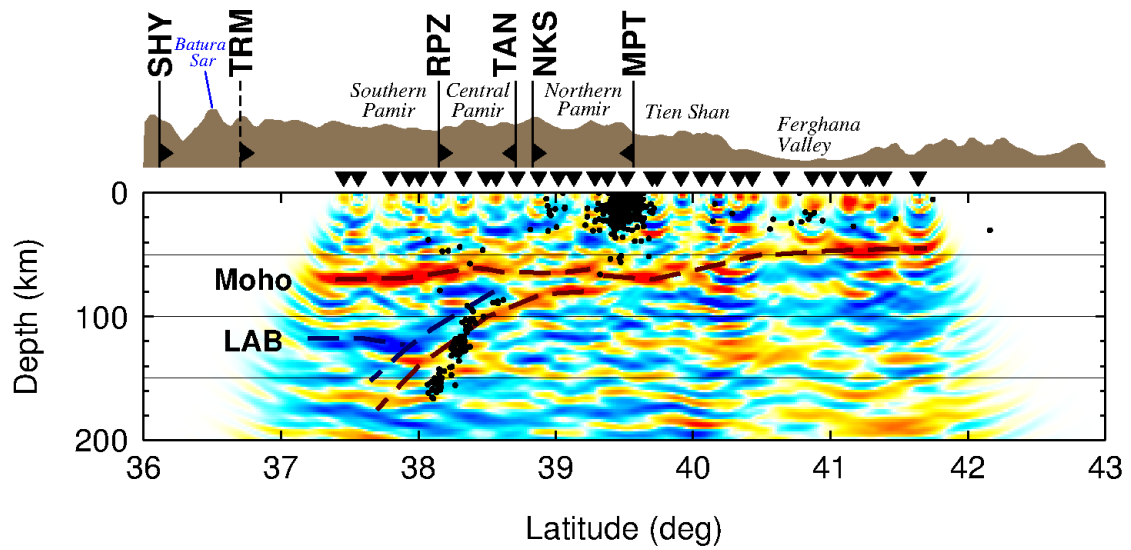
Supplementary Figure S5. Forward modeling of the LVZ phases has been performed for varying inclination angle (dip) of the LVZ. In the forward modeling process, the minimum thickness difference that can be distinguished visually by comparing synthetic and observed TRFs is 0.5 km, which is thus given as the error for each modeling. Assuming a dip of  $50^\circ \pm 10^\circ$  the LVZ thickness can be estimated to be  $10.5 \pm 1$  km.



Supplementary Figure S6. A: The TRF conversion phases of the LVZ at 6 stations as seen in a back azimuth range from  $30^\circ$  to  $110^\circ$ . For each station the leading negative phases were aligned at 0 s and stacked. The delay time between negative and positive conversions and corresponds to the thickness of the LVZ. Starting with the forward modeling at station P18, the thicknesses of the LVZ at the different stations were linearly extrapolated, taking into account the delay times at each station relative to the delay time at station P18. The resulting thicknesses are given at the left side of each trace. In this way, the LVZ thickness is estimated to be 10-15 km. B: Station map showing the positions of stations P17, MAD8, P18, P19, P20 and ALI8 as black filled triangles. Stations of the N-S profile are depicted as red triangles, the other stations of our network as gray triangles.



Supplementary Figure S7. Common conversion point (CCP) image along the N-S main profile with QRF and TRF components stacked together in each grid cell. For depths greater than 80 km CCP stacking was performed for 50° inclined conversion interfaces to image the LVZ, while for shallower depths horizontally layered converters were assumed to image the Moho at the right depth. The same processing on the CCP stacked data was applied as in Fig.2. Key: see Fig. 1.



Supplementary Figure S8. Depth section of QRFs along the main N-S profile converted to depth using a Fresnel-zone stacking algorithm. After the stacking a boxcar filter with 6 km diameter was applied. This image shows a clear negative signal south of the subducting LVZ labeled as the lithosphere-asthenosphere boundary (LAB). Key: see Fig. 1.