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1 Observations of guided waves from the Pamir seismic zone provide additional evidence for
2 the existence of subducted continental lower crust

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15 Abbreviated title: Guided waves from the Pamir seismic zone

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23 Keywords: Guided waves, Wave propagation, Continental crust, Central Asia

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26 Highlights

27

28 Guided waves from earthquakes within the Pamir seismic zone have been recognized

29 The guided waves occur as a secondary phase behind the first arrivals

30 The guided waves are generated by a low velocity layer about 10 km thick

31 The low velocity layer is caused by subducting lower continental crust

32 The velocities in the low velocity layer are partly due to the presence of melts

33

34

35 ABSTRACT

36

37 As part of the TIPAGE (Tien shan – PAmir GEodynamic program) project, passive
38 seismological observations were made along an approximately N-S profile crossing the Pamir
39 seismic zone for about one year. From these observations guided waves were recognized.
40 These guided waves occur as a single, continuous, secondary, compressional (P) wave phase
41 behind the first P-wave arrivals. An equivalent phase in the shear (S) wavefield is hardly
42 recognizable. Modelling of the phase shows that an approximately 10 km thick low velocity
43 zone (LVZ) between the Moho and about 160 km depth reproduces the guided waves as a
44 single, continuous phase much better than a 15-20 km thick LVZ. Modelling of the arrival
45 times of the guided waves reveals that a model with a P-wave velocity of 6.3 km/s above
46 about 100 km depth, and a velocity of 7.6 km/s between this depth and the deep cluster of
47 earthquakes at about 150 km depth provides the best fit to the observed travel-time data. One
48 plausible way to explain the low velocity of 6.3 km/s is to invoke the presence of melts in the
49 LVZ. Then, taking a velocity of 6.9 km/s for the lower crust being subducted, about 10-13%
50 melt is required to obtain a velocity of about 6.3 km/s in the LVZ between the Moho and
51 about 100 km depth. This would be in keeping with the estimated burial depths from xenoliths
52 of Gondwana terrane affinity brought to the surface in the southeastern Pamir around 11
53 million yr. ago. The present-day LVZ is interpreted to comprise continental lower crust.
54 Although guided waves are known to exist associated with subducted oceanic crust or fault
55 zones, this is the first time to the knowledge of the authors that guided waves have been
56 observed resulting from a LVZ associated with subducted continental lower crust.

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58

59 **1. Introduction**

60

61 Guided waves are generally associated with subduction zones and fault zones. Their
62 generation is due to total internal reflection within a low velocity zone associated with, for
63 example, a fault zone or a subduction zone (e.g. Catchings et al., 2016; Garth & Rietbrock,
64 2017). In some cases they form part of a distorted, dispersive first arrival phase (Abers, 2000;
65 Martin et al., 2003; Garth & Rietbrock, 2014a, b). In other cases they occur as a separate
66 secondary phase clearly separated from the first compressional (P) wave arrival (e.g. Martin et
67 al., 2005; Martin & Rietbrock, 2006; Ellsworth & Malin, 2011; Garth & Rietbrock, 2017;
68 Coulson et al., 2018). In yet other cases and especially when related to fault zones, they have
69 been recognized arriving later than the shear (S) wave phases (Li & Malin, 2008; Catchings et
70 al., 2016). They have been observed at several subduction zones around the Pacific, including
71 Alaska, the Aleutian Islands, the Kurile Islands, Japan, Mariana (Fukao et al., 1983; Hori,
72 1990; Abers, 2000; Garth & Rietbrock, 2014a, b; Coulson et al., 2018) and Chile (Martin et
73 al., 2003; Martin et al., 2005; Martin & Rietbrock, 2006; Garth & Rietbrock, 2017). They

74 have also been observed at some of Earth's major fault systems e.g. San Andreas fault system
75 (Cormier & Spudich, 1984; Ben-Zion & Malin, 1991; Li & Malin, 2008; Wu et al., 2010;
76 Ellsworth & Malin, 2011), North Anatolian fault (Ben-Zion et al., 2003), Dead Sea Transform
77 (Haberland et al., 2003). Guided waves have also been observed from the Vrancea
78 intermediate depth seismic zone in Romania (Bokelmann & Rodler, 2014), where dispersive
79 first P-wave arrivals have been recognized. The interpretation in this case is that the guided
80 waves have been generated in a low velocity zone associated with the subduction of oceanic
81 crust, now residing at uppermost mantle depths.

82

83 In this study, observations of guided waves from the Pamir seismic zone are presented. These
84 guided waves have been observed along the main profile of a passive seismological
85 experiment in the Pamir within the framework of the TIPAGE (Tien shan – PAmir
86 GEodynamic program) project (Mechie et al., 2012; Sippl et al., 2013a, b; Schneider et al.,
87 2013; Schurr et al., 2014a; Kufner et al., 2016; Li et al., 2018). In addition to the observations,
88 an analysis of the travel times and the amplitudes of the first arrivals and guided waves by 2-
89 D ray-tracing (Podvin & Lecomte, 1991) and full waveform, finite-differences seismograms
90 (Kelly et al., 1976) is presented and the results compared with those from other studies of the
91 data from the TIPAGE project.

92

93

94 **2. Geological / Tectonic Setting & Previous Work**

95

96 The Pamir – Hindu Kush seismic zone and the Vrancea seismic zone in Romania are the only
97 two intra-continental regions on Earth with significant seismicity at depths greater than 100
98 km. As mentioned above, guided waves have also been observed from the Vrancea seismic
99 zone, where Bokelmann & Rodler (2014) interpreted them to have been generated in a low
100 velocity zone associated with the subduction of oceanic crust. However, in the case of the
101 Pamir, subduction of oceanic material should have stopped with the final closure of the
102 Tethys ocean no later than 40 million yr ago (Yin & Harrison, 2000) and thus it is thought
103 that the low velocity zone already detected by a receiver function analysis (Schneider et al.,
104 2013) comprises lower continental crust.

105

106 The Pamir forms the northwestern portion of the Pamir-Tibet plateau (Fig. 1), the largest
107 plateau on Earth. The plateau has been formed by the India–Asia continent–continent
108 collision which is also responsible for the Himalayas, the highest mountain range on Earth
109 and deformation in large regions of Central and Eastern Asia (e.g. Tien Shan). Tibet and the
110 Pamir both stand 4-5 km high (Fielding, 1996). In Tibet this topography is supported by a 70–
111 80 km thick crust (Kind et al., 2002; Mechie et al., 2011; Mechie & Kind, 2013) and in the
112 Pamir it is supported by a 65-75 km thick crust (Mechie et al., 2012, Schneider et al., 2013).

113 Both Tibet and the Pamir have absorbed around 1800–2100 km of Cenozoic N-S shortening
114 (Le Pichon et al., 1992; Johnson, 2002). However, the N-S extent of the Pamir is about 500
115 km, compared with up to 1300 km for Tibet. Consequently, compared to Tibet, the Pamir has
116 either undergone more lateral extrusion of crust, more erosional unroofing or greater loss of
117 continental crust into the mantle.

118
119 The Pamir – Hindu Kush seismic zone would seem to be the obvious place to look for
120 continental crust being lost into the mantle in this region. In the past, some workers
121 understood the Pamir seismic zone to mark about 300 km southward subduction of Asian
122 crust and mantle lithosphere (Hamburger et al., 1992; Burtman & Molnar, 1993), although
123 other workers (Pegler & Das, 1998; Pavlis & Das, 2000) claimed that it represented
124 overturned northward subducting Indian lithosphere. In contrast, the Hindu Kush
125 intermediate-depth seismicity was interpreted to mark around 700 km of subducted Indian
126 lithosphere dipping at about 80° to the north (Negredo et al., 2007). East of the Pamir, Indian
127 lithosphere is ‘subducting’ beneath Tibet almost horizontally (Kumar et al., 2006; Jiménez-
128 Munt et al., 2008) and may underlie most of western Tibet (Li et al., 2008; Zhao et al., 2010).
129 More recently, the TIPAGE project, through a series of seismological experiments (Mechie et
130 al., 2012; Sippl et al., 2013a, b; Schneider et al., 2013; Kufner et al., 2016; Li et al., 2018),
131 has been able to show that the Pamir seismic zone is due to the southward subduction of
132 Asian continental lithosphere, including lower crustal material, and that the Hindu Kush
133 seismic zone is due to the northward subduction of Indian continental lithosphere. More
134 specifically, the southward subduction of Asian lower crust, which is thought to host the
135 Pamir seismic zone, was initiated about 10 million yr ago when Indian cratonic lithosphere
136 impinged on Asian cratonic lithosphere and, due to the already thickened Pamir crust pushing
137 Asian cratonic lithosphere down, caused it to delaminate and roll back (Kufner et al., 2016).
138 Receiver function analysis has shown that a 10-15 km thick low velocity zone marks the top
139 of the downgoing Asian lithosphere and provided evidence that this low velocity zone hosts
140 the intermediate-depth seismicity and is due to lower continental crust being taken down
141 together with the lithospheric mantle (Schneider et al., 2013). In the receiver function image,
142 the top of the low velocity zone is marked by a negative converter which, by definition,
143 indicates a decrease in velocity downwards (Schneider et al., 2013). Above the low velocity
144 zone, Sippl et al. (2013b) showed that a region with a P-velocity of around 7.1 km/s exists
145 down to a depth of about 100 km. This region was interpreted to be either re-equilibrated
146 felsic, upper crust or meta-stable intermediate, middle crust. Most recently, the Pamir – Hindu
147 Kush seismic zone has been proposed to be a single, contorted slab of Asian origin (Perry et
148 al., 2019). However, this hypothesis is at odds with the uppermost mantle part of the
149 seismicity dipping down to the north beneath the Hindu Kush (Pegler & Das, 1998; Kufner et
150 al., 2016).

151

152 The rocks that have been subducted southward may be exemplified by the xenoliths which
153 have been brought up by volcanic rocks about 11 million yr ago in the southeastern Pamir
154 (Hacker et al., 2005 and Fig. 1). These xenoliths comprise mainly eclogites and granulites and
155 a glimmerite. The protoliths for the eclogites were igneous rocks with mafic to felsic
156 compositions whereas the protoliths for the granulites were pelitic sediments (Hacker et al.,
157 2005). The xenoliths are thought to originate from the southern Pamir (Rutte et al., 2017;
158 Schaffer et al., 2017) which is thought to be the lateral equivalent of one of the Gondwana
159 terranes, namely the Lhasa terrane, in Tibet (Schwab et al., 2004). They are thought to have
160 been taken down to 90-100 km depth due to transient crustal drips and subduction erosion,
161 before being brought up to the surface (Hacker et al., 2005; Rutte et al., 2017). The crustal
162 rocks subducting at the present day as part of the Pamir seismic zone are thought to belong to
163 Asian cratonic lithosphere (Kufner et al., 2016) and thus they have no affinity with the
164 Gondwana terrane, from which the xenoliths were derived. However, as long as the rock
165 types are similar, then the processes that operated on the xenoliths in the past may very well
166 be occurring in the present-day Pamir seismic zone.

167

168

169 **3. Characteristics of Guided Waves, Background Model, Earthquake Selection &** 170 **Processing**

171

172 The guided waves observed in this study mainly occur as secondary arrivals in the P-
173 wavefield (Figs. 2-4). They cannot be observed nearly so well in the S-wavefield, although in
174 some cases it is possible that they are present. A dispersion analysis was carried out for the
175 data recorded by the vertical component at two stations (P08 & P09, Figs. 2-3) for two events
176 (302 & 304, Table 1). A frequency-time analysis was carried out following Dziewonski et al.
177 (1969), using 60 different filters centred between 0.67 and 10 Hz. The frequency-time plots
178 (Fig. 5) show that the guided waves are more prominent at lower frequencies whereas for
179 event 302 the first arrivals are more prominent at higher frequencies. For the guided waves
180 there is no smooth, consistent dependence of arrival time on frequency for any of the four
181 analysed traces, although the trace at station P09 from event 302 may show some dependence.
182 Although guided waves are dispersive in nature and this dispersion is often observed (e.g.
183 Abers, 2000; Martin et al., 2003; Ellsworth & Malin, 2011; Garth & Rietbrock, 2014a, b;
184 Bokelmann & Rodler, 2014, Coulson et al., 2018), there are cases where dispersion is not
185 observed (e.g. Haberland et al., 2003).

186

187 All the velocity models (Figs. 6 and S1-S2) that are utilized in this study for calculating the
188 travel times and amplitudes of the seismic waves incorporate, as the background model, a 2-D
189 cross-section close to the TIPAGE main profile through the 3-D model derived from local
190 earthquake tomography (Fig. 7 and cross-section D-D' in Fig. 2 of Sippl et al., 2013b). In

191 addition, the Moho and the boundaries of the low velocity zone hosting the Pamir seismic
192 zone were based in the initial test models on converters identified by a receiver function
193 analysis (red lines in Fig. 7 and Fig. 6 in Schneider et al, 2013). The boundaries of the low
194 velocity zone were slightly modified during the modelling process so that most of the
195 earthquakes occurred in the top half of the seismic zone.

196
197 Earthquakes which show guided waves and which are located close to the TIPAGE main
198 profile were utilized in this study. 13 earthquakes fulfilled these requirements (Table 1). The
199 waveforms of the different events recorded at the same station are often quite similar but there
200 are differences in detail. For example, although station P08 always shows a pickable guided
201 wave phase (Figs. 2-4), for events 318 and 325 the amplitude of the guided wave phase is
202 smaller than that of the first arrival whereas for the other events the amplitude of the guided
203 wave phase is larger than that of the first arrival. The earthquake hypocentres were derived by
204 relocating them, from the original catalogue provided by Sippl et al. (2013a), in the 3-D
205 model derived from local earthquake tomography (Sippl et al., 2013b), as a 2-D cross-section
206 through this model was utilized as the background model in the present study. The
207 earthquakes were then projected onto the TIPAGE main profile (Fig. 1). A comparison
208 between the theoretical 1st arrival times derived from the 3-D model of Sippl et al. (2013b),
209 the 2-D background model utilized in this study (Fig. 7) and one of the 2-D models
210 incorporating a low velocity zone (model 2 in Fig. 6) was made (Fig. S3). This comparison
211 shows that in the portion of the profile where the guided waves are observed, the travel-time
212 differences between the three models are generally less than 0.4 s. Power spectra were
213 calculated for two of the events (302 & 304, Table 1) which showed well-developed guided
214 waves. These spectra showed a peak between 1 and 2 Hz. Thus the record sections were band-
215 pass filtered between 0.1 and 2 Hz. Then the arrival times for both the first arrivals and the
216 guided waves were picked.

217

218

219 **4. Modelling of Travel Times of Guided Waves**

220

221 Synthetic seismograms were calculated using a finite-difference approximation of the wave
222 equation for 2-D heterogeneous elastic media by Kelly et al. (1976) with transparent boundary
223 conditions (Reynolds, 1978) and implemented by Sandmeier (1990). Firstly, synthetic
224 seismograms were calculated for the background model to examine if it would produce
225 guided waves. For this model and the models described in this and the following paragraph,
226 an explosive point source with a dominant frequency of 1.25 Hz in combination with a Fuchs-
227 Müller signal (Fuchs & Müller, 1971) was used. The model is stable up to a frequency of 5
228 Hz, as are all the models in this study. The background model did not produce any guided
229 wave energy. Then the converters bounding the low velocity zone identified by Schneider et

230 al. (2013) were added to the background model. This resulted in a 15-20 km wide zone
231 between the bounding converters (blue lines in Fig. 7). In addition, Schneider et al. (2013)
232 recognized that the velocity in the zone between the converters should be less than that of the
233 surrounding mantle material as the polarity of the upper converter meant that it represented a
234 velocity decrease with depth while the polarity of the lower converter meant that it
235 represented a velocity increase with depth. Thus, a P-velocity of 6.9 km/s was assigned to the
236 low velocity zone (LVZ) between the converters. This was based on the average for the basal
237 10-15 km of the crust from the interpretation of refracted and wide-angle reflected phases
238 along the profile (Mechie et al., 2012). Using this model synthetic seismograms were
239 calculated for a number of synthetic sources including sources inside the LVZ and sources
240 outside the LVZ, both above and below it. It was found that sources below the LVZ produced
241 little or no guided wave energy (Fig. S1, middle right). Sources within or above the LVZ
242 produced energy behind the first arrivals similar to that in the observed data. For sources far
243 enough above the LVZ, the strong secondary arrivals are at least in part due to strong
244 reflections from the base of the LVZ (see e.g. the distance range from 110-160 km in Fig. S1,
245 upper right). For sources within the LVZ, the secondary energy was often split into two
246 distinct phases (Fig. S1, middle left) while in the observed data the secondary phase looks like
247 one long continuous phase. This was taken to indicate that the LVZ was too wide. Thus the
248 width of the LVZ was reduced to about 10 km in the next model (black lines in Fig. 7).
249 Actually, forward waveform modelling of receiver functions for six stations in the vicinity of
250 the southern part of the profile revealed a thickness of 10-15 km for the LVZ (Schneider et
251 al., 2013).

252
253 Tests with synthetic sources for an approximately 10 km wide LVZ showed that the synthetic
254 secondary energy mimicked the observed secondary energy best in terms of looking like one
255 long continuous phase rather than two split phases, when the synthetic source was located in
256 the upper half of the LVZ (Fig. S1, lower right). The top boundary of the LVZ was thus
257 modelled so that the 13 real earthquakes would lie, if possible, in the upper half of the LVZ.
258 Without introducing a concave element to the top boundary of the LVZ, it was possible to
259 have all except the deepest earthquake of the middle cluster of events in the upper half of the
260 LVZ (Fig. 7). Further, the two boundaries of the approximately 10 km wide LVZ lie mainly
261 within the bounds defined by the converters identified by the receiver function analysis of
262 Schneider et al. (2013). With this configuration of boundaries for the approximately 10 km
263 wide LVZ in relation to the 13 real earthquakes, the one remaining parameter to be
264 determined is the velocity within the LVZ.

265
266 The velocity within the LVZ primarily controls the travel-time difference between the first
267 arrivals and the secondary guided wave energy travelling through the LVZ. Thus models were
268 constructed in which the velocity in the LVZ above about 130 km depth was varied in 0.1

269 km/s steps from 6.2 to 6.9 km/s for the shallow earthquake and from 6.3 to 6.9 km/s for all
270 other events. This results in velocity gradients along the dip direction of the LVZ for these
271 models and, in fact, for any of the models in this study of up to 0.05 km/s/km. As the power
272 spectra of the observed data showed a peak between 1 and 2 Hz, for these and all subsequent
273 models the dominant frequency of the explosive point source was 1.5 Hz and the resulting
274 bandwidth of the signal is from 0.8 to 3.5 Hz. At depths greater than about 160 km the LVZ
275 was constrained to disappear as the receiver function study of Schneider et al. (2013) showed
276 that the converters marking the top and bottom boundaries of the LVZ disappear at about this
277 depth and the study of Sippl et al. (2013a) showed that the seismicity also disappears at about
278 this depth. For each of the constructed models synthetic seismograms were calculated for the
279 13 earthquakes (Fig. 8). Then the travel times of the synthetic secondary guided wave arrivals
280 were picked and compared with the travel times of the observed secondary guided wave
281 arrivals. The picking was done on record sections with 15 km station spacing, which is
282 approximately the same as the station spacing in the observed data. Further, only phases that
283 have a significant amplitude compared to that of the first arrival were picked, as only such
284 phases could be picked in the observed data. This sometimes meant that the earliest guided
285 wave arrival was not picked as it was too weak, but a later more prominent arrival (Fig. 8c-d).
286 When comparing the absolute travel times of the observed and theoretical guided waves, the
287 model (model 1 in Fig. 6) with the smallest root mean square residual time (trms), was that in
288 which the LVZ had a velocity of 6.7 km/s above about 130 km depth (Fig. 9). In addition, the
289 travel times of the synthetic first arrivals were calculated using finite-differences ray-tracing
290 based on the eikonal equation (Vidale, 1988; Podvin & Lecomte, 1991; Schneider et al.,
291 1992). Then the differences between the synthetic secondary guided wave arrivals and first
292 arrivals were calculated and compared with the differences between the corresponding
293 observed arrivals. In this case, the model with the smallest root mean square residual time
294 (trms), was that in which the LVZ had a velocity of 6.6 km/s (model 1a) instead of 6.7 km/s
295 (model 1) above about 130 km depth.

296
297 During this analysis it was realized that the shallow earthquake (Fig. 4) required the smallest
298 average velocity (6.3 km/s), whereas the intermediate cluster of earthquakes (Fig. 3) required
299 a greater average velocity (6.4 – 6.6 km/s) and the deep cluster of earthquakes (Fig. 2)
300 required the largest average velocity (6.9 km/s or even larger). Thus a model (model 2 in Fig.
301 6) was made with varying velocity in the LVZ. Above the shallow earthquake a velocity of
302 6.3 km/s was assigned to the LVZ, whereas between the shallow earthquake and the deep
303 cluster of earthquakes a velocity of 7.6 km/s was assigned. This resulted in a trms which was
304 smaller than those for the cases where the LVZ had a velocity of 6.6 or 6.7 km/s above 130
305 km depth. However, this model only has one change in velocity between the shallow
306 earthquake and the deep cluster of earthquakes whereas, as mentioned above, the average
307 velocities required to fit the observed travel-time data imply that there may be an additional

308 change in velocity between the intermediate and deep clusters of earthquakes. For this reason
309 a further model (model 3 in Fig. 6) was made with varying velocity in the LVZ. In this case,
310 the velocity above the shallow earthquake was assigned a value of 6.4 km/s, the velocity
311 between the shallow earthquake and the intermediate cluster of earthquakes was assigned a
312 value of 7.2 km/s and the velocity between the intermediate and deep clusters of earthquakes
313 was assigned a value of 7.6 km/s. This resulted in a trms which was smaller than those for
314 models 1 or 1a but larger than that for model 2. Thus, from the point of view of the travel
315 times, model 2 has the smallest trms and is thus considered to be the best model. Snapshots
316 (Fig. 10) and an animation (Video S1) of waves propagating through this model for event 304
317 are presented. These snapshots and animation are representative of wave propagation in all
318 the models with an approximately 10 km wide LVZ, for events in the upper half of the LVZ.
319 A digital version of this model (Data Set S1) is also presented.

320
321

322 **5. Modelling of Amplitudes of Guided Waves**

323

324 The amplitudes of the first arrivals and the secondary guided wave phase were picked in the
325 observed data, filtered from 0.1 to 2 Hz, of the 13 earthquakes (see e.g. Figs. 2-4). For the
326 model (model 1) in which the LVZ had a velocity of 6.7 km/s above about 130 km depth and
327 for the two models (models 2 & 3) with varying velocities in the LVZ, the amplitudes of the
328 first arrivals and the secondary guided wave phase were picked in the synthetic data of the 13
329 earthquakes (see e.g. Fig. 8). The bandwidth of these synthetic data is from 0.8 to 3.5 Hz and
330 is thus similar to that of the observed data. Then the amplitude ratios of the secondary guided
331 wave phase to the first arrival phase were calculated for both the observed and synthetic data
332 and compared (Fig. 11). The differences in amplitude ratios between the observed and
333 synthetic data for all 13 earthquakes for each of the three models were calculated. Although
334 model 3 (Fig. 6) had the smallest root mean square (rms) amplitude ratio difference, the
335 difference between the three models is very small with all three models having a rms
336 amplitude ratio difference between 1.6 and 1.62. Thus amplitudes did not help in the present
337 study to decide which model is to be preferred. On the other hand, a more detailed study of
338 the amplitudes should probably also better approximate the radiation pattern of the real
339 sources rather than employing a simple explosive point source as in the present study.

340
341

342 **6. Discussion**

343

344 The fact that the earthquakes at 100-160 km depth in the Pamir seismic zone generate guided
345 waves is important, as it has implications for the physical properties of the seismic zone. The
346 presence of second arrivals behind the first arrivals suggests that a zone of low or high

347 velocity exists with fairly sharp transitions with respect to the wavelength, to the materials
348 above and below the zone. Although guided waves are really associated with low velocity
349 zones, when a model was made in which the low velocities were replaced by high velocities
350 of, for example, 8.8 km/s, second arrivals similar to the guided waves actually appeared
351 behind the first arrivals, even though these second arrivals are not actually guided waves (Fig.
352 S2). Interestingly, in the synthetic seismograms for event 304 for this model, the second
353 arrival phase does not separate from the first arrivals until 150-200 km distance and the time
354 difference between the two phases at 250 km distance is only about 2 s. This is in contrast to
355 the observed data where the guided wave phase appears as a separate phase at 80-100 km
356 distance and the time difference between the first arrivals and the guided waves at about 250
357 km distance is 5-6 s (see Fig. 3a). In the case of the Pamir, there are other lines of evidence
358 pointing to a low velocity zone (LVZ) rather than a high velocity zone. The first line of
359 evidence is from the receiver function analysis of Schneider et al. (2013), who imaged two
360 converters bounding the Pamir seismic zone. Additionally, as mentioned in section 4 above,
361 the two converters defined a LVZ in the region between them. The second line of evidence is
362 from the local earthquake tomography study of Sippl et al. (2013b), whose results are also
363 compatible with a LVZ in the vicinity of the Pamir seismic zone. From a geological point of
364 view it is also reasonable to expect a LVZ associated with the Pamir seismic zone, if the
365 lowermost crustal layer has stayed attached to the underlying downgoing mantle layer. This is
366 because the lowermost crustal layer will have a lower velocity than the underlying and
367 overlying mantle material.

368
369 When comparing the observations of guided waves made in this study with those made at
370 oceanic subduction zones the different geometries of the experimental setups come to mind.
371 At oceanic subduction zones where the situation is rather similar, with the oceanic crust acting
372 as a low velocity waveguide on top of the oceanic mantle lithospheric slab, guided waves are
373 only observed on land if there is a feature that allows them to leave the waveguide. In the case
374 of Chile this is a kink in the slab in one case (Martin et al., 2003; Martin & Rietbrock, 2006;
375 Garth & Rietbrock, 2017) and an equalization in seismic velocities in another case (Martin et
376 al., 2005). The equalization in seismic velocities occurs when the oceanic crust comes in
377 contact with continental crust, usually lower continental crust, which has similar seismic
378 velocities. In the case of Japan it is also a kink in the slab in one case (Garth & Rietbrock,
379 2014b) and an equalization in seismic velocities in other cases (Fukao et al., 1983; Hori,
380 1990). If this feature is not present, the guided waves follow the waveguide to the trench,
381 where usually no stations are located. In this study, there is a much better station coverage
382 around the waveguide since there is no ocean involved and the waveguide terminates into the
383 overlying lower crust.

384

385 From an analysis of receiver functions recorded by six of the stations above the Pamir seismic
386 zone, Schneider et al. (2013) determined that the LVZ is 10-15 km thick. From further
387 analysis at one of the stations (P18, Fig. 1), Schneider et al. (2013) concluded that the LVZ
388 has a S-wave velocity of about 4.16 km/s. Assuming a Poisson's ratio greater than 0.22, this is
389 equivalent to a P-wave velocity greater than 6.94 km/s. The depth for which this velocity
390 within the LVZ was obtained is about 100 km, due to the position of the station with respect
391 to the Pamir seismic zone and the incidence angles of the incoming waves. In the present
392 study a thickness of 10 km for the LVZ was found to better reproduce the appearance of the
393 guided waves as a single secondary phase than a thickness of 15-20 km. With respect to the P-
394 wave velocities at 100 km depth, the best models (model 1 or model 1a) with a uniform
395 velocity in the LVZ above about 130 km depth had a velocity of 6.6 or 6.7 km/s, whereas
396 model 3 (Fig. 6) had a velocity of 6.8 km/s, and model 2 (Fig. 6) with the smallest root mean
397 square residual time (trms) had a velocity of 7.0 km/s. Thus the model with the smallest trms
398 also has a velocity at 100 km depth in the LVZ which is closest to that derived from the
399 receiver function analysis, assuming a Poisson's ratio of 0.22 or greater.

400

401 From a local earthquake tomography study, Sippl et al. (2013b) found P-wave velocities of
402 around 7.1 km/s between 60 and 100 km depth on top of the high-velocity, south-dipping
403 lithospheric mantle slab. At 100 km depth model 2 (Fig. 6), with the smallest trms, has a
404 velocity in the LVZ which is closest to the value of 7.1 km/s obtained by Sippl et al. (2013b).
405 However, at shallower depths of, for example, about 70 km the models (model 1 or model 1a)
406 with a uniform velocity of 6.6 or 6.7 km/s in the LVZ above about 130 km depth have a
407 velocity closer to the value of 7.1 km/s obtained by Sippl et al. (2013b), as models 2 and 3
408 have velocities of 6.3 and 6.4 km/s in the LVZ at about 70 km depth. A similar argument can
409 be made if one compares the P-wave velocities in the LVZ derived in this study with the
410 average P-wave velocity of about 6.9 km/s for the basal 10-15 km of the crust from the
411 interpretation of refracted and wide-angle reflected phases along the TIPAGE main profile
412 (Mechie et al., 2012). The models (model 1 or model 1a) with a uniform velocity of 6.6 or 6.7
413 km/s in the LVZ above about 130 km depth show a difference of just 0.2 or 0.3 km/s with
414 respect to the average P-wave velocity of 6.9 km/s derived for the basal 10-15 km of the crust
415 by Mechie et al. (2012). Models 2 and 3, with laterally varying velocities, have velocities of
416 6.3 and 6.4 km/s above about 90 km depth in the LVZ. Only at about 100 km depth do models
417 2 and 3 have velocities in the LVZ closer to the average P-wave velocity of 6.9 km/s derived
418 for the basal 10-15 km of the crust by Mechie et al. (2012), than the models (model 1 or
419 model 1a) with a uniform velocity of 6.6 or 6.7 km/s in the LVZ above about 130 km depth.
420 One caveat in this case is that the velocity in the basal 10-15 km of the crust varies laterally
421 and then the velocity in the LVZ need not agree with the average P-wave velocity of 6.9 km/s
422 derived for the basal 10-15 km of the crust by Mechie et al. (2012) along the TIPAGE main
423 profile. In summary, a model with laterally varying velocities in the LVZ fits the travel times

424 of the guided waves or the travel-time differences between the guided waves and the first
425 arrivals best. However, a model with a uniform velocity of 6.6-6.7 km/s (model 1 or 1a) in the
426 LVZ above about 130 km depth may seem, from a first point of view, to fit better with other
427 geophysical results along the TIPAGE main profile and may appear to be the most reasonable
428 model from a geological point of view.

429
430 Inaccuracies in the hypocentres of the events are one possible source of error in determining
431 the travel times of the arrivals and hence the velocity distribution in the LVZ. In order to
432 study the effects of possible errors in the hypocentres of the events, the event 330 was moved
433 4.4 km to the south and 7 km deeper, so that it is again located just below the upper boundary
434 of the LVZ. Then synthetic seismograms were calculated for the relocated event in model 1.
435 First arrivals were calculated as described above using finite-differences ray-tracing and the
436 second arrivals of the guided wave phases were picked. Then the differences between the
437 synthetic secondary guided wave arrivals and first arrivals were compared with the
438 differences between the corresponding observed arrivals. The comparison shows that the trms
439 is 0.972 s. This value is smaller than the value of 1.399 s for the original hypocentre of the
440 event. It lies between the values of 1.062 s for a velocity of 6.5 km/s in the LVZ and 0.939 s
441 for a velocity of 6.4 km/s in the LVZ for the original hypocentre of the event. Thus moving
442 the event 4.4 km to the south and 7 km deeper results in a trms which is equivalent to that for
443 the original hypocenter for a velocity in the LVZ reduced by 0.2-0.3 km/s. This is a step in the
444 correct direction to reducing the overall misfit between the observed and synthetic data.
445 However, the skew in the first arrival times at the stations within about 50 km of the epicentre
446 is definitely noticeable, due to the small station spacing of about 15 km in the N-S direction
447 along the TIPAGE main profile (Fig. S4). This skew is unreasonable and thus it is concluded
448 that errors in the hypocentres are not a major factor in producing errors in the travel times of
449 the arrivals and especially the travel-time differences between the secondary guided wave
450 arrivals and first arrivals. Thus the problems in determining the velocities in the LVZ should
451 have other sources in addition to the possible, relatively modest ones introduced by the
452 inaccuracies of the hypocentres of the events.

453
454 There is perhaps one way to explain the variable velocities found in either models 2 or 3 in
455 the LVZ that is geologically reasonable. As mentioned above, Hacker et al. (2005) studied a
456 suite of xenoliths from the southeastern Pamir (Fig. 1) that were brought to the surface by
457 volcanic rocks about 11 million yr ago. This suite consists of eclogites for which the
458 protoliths were igneous rocks with mafic to felsic compositions and granulites for which the
459 protoliths were pelitic sediments. In the transformation of these rocks to eclogites and
460 granulites, high-pressure dehydration melting took place. The favoured scenario of Hacker et
461 al. (2005) was that the igneous and sedimentary rocks were being subducted when they were
462 metamorphosed to eclogites and granulites and that the melts released during this

463 transformation migrated back up the subduction zone, which would essentially be the LVZ in
464 this study. Assuming that this process is occurring in the slab presently subducting and that
465 the melts are concentrated above about 100 km depth (Fig. 12), then this might explain the
466 low average velocity in the LVZ required by the shallowest event, 330. Taking a P-wave
467 velocity of around 3.45 km/s at about 85 km depth for the melt (Stolper et al., 1981) and a P-
468 wave velocity of 6.9 km/s for the lower crust being subducted, then about 10-13% melt (e.g.
469 Watt et al., 1976) is required to obtain an average velocity of about 6.3 km/s (Fig. 12). This is
470 the average velocity which best fits the data from the shallowest event, 330 (Fig. 9). The
471 presence of melt in the LVZ may also help to explain why there are almost no earthquakes
472 between the Moho and 90 km depth in this part of the Pamir seismic zone (Sippl et al.,
473 2013a). It is also in agreement with high seismic attenuation being observed in this region
474 (Schurr et al., 2014b). That the melts concentrate in the LVZ above 100 km depth may be due
475 to the possible tensional forces which probably exist in this depth range as the lower crust is
476 bent down from the north into the Pamir seismic zone. Such forces are thought to exist and to
477 cause faulting in the oceanic crust as it enters a subduction zone (see e.g. Garth & Rietbrock,
478 2014a and references therein). There remains the question of why the melts remain confined
479 to the LVZ and do not migrate either into the lower crust above the Moho to the north of the
480 subducting slab or into the meta-stable middle crust of Sippl et al. (2013b) lying immediately
481 above the subducting lower crust (Fig. 12). The simplest explanation is that the melts are
482 lighter than the lower crust once it is part of the subducting slab but are heavier than both the
483 lower crust before it is part of the subducting slab and the meta-stable middle crust above the
484 subducting lower crust.

485

486

487 **7. Conclusions**

488

489 In this study, observations of guided waves from the Pamir seismic zone have been presented.
490 These guided waves occur as secondary P-wave arrivals behind the first P-wave arrivals.
491 Equivalent arrivals in the S-wavefield are much less readily observable. Four traces were
492 analysed for dispersion, but the guided wave phases only showed signs of any dispersion in
493 one of these cases.. A 10 km thick low velocity zone (LVZ) reproduces the guided waves as a
494 continuous, single phase much better than a 15-20 km thick LVZ. Along the TIPAGE main
495 profile, the LVZ exists down to about 160 km depth. Modelling of the velocities in the LVZ
496 shows that for models with a constant velocity in the LVZ above about 130 km depth, the
497 model (model 1 in Fig. 6) with a velocity of 6.7 km/s produced the smallest root mean square
498 residual time (trms) when modelling absolute travel times of the guided waves. In contrast,
499 the model with a velocity of 6.6 km/s produced the smallest trms when modelling the travel-
500 time differences between the guided waves and the first arrivals. However, among the models
501 tested, the smallest trms is realized by a model (model 2 in Fig. 6) with a velocity of 6.3 km/s

502 above about 100 km depth and a velocity of 7.6 km/s between the shallow event and the deep
503 cluster of earthquakes at about 150 km depth. Modelling of the amplitudes did not help to
504 discriminate further between the three best travel-time models (models 1, 2 and 3 in Fig. 6). If
505 a lateral change in composition of the lowermost crust is not responsible for a velocity
506 reduction from 6.9 km/s, as derived for the basal 10-15 km of the crust to the north of the
507 LVZ (Mechie et al., 2012), to 6.3 km/s in the LVZ above 100 km depth, then one plausible
508 way to explain the low value of 6.3 km/s is to invoke the presence of melts in the LVZ. Then,
509 taking a velocity of 6.9 km/s for the lower crust being subducted, about 10-13% melt is
510 required to obtain an average velocity of about 6.3 km/s in the LVZ between the Moho and
511 about 100 km depth. This would be in keeping with the results of Hacker et al. (2005), who
512 studied xenoliths with Gondwana terrane affinity brought to the surface in the southeastern
513 Pamir about 11 million yr. ago. Hacker et al. (2005) proposed that during the formation of
514 these eclogitic and granulitic xenoliths, high-pressure dehydration melting took place and that
515 these melts migrated back up the subduction zone. Although the present-day LVZ is
516 associated with the subduction of material with Asian affinity it is possible that the same
517 processes that operated in the Pamir about 11 million yr. ago are again presently active. Thus
518 the present-day LVZ is interpreted to comprise continental, lower crustal material. Although
519 there are a number of studies of guided waves resulting from LVZs associated with subducted
520 oceanic crust or fault zones, this is the first time to the knowledge of the authors that guided
521 waves have been observed resulting from a LVZ associated with subducted continental lower
522 crust.

523

524

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526

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535

536

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Table 1[Click here to download Table: casia055_tab1_rev.xls](#)

Table 1. Origin time, hypocentre and local magnitude M_L for each of the 13 events used in this study. The origin times and hypocentres have been derived by relocating the events in the 3-D model of Sippl et al. (2013b).

event	origin time year:day of year:hour:min:sec	longitude °E	latitude °N	depth km	M_L
302	2008:281:03:46:51.850	74.0620	38.1937	158.5	4.5
304	2008:304:10:52:17.430	73.9187	38.3563	125.8	3.8
314	2008:242:19:48:18.930	73.9827	38.1557	163.9	3.3
315	2009:081:00:17:19.730	74.1722	38.1535	165.0	3.0
317	2009:004:07:50:40.850	74.0317	38.1773	156.2	3.0
318	2008:246:16:52:45.190	74.0813	38.1922	156.1	3.9
319	2008:225:22:53:57.630	74.1098	38.1973	152.7	3.5
321	2009:098:16:42:28.800	74.1167	38.2305	148.5	3.7
325	2009:139:01:48:19.070	73.9970	38.3682	128.1	3.8
326	2008:231:14:13:44.180	73.9935	38.3705	126.9	3.8
327	2009:003:17:09:20.860	73.9092	38.3747	124.5	3.5
328	2009:057:07:50:49.150	73.9773	38.3677	129.6	5.2
330	2009:086:21:15:35.400	73.8485	38.4397	105.0	4.1

Figure 1

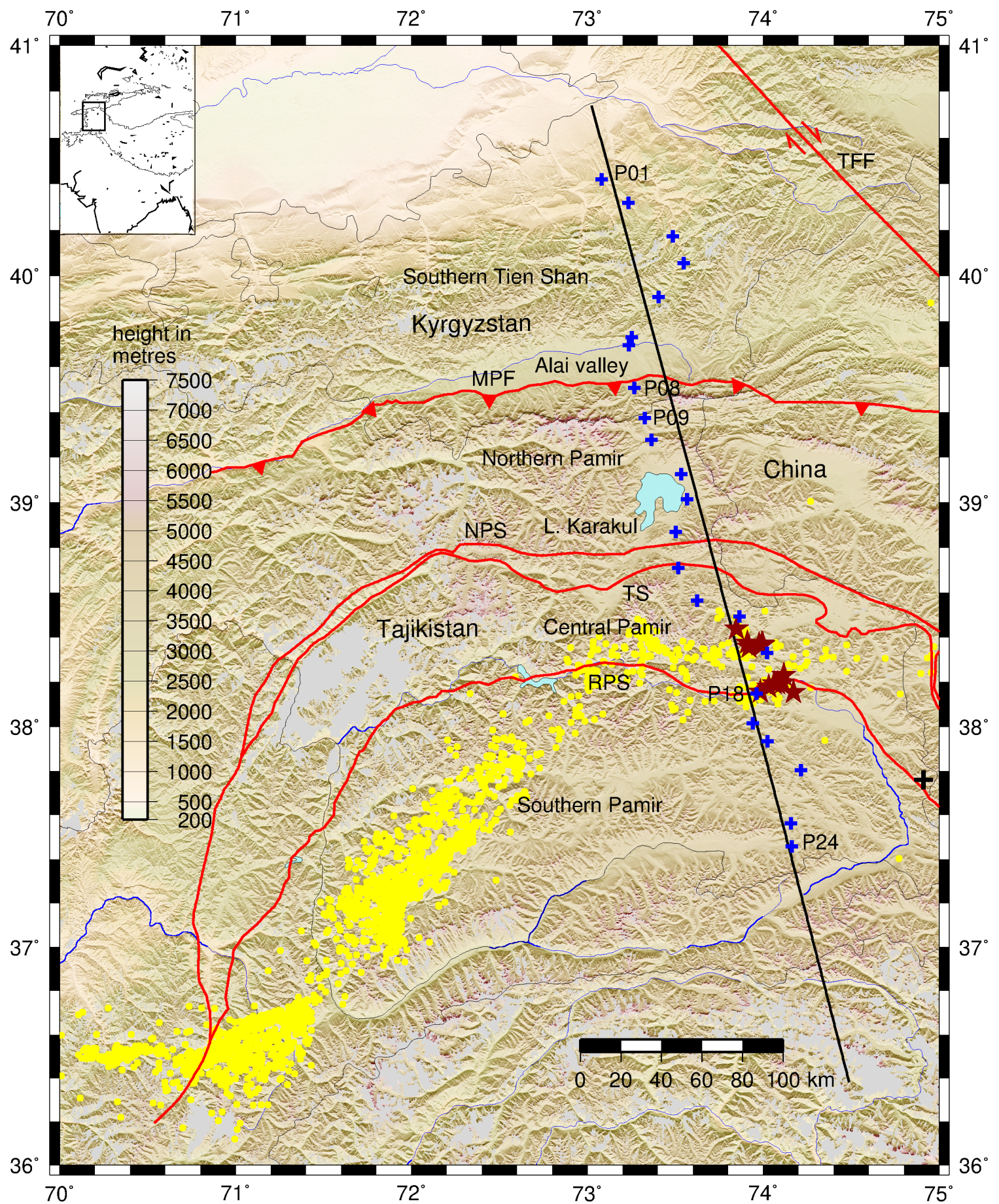


Fig. 1. Location map of the seismological stations (blue crosses) along the TIPAGE main profile (black line) and the 13 earthquakes (dark red stars) used in this study. The Pamir – Hindu Kush seismic zone delineated by earthquakes at depths greater than 90 km from the catalogue of Sippl et al. (2013a) is shown by yellow dots. The xenolith locality of Hacker et al. (2005) in the southeastern Pamir is marked by a black cross. The topography is from the Shuttle Radar Topography Mission (SRTM). In the inset the outline of Tibet, the Pamir and the Tien Shan, as defined by the 3000 m contour, is shown. Key: MPF – Main Pamir Fault, NPS – Northern Pamir Suture, TS – Tanymas Suture, RPS – Rushan Pshart Suture, TFF – Talas Fergana Fault.

Figure 2

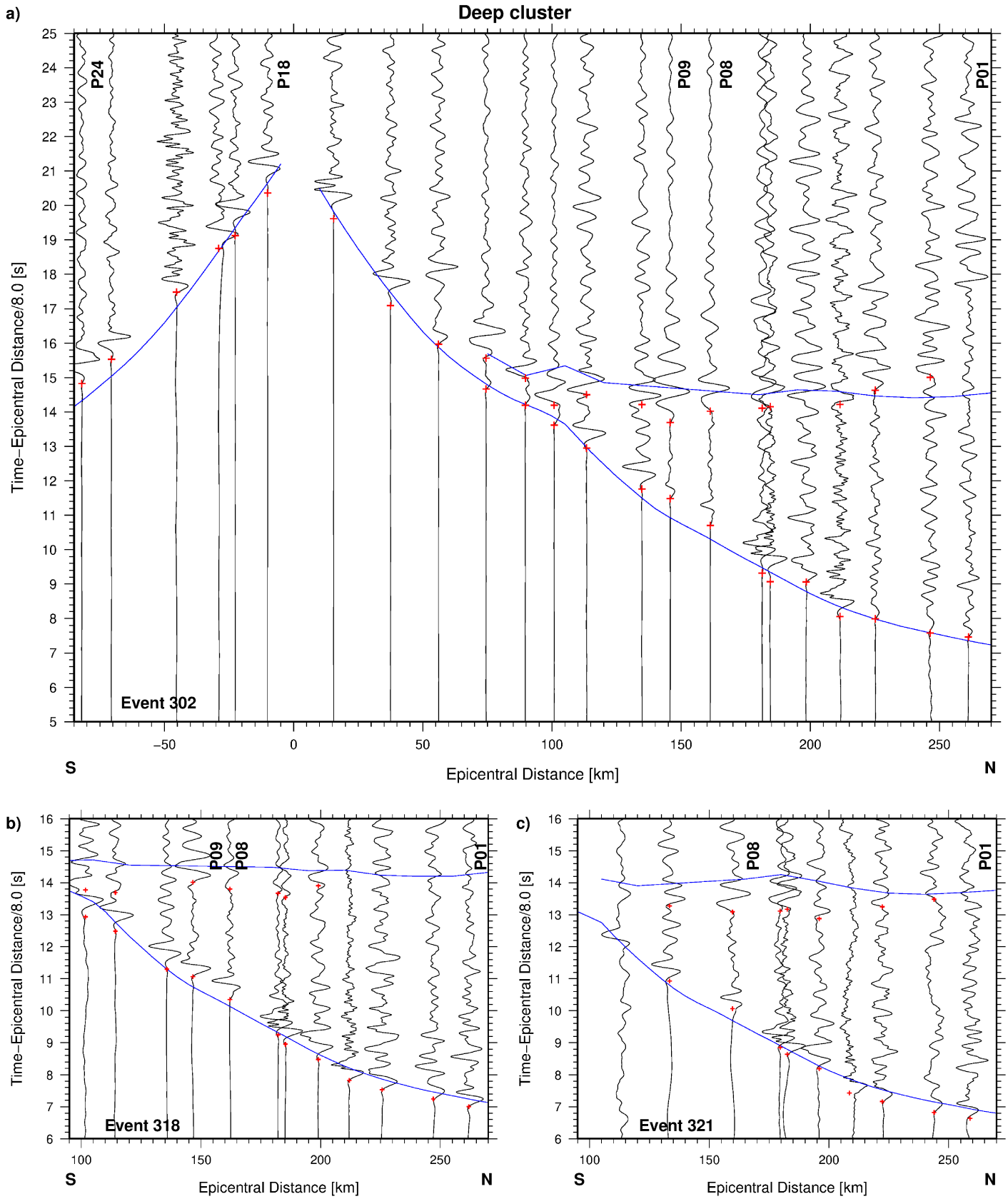


Fig. 2. Observations of guided waves from the deep cluster of earthquakes. Data from events a) 302, b) 318 and c) 321 are shown. The record sections show the vertical component of motion reduced with a velocity of 8 km/s. Each trace is normalized individually and is bandpass filtered from 0.1 to 2 Hz. The red crosses mark the picked 1st arrivals and 2nd arrivals of the guided waves. The blue lines mark the theoretical 1st arrivals and 2nd arrivals of the guided waves from model 2 shown in Fig. 6. The traces recorded by stations which are labelled in Fig. 1, including P08 and P09 for which the dispersion analysis was carried out, are also marked.

Figure 3

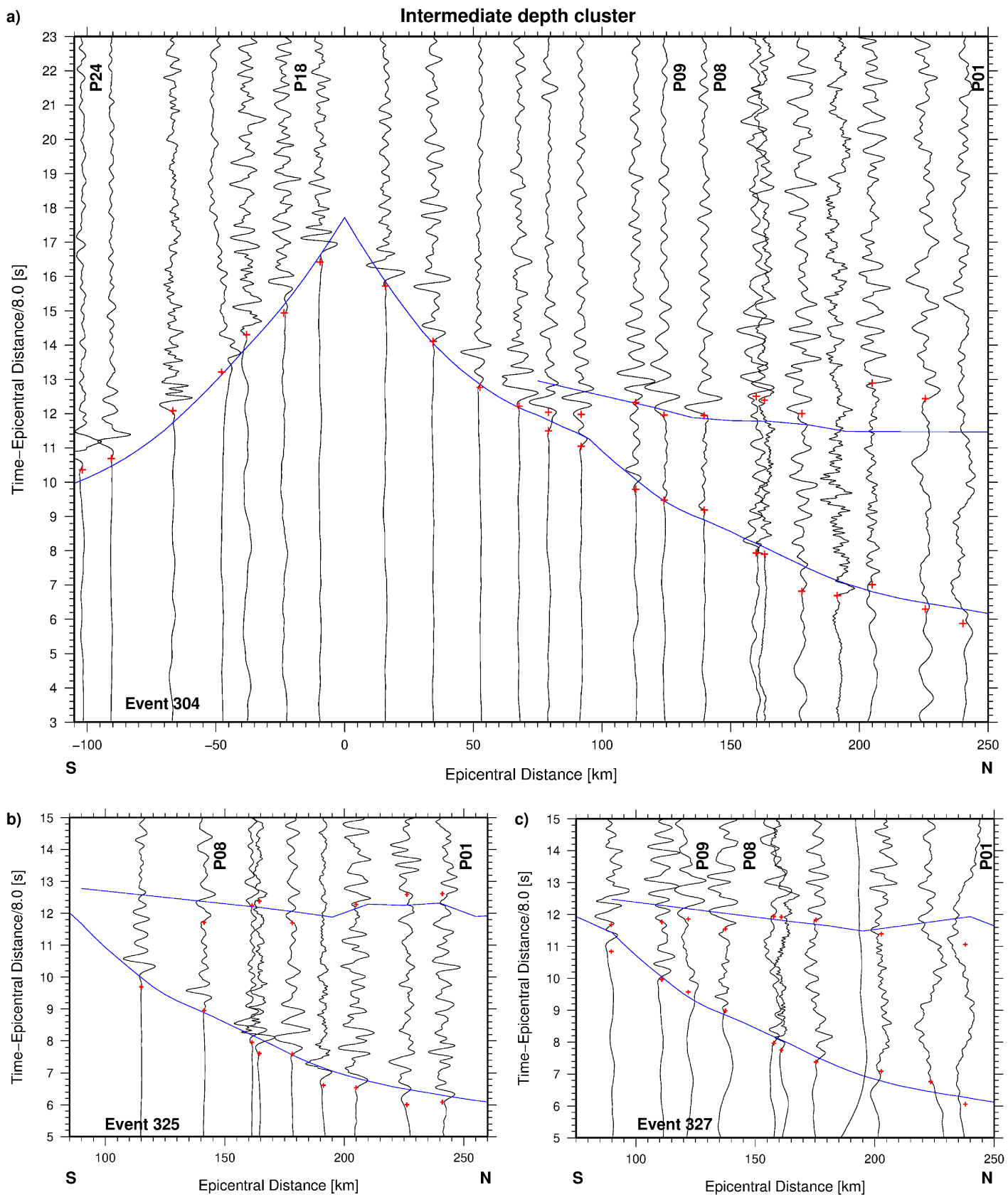


Fig. 3. Observations of guided waves from the intermediate cluster of earthquakes. Data from events a) 304, b) 325 and c) 327 are shown. The data are processed and presented as in Fig. 2.

Figure 4

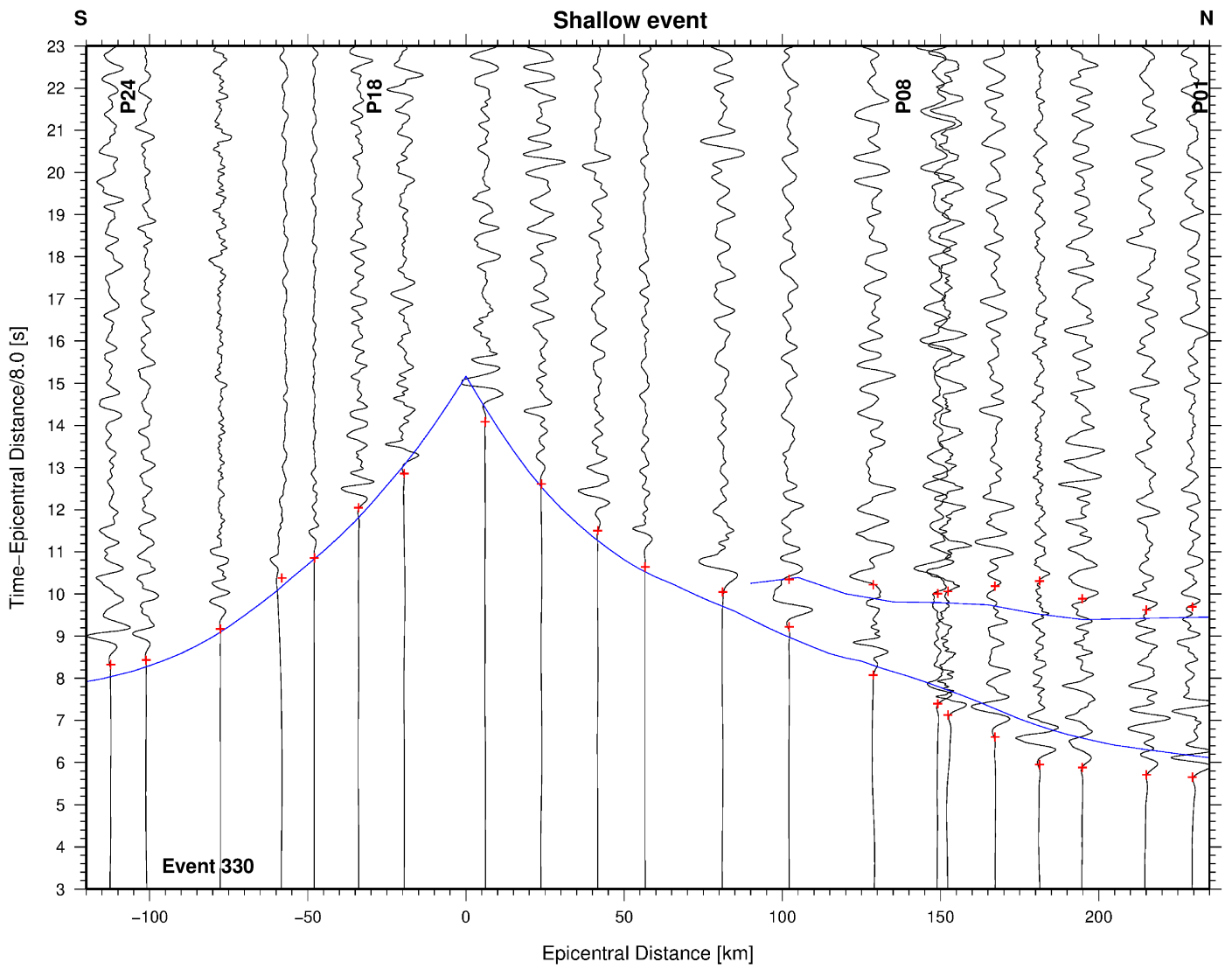


Fig. 4. Observations of guided waves from the shallow event, 330. The data are processed and presented as in Fig. 2.

Figure 5

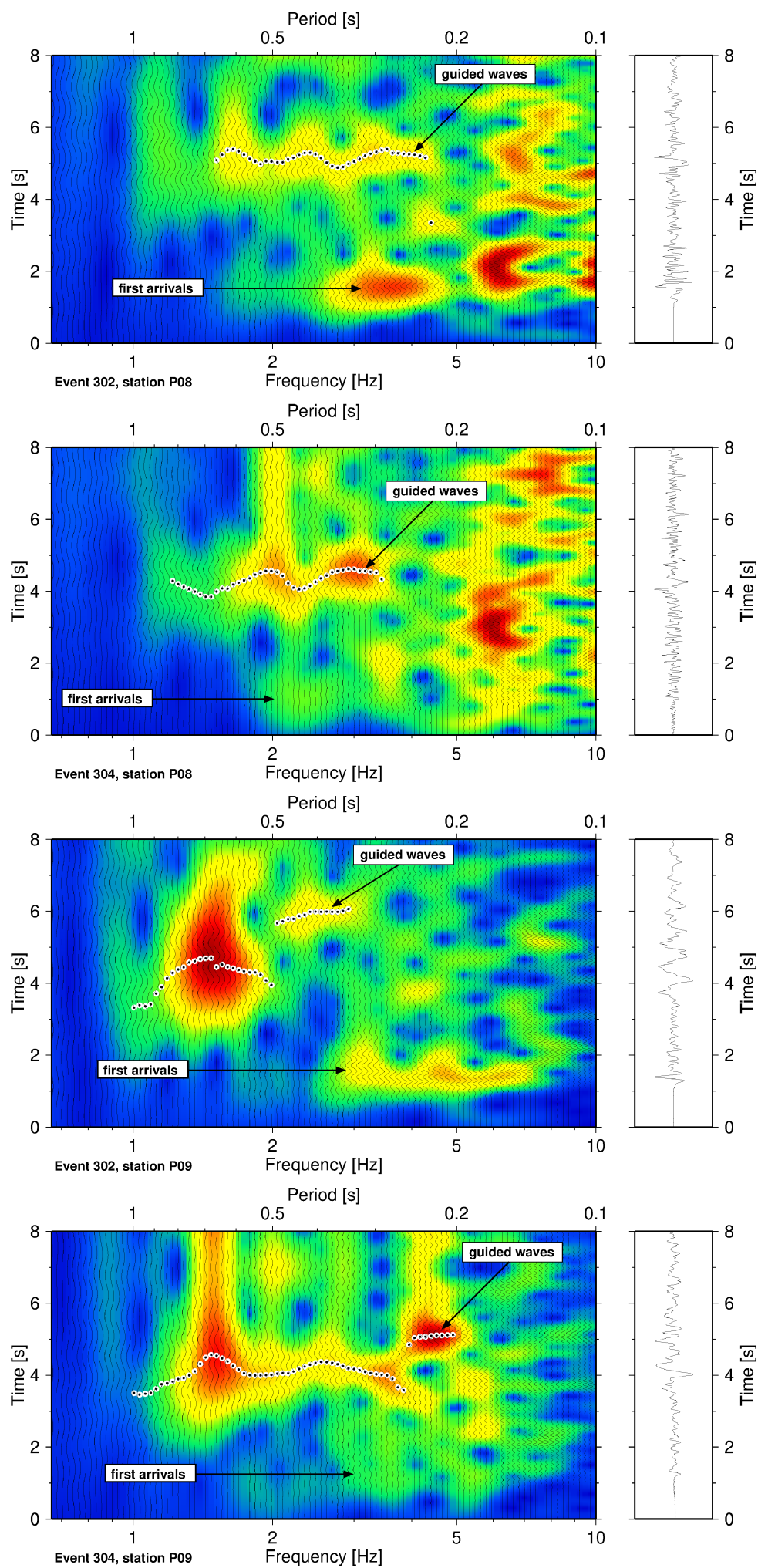


Fig. 5. Frequency (period) versus time plots for the P-wavefield recorded by the vertical component at two stations (P08 & P09, Figs. 2–3) for two events (302 & 304, Table 1). In each case, the analysed seismogram is shown to the right of the frequency–time plot and the dots mark the frequency–dependent arrival time of the guided waves. The locations of the events and stations are marked in Figs. 1, 6 and 7 and the relevant traces are also marked in Figs. 2a and 3a.

Figure 6

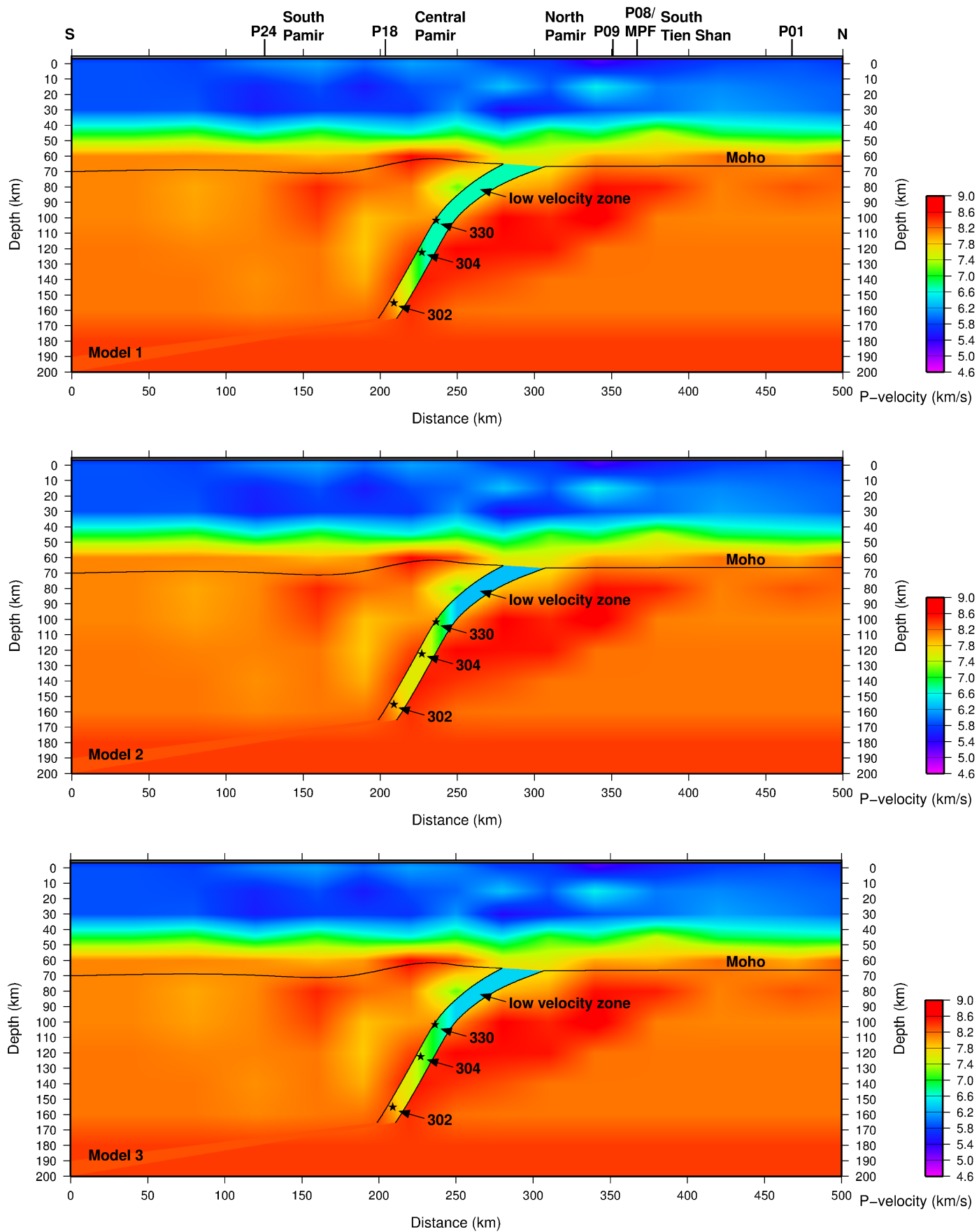


Fig. 6. Velocity models along the TIPAGE main profile (black line in Fig. 1) with a low velocity zone (LVZ) in the mantle producing guided waves. In model 1 (upper plot) the LVZ has a velocity of 6.7 km/s at depths shallower than about 130 km. In model 2 (middle plot) the velocities in the LVZ are such that the velocity above the shallow earthquake has a value of 6.3 km/s, and the velocity between the shallow earthquake and the deep earthquakes has a value of 7.6 km/s. In model 3 (lower plot) the velocity in the LVZ above the shallow earthquake was assigned a value of 6.4 km/s, the velocity between the shallow earthquake and the intermediate cluster of earthquakes was assigned a value of 7.2 km/s and the velocity between the intermediate and deep clusters of earthquakes was assigned a value of 7.6 km/s. The theoretical 1st arrivals and 2nd arrivals of the guided waves for model 2 are shown for various events in Figs. 2–4. The Moho between 60 and 70 km depth is after Schneider et al. (2013) and there is no vertical exaggeration. Stations labelled in Fig. 1 are also marked. Key: see Fig. 1.

Figure 7

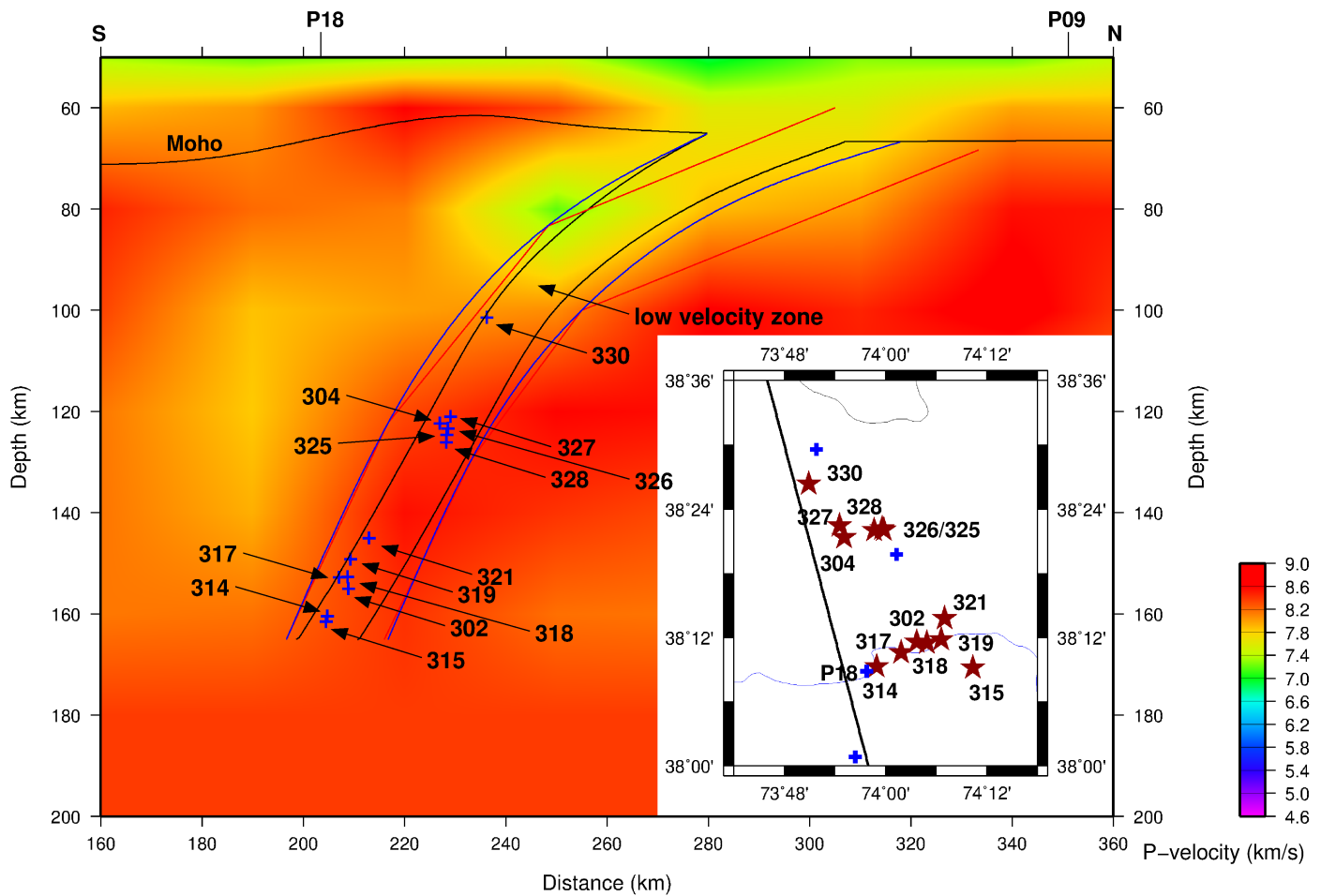


Fig. 7. The 13 events (see also large-scale inset map) from the Pamir seismic zone which showed guided waves and were used in this study in relation to the boundaries of the low velocity zone (LVZ) and the background velocity model (see also cross-section D-D' in Fig. 2 of Sippl et al., 2013b). The red lines are the boundaries of the LVZ shown in Schneider et al. (2013) while the blue and black lines are the boundaries of the LVZ used in this study (blue – initial, test models, black – later models) and also shown in Figs. 6 and S1–S2. The Moho between 60 and 70 km depth is after Schneider et al. (2013) and there is no vertical exaggeration. Stations labelled in Fig. 1 are also marked.

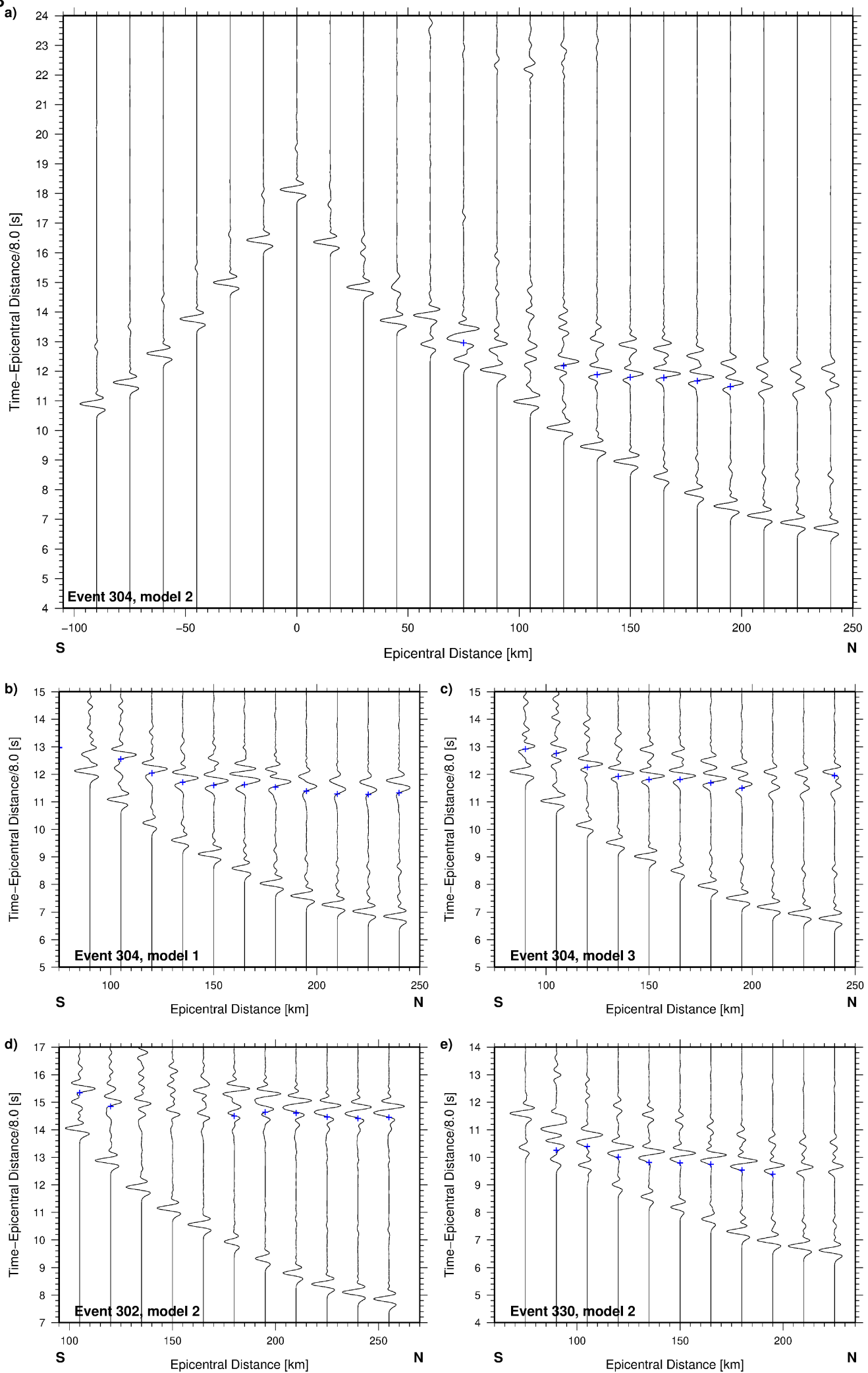
Figure 8

Fig. 8. Synthetic seismograms for event 304 for a) model 2, b) model 1 and c) model 3 and for events d) 302 and e) 330 for model 2. Models 1, 2 and 3 are described and shown in Fig. 6. The blue crosses mark the picked arrivals of the guided waves. The synthetic seismograms which show the vertical component of motion, were calculated for an explosive point source with a dominant frequency of 1.5 Hz. The record sections are reduced with a velocity of 8 km/s and each trace is normalized individually.

Figure 9

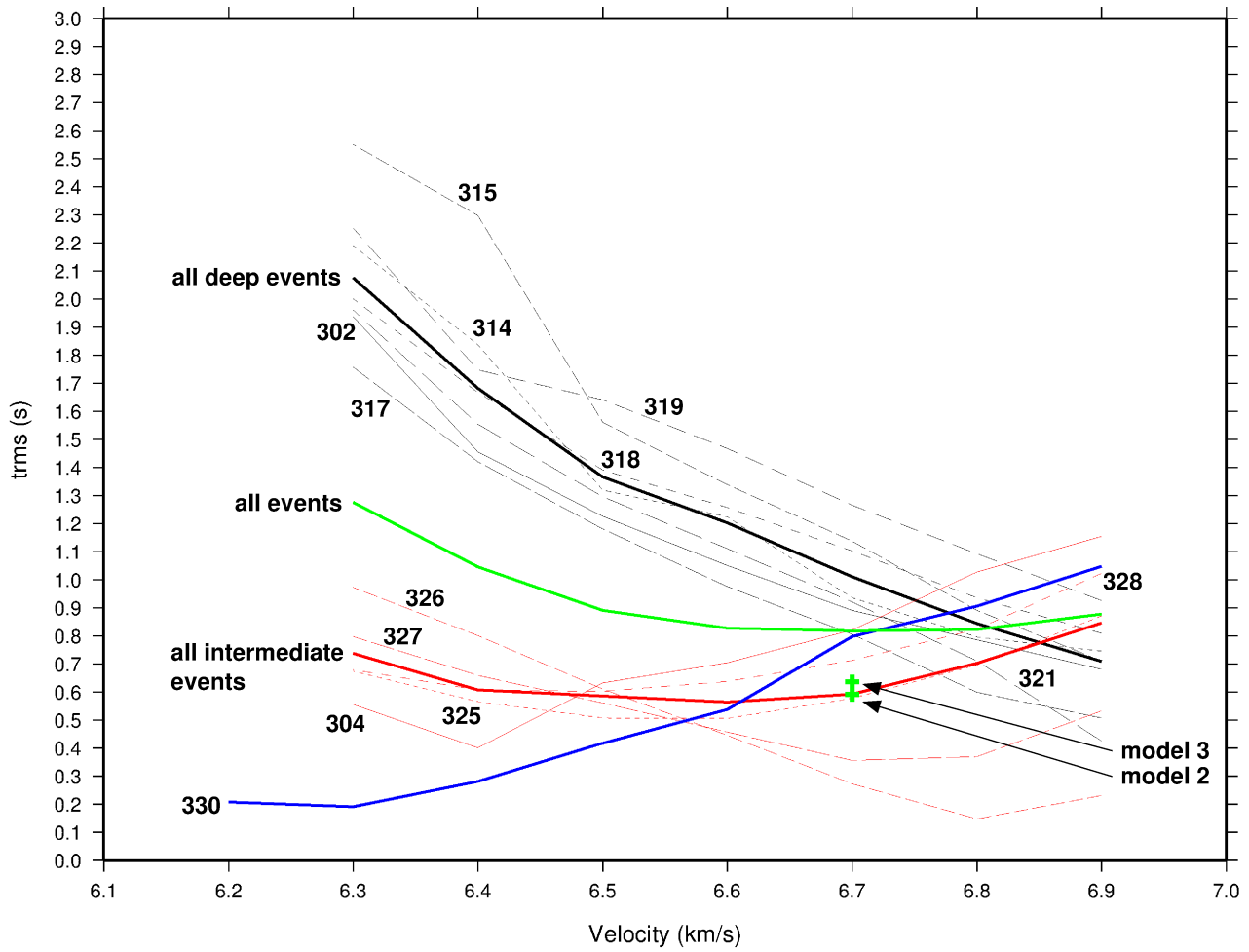


Fig. 9. Plot of root mean square travel-time residuals (trms) against the velocity in the low velocity zone at depths shallower than 130 km. For each of the 13 events, the event no. is plotted beside the line. Additionally, lines are shown for the five intermediate events (thick red line), the seven deep events (thick black line) and for all events (thick green line), and green crosses, plotted at the somewhat arbitrary velocity value of 6.7 km/s, are shown for models 2 and 3. For a description of models 2 and 3, see Fig. 6.

Figure 10

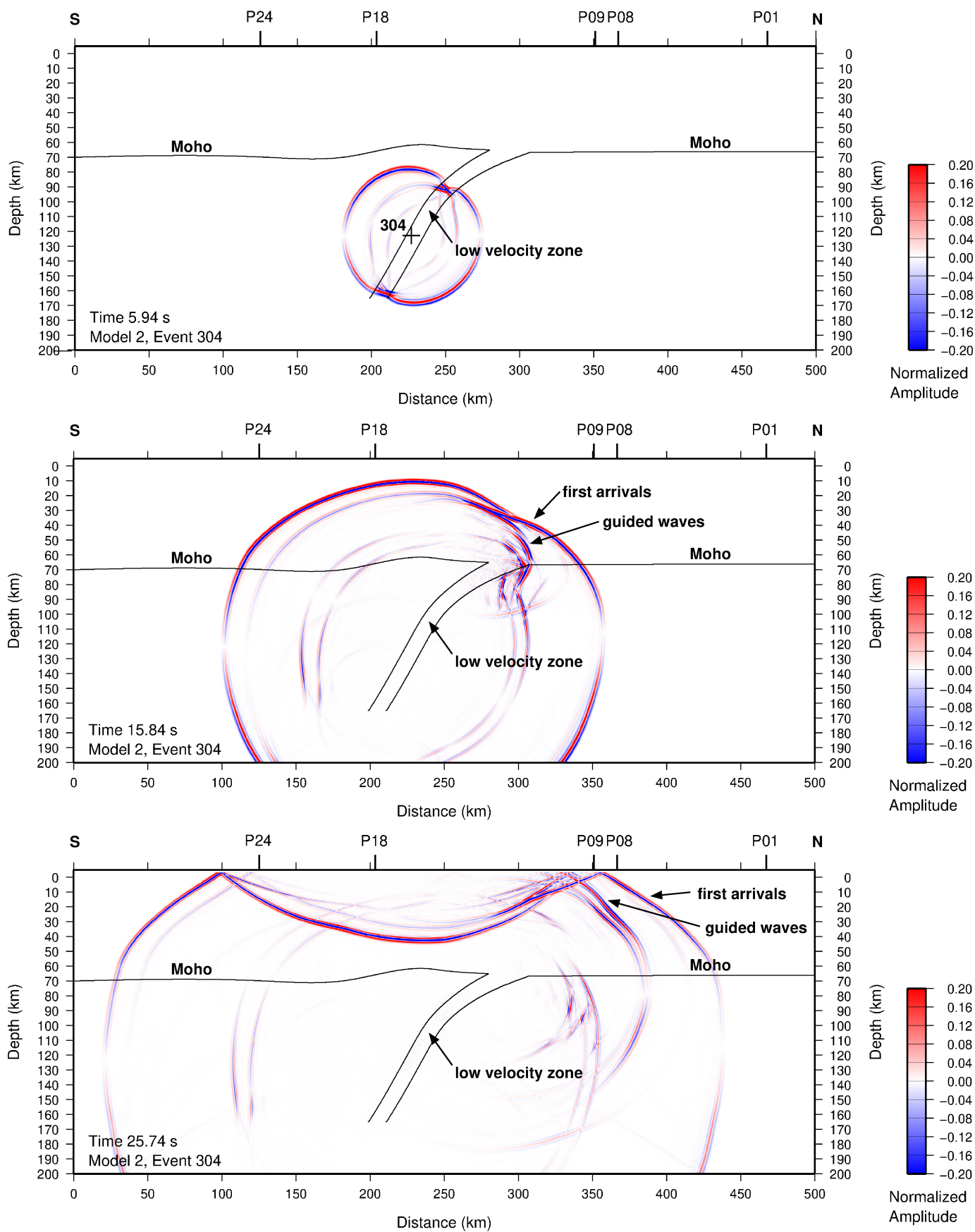


Fig. 10. Snapshots of waves propagating from event 304 through model 2 (Fig. 6). The snapshots are shown for 5.94 s (top), 15.84 s (middle) and 25.74 s (bottom) after the origin time of the event. The boundaries of the low velocity zone are marked as is the Moho between 60 and 70 km depth after Schneider et al. (2013). There is no vertical exaggeration. Stations labelled in Fig. 1 are also marked. An animation of this example is shown as Video S1.

Figure 11

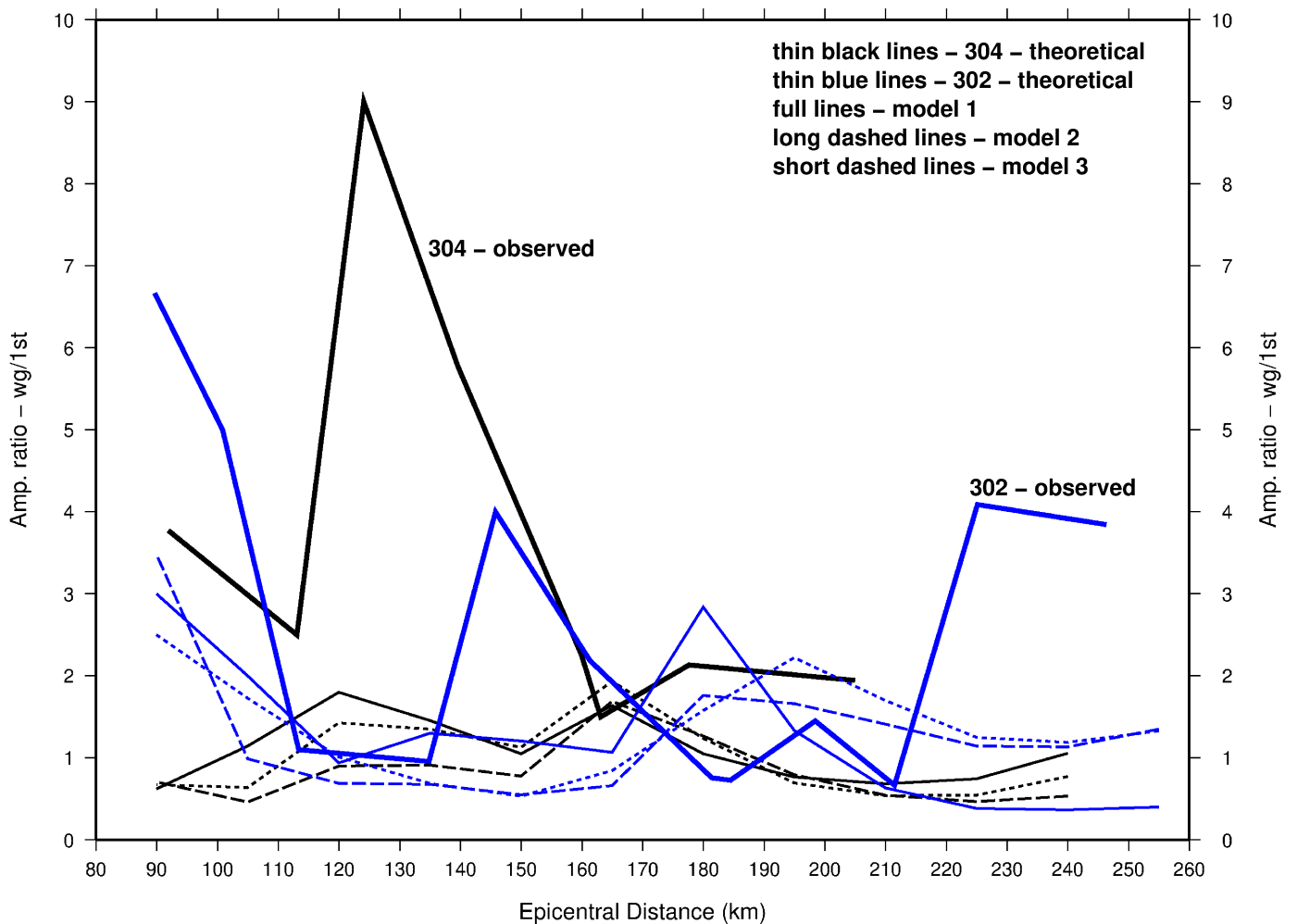


Fig. 11. Plot of amplitude ratios between the guided waves and the 1st arrivals (Amp. ratio - wg/1st) against the epicentral distance for the two events 302 and 304 and for the models 1-3 defined in Fig. 6. The theoretical amplitude ratios from models 1-3 are marked by thin blue lines for event 302 and thin black lines for event 304, while the observed amplitude ratios for events 302 and 304 are marked by thick blue and black lines respectively. Large amplitude ratios mean high amplitude guided waves with respect to the amplitudes of the 1st arrivals.

Figure 12

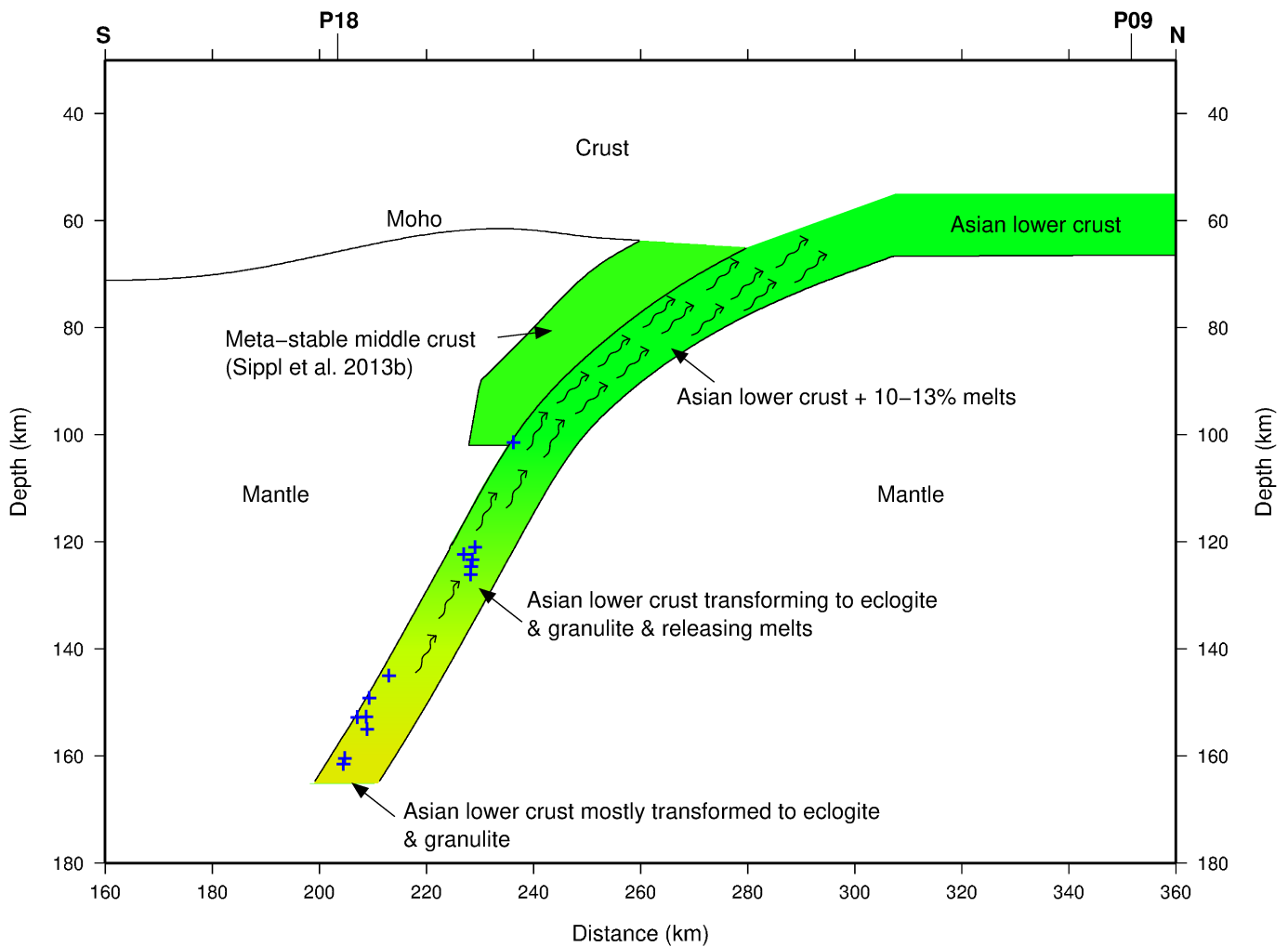


Fig. 12. Schematic diagram of the subduction of Asian crust within the Pamir seismic zone and the transformation of the lower crust to eclogite and granulite accompanied by the release of melts. The meta-stable middle crust is shown after Sippl et al. (2013b). Blue crosses mark the locations of the 13 events showing good guided waves used in this study. The melts are thought to be confined to the zone where the lower crust is subducting for density reasons.