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1     **Role of erosion and isostasy in the Cordillera Blanca uplift: insights**  
2             **from landscape evolution modeling (northern Peru, Andes)**

3  
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7  
8     **Highlights:**

9     - Inversion of the landscape evolution coupled with thermochronological data provides  
10     constraints on erosion efficiency factor, uplift rates and geothermal gradient

11    - Isostatic effect of eroding a denser rock mass represent a not negligible contribution to  
12    the Cordillera Blanca uplift on a < 5 Ma time scale

13    - Cordillera Blanca drainage divide location is controlled by initial drainage network  
14    rather by maximum uplift rates and precipitation distribution

15  
16    **Abstract**

17    The processes driving uplift and exhumation of the highest Peruvian peaks (the  
18    Cordillera Blanca) are not well understood. Uplift and exhumation seem closely linked to  
19    the formation and movement on the Cordillera Blanca normal fault (CBNF) that delimits  
20    and shapes the western flank of the Cordillera Blanca. Several models have been  
21    proposed to explain the presence of this major normal fault in a compressional setting,  
22    but the CBNF and the Cordillera Blanca recent rapid uplift remain enigmatic. Whereas  
23    the Cordillera Blanca morphology demonstrates important erosion and thus a  
24    significant mass of rocks removal, the impact of erosion and isostasy on the evolution of  
25    the Cordillera Blanca uplift rates has never been explored. We address the role of  
26    erosion and associated flexural rebound in the uplift and exhumation of the Cordillera

27 Blanca with numerical modeling of landscape evolution. We perform inversions of the  
28 broad features of the present-day topography, total exhumation and  
29 thermochronological data using a landscape evolution model (FastScape) to provide  
30 constraints on the erosion efficiency factor, the uplift rate and the temperature gradient.  
31 Our results evidence the not negligible contribution of erosion and associated flexural  
32 rebound to the uplift of the Cordillera Blanca and allow us to question the models  
33 previously proposed for the formation of the CBNF.

34

35 **Keywords:** Rock uplift; Isostatic effect of eroding a denser rock mass; Normal fault;  
36 Low-temperature thermochronology; Numerical modeling of the landscape evolution;  
37 Cordillera Blanca; Peruvian Andes

38

## 39 **1. Introduction**

40 In mountain ranges, surface uplift is usually assumed to be the result of shortening and  
41 crustal thickening. Surprisingly, in northern Peru, uplift of the footwall of an active  
42 normal fault is responsible for the formation of the highest Peruvian summits in the  
43 Cordillera Blanca (Fig. 1). Several models have been proposed [*Dalmayrac and Molnar,*  
44 *1981; McNulty and Farber, 2002*] to explain this unusual situation, but the processes  
45 driving both the Cordillera Blanca uplift and extensional deformation along the  
46 Cordillera Blanca normal fault (CBNF) remain poorly constrained. The CBNF trends  
47 parallel to the Andean range and is the most spectacular normal fault in the Andes [Fig.  
48 *1; Margirier et al., 2017*]: the CBNF is ~200 km long and shows ~7 km of vertical offset  
49 in total [*Margirier et al., 2016*], it has been active since ~5.4 Ma [*Bonnot, 1984; Giovanni,*  
50 *2007*]. The CBNF is located above the Peruvian flat-slab [*Barazangi and Isacks, 1976*], a  
51 section of the convergent plate boundary between the Nazca Plate and the South

52 American Plate characterized today by near-horizontal subduction geometry. The  
53 Cordillera Blanca and the Cordillera Negra form, respectively, the hanging wall and the  
54 footwall of the CBNF (Fig. 1).

55 The Cordillera Blanca fast exhumation rate ( $\sim 1$  mm/yr) has been previously linked to  
56 motion on the CBNF [e.g., *Bonnot, 1984; McNulty and Farber, 2002; Giovanni, 2007;*  
57 *Margirier et al., 2015*]. New thermobarometry data and erosion rates reconstruction  
58 based on thermochronological data indicate a recent, i.e. Early Pleistocene, increase in  
59 erosion rate in the Cordillera Blanca [ $\sim 2-0$  Ma; *Margirier et al., 2016*]. *Margirier et al.*  
60 [2016] suggested that an important isostatic contribution from glacial erosion may  
61 explain the recent exhumation of the Cordillera Blanca batholith. Indeed, the removal of  
62 such a mass of material represents a significant upward unloading on the lithosphere,  
63 which should drive substantial flexural uplift. This unloading and flexural uplift would  
64 have also generated large differential stresses in the lithosphere, which could have  
65 caused the reactivation of pre-existing structures such as the CBNF. Previous studies  
66 demonstrated that the flexural uplift driven by alpine-type valley incision could reach  
67 rates similar to those caused by tectonic processes [*Montgomery, 1994; Small and*  
68 *Anderson, 1995; Cederbom et al., 2004; Stern et al., 2005*]. Recently *Braun et al.* [2014]  
69 proposed that erosion-driven isostatic rebound should scale with the density of surface  
70 rocks: denser rocks, such as a granitic body intruded in sedimentary rocks, rebound and  
71 therefore are exhumed faster than the surrounding less dense rocks.

72 The rapid uplift of the Cordillera Blanca, the large volume of eroded rocks since the  
73 emplacement of the Cordillera Blanca batholith and its location in the footwall of an  
74 active normal fault in a compressive plate boundary, make the Cordillera Blanca the  
75 perfect place to question the nature and efficiency of potential feedbacks between  
76 erosion and uplift along the CBNF. The aim of this paper is thus (i) to test whether the  
77 increase of erosion rate suggested for the last 2 Ma in the Cordillera Blanca [*Margirier et*

78 *al.*, 2016] could be due to an increase of rock uplift rates since 2 Ma rather than a change  
79 of climate and/or erosion process and (ii) to quantify the importance of isostatic  
80 rebound associated with valley incision and erosion of denser rocks to explain the uplift  
81 of the Cordillera Blanca and to test the adequacy of a flexure-driven model in such a  
82 setting. To address this issue, we have attempted to model landscape evolution in the  
83 Cordillera Blanca using a landscape evolution model (LEM), in this case based on the  
84 FastScape algorithm [*Braun and Willett, 2013*].

85

## 86 **2. Context**

### 87 2.1 Geologic and climatic context

88 The Cordillera Blanca hosts the highest Peruvian summits with a cluster of 6000 m  
89 peaks (Fig. 1). It hosts a large 14–5 Ma granitic pluton [zircon U-Pb; *Mukasa, 1984*;  
90 *Giovanni, 2007*] emplaced at ~3 km depth into deformed Jurassic sediments [*Margirier*  
91 *et al.*, 2016]. The Cordillera batholith is elongated (150 × 15 km) and trends parallel to  
92 the Andean range (Fig. 1A). Based on apatite fission-tracks and (U-Th/He) dating several  
93 studies gave estimations of exhumation rates ranging between 1 and 2 mm/yr in the  
94 central part of the Cordillera Blanca for the last 3-4 Myr [*Giovanni, 2007*; *Hodson, 2012*;  
95 *Margirier et al.*, 2015]. This exhumation phase is likely associated to rock uplift and  
96 CBNF activity. On a shorter time scale (30-0 ka), the uplift rates have been constrained  
97 by <sup>10</sup>Be dating of scarps along the CBNF [*Siame et al.*, 2006] and geomorphic features  
98 (moraines) displaced by the fault [*Schwartz, 1988*; *Gérard et al.*, in prep]. The vertical  
99 slip rates decrease from north to south ranging from 5.1 ± 0.8 mm/yr to 0.6 ± 0.2 mm/yr  
100 [*Schwartz, 1988*; *Siame et al.*, 2006; *Margirier et al.*, 2017; *Gérard et al.*, in prep].  
101 Whereas the higher peaks in the Cordillera Blanca are located close to the CBNF and  
102 should therefore correspond to the region of maximum tectonic uplift, the drainage

103 divide is located in the eastern part of the Cordillera Blanca, ~15 km away from the  
104 CBNF (Fig. 1B). The batholith is deeply incised by deep U-shaped valleys resulting from  
105 recent glacial erosion. Several other glacial landforms (moraines, roches moutonnées)  
106 are evidence of the large extension of the glaciations that shaped the Cordillera Blanca  
107 morphology [Farber *et al.*, 2005]. Today, the Cordillera Blanca summits form the Andean  
108 drainage divide [Wise and Noble, 2003] and act as an orographic barrier to moisture  
109 coming from the Amazon basin [Fig. 1B; Bookhagen and Strecker, 2008]. Even if the most  
110 important orographic effect of the Andes is controlled by the lower relief of the sub-  
111 Andes, the Cordillera Blanca high elevation prevents moisture from reaching the  
112 western flank of the Andes, resulting in wetter climatic conditions in the Cordillera  
113 Blanca (mean rainfall ~1.5 m/yr) than in the Cordillera Negra (mean rainfall ~0.5  
114 m/yr), farther to the west [Bookhagen and Strecker, 2008].

115 The morphology of the Cordillera Negra is that of a 4500 m high plateau incised by 1–2  
116 km deep valleys along its western flank (Fig. 1). The range hosts Cretaceous and  
117 Paleogene plutons [73–48 Ma; Beckinsale *et al.*, 1985] intruded into Jurassic sediments.  
118 Neogene volcano-sedimentary deposits cap the Cordillera Negra [54–15 Ma Calipuy  
119 Formation; Cobbing *et al.*, 1981]. A regional surface uplift associated with the subduction  
120 geometry change and dynamic topography process has been evidenced in the Cordillera  
121 Negra from 15 Ma [Eakin *et al.*, 2014; Margirier *et al.*, 2015]. Rare and discrete moraines,  
122 only seen above ~4200 m, indicate limited ice cover and resulting glacial erosion in the  
123 Cordillera Negra [Bonnot, 1984].

124

## 125 2.2 Paleogeography

126 Several studies have documented the Late Miocene paleogeography of the Cordillera  
127 Blanca region [Wise and Noble, 2003; Giovanni, 2007; Hoorn *et al.*, 2010]. Based on  
128 pollen analyses, Hoorn *et al.* [2010] constrained the elevation in the Cordillera Blanca

129 region to be ~4 km in the Middle Miocene. However, at that time, the Cordillera Negra  
130 topography formed the main drainage divide [*Wise and Noble, 2003*]. Based on the  
131 location, age and stratigraphy of volcanic deposits in the Callejón de Huaylas Basin that  
132 separates the two ranges, *Wise and Noble [2003]* and *Giovanni [2007]* suggested that a  
133 depression already existed between the Cordillera Blanca and the Cordillera Negra at  
134 the end of the Late Miocene. In addition, based on  $\delta^{18}\text{O}$  analyses of paleolake deposits,  
135 *Giovanni et al. [2010]* showed that highest elevations in the Callejón de Huaylas Basin  
136 were attained in the latest Miocene. Little is known, however, about the topography of  
137 the Cordillera Blanca before the emplacement of the batholith.

138

### 139 **3. Landscape evolution model: FastScape**

#### 140 3.1 Model

141 We used the FastScape algorithm [*Braun and Willett, 2013*] to solve the stream power  
142 law to predict landscape evolution following a set of tectonic forcing (uplift) and initial  
143 topography (geomorphic setting). Because of the optimum ordering of the nodes, the  
144 algorithm is implicit in time and computationally very efficient, requiring only  $O(n)$   
145 operations where  $n$  is the number of points used to discretize the topography.  
146 Consequently, FastScape can be used repetitively, even if using a very high spatial  
147 discretization, to attempt to reproduce the main features of present-day topography  
148 and, using a simple 1D thermal model, predicted cooling ages (apatite fission-tracks and  
149 (U-Th/He) ages which can be related to erosion) in the hope of deducing a plausible  
150 first-order topographic and uplift history. To do this, we coupled FastScape to the  
151 neighborhood algorithm (NA) [*Sambridge, 1999a; Sambridge, 1999b*] to “invert” the  
152 thermochronological ages, the barometric constraints and the known, final topography  
153 in order to find the best-fitting values of several unknown parameters, including the

154 erosion efficiency ( $K_f$ ), the elastic thickness ( $T_e$ ), the uplift rates ( $U_1$ ) at several periods in  
155 the past and the temperature at the base of the model ( $T_{max}$ ). Using NA, we first carried  
156 out a large number of FastScape runs with parameter values randomly selected between  
157 fixed limits. For each run the model predictions were compared to the data and a misfit  
158 function was estimated, which was used to select a new set of model parameters to be  
159 tested. This procedure was repeated several times until an optimum combination of  
160 parameter values was found and, more importantly, the shape of the misfit function  
161 could be mapped. This allows us to analyze and discuss the range of model parameters  
162 where the model prediction fits the observations, as well as the sensitivity of the model  
163 predictions to the model parameters.

164

### 165 3.2 Erosion and isostasy

166 To represent surface erosion, we solve the stream power law that is commonly used to  
167 parameterize bedrock incision by rivers in steep mountainous terrains:

168

$$169 \quad (1) \quad \frac{\partial h}{\partial t} = U - K A^m S^n$$

170

171 where  $h$  is topographic height,  $t$  is time,  $U$  is uplift rate,  $A$  is drainage area and  $S$  is slope  
172 in the direction of water flow, taken here to be the steepest path between a point on the  
173 landscape and any among its eight neighbors. Parameters  $K$ ,  $m$  and  $n$  are poorly  
174 constrained constants that mostly depend on lithology and climate. We will use  $m = 0.4$ ,  
175  $n = 1$  and vary  $K$  around a mean value of  $1.5 \times 10^{-5} \text{ m}^{1-2m} \cdot \text{yr}^{-1}$  (see *Croissant and Braun*  
176 [2013] for a discussion on the value of these parameters). We also assume that  
177 hillslopes are subject to mass transport at a rate that is simply linearly proportional to  
178 topographic slope. Assuming mass conservation, this leads to a law for the rate of



179 change of topography on hillslopes that is proportional to the curvature of topography.  
180 Combining this with the stream power law leads to the following evolution equation for  
181 the surface topography:

182

$$183 \quad (2) \quad \frac{\partial h}{\partial t} = U - K A^m S^n + K_D \left( \frac{\partial^2 h}{\partial x^2} + \frac{\partial^2 h}{\partial y^2} \right)$$

184

185 where  $K_D$  is a transport coefficient or diffusivity. We will assume an arbitrary value of  $K_D$   
186 = 0.3 m<sup>2</sup>/yr. We realize that using the stream power law and a simple diffusion-like  
187 representation of hillslope processes may seem inappropriate to model topographic  
188 evolution in a glaciated landscape. We justify our choice to use the stream power law by  
189 the fact that (i) no algorithm exists to solve the glacial erosion equations in an efficient  
190 manner that would allow for the type of inversion we have conducted here, (ii) we are  
191 interested in the effect of denser rocks erosion and the resulting isostatic rebound on  
192 the range-scale shape of the Cordillera Blanca and not in the details of the landform, and  
193 (iii) for both fluvial and glacial erosion processes, erosion rate is primarily controlled by  
194 slope and the geometry of a first order drainage system.

195

196 To account for the isostatic rebound associated with surface erosion, we also solve the  
197 following bi-harmonic equation representing the flexure of a thin elastic plate subject to  
198 surface loading/unloading:

199

$$200 \quad (3) \quad D \left( \frac{\partial^4 w}{\partial x^4} + 2 \frac{\partial^4 w}{\partial x^2 \partial y^2} + \frac{\partial^4 w}{\partial y^4} \right) = \Delta \rho g w + \rho_s g \Delta h$$

201

202 where  $w$  is the surface displacement associated with the isostatic adjustment of an  
203 increment in erosion  $\Delta h$ ,  $D$  is flexural,  $\rho_s$  is the density of surface rocks and  $\Delta \rho$  is the

204 density contrast between asthenospheric density,  $\rho_a$  and surface rock density [*Turcotte*  
205 *and Shubert, 1982*].

206

207 A short word of explanation is necessary to understand how we have coupled Equation  
208 (2) and Equation (3). As explained above, we assume that the imposed uplift rate does  
209 not contribute to the load applied to the thin elastic plate representing the lithosphere;  
210 only the eroded material does. This means that we can rewrite Equation (2) in the  
211 following form:

$$\frac{\partial h}{\partial t} = U - F(e)$$

212 where  $e$  is erosion rate:

$$e = KA^m S^n - K_D \left( \frac{\partial^2 h}{\partial x^2} + \frac{\partial^2 h}{\partial y^2} \right)$$

213 and  $F(e)$  is the rate of isostatic response to this erosion rate.  $U$  is the uplift rate that  
214 would result in the absence of erosion. We can make this equation more readily  
215 understandable, if we assume local isostasy, in which case, it becomes:

$$\frac{\partial h}{\partial t} = U - \left( 1 - \frac{\rho_s}{\rho_a} \right) e$$

216 We see that when erosion rate is nil, the rate of surface uplift is  $U$  but that the steady  
217 state erosion rate,  $e_{ss}$ , i.e. corresponding to  $\partial h / \partial t = 0$ :

$$e_{ss} = U / \left( 1 - \frac{\rho_s}{\rho_a} \right)$$

218 is larger than  $U$ . Consequently, during a numerical experiment, we expect to see the  
219 erosion rate progressively increase to values that can reach 5-6 times the imposed uplift  
220 rate. Flexure should dampen this progressive increase of the erosion rate as a function  
221 of time.

222

223 3.3 Age predictions

224 As stated earlier, we will constrain our inversion by testing model predictions against  
225 observed cooling ages obtained for a series of samples collected in the Cordillera Blanca  
226 [Montario, 2001; Giovanni, 2007; Hodson, 2012; Margirier et al., 2015]. At each point of  
227 the model where we wish to predict a cooling age, we solve the 1D heat transport  
228 equation by conduction and advection:

229

$$230 \quad (6) \quad \frac{\partial T}{\partial t} + v \frac{\partial T}{\partial z} = \kappa \frac{\partial^2 T}{\partial z^2}$$

231

232 where  $\kappa$  is thermal diffusivity and  $v$  is the erosion rate predicted by the landscape  
233 evolution model. We neglect the effect of radiogenic heat production and assume that  
234 the top and base of the crust/lithosphere are held at constant temperature:  $T(z=0)=0$   
235 and  $T(z=L)=TL$ . The thermal history obtained by solving this equation through time is  
236 then used to compute cooling ages by using standard methods developed to simulate the  
237 annealing of fission tracks and the diffusion of He in apatite [Braun et al., 2006].

238

239 The thermochronological dataset consists of apatite fission-track (AFT) ages and (U-  
240 Th)/He (AHe) mean ages [Montario, 2001; Giovanni, 2007; Hodson, 2012; Margirier et  
241 al., 2015]. The AFT and AHe thermochronological systems record the cooling histories of  
242 rocks below 120°C and 80°C, respectively; at a given location subject to steady  
243 exhumation, AFT age should be older than AHe age. For our inversions we exclude AFT  
244 ages that are older than the age of the emplacement of the Cordillera Blanca batholith  
245 [ $\sim 7$  Ma; Mukasa, 1984; Giovanni, 2007] and the AHe ages that are older than AFT age for  
246 a same sample. Indeed, the AHe ages are scattered, raising the question of their  
247 reliability [Margirier et al., 2015].

248

249 The misfit function used to guide the inversion procedure is defined by:

$$(7) \quad \mu = \frac{1}{N} \sqrt{\sum \frac{(Age_{oi} - Age_{pi})^2}{\partial \Delta Age^2}} + \frac{|Hmax_o - Hmax_p|}{\Delta Hmax} + \frac{|Vmax_o - Vmax_p|}{\Delta Vmax}$$

251 where  $N$  is the number of thermochronological ages,  $Age_{oi}$  and  $Age_{pi}$  the observed and  
252 predicted ages and  $\Delta Age_i$  the uncertainty on the observed ages.  $Hmax_o$  and  $Hmax_p$  are  
253 the observed (present-day) and predicted maximum topography of the Cordillera Blanca  
254 (6600 m) and  $\Delta Hmax$  its regional variability (200 m).  $Vmax_o$  and  $Vmax_p$  are the observed  
255 (present-day) and total exhumed volume of the Cordillera Blanca ( $1 \times 10^{13}$  m<sup>3</sup>), which  
256 we derive from amphibole barometry data [Margirier *et al.*, 2016], and  $\Delta Vmax$  its  
257 assumed uncertainty ( $5 \times 10^{12}$  m<sup>3</sup>).

258

### 259 3.4 Parameterization

260 For each inversion, we ran 15000 models; during the first iteration, we sampled 3000  
261 models to explore the parameter space, then we performed 15 iterations of 800 models  
262 with 400 cells resampled. The runs were performed on a  $200 \times 200$  km square domain  
263 (Fig. 2), discretized by  $500 \times 500$  regularly space nodes. All model runs last 7 Ma, which  
264 is approximately equivalent to the age of the Cordillera batholith emplacement (see  
265 geochronological data compilation in Margirier *et al.*, 2016). We chose to take the  
266 geological ages as references for the modelisations, thus the model runs begin at 7 Ma  
267 and end at 0 Ma. We fixed the time step length to 1 ka (other model parameters are  
268 detailed in the Table A1). We used an arbitrary initial topography including both a  
269 “proto Cordillera Negra” (3500 m) and, for inversions 1b and 2, a “proto Cordillera  
270 Blanca” (3000 m), following estimates from Wise and Noble [2003] and Hoorn *et al.*  
271 [2010]. We imposed a uniform precipitation rate of  $\sim 0.5$  m/yr in the Cordillera Negra  
272 and  $\sim 1.0$  m/yr in the Cordillera Blanca in accordance with present-day estimates from  
273 Bookhagen and Strecker [2008]. According to the Cordillera Blanca batholith lithology

274 and shape [granite/granodiorite with a typical density  $\sim 2800 \text{ kg/m}^3$ ; *Petford and*  
275 *Atherton, 1992; Sharma, 1997*] and the nature of the surroundings sedimentary rocks  
276 [alternating sandstone and marl with sporadic limestone with an estimated density  
277  $\sim 2400 \text{ kg/m}^3$ ; *Bonnot, 1984; Sharma, 1997*], we included a  $150 \times 20 \text{ km}$  ellipsoidal  
278 intrusion that is characterized by an anomalously high density ( $400 \text{ kg/m}^3$  heavier than  
279 the surrounding rocks). The top of the intrusion is initially at  $1.5 \text{ km}$  beneath sea level  
280 ( $\sim 3 \text{ km}$  depth). This is constrained by amphibole barometry data obtained from the roof  
281 of the Cordillera Blanca batholith [*Margirier et al., 2016*]. We fixed a regional uplift rate  
282  $U_0 = 0.2 \text{ mm/yr}$  in agreement with the thermal history provided by *Margirier et al.*  
283 [2015] for the Cordillera Negra. In addition to the widespread uplift ( $U_0$ ) we simulated  
284 the presence of an active fault by imposing a steep gradient in vertical displacement rate  
285 along a north-south boundary located a distance of  $100 \text{ km}$  from the left side of the  
286 model (Fig. 2). Along the boundary the uplift rate is equal to  $U_1$  and it decreases linearly  
287 to reach a value of  $0$  at  $25 \text{ km}$  away from the boundary. The uplift history along this  
288 boundary is divided into 3 successive episodes based on observations concerning the  
289 initiation of the CNBF and an apparent increase in erosion rate evidenced by *Margirier*  
290 *et al.* [2016]. From  $7$  to  $5.4 \text{ Ma}$  (before the CBNF initiation) the uplift rate ( $U_1$ ) is equal to  
291  $0$ ; between  $5.4$  and  $2 \text{ Ma}$ , the uplift rate is constant (and equal to  $U_1$ ); between  $2 \text{ Ma}$  and  
292 the present, the uplift rate is increased by a factor  $f_{U1}$ . *Perez-Gussinye et al.* [2009]  
293 estimated elastic thickness ( $T_e$ ) to be in the range  $[0-10] \text{ km}$ , which are relatively low  
294 values for continental regions; we use a value of  $3 \text{ km}$  in inversions where the elastic  
295 thickness is fixed. The four parameters that we want to constrain by inversion of the  
296 thermochronological ages and the range-scale present-day topography and eroded  
297 volume in our inversions, are the erosion efficiency ( $K_f$ ), the mean uplift rate ( $U_1$ ), the  
298 magnitude of the recent increase of uplift rate in the Cordillera Blanca ( $f_{U1}$ ), the

299 temperature at the base of the model ( $T_{max}$ ) and, in a third set of inversions, the elastic  
300 thickness of the lithosphere ( $T_e$ ).

301

## 302 **4. Results**

### 303 4.1 Role of tectonics, erosion and initial topography

304 Here we aim to test (i) if the increase of erosion rate suggested for the last 2 Ma in the  
305 Cordillera Blanca [Margirier *et al.*, 2016] could be due to an increase in uplift rates since  
306 2 Ma rather than a change of climate and/or erosion process (glacial erosion vs fluvial  
307 erosion) and (ii) the role of initial topography in the present day Cordillera Blanca  
308 drainage divide location. In these inversions, the four parameters that we wanted to  
309 constrain by inversion of the thermochronological ages, the total exhumation and the  
310 range-scale present-day topography are the erosion efficiency ( $K_f$ ), the mean uplift rate  
311 ( $U_1$ ), the magnitude of the recent increase of uplift rate in the Cordillera Blanca ( $f_{U1}$ ) and  
312 the temperature at the base of the model ( $T_{max}$ ).

313 We performed two inversions, one without an initial topography in the Cordillera Blanca  
314 and a second one with a “proto Cordillera Blanca”, in order to evaluate the  
315 paleogeography and inherited drainage network influence on the location of the present  
316 day drainage divide in the Cordillera Blanca.

317

#### 318 4.1.1 Inversion 1a (without a “proto Cordillera Blanca”)

319 Figure 3 presents the results of inversion 1a (without a “proto Cordillera Blanca”) as  
320 values of the misfit function,  $\mu$ , displayed in parameter space. In each of the two panels,  
321 each dot represents a single model run and the color scale refers to the value of the  
322 misfit of the model (from blue for high misfit value or poor fit to the data, to red for low

323 misfit value or good fit to the data). In each of the panels, the red star indicates the  
324 position of the best-fitting model in parameter space.

325 The mean uplift rate ( $U_1$ ), the amplitude of its recent increase ( $f_{U1}$ ) and the erodibility  
326 ( $K_f$ ) are all well constrained by the thermochronological data (Fig. 3) although a trade-  
327 off exists (as shown in figure 3B) between  $U_1$  and  $f_{U1}$ . These two parameters are clearly  
328 correlated with higher mean uplift rate requiring a smaller recent increase in uplift rate  
329 and vice-versa. The arcuate shape of the region of minimum misfit in Figure 3B can also  
330 be interpreted by stating that the product  $U_1 \times f_{U1}$  (or in other words the uplift rate  
331 between 2 and 0 Ma) is very well constrained by the data at a value of approximately 1.8  
332 mm/yr (the best fitting parameters are indicated by a red star on figures 3A and 3B).  
333 The temperature at the base of the model is not as well constrained (Fig. 3A) although  
334 values between 650 and 850°C yield the lowest misfit values. These values should be  
335 interpreted as suggesting best fitting surface geothermal gradient values of 21 to  
336 28°C/km. The inversion results suggest that the data is best explained by values of  $K_f =$   
337  $1.9 \times 10^{-5} \text{ m}^{1-2\text{m}}.\text{yr}^{-1}$ ,  $U_1 = 0.80 \text{ mm/yr}$ ,  $f_{U1} = 1.8$  and  $T_{max} = 677^\circ\text{C}$ . The set of the good  
338 fitting value of  $f_{U1}$  (misfit < 0.6) indicates that a substantial increase in erosion rate in the  
339 recent past is necessary to explain the age dataset and the total exhumation. The best  
340 value for the basal temperature ( $T_{max} = 677^\circ\text{C}$ ) corresponds to a geothermal gradient of  
341  $\sim 23^\circ\text{C/km}$ , which is consistent with *Henry and Pollack* [1988] terrestrial heat flow  
342 measurements.

343 The misfit is 0.53 for the best fitting parameters model, suggesting that the model can  
344 reproduce the data (present-day topography, total exhumation and thermochronological  
345 ages, Fig. 3). In addition, the thickness of rock eroded in the model after 7 Ma  
346 reproduces the total eroded volume we fixed to correspond with field observations and  
347 granite emplacement depth obtained by *Margirier et al.* [2016]. However, the apatite  
348 fission track ages are not well reproduced by the best fitting parameters model, the

349 apatite (U-Th)/He ages predicted by the model are too young (Fig. 3C, D). Moreover, the  
350 best fitting parameters model does not reproduce all characteristic features of the  
351 present-day Cordillera Blanca topography. Notably, the drainage divide of the predicted  
352 topography is located close to the CBNF, which is not the case in the Cordillera Blanca  
353 (Figs. 1, 4).

354

#### 355 *4.1.2 Inversion 1b (with a “proto Cordillera Blanca”)*

356 We performed a second inversion (inversion 1b) using the same constraints as for  
357 inversion 1a but with a finite initial topography in the Cordillera Blanca (see initial  
358 topography on Figure 2). Figure 5A, B presents the results of the inversion as values of  
359 the misfit function,  $\mu$ , displayed in parameter space. The misfit is 0.44 for the best fitting  
360 parameters model, indicating that the model can reproduce the present-day topography,  
361 total exhumation and thermochronological ages (Fig. 5C, D) as well as the best model of  
362 inversion 1a.

363 The mean uplift rate ( $U_1$ ), the amplitude of its recent increase ( $f_{U1}$ ) and the erosion  
364 efficiency ( $K_f$ ) are all well constrained by our data. Similarly to inversion 1a,  $U_1$  and  $f_{U1}$   
365 are correlated with higher mean uplift rate requiring a smaller recent increase in uplift  
366 rate and vice-versa. The arcuate shape of the region of minimum misfit in Figure 5B can  
367 also be interpreted by stating that the product  $U_1 \times f_{U1}$  is very well constrained by the  
368 data at a value of approximately 1.0 mm/yr. The temperature at the base of the model  
369 ( $T_{max}$ ) is not as well constrained (Fig. 5A, B) although values between 850 and 1050°C  
370 yield the lowest misfit values. These values should be interpreted as suggesting best  
371 fitting surface geothermal gradient values of 28 to 35°C/km. The inversion result  
372 suggests that the data is best explained by values of  $K_f = 1.3 \times 10^{-5} \text{ m}^{1-2\text{m}}.\text{yr}^{-1}$ ,  $U_1 = 0.52$   
373 mm/yr,  $f_{U1} = 2.0$  mm/yr and  $T_{max} = 945^\circ\text{C}$ . The best value for the basal temperature ( $T_{max}$   
374 = 945°C) corresponds to a geothermal gradient of  $\sim 32^\circ\text{C}/\text{km}$ .



375 Figure 6 presents the topographic evolution and the uplift rates predicted in the  
376 Cordillera Blanca using the best fitting parameters. Predicted uplift rates obtained at the  
377 end of the model are higher than the imposed uplift rate due to isostasy (Figs. 5, 6).  
378 Considering the low elastic thickness of the lithosphere used in this model ( $T_e = 3 \text{ km}$ ),  
379 according to existing estimate of the elastic thickness of the lithosphere in the Cordillera  
380 Blanca region [0-10 km; *Perez-Gussinye et al., 2009*], the flexural rebound is substantial.  
381 As explained in the method section, if we chose to ignore the flexural rebound due to  
382 erosion, the uplift rates would correspond to the vertical displacement (i.e. tectonic  
383 forcing) imposed on the fault ( $U_1 = 0.52 \text{ mm/yr}$  and  $U_1 \times f_{U1} = 1.0 \text{ mm/yr}$ ) but if we  
384 consider the flexural rebound associated with the erosion, the uplift rates significantly  
385 increase with time (up to  $\sim 2 \text{ mm/yr}$ ; Fig. 6). The predicted uplift rates at the end of the  
386 model are consistent with the Quaternary vertical slip rates on the CBNF ranging from  
387  $5.1 \pm 0.8 \text{ mm/yr}$  to  $0.6 \pm 0.2 \text{ mm/yr}$  [*Schwartz, 1988; Siame et al., 2006; Margirier et al.,*  
388 *2017*] and with  $^{10}\text{Be}$  catchment-wide erosion rates in the Cordillera Blanca [mean of  $0.7$   
389  $\text{mm/yr}$ ; *Hodson, 2012*].

390

#### 391 4.1.3 Comparison between the results of the two inversions

392 Best fitting parameters for inversion 1b are in the same order of magnitude as for  
393 inversion 1a. However, we obtained a larger value for the product  $U_1 \times f_{U1}$  for the  
394 inversion 1a, indicating larger total uplift in the Cordillera Blanca since 5.4 Ma. This can  
395 be easily explained by the lower initial elevation in the inversion without an initial  
396 topography; thus to reach the present day elevation of the Cordillera Blanca, the total  
397 uplift has to be larger.

398 The most striking difference between these two inversions is that without a “proto  
399 Cordillera Blanca” (inversion 1a), even if the predicted elevations are well correlated  
400 with the observed topographic bulge in the Cordillera Blanca, the model still does not

401 reproduce the present-day location of the main drainage divide (Fig. 7). Notably, for the  
402 model without initial relief in the Cordillera Blanca the drainage divide is located close  
403 to the CBNF, which is not the case for the present-day topography. The existence of an  
404 early relief, the “proto Cordillera Blanca”, probably controls the early drainage network  
405 and localizes the drainage divide farther east in the present day Cordillera Blanca. Our  
406 results suggest that the presence of an initial relief is needed to localize the drainage  
407 divide at its present position (eastern part of the Cordillera Blanca), although we cannot  
408 preclude a different scenario.

409

#### 410 4.2 Role of the granite

411 In the Cordillera Blanca the granite reached the surface and starts to be eroded at ~3 Ma  
412 [Bonnot, 1984]. This enables us to explore the influence of eroding denser rocks on the  
413 evolution of topography and uplift rates on a Ma time scale. We performed additional  
414 inversions (2a and 2b) of the topography, the total exhumation and  
415 thermochronological ages to assess the role of the intrusion in the recent increase of  
416 uplift rates in the Cordillera Blanca and in high topography building. In these two  
417 inversions the initial topography includes a proto Cordillera Blanca and we fixed  $f_{U1} = 1$   
418 (no uplift increase after 2 Ma) in order to test if the increase of the uplift rate due to the  
419 batholith erosion could explain by itself the increase of erosion rates observed by  
420 [Margirier *et al.*, 2016]. The inversion 2a didn't include an intrusion. Then, the inversion  
421 2b included an intrusion that is characterized by an anomalously high density (400  
422 kg/m<sup>3</sup> heavier than the surrounding rocks). In these inversions the parameters that we  
423 aimed to constrain were  $K_f$ ,  $T_e$ ,  $U_1$ , and  $T_{max}$ . Figure 8 presents the results of the  
424 inversions 2a and 2b as values of the misfit function,  $\mu$ , displayed in parameter space.

425

##### 426 4.2.1 Inversion 2a (without contrasting density intrusion)

427 The misfit of inversion 2a is 0.42 for the best fitting parameters model, indicating that  
428 the model can reproduce the present-day topography, total exhumation and  
429 thermochronological ages (Fig. 8D). In this inversion without intrusion, the mean uplift  
430 rate ( $U_1$ ) and the erodability ( $K_f$ ) are also well constrained whereas the temperature at  
431 the base of the model ( $T_{max}$ ) and the elastic thickness of the crust ( $T_e$ ) are not as well  
432 constrained (Fig. 8A, B, C), although values between 750 and 1050°C and 3 and 7 km  
433 respectively yield the lowest misfit values. These values should be interpreted as  
434 suggesting best fitting surface geothermal gradient values of 25 to 35 °C/km. The data is  
435 best explained by very similar values of  $K_f = 1.2 \times 10^{-5} \text{ m}^{1-2\text{m}}.\text{yr}^{-1}$ ,  $T_e = 5.6 \text{ km}$ ,  $U_1 = 0.86$   
436  $\text{mm/yr}$ , and  $T_{max} = 1001^\circ\text{C}$ . The best value for the basal temperature ( $T_{max} = 1001^\circ\text{C}$ )  
437 corresponds to a geothermal gradient of  $\sim 33^\circ\text{C/km}$ .

438

#### 439 *4.2.2 Inversion 2b (with contrasting density intrusion)*

440 The lowest misfit is 0.42 for inversion 2b, suggesting that the model can also reproduce  
441 the present-day topography, total exhumation and thermochronological ages (Fig. 8). In  
442 this inversion that includes a dense intrusion, the mean uplift rate ( $U_1$ ) and the erosion  
443 efficiency ( $K_f$ ) are all well constrained (Fig. 8E, F, G). The temperature at the base of the  
444 model ( $T_{max}$ ) and the elastic thickness of the crust ( $T_e$ ) are not as well constrained  
445 although values between 750 and 1050°C and 3 and 5 km respectively yield the lowest  
446 misfit values (Fig. 8E, F). These  $T_{max}$  values should be interpreted as suggesting best  
447 fitting surface geothermal gradient values of 25 to 35°C/km.  $U_1$  and  $T_e$  are clearly  
448 correlated with smaller elastic thickness of the crust requiring smaller uplift rate and  
449 vice-versa. The inversion result suggests that the data is best explained by values of  $K_f =$   
450  $1.2 \times 10^{-5} \text{ m}^{1-2\text{m}}.\text{yr}^{-1}$ ,  $T_e = 3.3 \text{ km}$ ,  $U_1 = 0.69 \text{ mm/yr}$ , and  $T_{max} = 954^\circ\text{C}$ . The best value for  
451 the basal temperature ( $T_{max} = 954^\circ\text{C}$ ) corresponds to a geothermal gradient of  
452  $\sim 32^\circ\text{C/km}$ .

453

#### 454 4.2.3 Comparison between the results of the two inversions

455 For these two inversions the predicted elevations are well correlated with the observed  
456 topographic bulge in the Cordillera Blanca. The models also reproduce the present-day  
457 location of the main drainage divide and the thermochronological ages (Fig. 8D, H). The  
458 best fitting  $K_f$ ,  $T_e$  and  $T_{max}$  are not significantly different for the best models 2a and 2b.  
459 For both inversions the best fitting  $T_e$  (~4 km) is consistent with *Perez-Gussinye et al.*  
460 (2009) predictions at a larger scale. However, the good fitting values of  $U_1$  (misfit < 0.6)  
461 are significantly smaller for the inversion 2b (with an intrusion) than 2a, especially if the  
462 elastic thickness decreases below 4 km (Fig. 7). Figure 9 shows the predicted  
463 topography and uplift rate at the end of the best fitting parameters models for inversion  
464 2a (without intrusion) and 2b (with an intrusion). Taking isostasy into account, the two  
465 best fitting parameters models reach a maximum uplift rate of 1.7 mm/yr at the end of  
466 the run. The maximum uplift rates obtained from our models at the end of the runs are  
467 similar to the Quaternary vertical slip rates documented along the CBNF and are in the  
468 same order of magnitude than  $^{10}\text{Be}$  catchment-wide erosion rates obtained in the  
469 Cordillera Blanca [*Schwartz, 1988; Siame et al., 2006; Hodson, 2012; Margirier et al.,*  
470 *2017*]. Since the predicted maximum uplift rates are similar at the end of the two models  
471 which take isostasy into account (1.7 mm/yr; Fig. 9C, D), the difference between the  
472 good fitting  $U_1$  values (misfit < 0.6) can be explained by the more important contribution  
473 of isostasy for the model including a dense intrusion.

474

## 475 **5. Discussion**

### 476 5.1 Paleogeography

477 The substantial surface uplift and resulting erosion in the Cordillera Blanca make it  
478 difficult to study the paleogeography in this area based on remnants geomorphological  
479 features. However, the drainage network geometry and the drainage divide location  
480 provide information on the topography of the Cordillera Blanca in the past and its  
481 evolution. Notably, the Cordillera Blanca drainage divide is located in the eastern part of  
482 the range whereas both the higher peaks and higher uplift rates are located in the  
483 western flank of the Cordillera Blanca [Margirier *et al.*, 2016]. But the precipitation  
484 gradient, with higher mean precipitation on the eastern flank of the Cordillera Blanca,  
485 should favor the location of the drainage divide further west. Our results suggest that  
486 the location of the drainage divide in the eastern part of the Cordillera Blanca is  
487 controlled by an initial topography in the Cordillera Blanca area and an inherited  
488 drainage network. Based on these results we propose that a proto-Cordillera Blanca  
489 existed before the CBNF initiation ~5 Myr ago. The presence of such a relief is in  
490 agreement with *Giovanni et al.* [2010] and *Wise and Noble* [2003], which already  
491 proposed that the Callejón de Huaylas was a topographic depression during the Late  
492 Miocene. *Wise and Noble* [2003] suggested that at that time, before the Cordillera Blanca  
493 exhumation, the Cordillera Negra corresponded to the drainage divide. However, the  
494 presence of a proto-Cordillera Blanca calls into question the position of the Andean  
495 drainage divide during the Miocene.

496

## 497 5.2 The Cordillera Blanca uplift: role of granite erosion and flexural rebound

498 Our results show a difference in imposed uplift rates for the best fitting models of the  
499 inversions 2a and 2b that do not or do include a granite ( $U_1=0.86$  mm/yr versus  $U_1=0.69$   
500 mm/yr; Fig. 8). In addition, our models do not take into account the presence of the  
501 CBNF for the isostatic response calculation despite the existence of a 200 km-long  
502 crustal fault [Dalmayrac and Molnar, 1981; McNulty and Farber, 2002]. If, to reflect the

503 regional geology, we had used a broken plate approximation to calculate the isostasy  
504 contribution to uplift and erosion, the isostatic response would, potentially, have been  
505 more important. In addition, even when considering the best fitting value for  $T_e$  the role  
506 of the granite seems to be small; if we consider the range of acceptable values for  $T_e$ , we  
507 notice that smaller values of  $T_e$  would increase the importance of isostasy, which, in  
508 turn, would require smaller valued of imposed uplift,  $U_1$  (Fig. 8E). For example, if we  
509 consider the best fitting models or the models involving a 3 km elastic thickness which  
510 also show low misfit values, the isostasy contributes c.a. 20% of the total rock uplift (Fig.  
511 8B, F). Considering that for the best fitting parameters model of the inversion 2b, the  
512 granite reaches the surface approximately 3 Ma before the end of the model run, our  
513 results indicate that the isostatic effect of eroding a denser rock mass is not negligible  
514 even on a  $< 5$  Ma time scale. Finally, we note that the inversion that includes a denser  
515 granite (inversion 2b) converges toward smaller values for the imposed uplift rate,  $U_1$ ,  
516 than the inversion which doesn't include an intrusion (inversion 2a) implying that  
517 erosion of the Cordillera Blanca dense intrusion may substantially contribute to the  
518 present-day higher uplift rates as proposed by *Braun et al.* [2014].

519 Interestingly, our results suggest that the recent increase in exhumation rate  
520 documented by *Margirier et al.* [2016] might not be the result of a change in tectonic  
521 since 2 Ma or climate forcing, but the result of the progressive flexural isostatic response  
522 of the area to erosion. At the onset of uplift, slopes are relatively low, erosion is limited  
523 to the sides of the uplifting region and the weight of uplifted material acts as a negative  
524 feedback to further uplift; as the uplifting region becomes affected by surface processes  
525 (fluvial erosion in our model, but more likely glacial erosion in the Cordillera Blanca) the  
526 resulting erosional unloading reduces the negative feedback and uplift and erosion rates  
527 increase until a steady-state situation is reached between uplift, erosion and isostasy.

528 The time scale over which this steady-state balance is reached must be of the order of a  
529 few million years, at least [*Whipple and Meade, 2006*].

530 From the time when the Cordillera Blanca summits reached high elevation, the  
531 Cordillera Blanca itself has acted as an orographic barrier for moisture carried from the  
532 Amazon basin [*Montgomery et al., 2001; Bookhagen and Strecker, 2008*]. The induced  
533 rain shadow included in our model prevents erosion of the Cordillera Negra, and results  
534 in higher erosion rates in the Cordillera Blanca [i.e. *Montgomery and Brandon, 2002*] in  
535 the recent past. It is also likely that Quaternary glaciations [ $\sim 1.5\text{-}0$  Ma; *Farber et al.,*  
536 *2005; Smith et al., 2005*] have increased erosion rates through the formation of deeply  
537 incised U-shaped valleys, as evidenced by *Montgomery* [2002] for the Olympic  
538 Mountains on the Pacific coast of the North America and suggested by *Margirier et al.*  
539 [2016] for the Cordillera Blanca. Finally, the recent glacial retreat [since  $\sim 21$  ka; *Seltzer*  
540 *et al., 2002; Farber et al., 2005; Smith et al., 2005*] could have induced a flexural rebound  
541 that both increased uplift rates in the Cordillera Blanca and slip-rates on the CBNF [e.g.,  
542 *Hetzl and Hampel, 2005; Hampel et al., 2007*].

543

### 544 5.3 Models for the CBNF

545 Two models have been proposed to explained extension on the CBNF. *Dalmayrac and*  
546 *Molnar* [1981] proposed that the fault is the result of a gravitational collapse of the  
547 thickened crust, whereas *McNulty and Farber* [2002] suggested that the subduction of  
548 the buoyant Nazca Ridge below the Cordillera Blanca drove the footwall uplift. However,  
549 there is still no consensus on a model to explain the CBNF and the processes driving the  
550 Cordillera Blanca uplift. *Margirier et al.* [2015] already provided evidence for  
551 exhumation in the Cordillera Negra since 15 Ma, which is likely to be related to a  
552 regional surface uplift of the Western Andes [*McLaughlin, 1924; Farrar and Noble, 1976,*  
553 *Myers, 1976; Wipf, 2006; Hoorn et al., 2010*]. The 1300 m-thick sedimentary filling of the

554 Callejón de Huaylas suggests only low subsidence of the basin since  $\sim 5.4$  Ma [Bonnot,  
555 1984; Giovanni *et al.*, 2010] even if the vertical displacement on the CBNF is estimated to  
556 be  $\sim 7$  km [Margirier *et al.*, 2016]. In addition, our models of the landscape evolution  
557 provide new constraints on uplift rates in the Cordillera Blanca region. We suggest that  
558 the Cordillera Blanca has been uplifted in relation to the Cordillera Negra during the  
559 past 7 Ma. Observations clearly suggest that the displacement on the CBNF is mostly the  
560 result of the Cordillera Blanca uplift rather than subsidence of the hanging wall. The  
561 rock uplift evidenced in this study for the Cordillera Blanca (7-0 Ma) and proposed in  
562 the Cordillera Negra [15-0 Ma; Margirier *et al.*, 2015] is not compatible with the collapse  
563 model proposed by Dalmayrac and Molnar [1981]. Moreover, new reconstructions of the  
564 timing (15-11 Ma) and location of the initial Nazca Ridge subduction [Hampel, 2002;  
565 Rosenbaum *et al.*, 2005; Antonijevic *et al.*, 2015] are not compatible with the timing and  
566 location of normal faulting in the Cordillera Blanca [5.4 Ma; Bonnot, 1984; Giovanni *et al.*,  
567 2010]. Indeed, even if the Nazca Ridge subduction initiate at  $11^\circ\text{S}$ , the latitude of the  
568 Cordillera Blanca between 15 and 11 Ma, at  $\sim 5$  Ma, the time of initiation of normal  
569 faulting the ridge is farther south ( $\sim 13^\circ\text{S}$ ) and at the present day the ridge is at  $15^\circ\text{S}$   
570 whereas the CBNF is still active. This suggests that the Nazca Ridge did not trigger the  
571 Cordillera Blanca uplift as proposed by McNulty and Farber [2002]. Our landscape  
572 evolution models do not permit to test the role of geodynamic processes such as slab  
573 flattening on the CBNF initiation and Cordillera Blanca uplift. However, the subduction  
574 geometry and change of the mechanical coupling on the subduction interface may have  
575 had an influence on the tectonic regime of the upper plate as suggested by Margirier *et*  
576 *al.* [2017] for the Cordillera Blanca region. After the initiation of normal faulting, when  
577 the Cordillera Blanca batholith started to be eroded, since  $\sim 3$  Ma, the erosion of denser  
578 rocks is likely to have trigger an increase of the uplift rates along the CBNF.

579



## 580 **6. Conclusions**

581 Our study provides new constraints on the erosion efficiency, elastic thickness of the  
582 lithosphere, temperature gradient in the crust and uplift rates in the Andes of northern  
583 Peru. The absolute rock uplift rates obtained at the end of the models for the Cordillera  
584 Blanca (ranging from 1.5 to 2.5 mm/yr) are coherent with Quaternary slip rates  
585 documented on the CBNF [ $5.1 \pm 0.8$  mm/yr to  $0.6 \pm 0.2$  mm/yr; *Schwartz, 1988; Siame et*  
586 *al., 2006; Margirier et al., 2017; Gérard et al., in prep*]. Our results suggest an  
587 acceleration of rock uplift in the Cordillera Blanca at 2 Ma in agreement with the  
588 increase in erosion rate evidenced by *Margirier et al. [2016]*. We show, however, that  
589 such an increase in erosion rate may be the result of erosional unloading and isostasy  
590 rather than being related to a change in either tectonic or climatic forcing. We also show  
591 that this acceleration may have been amplified by the unroofing of a dense granitic  
592 intrusion.

593 Based on the present-day drainage divide location in the Cordillera Blanca and the  
594 results of our modeling, we propose that the Callejón de Huaylas was already a  
595 depression before the Cordillera Blanca batholith emplacement and that a proto  
596 Cordillera Blanca already existed at that time.

597 Finally, in the light of our modeling of the landscape evolution, we show that the two  
598 models previously published for the CBNF [e.g., gravitational collapse of the thickened  
599 crust / footwall uplift due to the subduction of the Nazca Ridge; *Dalmayrac and Molnar,*  
600 *1981; McNulty and Farber, 2002*] are inconsistent with the thermochronological data,  
601 geodynamic and geologic context. Further investigations are needed to assess if the  
602 Quaternary extensional tectonics have been triggered by flat slab subduction, which may  
603 also have triggered at the same time the uplift of the Cordillera Blanca. We suggest,  
604 however, that, in this context, the construction of high-relief topography was strongly

605 influenced by the isostatic flexural rebound and, to some degree, by the exhumation of  
606 the dense Cordillera Blanca granite.

607

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614

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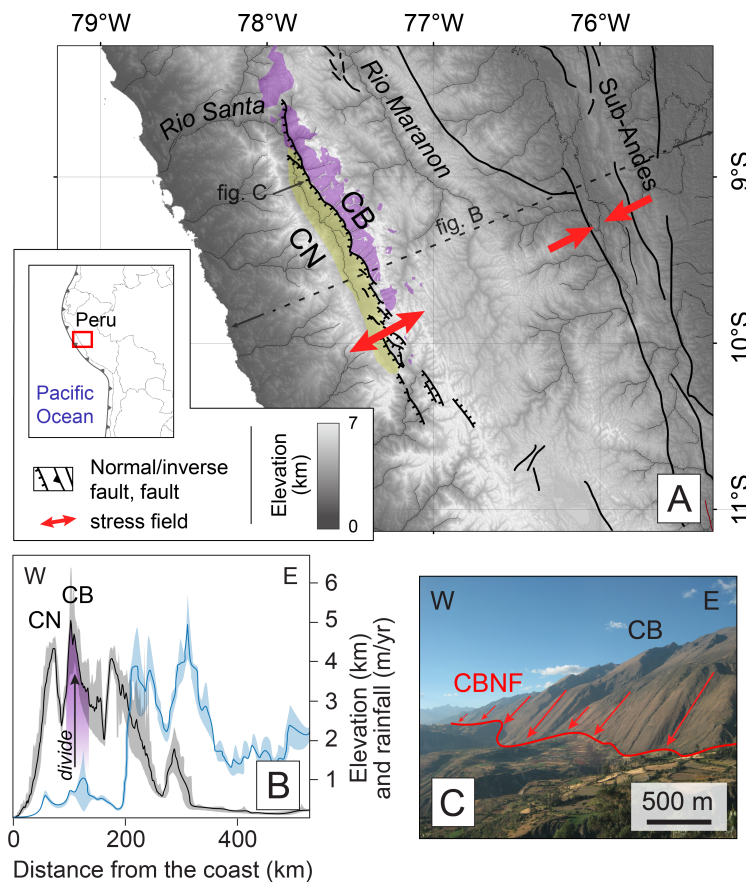
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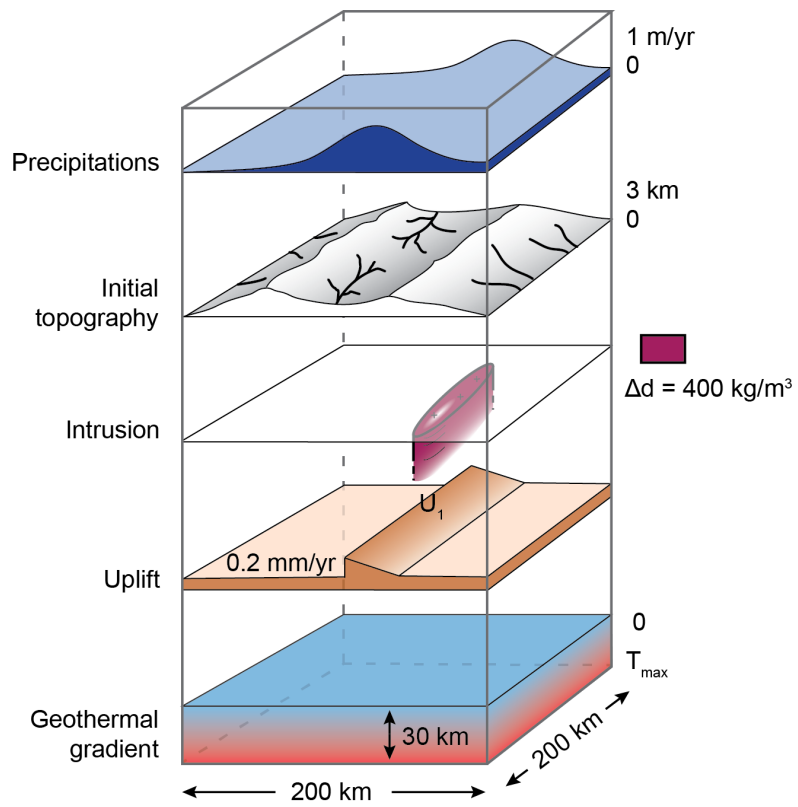
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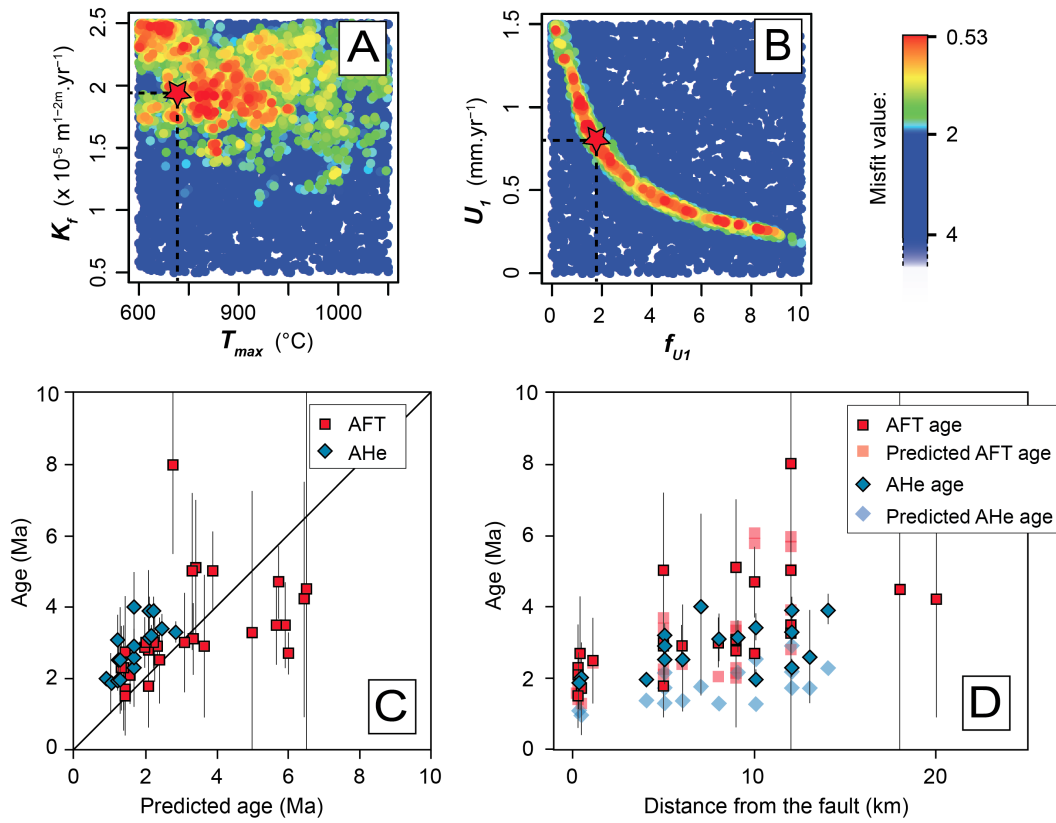
769 Figure 1: A) Topographic map of the northern Peru showing active tectonic features (Neotectonic Open Database,  
770 neotec-opendata.com), location of the Cordillera Blanca (CB, the batholith is highlighted in purple), the Cordillera  
771 Negra (CN) and the Callejón de Huaylas (yellow). Inset shows map location within the South America. B) E-W cross-  
772 section of the Andes at the latitude of the Cordillera Blanca both the topography (black line) and the rainfall (blue  
773 line) are represented [modified from Bookhagen and Strecker, 2008]. On the topographic cross-section the Cordillera  
774 Blanca batholith is represented in purple and a black arrow points the drainage divide location. C) Photograph of the  
775 CBNF showing the 1 km high triangular facets along the active fault scarp.

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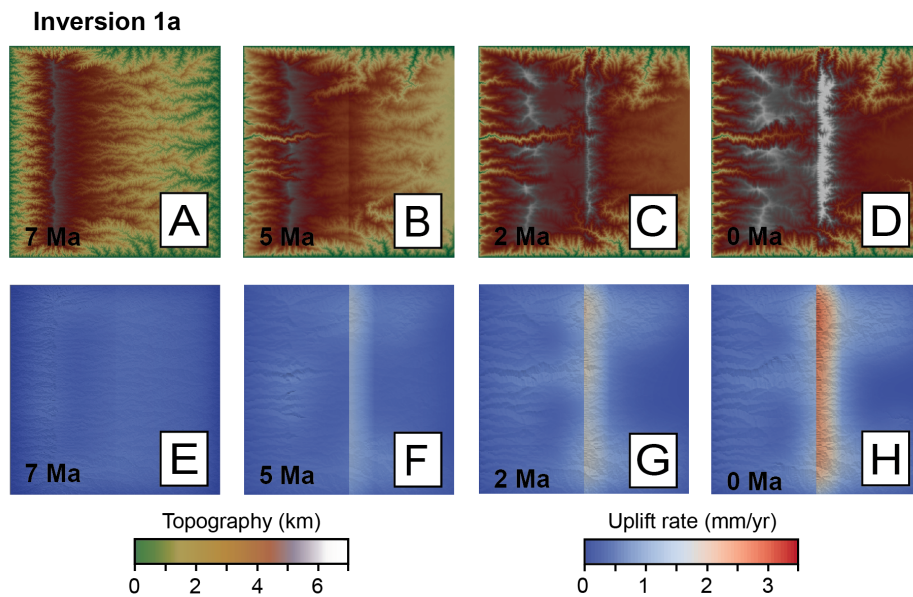
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 778 Figure 2: Parameterization of the models.  
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**Inversion 1a** without a «proto Cordillera Blanca»



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Figure 3: Results of the NA inversion 1a. A, B) Scatter plots colored by likelihood values for erosion efficiency ( $K_f$ ), basal temperature ( $T_{max}$ ), uplift rate ( $U_1$ ) and factor of uplift rate increase for the last 2 Ma ( $f_{U1}$ ). The most likelihood model is indicated with a red star. C) Comparison between observed ages and synthetic ages predicted by the best-fit model for apatite fission-track ages and apatite (U-Th)/He ages. D) Observed ages and synthetic ages vs. distance to the Cordillera Blanca normal fault.

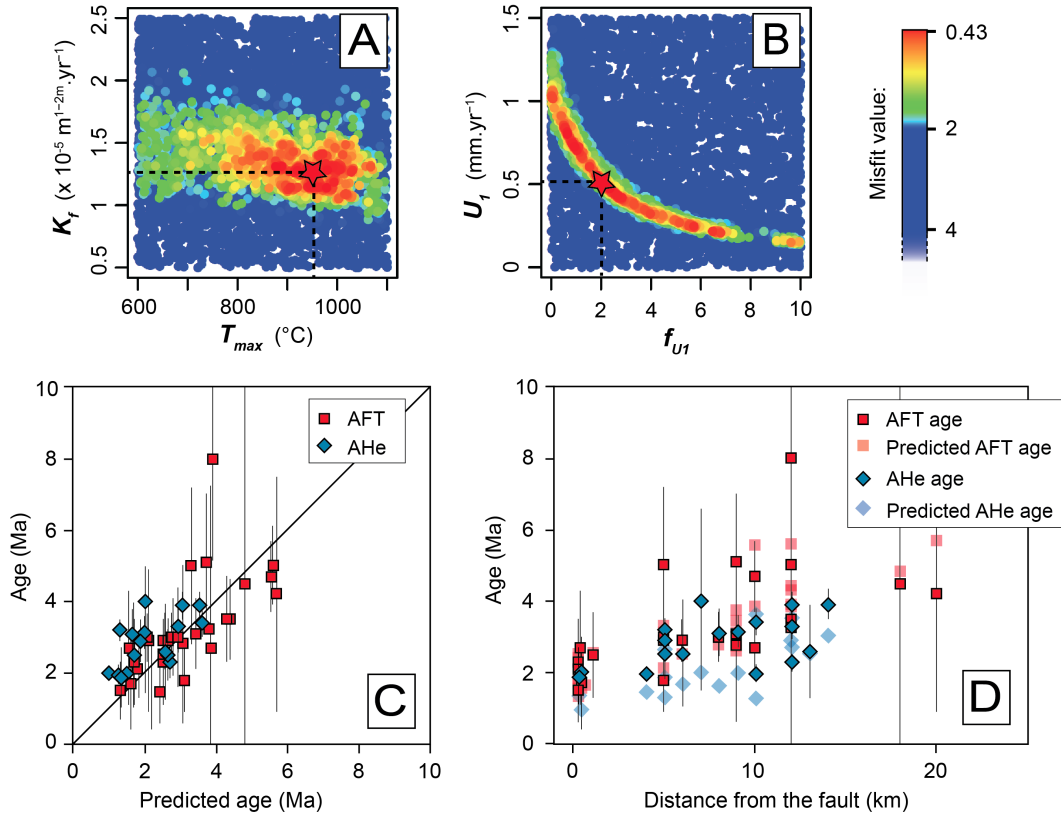


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Figure 4: Evolution of the topography and rock uplift rate for the best fitting parameter model of inversion 1a. A-D) Topography, E-H) Rock uplift rate.

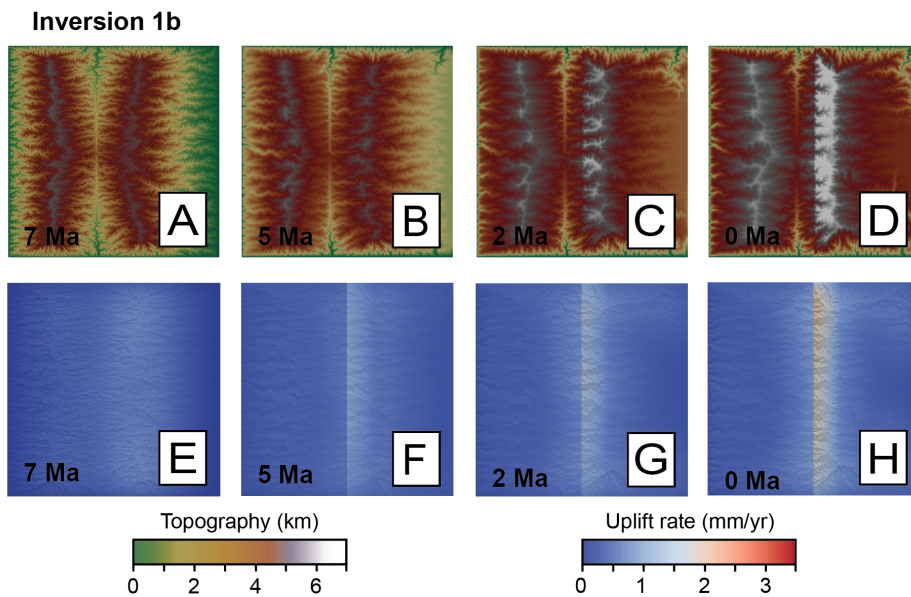


**Inversion 1b** with a «proto Cordillera Blanca»



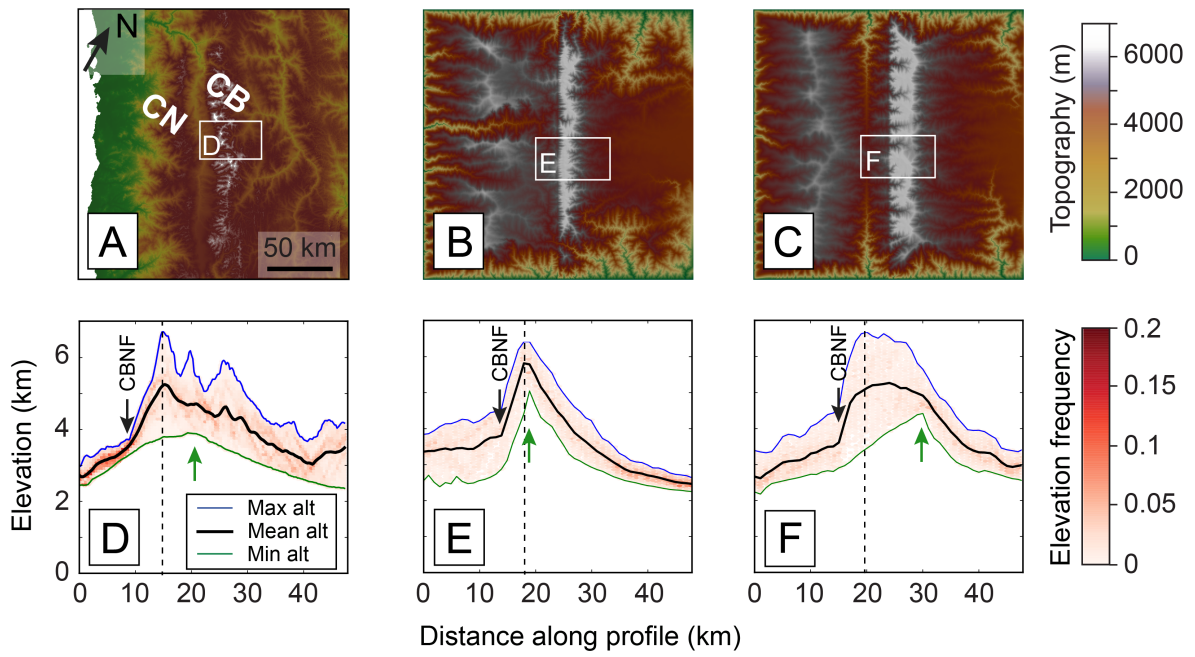
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Figure 5: Results of the NA inversion 1b. A, B) Scatter plots colored by likelihood values for erosion efficiency ( $K_f$ ), basal temperature ( $T_{max}$ ), uplift rate ( $U_1$ ) and factor of uplift rate increase for the last 2 Ma ( $f_{U1}$ ). The most likelihood model is indicated with a red star. C) Comparison between observed ages and synthetic ages predicted by the best-fit model for apatite fission-track ages and apatite (U-Th)/He ages. D) Observed ages and synthetic ages vs. distance to the Cordillera Blanca normal fault.



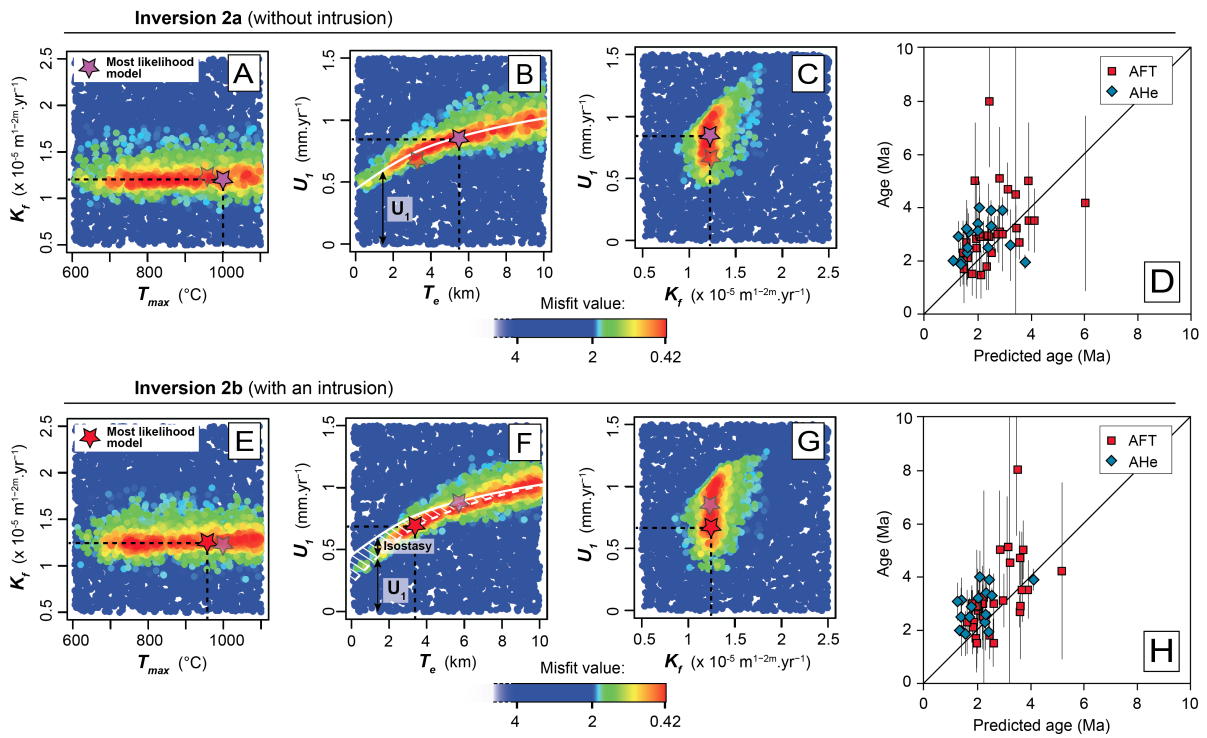
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Figure 6: Evolution of the topography and rock uplift rate for the best fitting parameter model of inversion 1b. A-D) Predicted topography, E-H) Map of the predicted rock uplift rate.

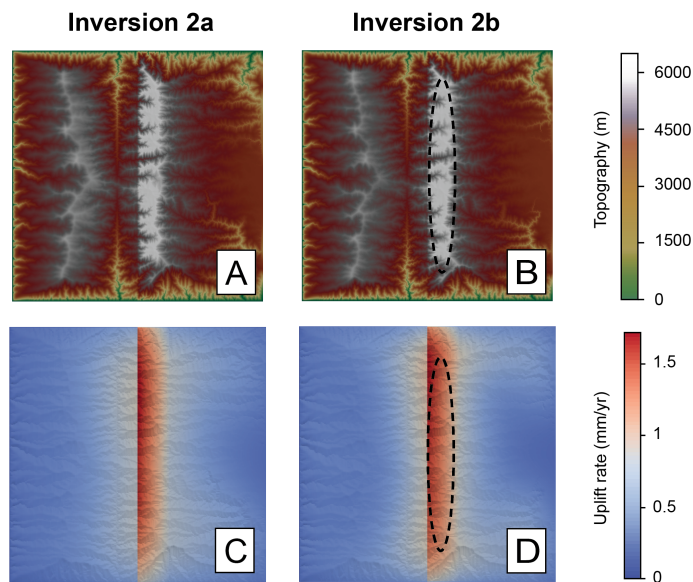


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Figure 7: DEM and swath profiles through the Cordillera Blanca region and modeled topography. A) DEM of the Cordillera Blanca region. B) DEM of the best model of inversion 1a, which involves a fault, a granitic batholith, a flexural rebound and a “proto” Cordillera Negra. C) DEM of the best model of inversion 1b with both a “proto” Cordillera Negra and a “proto” Cordillera Blanca. D-F) Swath profiles through the Cordillera Blanca and from the two models showing elevation (mean, maximum and minimum), the drainage divide position (green arrow), the CBNF location (black arrow) and the maximum elevation (vertical dot line).



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 817 Figure 8: Results of the NA inversions 2a and 2b as scatter plots colored by likelihood values (from blue for high misfit  
 818 value or poor fit to the data, to red for low misfit value or good fit to the data) for erosion efficiency ( $K_f$ ), basal  
 819 temperature ( $T_{max}$ ), uplift rate ( $U_i$ ) and elastic thickness of the crust ( $T_e$ ). The most likelihood models are indicated  
 820 with purple stars (inversion 2a) and red stars (inversion 2b). A-C Scatter plots and D) comparison between observed  
 821 thermochronological ages and synthetic ages predicted by the best-fit model for the inversion 2a (without intrusion).  
 822 E-G Scatter plots for inversion 2b (with an intrusion). H) Comparison between observed thermochronological ages  
 823 and synthetic ages predicted by the best-fit model for the inversion 2b.  
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 827 Figure 9: Map of the topography at the end of the model A) without granite and B) with a dense granite. Map of the  
 828 uplift rates at the end of the model C) without granite and D) with a dense granite. The oval dotted line indicates the  
 829 location of the granite.