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1 **The depositional flux of meteoric cosmogenic ¹⁰Be from modeling and**
2 **observation**

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14 **Abstract:** Meteoric cosmogenic ^{10}Be is a powerful tracer to quantify dates and rates of Earth surface
15 processes over timescales of 10^3 - 10^5 yrs. A prerequisite for its applications is knowledge of the flux at
16 which ^{10}Be , produced in the atmosphere, is delivered to the Earth surface. Four entirely independent
17 approaches are available to quantify this flux: 1) General Circulation Models (GCM) combined with ^{10}Be
18 production functions and aerosol dynamics; 2) ^{10}Be in precipitation collections; 3) ^{10}Be inventories in
19 dated soil profiles; 4) riverine ^{10}Be exported in solid and dissolved forms. We compiled and reprocessed
20 published globally distributed ^{10}Be flux data from each of these methods and compared them with each
21 other after normalization to a common atmospheric production rate. Based on precipitation records, we
22 propose a simple framework to discriminate between two delivery effects on ^{10}Be fluxes. In the additive
23 effect water vapor and ^{10}Be are continuously accumulating during long-distance transport, leading to an
24 increase in ^{10}Be flux with precipitation rate. In the dilution effect, the ^{10}Be flux is delivered from proximal
25 vapor sources, limited by the rate of ^{10}Be introduction from the stratosphere and independent of
26 precipitation rate. Both effects are mostly present in combination, and the relative weight of either effect
27 depends on vapor condensation rate and on the ratio of vapor condensation area to precipitation area. A
28 comparison between precipitation-derived fluxes and GCM-derived fluxes shows that half of the
29 precipitation estimates are >2 times greater than GCM-derived fluxes. By comparison, soil- and GCM-
30 derived fluxes agree within a factor of 2 for more than half ($\sim 57\%$) of the dataset, and the remaining soil
31 estimates ($\sim 43\%$) are much lower than GCM-derived fluxes. 71% of ^{10}Be flux estimates from riverine
32 export using ^{10}Be (meteoric)/ ^9Be ratios also agree with GCM-derived fluxes within a factor of 2. We
33 explain the precipitation-derived fluxes that commonly exceed all other estimates by short-term
34 stochasticity in precipitation events that might introduce a measurement time-interval bias towards higher
35 fluxes. This bias is not present over longer-term (10^3 - 10^5 yrs) flux estimates like those from soil profiles.
36 Soil-derived fluxes might still present an underestimation when retention of ^{10}Be in soil is incomplete. We
37 recommend producing more ^{10}Be depositional flux data from soil inventories with full Be retention, as
38 these generate in our view the most relevant estimates for applications of meteoric ^{10}Be on millennial-
39 scale Earth surface processes.

40 **Keywords:** cosmogenic nuclides, meteoric ^{10}Be , general circulation model (GCM), production rates,
41 Earth surface processes, geochronology

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43 **1. Introduction**

44 When galactic cosmic rays impinge upon the atmosphere, they induce spallation reactions in oxygen
45 and nitrogen atoms that produce meteoric cosmogenic nuclide beryllium-10 (^{10}Be) that is radioactive with
46 a half-life of 1.39 Myr (Chmeleff et al., 2010). Meteoric ^{10}Be is scavenged from the atmosphere primarily
47 by precipitation after adhering to aerosols. When delivered to the Earth's surface it binds tightly to soil
48 particles provided that the pH of soil solutions is circum-neutral. The concentration of meteoric ^{10}Be
49 measured in Earth surface materials is thus controlled by the following processes: ^{10}Be production and
50 redistribution in the atmosphere, ^{10}Be delivery to the surface, ^{10}Be retention in the substrate, and substrate
51 stability with regards to erosion (Willenbring and von Blanckenburg, 2010b). As such, meteoric ^{10}Be has
52 been used as a tracer for a wide range of Earth surface processes over timescales of 10^3 - 10^5 yrs, including
53 determining soil residence time (Bacon et al., 2012; Pavich et al., 1986) and ages of sediment archives
54 (Egli et al., 2010; Lebatard et al., 2010), tracing soil movement (Jungers et al., 2009; Mckean et al., 1993),
55 quantifying rates of erosion and weathering over hillslope scale (Maher and von Blanckenburg, 2016) and
56 over basin scale (Brown et al., 1988; Portenga et al., 2019; von Blanckenburg et al., 2012; Wittmann et al.,
57 2015), and also the expression of these processes back through time (von Blanckenburg et al., 2015;
58 Willenbring and von Blanckenburg, 2010a). In order to quantify these dates and rates of Earth surface
59 processes using meteoric ^{10}Be , knowledge of the ^{10}Be depositional flux is a key prerequisite.

60 ^{10}Be production in the atmosphere depends on the primary cosmic ray flux and thus on solar activity
61 and geomagnetic field strength (Masarik and Beer, 2009). After production, ^{10}Be is well-mixed in the
62 stratosphere (~ 1 yr) and hydrolyzed to surface-reactive $\text{Be}(\text{OH})_2$. After adsorption to atmospheric
63 aerosols, ^{10}Be is delivered to Earth's surface by wet and dry deposition. The delivery thus depends on

64 atmospheric aerosol loading and climate, which can significantly modify the temporal and spatial
65 distribution of ^{10}Be depositional fluxes (Field et al., 2006; Heikkilä et al., 2013b). Over recent decades,
66 substantial efforts on modeling and observation have been conducted to advance our knowledge of ^{10}Be
67 deposition over a range of temporal and spatial scales. These efforts include four independent approaches.

68 1) Atmospheric production and delivery models (Heikkilä et al., 2013a). In order to study production,
69 atmospheric transport and deposition of meteoric ^{10}Be , physics-based ^{10}Be production functions (Masarik
70 and Beer, 2009) and aerosol physics-chemistry (Stier et al., 2005) have been coupled with atmospheric
71 general circulation models (GCM) (Schmidt et al., 2006; Stier et al., 2005). These integrated models have
72 revealed the spatio-temporal pattern of ^{10}Be depositional flux and its response to variations in solar
73 activity, geomagnetic field strength (Field et al., 2006; Heikkilä et al., 2008b), and climate, including
74 atmospheric mixing and precipitation (Heikkilä et al., 2013b; Heikkilä and Smith, 2013).

75 2) Precipitation collections. Over contemporary timescales, ^{10}Be measurements in rainfall and snow
76 in combination with precipitation rates yield local depositional fluxes. Based on such data, previous
77 studies investigated the effects of latitude (Graly et al., 2011), altitude (Heikkilä et al., 2008a),
78 stratosphere-troposphere exchange (Graham et al., 2003), precipitation rate (Willenbring and von
79 Blanckenburg, 2010b), and aerosol content (Mann et al., 2011) on ^{10}Be depositional fluxes.

80 3) Soil profiles. Over millennial timescales, inventories of meteoric ^{10}Be in non-eroding soil profiles
81 of known ages, developed on river terraces, moraine deposits, or bedrock, provide site-specific
82 depositional fluxes that average over 10^3 to 10^5 years when assuming full retention (Dixon et al., 2018;
83 Ouimet et al., 2015; Reusser et al., 2010). If ^{10}Be retention is incomplete in soil profiles, this approach
84 provides a lower-limit estimate (Maher and von Blanckenburg, 2016; Schoonejans et al., 2017).

85 4) Riverine fluxes. Over erosion and weathering timescales (e.g. 10^3 - 10^4 yrs), the atmospheric input
86 flux of ^{10}Be is expected to be balanced by the riverine sedimentary and dissolved ^{10}Be flux (Wittmann et

87 [al., 2015](#)). Assuming such steady state, the ^{10}Be flux from riverine export thus reflects the ^{10}Be
88 depositional flux over this long timescale.

89 Previous studies, including [Willenbring and von Blanckenburg \(2010b\)](#) and [Graly et al. \(2011\)](#),
90 investigated the fluxes derived from different approaches by comparing modeled with measured ^{10}Be
91 depositional fluxes on the one hand, and short-term (precipitation-derived) with long-term (soil-derived)
92 ^{10}Be depositional fluxes in the other hand, respectively. In both types of comparison, the overall
93 agreement between approaches was encouraging. However, in recent studies ([Dixon et al., 2018](#); [Ouimet](#)
94 [et al., 2015](#)), the ^{10}Be depositional fluxes obtained from dated soil profiles clearly differed from flux
95 estimates by GCM ([Heikkilä and von Blanckenburg, 2015](#)), and also from a fitting equation based on
96 precipitation ^{10}Be records ([Graly et al., 2011](#)). The lack of agreement between these estimates has been
97 explained by local climatic variability ([Heikkilä et al., 2013a](#)), lack of suitability of soil inventories, or
98 failure to adequately account for temporal variations in cosmic ray fluxes.

99 An understanding of the climatic control on ^{10}Be depositional flux, meaning the relationship between
100 ^{10}Be concentration, flux, and precipitation rate, is essential for upscaling contemporary ^{10}Be fluxes in
101 space and time. In this regard, two contrasting views exist. One observation holds that the ^{10}Be
102 depositional flux is positively correlated with precipitation rate, and that this precipitation dependence can
103 explain the variations of ^{10}Be inventories in some soil profiles ([Dixon et al., 2018](#); [Graly et al., 2011](#)).
104 Other studies show no correlation between ^{10}Be flux and precipitation, and regard the ^{10}Be flux to be
105 independent of precipitation ([Ouimet et al., 2015](#); [Willenbring and von Blanckenburg, 2010b](#)). To account
106 for these disparate views, [Willenbring and von Blanckenburg \(2010b\)](#) proposed two ^{10}Be delivery effects:
107 the “dilution effect” and the “additive effect”. In the dilution effect, precipitation is delivered from moist
108 air mass with proximal vapor sources where moisture has high turnover rates, such that the delivery of
109 meteoric ^{10}Be is ultimately limited by the rate at which it is introduced into moist air mass by
110 stratospheric production. Hence the ^{10}Be concentration in precipitation can be “diluted” ([Willenbring and](#)
111 [von Blanckenburg, 2010b](#)). In the additive effect, water vapor is continuously accumulating in moist air

112 mass during long-distance transport, and so is ^{10}Be , such that the ^{10}Be depositional flux will increase with
113 precipitation rate (Graly et al., 2011). However, the competition between both effects is still unclear and
114 requires a comprehensive comparison between flux estimates across spatial and temporal scales.

115 In this study, we compiled literature data and recalculated ^{10}Be fluxes recorded in different kinds of
116 materials including precipitation, soil profiles and riverine solid and solute loads, and compared them to
117 modeled (GCM-derived) ^{10}Be depositional fluxes given the wide spatial coverage (i.e. global scale) of the
118 GCM-derived dataset. The main aim is to evaluate the reliability of ^{10}Be depositional fluxes from each
119 approach and the associated bias in flux derivation. This study advances the understanding of meteoric
120 ^{10}Be delivery and deposition over a range of spatial and temporal scales, and provides new insights on
121 method-specific uncertainties on ^{10}Be flux estimates which shall benefit future Earth surface applications
122 using meteoric ^{10}Be .

123 **2. Methods: compilation of modeled and measured ^{10}Be fluxes and assessment of delivery effects**

124 *2.1. GCMs with ^{10}Be production functions and aerosol dynamics*

125 To investigate the spatio-temporal pattern of ^{10}Be deposition, two GCMs, including GISS (Goddard
126 Institute for Space Studies) ModelE (Field et al., 2006) and ECHAM5 (European Centre for Medium-
127 Range Weather Forecasts-Hamburg Model 5)-HAM (Heikkilä et al., 2013a), were previously applied by
128 incorporating meteoric ^{10}Be production functions in the atmosphere (Masarik and Beer, 2009) and aerosol
129 physics-chemistry. Comparing both models, the GISS ModelE with 20 vertical layers and a horizontal
130 resolution of $4^\circ \times 5^\circ$ (Field et al., 2006) shows a coarser spatial resolution than the ECHAM5-HAM that
131 includes 31-39 vertical layers and a horizontal resolution of $2.8^\circ \times 2.8^\circ$ (Heikkilä et al., 2008b). It has been
132 suggested (Heikkilä et al., 2013a) that the lower spatial resolution of GISS ModelE (Field et al., 2006)
133 might bias the modeling of atmospheric mixing and advection of ^{10}Be , leading to artificial latitudinal
134 differences in the response of ^{10}Be deposition to production changes (Heikkilä et al., 2008b). We thus
135 focus on the modeled ^{10}Be depositional flux derived from the ECHAM5-HAM (Fig. 1).

136 We used published results of the ECHAM5-HAM model based on a ^{10}Be production rate calculated
137 for a modern solar modulation factor (ϕ in MeV) of 501.76 MeV (Fig. 2). To account for the climate-
138 related uncertainty introduced by spatial variations in ^{10}Be delivery over the Holocene, we used two
139 different model outputs: one for modern (“industrial”, time span: 1977-2006) climatic conditions
140 (Heikkilä and Smith, 2013) and the other for the early Holocene (“pre-industrial”, time span: 12-10 kyr
141 BP) conditions (Heikkilä et al., 2013b). The difference in the spatial pattern of modeled ^{10}Be depositional
142 fluxes between both runs is our estimate of flux uncertainty (Heikkilä and von Blanckenburg, 2015). Here,
143 the modeled fluxes of both runs under modern production conditions were rescaled to the average
144 Holocene ^{10}Be production rate (Fig. 2). The resulting distribution map (Fig. 1) shows the average flux of
145 both rescaled model runs.

146 2.2. Precipitation collections

147 We compiled ^{10}Be data from precipitation collectors ($[^{10}\text{Be}]_{\text{rain}}$, in at/m^3) and the corresponding
148 precipitation rate (P , in $\text{m}^3/\text{m}^2/\text{yr}$) to calculate the ^{10}Be depositional flux (F_{met} , in $\text{at}/\text{m}^2/\text{yr}$):

$$149 \quad F_{\text{met}} = [^{10}\text{Be}]_{\text{rain}} \times P \quad (1)$$

150 The analytical uncertainty of $[^{10}\text{Be}]_{\text{rain}}$ was propagated into the uncertainty of F_{met} . To obtain an
151 annual ^{10}Be depositional flux from measured monthly data, we calculated the monthly-precipitation-
152 weighted ^{10}Be concentration and multiplied it by the annual precipitation rate. The derived flux is the sum
153 of three components: wet deposition (scavenging of $^{10}\text{Be}(\text{OH})_2$ by precipitation), dry deposition (delivery
154 by friction-based aerosol deposition), and the recycled fraction (^{10}Be carried by dust). This precipitation
155 dataset covers continental settings (Fig. 1), including Switzerland (Heikkilä et al., 2008a), India
156 (Somayajulu et al., 1984), and USA (Monaghan et al., 1986), and islandic settings, including Trinidad
157 (Brown et al., 1992), Kikai Island (Japan) (Maejima et al., 2005) and New Zealand (Graham et al., 2003).
158 Records with an observation period < 9 months (see details in Supplementary dataset) are considered
159 unrepresentative over the annual scale as they might be biased by seasonal cycles (Heikkilä and Smith,

160 2013), and thus these ^{10}Be data are shown in figures for completeness but excluded from the discussion of
161 fluxes.

162 We applied normalization procedures that account for variations in the flux of galactic cosmic rays
163 and analytical differences, to compare between datasets of different observation periods and derived from
164 different methods (Fig. 2). 1) All ^{10}Be concentrations were normalized to the AMS (accelerator mass
165 spectrometry) standard KNSTD07 (Nishiizumi et al., 2007). 2) Precipitation records that are subject to
166 inter-annual variations in solar modulation were normalized to a common modern solar modulation factor
167 of 501.76 MeV. 3) The contemporary precipitation-derived fluxes were further normalized to the average
168 Holocene ^{10}Be production rate by multiplying them with ~ 1.23 which is the ratio of Holocene relative to
169 modern production and encompasses variations in solar modulation and magnetic field strength
170 (Steinhilber et al., 2012) (Fig. 2).

171 2.3. Soil profiles

172 Over a timescale integrated by most geomorphic processes (10^3 - 10^5 yrs), the site-specific ^{10}Be
173 depositional flux can be calculated from the total ^{10}Be inventory (at/m^2) measured in soil profiles.
174 Prerequisites are a known age of the onset of soil formation (t , in yr) and full retention of meteoric ^{10}Be .
175 In the case of an alluvial substrate any meteoric ^{10}Be ($[^{10}\text{Be}]_{\text{inh}}$, in at/kg) inherited prior to deposition that
176 is apparent at the bottom of profiles should be subtracted. The equation for calculating F_{met} as modified
177 from Egli et al. (2010) is:

$$178 \quad F_{\text{met}} = \sum_{i=1}^n \left[\left([^{10}\text{Be}]_{\text{s},i} - [^{10}\text{Be}]_{\text{inh},i} \right) \times \rho_i \times z_i \times V_{f,i} \right] \times \lambda / (1 - e^{-\lambda t}) \quad (2)$$

179 where $[^{10}\text{Be}]_{\text{s},i}$ is the soil ^{10}Be concentration (at/kg) at i^{th} horizon, ρ_i the soil bulk density (kg/m^3), z_i
180 the thickness of the i^{th} horizon (m), $V_{f,i}$ the volume percentage (%) of the fine earth fraction (<2 mm)
181 containing ^{10}Be at the i^{th} horizon, λ the decay constant of ^{10}Be ($5 \times 10^{-7} \text{ yr}^{-1}$). If no V_f data was reported in
182 literature because the fine fraction dominates, V_f of 100% is used. The uncertainty of the resulting F_{met} is
183 propagated from the individual uncertainties of the ^{10}Be measurement, the age determination, 10% for soil

184 density measurements, and 5% for V_f (Egli et al., 2010). We compiled ^{10}Be data from soil profiles (Fig. 1)
 185 located in continental settings, including European Alps (Egli et al., 2010) and the USA (Balco, 2004;
 186 Ouimet et al., 2015), and in islandic settings, including Hawaii (Dixon et al., 2018) and New Zealand
 187 (Reusser et al., 2010). We did not compile or show data from other archives like coral reefs (Maejima et
 188 al., 2005) or old soil profiles (age >200 ka) (Pavich et al., 1986) here because of ^{10}Be contamination by
 189 seawater and a high potential for surface erosion and incomplete retention, respectively. We applied 1)
 190 the AMS normalization same as above and 2) a millennial-scale variability normalization, depending on
 191 the integration timescale of each soil profile, to the average Holocene ^{10}Be production rate for comparison
 192 with other approaches (Fig. 2).

193 Some of the compiled profiles are considered to be affected by significant erosional loss, sampling
 194 depth not sufficient to quantify the ^{10}Be inheritance, or excessive ^{10}Be inheritance. These are shown in
 195 figures for completeness but excluded from the discussion of fluxes (see details in Supplementary dataset).

196 2.4. Riverine export

197 Over the river basin scale, when assuming steady state of ^{10}Be fluxes over weathering and erosion
 198 timescales, the basin-averaged ^{10}Be depositional flux can be derived from the ^{10}Be flux exported by rivers
 199 (Brown et al., 1988). However, a major obstacle arises from the grain size dependence of ^{10}Be
 200 concentrations in river sediment as Be adsorption depends on particle surface area (Singleton et al., 2016;
 201 Wittmann et al., 2012), more specifically, on adsorption site density (Maher and von Blanckenburg, 2016).
 202 As such, we adopted the isotopic ratio of ^{10}Be to ^9Be which is independent of mineralogical sorting,
 203 sorption capacity, and Be retentivity (Wittmann et al., 2012). We rearranged the equation for calculating
 204 denudation rates using ^{10}Be (meteoric)/ ^9Be ratios (von Blanckenburg et al., 2012) to solve for F_{met} :

$$205 \quad F_{\text{met}} = D \times \left(\frac{^{10}\text{Be}}{^9\text{Be}} \right)_{\text{react}} \times [^9\text{Be}]_{\text{parent}} \times f_{\text{react+diss}}^9 \times (1 - e^{-\lambda t_s}) \quad (3)$$

206 where D (kg/m²/yr) is the independently known catchment-wide denudation rate, $(^{10}\text{Be}/^9\text{Be})_{\text{react}}$ is the
 207 concentration ratio of ^{10}Be (at/kg) to ^9Be (mg/kg) measured in the reactive fraction of river sediments

208 (adsorbed onto or precipitated in secondary minerals). $[^9\text{Be}]_{\text{parent}}$ (mg/kg) comprises the ^9Be present in the
209 parent bedrock prior to weathering, determined by assuming a value of the average upper continental
210 crust (2.5 ± 0.5 mg/kg) (von Blanckenburg et al., 2012) or by measuring local bedrock samples in small
211 catchments (Dannhaus et al., 2018; Deng et al., 2020). $f_{\text{react+diss}}^9$ represents the mobile flux fraction of ^9Be
212 that is released from primary minerals during weathering and measured in the sediment reactive fraction
213 and dissolved fraction. Radioactive decay of ^{10}Be is negligible if the timescale of sediment transfer and
214 storage (t_s) is short compared to its half-life.

215 We compiled literature meteoric ^{10}Be and ^9Be data of river sediment with *in situ* cosmogenic ^{10}Be
216 data nearby, which was used as an independent estimate of D in Eq. (3). This dataset covers a wide spatial
217 range (Fig. 1), including the Amazon River basin in South America (Wittmann et al., 2015), the Potomac
218 River basin in USA (Portenga et al., 2019), the Zhuoshui River in Taiwan Island (Deng et al., 2020), the
219 Ganga River basin in South Asia (Rahaman et al., 2017), and small catchments in the Slavkov Forest of
220 Czech Republic (Dannhaus et al., 2018). The reported water pH was >7 , suggesting a low dissolved Be
221 flux, in most of studied rivers except for catchments in the Czech Republic, where the dissolved Be loss
222 was corrected to calculate $f_{\text{react+diss}}^9$ (Dannhaus et al., 2018). All these ^{10}Be measurements were based on
223 the standard KNSTD07 or equivalents, and the used rock density was normalized to 2.65 g/cm 3 .

224 2.5 Discrimination of ^{10}Be delivery effects

225 We discriminate between additive and dilution effects using a simple mass balance framework of
226 water vapor and ^{10}Be in the “source” (condensation) and the “sink” (precipitation). Details of mass
227 balance equations are provided in Appendix 1. In general, the ^{10}Be concentration of wet deposition is
228 described by the ratio of the ^{10}Be production flux from the upper atmosphere (F_{prod} , in at/m 2 /yr) to the
229 condensation rate in the vapor source area (C , in m 3 /m 2 /yr) (Fig. 3). The difference in F_{met} between
230 additive and dilution effects lies in the travel distance of moist air mass, illustrated by the ratio of surface
231 area of condensation (A_c , in m 2) to area exposed to precipitation (A_p , in m 2). A_c/A_p ratios close to unity
232 and $\gg 1$ indicate dominance of dilution effect and additive effect, respectively.

233 In a diagram of $[^{10}\text{Be}]_{\text{wet}}$ versus $1/P$, a dilution effect is illustrated by an A_c/A_p of 1 and co-variation
234 of P with C , resulting in a slope ($\text{at}/\text{m}^2/\text{yr}$) equal to the ^{10}Be production flux F_{prod} , i.e. a constant F_{met} (Fig.
235 3a) (Willenbring and von Blanckenburg, 2010b). In contrast, a pure additive effect results in a zero-slope
236 relationship in that diagram where differences in P result from differences in A_c/A_p , and the y-intercept
237 (at/m^3) is equal to F_{prod}/C , resulting in a constant $[^{10}\text{Be}]_{\text{wet}}$ for a given C (Fig. 3c).

238 In practice, the variable behaviors of moist air masses carrying both vapor and ^{10}Be will lead to a
239 combination of delivery effects. When both A_c/A_p and C vary over the observation timescale, a mixture of
240 dilution and additive effects may result. In this case, the ^{10}Be flux can be separated into two components:
241 one component with unity A_c/A_p ratio (stable vapor source) and constant ^{10}Be flux (dilution component),
242 and the other with constant condensation rate and thus constant $[^{10}\text{Be}]_{\text{wet}}$ (additive component). In the
243 $[^{10}\text{Be}]_{\text{wet}}$ versus $1/P$ diagram the dilution component is identified from the slope (Fig. 3) and the additive
244 component from the y-intercept. For a mixture of dilution and additive components, the proportion of the
245 additive component (f_{add} , %) relative to the total ^{10}Be flux can be calculated as:

$$246 \quad f_{\text{add}} = (\text{intercept} \times P) / (\text{intercept} \times P + \text{slope}) \times 100\% \quad (4)$$

247 This relationship shows that when P gets very high and there is a y-intercept resulting from the
248 additive effect, $\text{intercept} \times P \gg \text{slope}$ and f_{add} will be close to 100%. The framework developed also
249 allows taking dry and dust ^{10}Be components into consideration. In their presence the slope in the $[^{10}\text{Be}]_{\text{wet}}$ -
250 $1/P$ diagram (Figs. 3a-c) will steepen, because dry-dust ^{10}Be depositional fluxes are independent of
251 precipitation rate and contribute as a component diluted by precipitation. Although mixing between
252 additive and dilution effects can be shown by modern precipitation records (Eq. (4)), this mixing process
253 also presents in millennial-scale records and may vary with time.

254 **3. Results: latitudinal distribution of ^{10}Be depositional fluxes**

255 All fluxes compiled from different approaches are shown as normalized to a common Holocene
256 production rate in Fig. 4, and both the original and the re-normalized ¹⁰Be fluxes are provided in
257 [Supplementary dataset](#).

258 From the GCM output, the Holocene depositional fluxes, averaged over a given latitude band, are
259 within a relatively narrow range for latitudes from 0° to 50°, ranging from 0.88×10⁶ to 1.55×10⁶ at/cm²/yr.
260 The peak of GCM-derived flux is located at ~40° and can be explained by a higher air mass contribution
261 from the stratosphere, which is the major reservoir (~65% of the total) of meteoric ¹⁰Be ([Heikkilä et al.,
262 2013a](#)).

263 From precipitation collections we recalculated ¹⁰Be depositional fluxes based on Eq. (1).
264 Contemporary ¹⁰Be depositional fluxes from reliable records (filled symbols) generally increase from low
265 to high latitudes, with a range of 0.42 to 5.30×10⁶ at/cm²/yr between latitudes of 10.4° and 47.3°. The
266 reported percentage of recycled ¹⁰Be from eolian dust, calculated by different methods for the dataset in
267 New Zealand (35-50°N), Switzerland (46-47°N) and USA (30-50°N), varies from 6.2% to 35% (see
268 [Table S1 in Supplementary materials](#)). Given the differences in dust correction methods employed and the
269 substantial uncertainties resulting thereof, total ¹⁰Be depositional fluxes are reported instead of fluxes
270 corrected for recycled fractions. As such, these estimates can be directly compared to fluxes from soil
271 profiles and riverine export where the recycled fraction is also included.

272 We also recalculated the ¹⁰Be flux for comparison based on an empirical equation for estimating
273 primary ¹⁰Be fallout derived from data fitting between precipitation rate and precipitation ¹⁰Be fluxes at
274 mid-low latitudes ([Graly et al., 2011](#)):

$$275 \quad F_{\text{met}} = P \times (1.44 / (1 + \text{EXP}((30.7 - L) / 4.36)) + 0.63) \times 10^{10} \quad (5)$$

276 where L is latitude (°) and P is precipitation rate (m³/m²/yr). The recycled ¹⁰Be component (dust
277 contribution) was subtracted here. The derived fluxes, originally rescaled to global production rate based
278 on a ϕ of 700 MeV ([Graly et al., 2011](#)), were re-normalized to an average Holocene rate following

279 normalization steps 2) and 3) in Section 2.2 for comparison with other approaches. The latitudinal trend
280 of precipitation collections is mirrored by this fitting equation (Graly et al., 2011) as expected (correlation:
281 $R = 0.60$, $p < 0.001$) given that this fitting equation is derived from the same dataset.

282 Based on Eq. (2), we recalculated ^{10}Be depositional fluxes from soil profiles. These millennial-scale
283 fluxes from reliable records (filled symbols) vary from 0.12×10^6 to 4.07×10^6 at/cm²/yr. No clear trend
284 with latitude is apparent from soil-derived fluxes as observations exist in only few latitudinal bands. Low
285 soil pH (<4) may lead to non-negligible loss of dissolved ^{10}Be (Maher and von Blanckenburg, 2016) and
286 thus may bias the soil-derived ^{10}Be flux to lower values. According to those studies providing soil pH
287 data, pH of representative soil profiles were between 6 and 8 in Hawaii (~20°N) (Dixon et al., 2018),
288 above 5 in the USA (40-45°N) (Ouimet et al., 2015), and between 3 and 5 in the European Alps (45-50°N)
289 (Egli et al., 2010). Hence, some soil-derived ^{10}Be fluxes in the European Alps might only provide a lower
290 limit. In general, for a given latitudinal band the precipitation-derived fluxes are higher than the soil-
291 derived estimates.

292 Based on Eq. (3), ^{10}Be fluxes over a larger spatial scale, that is the river basin-scale, are evaluated.
293 Riverine ^{10}Be fluxes derived from ^{10}Be (meteoric)/ ^9Be ratios increase with latitude (1° to 30°) and range
294 from 0.24 to 5.74×10^6 at/cm²/yr, and become quite scattered at 38-40° (the Potomac Basin) with a mean
295 value of 1.84×10^6 at/cm²/yr. At a higher latitude of ~50° (Czech Republic), riverine ^{10}Be fluxes decrease
296 to 0.44 - 1.18×10^6 at/cm²/yr. Riverine fluxes from the single isotopic system (meteoric ^{10}Be) are also
297 provided in [Supplementary dataset](#), but are not discussed below.

298 **4. Discussion**

299 *4.1. Precipitation collections: dilution or additive effect?*

300 Based on the published precipitation ^{10}Be dataset and the framework developed in Section 2.5, we
301 can investigate the delivery effect on ^{10}Be depositional flux over monthly and annual scales. For New
302 Zealand (2 years of observation) and Switzerland (7 years of observation), monthly-measured ^{10}Be data

303 are grouped based on sampling seasons and fitted separately (Fig. 5). The slopes are distinct between
304 seasons. The higher slopes in spring and summer are consistent with the seasonal cycles of stratosphere-
305 troposphere exchange, because more ^{10}Be -enriched stratospheric air is injected into the lower atmosphere
306 in spring and summer (Graham et al., 2003; Heikkilä et al., 2008a). Both in New Zealand and in
307 Switzerland the intercept and the slope show non-zero values, indicating a mixture of additive and
308 dilution components. To derive a representative f_{add} over a certain timescale (seasonal or annual scale), we
309 used the median precipitation rate of all measurements over this period for P in Eq. (4). For New Zealand
310 the resulting f_{add} shows a clear seasonal trend, with lower f_{add} in spring and summer (49-61%) than in
311 autumn and winter (74-87%). For Switzerland, the resulting f_{add} is also lower in spring and summer (59-
312 64%) than in autumn and winter (73-79%). The weaker additive effect in spring and summer can be
313 explained by more frequent occurrence of convective precipitation (dominated by the dilution effect)
314 during this period.

315 The dataset where ^{10}Be fluxes integrated from annual records are available can be used to detect the
316 delivery effect over a larger spatial range (Fig. 6). In mid-latitudes ($30\text{-}50^\circ$) a mixture of additive and
317 dilution components exists in islandic (New Zealand) and continental (USA and Switzerland) settings,
318 and the f_{add} are 59% and 62%, respectively. Note that the dilution effect (the slope of the fit line in Fig. 6)
319 cannot be caused solely by dry and recycled ^{10}Be components. The range of published recycled ^{10}Be
320 fluxes estimated in the two settings is $0.11\text{-}0.33 \times 10^6$ at/cm²/yr (Table S1 in Supplementary materials),
321 and the maximum ^{10}Be dry depositional flux derived from the preindustrial GCM model run is $\sim 0.19 \times 10^6$
322 at/cm²/yr (Heikkilä et al., 2013b). As such, the contribution of dry and recycled fluxes ($0.30\text{-}0.52 \times 10^6$
323 at/cm²/yr) is much lower than the slope of $0.94\text{-}1.00 \times 10^6$ at/cm²/yr in the two settings (Fig. 6). In Kikai
324 Island at $\sim 28^\circ$ the relatively low slope and comparatively high intercept indicate the dominance of an
325 additive effect, perhaps caused by the high precipitation rate within the observation period. For the dataset
326 at $10\text{-}30^\circ$ (mainly from India), the negative intercept can be explained with F_{prod} (Eq. (A3)) differing

327 between sampling stations, as this dataset shows a wide range in latitude and a positive correlation
328 ($R^2=0.60$) between latitude and ^{10}Be depositional flux.

329 Overall, our re-analysis of precipitation ^{10}Be data shows that a mixture of additive and dilution
330 components could be present in both continental and islandic settings. Additionally, the proportion of the
331 additive component seems to decrease with a longer integration timescale. Based on the ECHAM5-HAM
332 output, the ^{10}Be deposition variability is dominated by precipitation rate over the seasonal scale, but when
333 averaging out the seasonal cycle on multiannual scale, the ^{10}Be deposition mainly reflects ^{10}Be production
334 variability rather than precipitation rate (Heikkilä and Smith, 2013), suggesting a larger dilution
335 component and lower f_{add} . Our analysis is broadly consistent with these modeling results, showing that
336 annual-scale f_{add} (59% and 62%) in both New Zealand and Switzerland is at the lower limit of the
337 seasonal f_{add} range (49-87% and 59-79%, respectively). The potential reason is that the A_s/A_p ratio over
338 multiannual scale may be more stable due to averaging of seasonal variability in vapor sources.
339 Nevertheless, longer time-series precipitation ^{10}Be records (e.g. decadal-scale) are needed to confirm this
340 temporal variation in f_{add} . In any case, this analysis highlights that the dilution effect does not necessarily
341 dominate global ^{10}Be precipitation records (Willenbring and von Blanckenburg, 2010b). Also, the
342 empirical equation derived from data fitting between F_{met} and P (Graly et al., 2011) may not result in
343 representative fluxes, because such treatment implicitly assumes a sole additive effect which is mostly not
344 the case.

345 *4.2. Comparing precipitation- with GCM- derived ^{10}Be fluxes*

346 We next compare ^{10}Be fluxes determined from precipitation, soil, and river sediment with those
347 derived from GCM modeling (Fig. 7). We do so not because we believe the GCM-derived dataset serves
348 as a benchmark that all other estimates should match, but rather, because of its global spatial coverage
349 making it the only dataset to allow comparison with all other datasets.

350 The precipitation-derived fluxes are plotted against GCM-derived fluxes as shown in Fig. 7a. In
351 general, no clear correlation between precipitation- and GCM- derived fluxes is apparent, and the range in
352 precipitation records ($0.42\text{-}5.30\times 10^6$ at/cm²/yr) is much larger than that in GCM-derived fluxes (0.66-
353 2.20×10^6 at/cm²/yr). Almost half of precipitation records are >2 times greater than the corresponding
354 GCM-derived fluxes; 43% of precipitation-derived data are 1-2 times of GCM-derived fluxes, and only 8%
355 of precipitation-derived data is lower than the GCM-derived fluxes (Fig. 7a). This general pattern of
356 significantly higher precipitation-derived fluxes ($p < 0.001$ from a two-sided Wilcoxon signed-rank test)
357 is independent of location (Fig. 7a). The large variability of precipitation-derived fluxes is consistent with
358 their sensitivity to short-term ($10^0\text{-}10^1$ yr) local variations in atmospheric dynamics, such as vapor sources
359 and stratosphere-troposphere exchange. Hence, observations of only 1-2 years may not capture the
360 millennial-scale ¹⁰Be depositional signal required for geomorphic applications.

361 4.3. Comparing soil- with GCM- derived ¹⁰Be fluxes

362 Over millennial timescale, the relation between soil-derived and GCM-derived fluxes for each
363 latitudinal zone is shown in Fig. 7b. Most of flux data in the “excluded” series yield lower values than
364 other soil data or corresponding GCM-derived fluxes. The potential deficits in ¹⁰Be fluxes can be
365 explained by incomplete sampling of the depth profile or ¹⁰Be loss through surface erosion. For each
366 reliable data series, no correlation between fluxes derived from both methods can be observed, and the
367 range of soil-derived fluxes ($0.11\text{-}4.07\times 10^6$ at/cm²/yr) is larger than that of GCM-derived fluxes (1.08-
368 2.55×10^6 at/cm²/yr). Specifically, GCM-derived data averages over a large grid size ($\sim 10^2$ km) and may
369 not fully capture the distinct local variability of ¹⁰Be delivery in Hawaii where precipitation rates increase
370 by one order of magnitude within a short distance (~ 20 km) (Dixon et al., 2018). The ratio of soil-derived
371 flux to GCM-derived flux is generally lower than 1 ($p < 0.05$) and varies within a narrow range between
372 0.48 and 0.90 for average value of each data series. Also, it is worth noting that more than half (57%) of
373 the soil-derived fluxes agree with GCM-derived fluxes within a factor of 2.

374 4.4. Comparing riverine with GCM-derived ¹⁰Be fluxes

375 Over the river basin scale, riverine export fluxes derived from ^{10}Be (meteoric)/ ^9Be ratios are
376 compared with GCM-derived meteoric ^{10}Be fluxes (Fig. 7c). The riverine output of meteoric ^{10}Be agrees
377 with atmospheric input within a factor of ~ 2 for each data series, ranging from 0.52 (on average) in the
378 Amazon basin to 1.78 (on average) in the Ganga basin (Fig. 7c). Reasons for a weaker agreement for
379 some samples comprise, for example: 1) coarse-grained quartz used for *in situ* ^{10}Be measurement (to
380 determine D in Eq. (3)) may be derived from a different sediment source compared to fine-grained
381 sediment used for ^{10}Be (meteoric)/ ^9Be measurement, or multiple grain size fractions analyzed for ^{10}Be
382 (meteoric)/ ^9Be differ in source even at the same location (Wittmann et al., 2015); 2) In basins where no
383 $[^9\text{Be}]_{\text{parent}}$ was measured directly and the drainage area is not large enough to average out lithological
384 heterogeneity (e.g. small sub-basins of $<100 \text{ km}^2$ in the Potomac River, USA), the resulting riverine
385 fluxes could be biased by variations in $[^9\text{Be}]_{\text{parent}}$. Nevertheless, although the riverine ^{10}Be fluxes are
386 generally lower than GCM-derived values ($p < 0.001$), the majority (71%) of the two estimates indeed fall
387 within a close range (between 2:1 and 1:2 lines) (Fig. 7c).

388 4.5. Evaluation of ^{10}Be depositional fluxes from multiple approaches

389 To provide a more in-depth comparison between different approaches, we selected three regions
390 where both precipitation- and soil- derived ^{10}Be depositional fluxes exist within a narrow latitudinal band
391 such that both can be compared between each other and with GCM-derived fluxes (Fig. 8). The riverine
392 data is not included as its spatial scale by far exceeds the more local scale over which precipitation or soil
393 ^{10}Be data integrate. The selected regions include New Zealand (36-46°S), the interior USA (40-45°N) and
394 the European Alps (46-47°N), covering continental and islandic settings. All flux data were normalized to
395 an average Holocene ^{10}Be production rate.

396 Over the annual scale, precipitation-derived ^{10}Be fluxes from the three regions are in close agreement
397 and provide a mean flux of $3.01 \times 10^6 \text{ at/cm}^2/\text{yr}$. The relative standard deviation of ^{10}Be fluxes varies from
398 14-17% (New Zealand and European Alps) to 31% (interior USA). Over the millennial scale, the soil-
399 derived ^{10}Be fluxes are generally lower than the precipitation-derived ^{10}Be fluxes (Fig. 8), and amount to

400 1.28×10^6 at/cm²/yr on average. On a similar (Holocene) timescale, GCM-derived ¹⁰Be fluxes, with a
401 mean value of 1.76×10^6 at/cm²/yr in the three regions, fall between precipitation- and soil- derived fluxes.
402 Overall, the ratio of average precipitation-derived flux to average GCM-derived flux varies from 1.4 to
403 1.9. The ratio of average soil-derived flux to average GCM-derived flux falls within a range of 0.5-1.1.

404 Bearing in mind that all data were normalized to ¹⁰Be production rates with identical solar
405 modulation and magnetic field strength, the fact that ¹⁰Be depositional fluxes from precipitation
406 collections are highest might be caused by a short-term, transient enhancement of ¹⁰Be delivery that is
407 averaged-out over longer timescales (e.g. millennia). Such enhancement may potentially be caused by a
408 transient increase in atmospheric aerosol loading (e.g. from Be scavenging) today due to anthropogenic
409 emission (Dentener et al., 2006). However, this scenario contradicts the findings from GCM modeling,
410 where the difference in GCM-derived fluxes between industrial and early Holocene runs at each site are
411 relatively small (e.g. <~40%), despite of much higher (~5 fold) aerosol loading used in the industrial run
412 (Heikkilä and Smith, 2013). The reason for the low forcing of aerosol loading on ¹⁰Be delivery is that the
413 atmospheric sulfate burden exceeds that of ¹⁰Be by several orders of magnitude under both industrial and
414 pre-industrial conditions (Field et al., 2006; Heikkilä et al., 2013a), and hence ¹⁰Be delivery is production-
415 rather than sulfate-scavenging limited. We can therefore discount this explanation.

416 Another possibility is that the modern enhancement of ¹⁰Be delivery is a bias introduced through a
417 dependence of precipitation on observational interval. In fact, it was suggested that the precipitation rate
418 depends on measurement interval due to hiatuses (i.e. periods without precipitation) following a power
419 law distribution (Wilkinson, 2015). Over a measurement interval of several years or less, precipitation
420 rate and measurement interval scale by a negative power law, that is, higher precipitation rates are
421 apparent over shorter measurement intervals. This power law behavior emerges from the stochastic nature
422 of synoptic weather systems (Wilkinson, 2015). In contrast, over longer timescales (e.g. decades or
423 longer), the precipitation rate gradually settles to a constant value that reflects the long-term regional
424 climate, and is independent of measurement interval (Wilkinson, 2015). In other words, this measurement

425 interval bias is not present over a larger spatial or longer temporal span (Sadler and Jerolmack, 2015),
426 such as in GCM and soil profiles, where precipitation rates are deterministic rather than stochastic.

427 We explore how such measurement interval bias on precipitation rates affects ^{10}Be delivery. This
428 bias will only affect the additive component of the ^{10}Be flux because the dilution component does not
429 depend on precipitation rate. Because of the higher precipitation rate over shorter intervals (10^0 - 10^1 yr)
430 (Wilkinson, 2015), and if this results from long-distance transport of vapor, the additive ^{10}Be component
431 f_{add} (Eq. (4)) will increase accordingly and a higher delivery flux of ^{10}Be will result. Conversely, the
432 measurement interval bias also offers a hypothesis for the lower soil-derived ^{10}Be fluxes compared to
433 precipitation-derived fluxes: the additive component is suggested to decrease with increasing integration
434 time-interval (Section 4.1) and thus should account for a lower proportion in soil-derived ^{10}Be fluxes
435 integrating over 10^4 - 10^5 yrs. When corrected for temporal changes of ^{10}Be production, the dilution
436 component of the ^{10}Be flux is expected to be time-invariant, and will correspond to the dilution flux
437 contained in precipitation collections. In practice, we can use this assumption to estimate f_{add} over the
438 timescale integrated by soil profiles. The additive component of the ^{10}Be depositional flux determined on
439 soils is calculated as the difference between total soil-derived flux and the dilution component of
440 precipitation-derived flux (i.e. slope values in Fig. 6). The resulting proportions of the additive component
441 (f_{add}) for soil profiles in New Zealand (42%) and USA (31%) are indeed much lower than those derived
442 from precipitation records in Fig. 6 (59-62%). In the European Alps the total soil-derived flux (0.76×10^6
443 $\text{at}/\text{cm}^2/\text{yr}$) is even lower than the precipitation-derived dilution component (1.00×10^6 $\text{at}/\text{cm}^2/\text{yr}$, Fig. 6).
444 This deficit might hint at a very small additive effect combined with potential loss of soil ^{10}Be inventory
445 because of the low soil pH (3-5).

446 The conclusion of these considerations is that, as most geomorphologic processes integrate over
447 millennial scale, ^{10}Be fluxes from soil profiles rather than precipitation collections provide the most
448 relevant estimate, given their potential to record a representative additive effect over millennia.

449 Finally, we explore why GCM-derived fluxes commonly exceed soil-derived fluxes (Fig. 8). One
450 explanation is that soil-derived ^{10}Be depositional flux provides an underestimate, because the assumption
451 of full ^{10}Be retention may be invalid (Graly et al., 2011; Maher and von Blanckenburg, 2016). Another
452 explanation is that GCMs provide overestimated fluxes because the production function (Masarik and
453 Beer, 2009) adopted by the GCM (Heikkilä and Smith, 2013) may overestimate (e.g. ~15% higher under
454 a ϕ of 550 MeV) the global production rate of ^{14}C (Kovaltsov et al., 2012), and presumably also ^{10}Be .
455 Since the design of these models, production rates of cosmogenic nuclides by heavier nuclei (atom
456 number ≥ 2) have been revised downwards (Kovaltsov et al., 2012). To conclude, we suggest that the
457 “true” ^{10}Be depositional flux representative for the millennial scale lies between soil-derived and GCM-
458 derived fluxes reported here.

459 5. Conclusions

460 Our re-analysis of ^{10}Be flux estimates from four independent methods provides important conceptual
461 insights for the use of meteoric ^{10}Be in Earth surface studies. These are:

462 1) Contemporary ^{10}Be flux records (precipitation collections) are subject to both a dilution effect and
463 an additive effect. The weight of either effect depends on the area ratio of vapor condensation to
464 precipitation and the prevailing condensation rate. Based on existing precipitation records integrating over
465 the seasonal scale, we estimate the proportion of the additive component to be higher in autumn and
466 winter (up to 87%); over the annual scale, an additive component of ~60% is found in islandic and
467 continental settings. Because only the additive effect depends on precipitation rate, the empirical fitting
468 between ^{10}Be fluxes and precipitation rates assuming a sole additive effect likely overestimates the
469 control of precipitation on ^{10}Be delivery.

470 2) ^{10}Be flux estimates that average over 10^3 to 10^5 years, like those from soil profiles and river fluxes,
471 agree with GCM-derived fluxes within a factor of 2, but can be less than half of the modern ^{10}Be fluxes
472 from precipitation collections. A possible negative power-law relationship between precipitation rates and

473 measurement interval over short timescales (10^0 - 10^1 yr), combined with a ^{10}Be additive effect, can explain
474 higher ^{10}Be fluxes in short-term precipitation collections. Over a longer timescale (10^4 - 10^5 yrs) the
475 additive component can decrease to 30-40%. Because most geomorphologic processes integrate over
476 millennial timescales, ^{10}Be fluxes from soil profiles, in contrast to precipitation collections, likely provide
477 the most relevant estimate, given their representative estimation of the additive effect over millennia.

478 3) GCM-derived ^{10}Be fluxes commonly exceed soil-derived fluxes. The soil-derived fluxes might
479 represent an underestimate due to potential partial retention of ^{10}Be . GCM on the other hand might
480 overestimate the flux by using outdated spectra of the heavier nuclei component of cosmic rays. We
481 conclude that the best estimate for the millennial-scale ^{10}Be depositional flux lies between soil- and
482 GCM- derived fluxes.

483 For producing new ^{10}Be flux data, we recommend that more efforts should be spent on soil profiles
484 as published data are still rare, and many of these do not meet the stringent requirements for flux
485 calculation. The chosen soil profiles should meet the following criteria: minimum surface erosion, high
486 soil pH, low ^{10}Be inheritance, and sufficient sampling depth for quantification of the inheritance. If these
487 conditions can be met, ^{10}Be flux estimates from soil profiles are the method of choice.

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493 **Appendix 1 A mass balance framework for ^{10}Be delivery effects**

494 The total ^{10}Be deposition is comprised of wet deposition of $^{10}\text{Be}(\text{OH})_2$ scavenged by precipitation,
495 dry deposition of $^{10}\text{Be}(\text{OH})_2$ by friction-based aerosol deposition, and recycled ^{10}Be attached to dust.
496 When neglecting dry and recycled components, we can balance water vapor and ^{10}Be between vapor

497 “source” (condensation) and “sink” (precipitation). We conceptually describe the vapor accumulation to
 498 be the product of the water condensation rate C ($\text{m}^3/\text{m}^2/\text{yr}$) and the surface area of condensation A_c (m^2),
 499 which is balanced by precipitation as calculated by the product of the precipitation rate P ($\text{m}^3/\text{m}^2/\text{yr}$) and
 500 the area exposed to precipitation A_p (m^2):

$$501 \quad C \times A_c = P \times A_p \quad (\text{A1})$$

502 and thus the precipitation rate P is

$$503 \quad P = C \times (A_c/A_p) \quad (\text{A2})$$

504 The ^{10}Be production flux from the upper atmosphere is F_{prod} ($\text{at}/\text{m}^2/\text{yr}$), and the total flux of ^{10}Be
 505 accumulating in a given moist air mass is proportional to the area A_c where vapor also condenses. The
 506 ^{10}Be concentration of wet deposition ($[^{10}\text{Be}]_{\text{wet}}$, in at/m^3) in the condensation area is

$$507 \quad [^{10}\text{Be}]_{\text{wet}} = (F_{\text{prod}} \times A_c)/(C \times A_c) = F_{\text{prod}}/C \quad (\text{A3})$$

508 In the dilution effect, both vapor condensation and ^{10}Be accumulation take place over the same area
 509 as precipitation due to rapid water turnover (e.g. convective precipitation), and thus $A_c = A_p$. Hence, the
 510 meteoric delivery flux is:

$$511 \quad F_{\text{met}} = [^{10}\text{Be}]_{\text{wet}} \times P = F_{\text{prod}} \quad (\text{A4})$$

512 In the additive effect, in contrast, moist air mass accumulating both vapor and ^{10}Be is advected over
 513 extended distances (e.g. cyclonic precipitation), and $A_c > A_p$. The meteoric delivery flux is:

$$514 \quad F_{\text{met}} = [^{10}\text{Be}]_{\text{wet}} \times P = F_{\text{prod}} \times (A_c/A_p) \quad (\text{A5})$$

515 Thus, the difference in F_{met} between additive and dilution effects is whether the A_c/A_p ratio (related
 516 to the travel distance of moist air mass) controls the precipitation rate (Eq. (A2)) and ^{10}Be accumulation in
 517 moist air mass. Eqs. (A4-A5) form the base for a practical evaluation of the ^{10}Be delivery effect from
 518 combined measurements of ^{10}Be concentration in precipitation and the precipitation rate.

519

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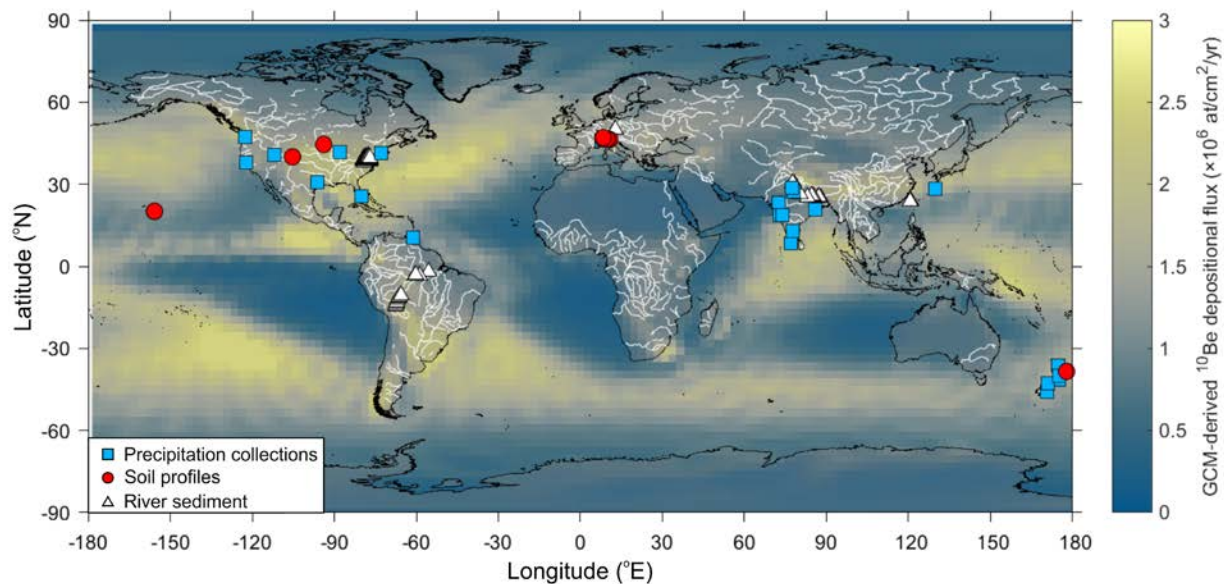
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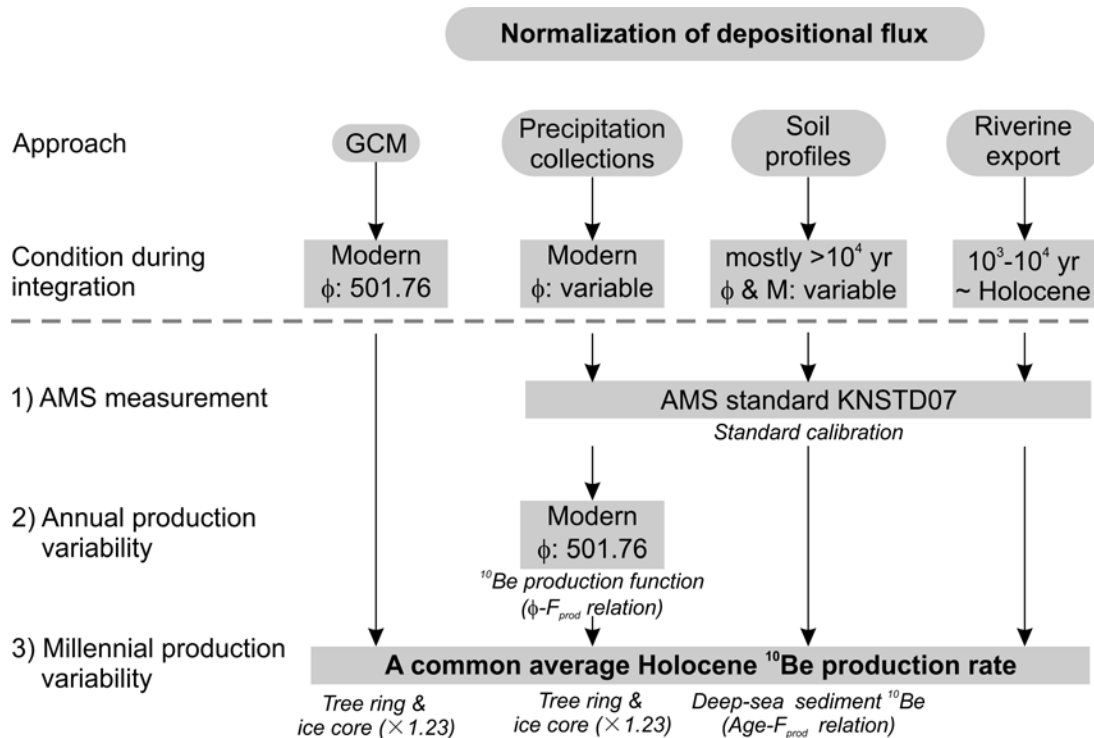
655

656 **Figures**



657
658 Fig. 1 The spatial distribution of flux data measured in precipitation collections (N=46), soil profiles
659 (N=23) and river sediment (N=90). Data sources are described in detail in Section 2 and the flux data are
660 provided in [Supplementary dataset](#). The distribution map shows the GCM-derived ^{10}Be depositional flux
661 from the ECHAM5-HAM model with the average ^{10}Be production and climatic conditions of the
662 Holocene ([Heikkilä and von Blanckenburg, 2015](#)).

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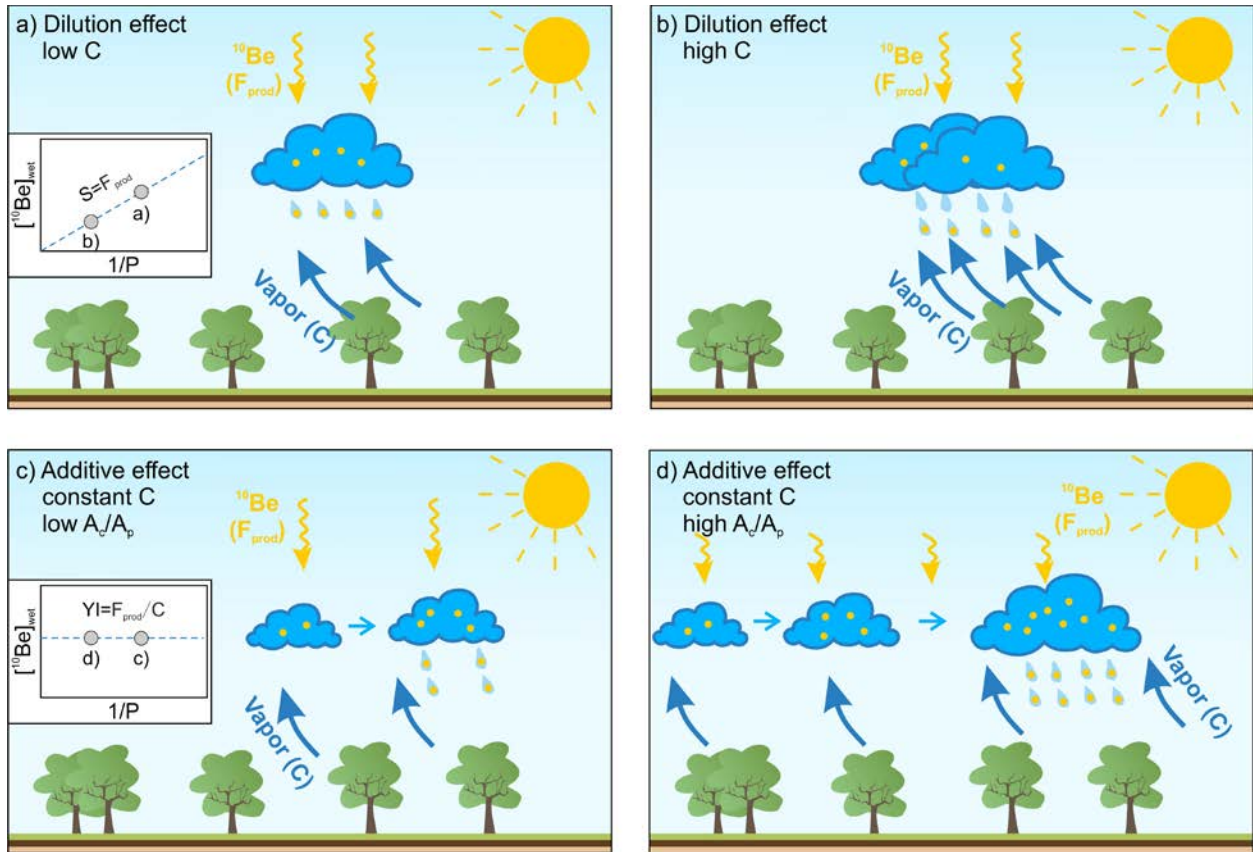
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665 Fig. 2 Procedures for normalization of ^{10}Be depositional fluxes that account for analytical differences (1)
 666 and temporal variations in the flux of galactic cosmic rays (2-3). The aim of these procedures is to
 667 normalize all fluxes compiled from four approaches to a common Holocene ^{10}Be production rate from the
 668 upper atmosphere (F_{prod} , in $\text{at}/\text{m}^2/\text{yr}$). ϕ is the solar modulation factor (MeV) and M the geomagnetic field
 669 strength. 1) ^{10}Be concentrations are normalized to a common AMS standard (Nishiizumi et al. (2007). 2)
 670 Given that ^{10}Be is well-mixed in the stratosphere after production (Heikkilä et al., 2013a; Heikkilä et al.,
 671 2008b), we perform a global normalization using the modeled relationship between ϕ and global-averaged
 672 production rate (Masarik and Beer, 2009). For precipitation collections, each annual depositional flux is
 673 normalized using the average ϕ within the observation period to the modern ϕ of 501.76 MeV. 3)
 674 Depending on the integration timescale of each approach, fluxes are normalized to an average Holocene
 675 production rate using a variety of proxy records of ^{10}Be production that combine variations of both solar
 676 modulation and magnetic field strength (See Fig. S1 in Supplementary materials). The ratio of average
 677 Holocene production rate to the modern value at a ϕ of 501.76 MeV (~ 1.23) is derived from ^{10}Be
 678 production proxy records from tree rings and ice cores (Steinhilber et al., 2012). The global ^{10}Be

679 production curve reconstructed from deep-sea sediments over longer term (since 250 ka) is sourced from
680 [Christl et al. \(2010\)](#). The riverine flux estimates were left un-normalized for F_{prod} variations as their
681 integration time scale is unknown and depends on denudation rate and infiltration depth of meteoric ^{10}Be
682 ([Willenbring and von Blanckenburg, 2010b](#)). It is considered to be typically 10^3 - 10^4 yrs, which is within
683 the Holocene reference timescale.

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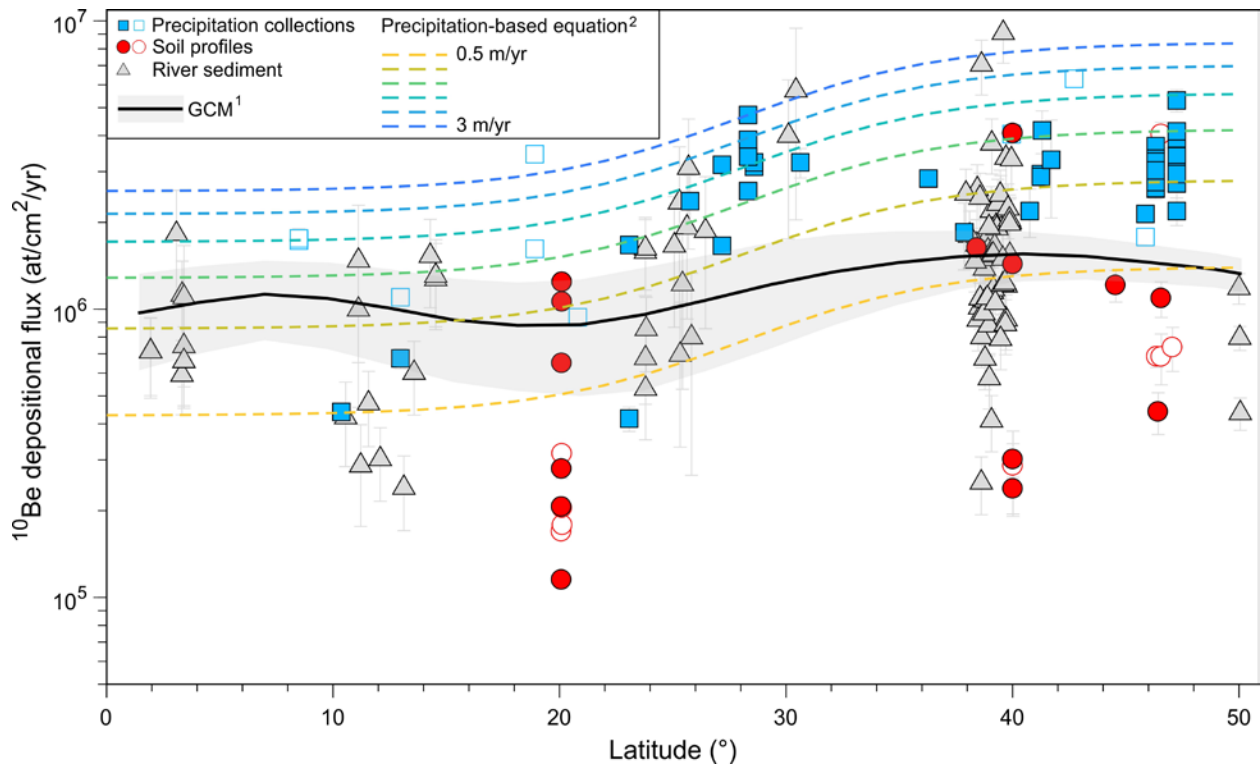
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687 Fig. 3 Conceptual framework of ^{10}Be delivery. a) and b): the dilution effect is in operation when the ratio
 688 of vapor source area (A_c) to area of precipitation event (A_p) is roughly unity, meaning that the vapor is of
 689 local origin and also no ^{10}Be is advected from further distances. In that case the precipitation rate (P , blue
 690 droplets) is equal to the condensation rate (C), and the ^{10}Be flux (yellow dots in droplets) is constant, but
 691 the ^{10}Be concentration depends on the amount of dilution by vapor. The inset in a) shows the relation of
 692 $[^{10}\text{Be}]_{\text{wet}}$ to $1/P$ for the dilution effect, and the slope (S , in $\text{at}/\text{m}^2/\text{yr}$) is equal to the atmospheric production
 693 rate of ^{10}Be (F_{prod}). c) and d): the additive effect is in operation when the condensation rate C is constant,
 694 but both the precipitation rate P and the amount of ^{10}Be accumulated scale with A_c/A_p . The inset in c)
 695 shows the relation between $[^{10}\text{Be}]_{\text{wet}}$ and $1/P$ for the additive effect (neglecting dry and recycled ^{10}Be
 696 components). Its y-intercept (YI , in at/m^3) is set by F_{prod}/C , i.e. the production flux of ^{10}Be and the water
 697 condensation rate. This schematic plot is modified from [Willenbring and von Blanckenburg \(2010b\)](#).

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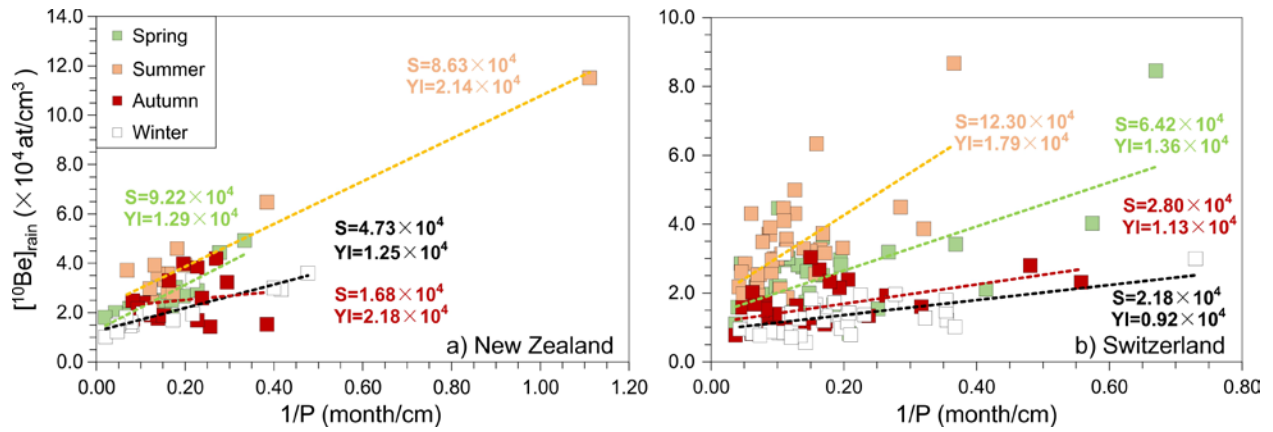


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700 Fig. 4 Latitudinal distribution of measured and modeled ^{10}Be depositional fluxes derived from the
 701 approaches compiled in this study. The latitude shown on the X-axis equally refers to the southern and
 702 northern hemispheres, respectively. Note that the latitudes of river sediments are those of sampling
 703 locations. 1) GCM results of the average Holocene ^{10}Be production and climatic conditions are sourced
 704 from [Heikkilä and von Blanckenburg \(2015\)](#), and the average flux (black line) and the standard deviation
 705 (gray shade) at each latitudinal zone are shown. 2) Results from precipitation-based fitting equation (Eq.
 706 (5)) are from [Graly et al. \(2011\)](#). For comparison between different approaches, all flux data were
 707 normalized to an average Holocene ^{10}Be production rate (see [Fig. 2](#)). For each dataset, only filled symbols
 708 are considered representative and thus used for the discussion of fluxes, whereas flux data shown by open
 709 symbols are considered potentially unrepresentative and only shown for completeness.

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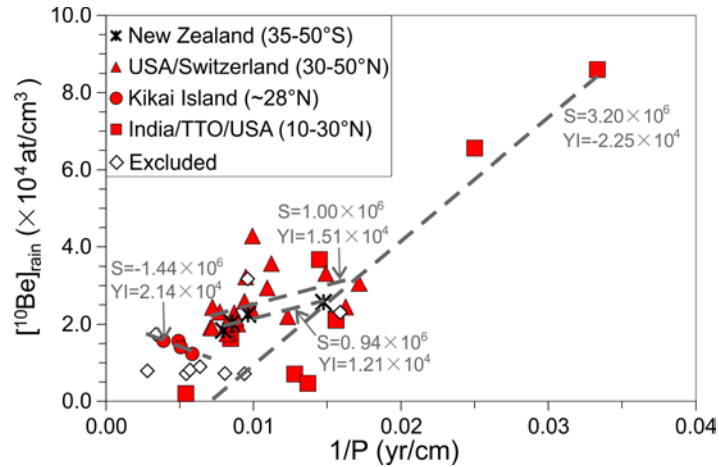
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713 Fig. 5 Plot of the inverse of monthly precipitation rate ($1/P$) and $[^{10}\text{Be}]_{\text{rain}}$ in (a) New Zealand and (b)

714 Switzerland based on monthly records. The slope (S) and y-intercept (YI) are shown for each fit.

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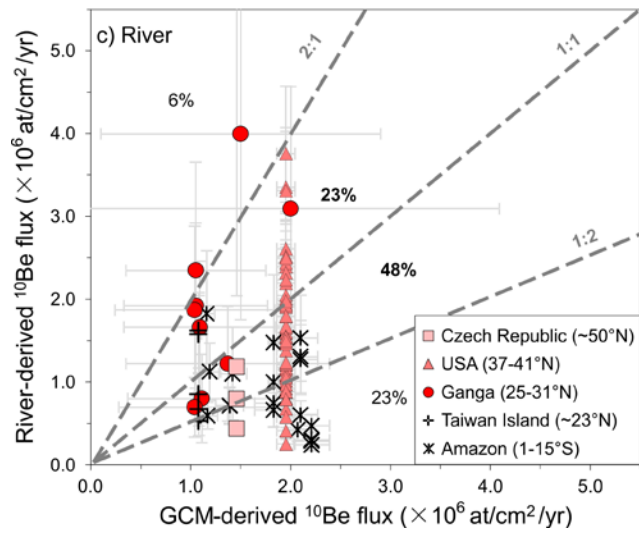
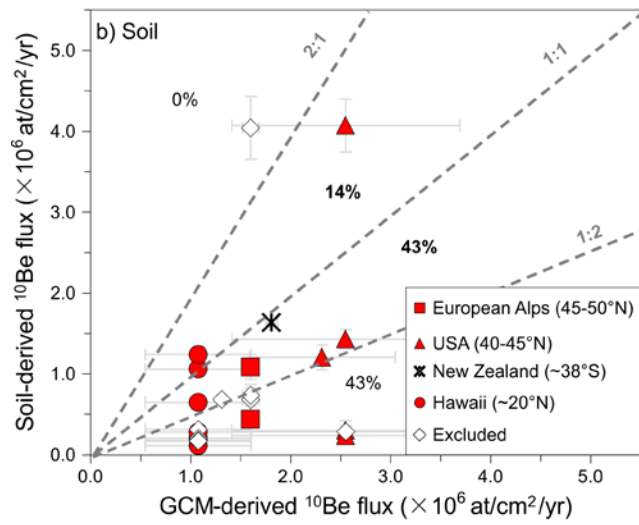
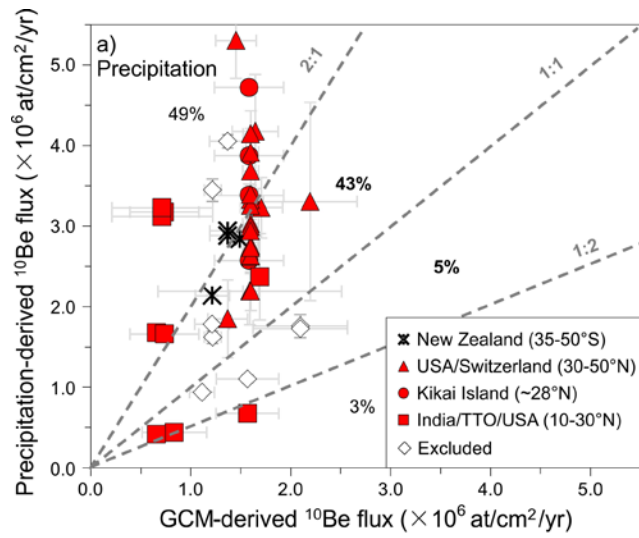
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718 Fig. 6 Plot of inverse of annual precipitation rate ($1/P$) and mean $[^{10}\text{Be}]_{\text{rain}}$ based on annual records (data
 719 in [Supplementary dataset](#)). The slope (S) and y-intercept (YI) are shown for each fit. Series
 720 “India/TTO/USA (10-30° N)” includes precipitation data from India (majority), Trinidad (shortened as
 721 TTO) and USA at 10-30° N. Note that data from northern India were excluded in a previous data
 722 compilation ([Graly et al., 2011](#)) as they were assumed to be affected by significant but unquantified dust
 723 ^{10}Be flux. In contrast, we still include such dataset here because recent modeling studies on global dust
 724 distribution ([Mahowald et al., 2005](#)) show that the dust flux in northern India is not outstandingly high (5
 725 to 20 $\text{g}/\text{m}^2/\text{yr}$). The records with a collection period of less than 9 months are considered unrepresentative
 726 and marked as “Excluded” (open symbols). The uncertainties of $[^{10}\text{Be}]_{\text{rain}}$ are generally lower than the
 727 symbol size.

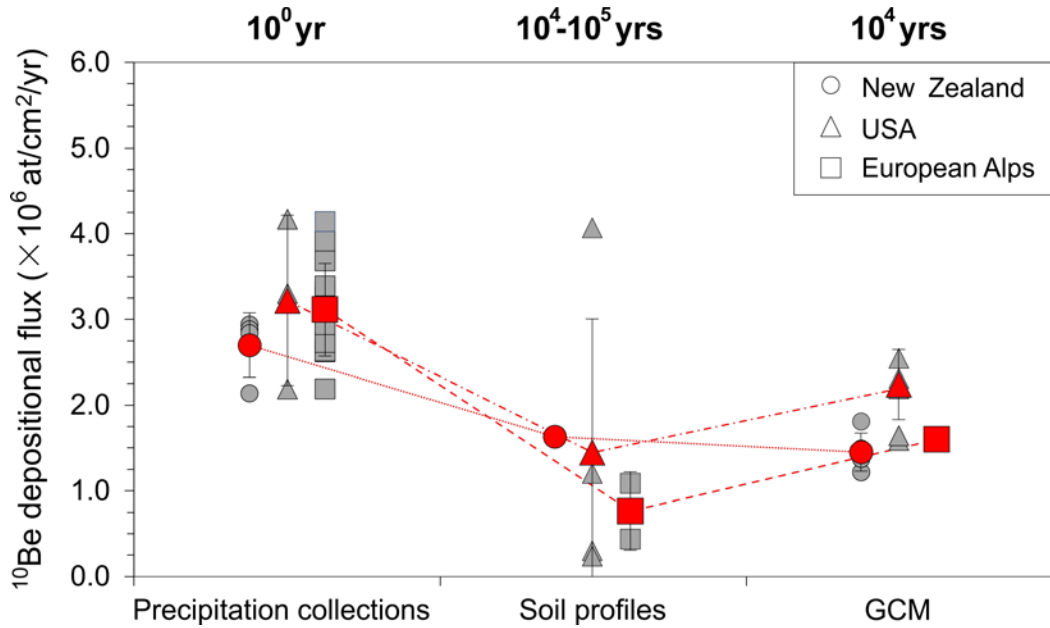
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730 Fig. 7 Comparison between GCM-derived ^{10}Be depositional fluxes and precipitation- (a) or soil- (b)
731 derived ^{10}Be depositional fluxes and between basin-averaged ^{10}Be depositional fluxes from GCM and
732 riverine export ^{10}Be fluxes (c). Note that error bars of ^{10}Be fluxes can be smaller than the symbol size for
733 some locations. The GCM-derived flux is based on the average Holocene model of ECHAM5-HAM
734 (Heikkilä and von Blanckenburg, 2015). The corresponding flux uncertainty is the difference between
735 modern and early Holocene model runs, or standard deviation among adjacent grid cell values (applicable
736 for some riverine data, see details in [Supplementary dataset](#)). Precipitation- and soil- derived ^{10}Be flux
737 data are normalized to an average Holocene production rate for comparison. The “Excluded” series in a)
738 is the same as [Fig. 6](#). “Excluded” series in b) includes profiles with significant erosional loss, incomplete
739 sampling or high ^{10}Be inheritance (see details in [Supplementary dataset](#)). Only riverine fluxes derived
740 from ^{10}Be (meteoric)/ ^9Be ratios are shown in (c). Three riverine flux data (range: $5.74\text{-}9.09 \times 10^6$
741 $\text{at}/\text{cm}^2/\text{yr}$) are out of the y-axis range and not shown in [Fig. 7c](#), but they are included in the statistics. Each
742 number in the figure represents the percentage of dataset falling in the corresponding domain (divided by
743 the dashed lines).

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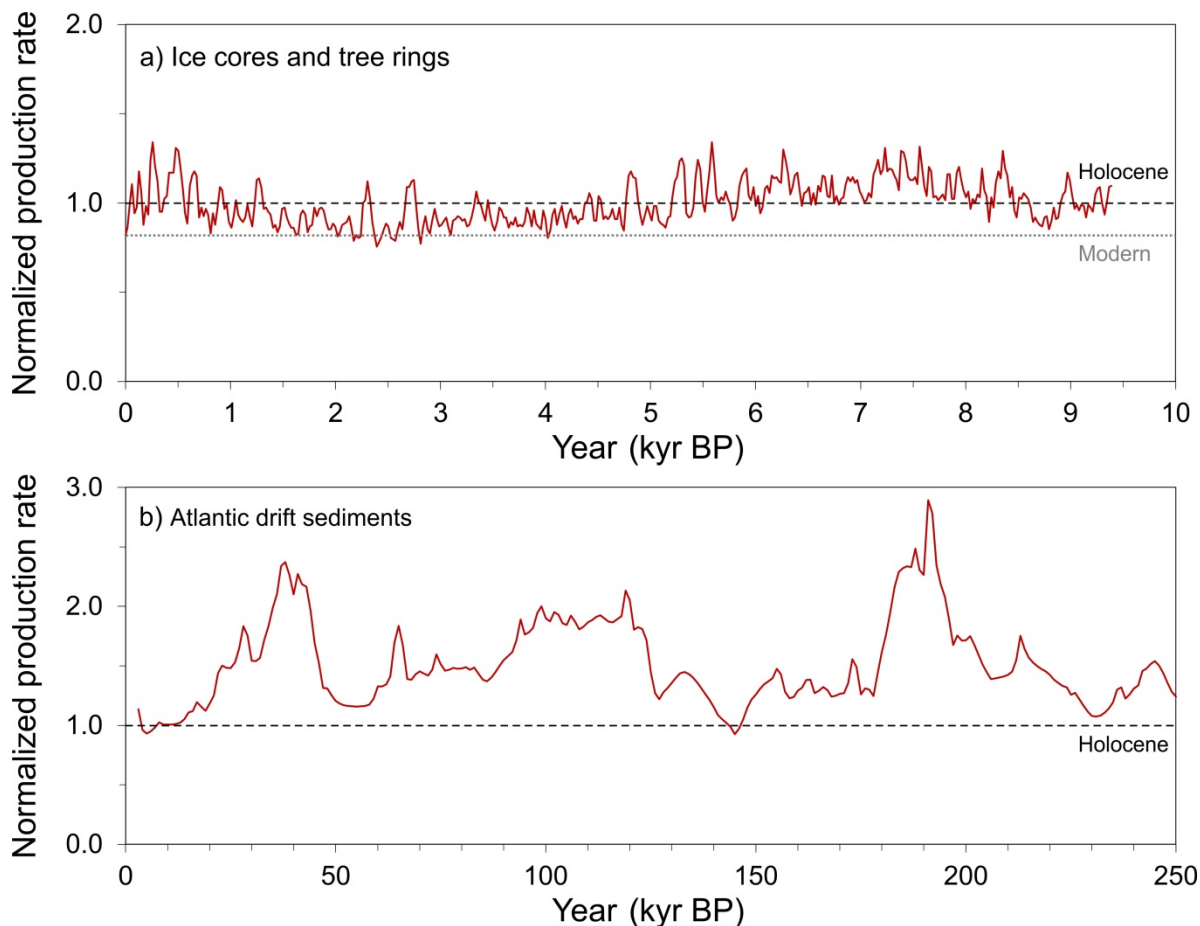
747 Fig. 8 Comparison of ¹⁰Be depositional fluxes derived from three approaches across timescales. The gray
 748 symbols are individual data, and the red symbols are average values derived thereof with an uncertainty
 749 of one standard deviation. “Precipitation collections” and “Soil profiles” are measurements over
 750 contemporary and millennial timescales, respectively. “GCM” is a modeled dataset derived from
 751 ECHAM5-HAM output over average Holocene conditions (Heikkilä and von Blanckenburg, 2015). The
 752 integration timescale of each method is provided at the top. All flux data are normalized to the average
 753 Holocene ¹⁰Be production rate. Dashed red lines connect mean values of every approach applied in the
 754 same region.

755

756

Supplementary materials for Deng et al. (2020, EPSL)

757 1. Reconstruction of paleo- ^{10}Be production rates



758

759 **Fig. S1** Temporal variations of the global-averaged ^{10}Be production rates reconstructed from a) ice cores
760 and tree rings in the Holocene (Steinhilber et al., 2012); b) highly accumulating Atlantic drift sediments
761 since 250 ka (Christl et al., 2010). Each curve is normalized to its average production rate in the Holocene
762 (see the black dashed line). Note that the average ^{10}Be production rate in the Holocene (a) is ~ 1.23 times
763 that of the modern value (gray dot line). All ^{10}Be fluxes in this study are normalized to the average
764 Holocene production rate by dividing by the average normalized production rates during the integration
765 timescale of the archive, upon which the normalization will be done using either curve a) (e.g. for
766 precipitation records) or b) (e.g. for soil profiles).

767

768 **2. Estimation of recycled ^{10}Be flux**

769 **Table S1** Recycled ^{10}Be flux estimate in New Zealand, Switzerland and USA. The range is $1.10\text{-}3.28 \times$
 770 10^5 at/cm²/yr.

Region	Station	Recycled ^{10}Be %	Total ^{10}Be depositional flux ³ at/cm ² /yr	Recycled ^{10}Be flux at/cm ² /yr	Data source
New Zealand	Gracefield	9.5% ¹	2.51E+06	2.38E+05	Graham et al. (2003)
New Zealand	Leigh	9.5% ¹	2.40E+06	2.28E+05	Graham et al. (2003)
New Zealand	Dunedin	16.1% ¹	1.81E+06	2.91E+05	Graham et al. (2003)
Switzerland	Dubendorf	7.5% ¹	2.31E+06	1.73E+05	Heikkilä et al. (2008)
Switzerland	Jungfrauoch	6.2% ¹	2.21E+06	1.37E+05	Heikkilä et al. (2008)
USA	Tacoma	8% ²	3.58E+06	1.20E+05	Monaghan et al. (1986)
USA	Berkeley	21% ²	1.25E+06	1.10E+05	Monaghan et al. (1986)
USA	Salt Lake City	29% ²	1.47E+06	1.80E+05	Monaghan et al. (1986)
USA	College Station	26% ²	2.17E+06	2.37E+05	Monaghan et al. (1986)
USA	Argonne	35% ²	2.23E+06	3.28E+05	Monaghan et al. (1986)
USA	Miami	26% ²	1.60E+06	1.75E+05	Monaghan et al. (1986)
USA	New Haven	20% ²	2.82E+06	2.37E+05	Monaghan et al. (1986)

771 Note: 1. The percentage of recycled ^{10}Be is calculated using $^{10}\text{Be}/^7\text{Be}$ ratios by [Graly et al. \(2011\)](#).

772 2. The percentage of recycled ^{10}Be is calculated by multiplying ^{238}U -derived dust flux by assumed dust
 773 ^{10}Be concentration. However, in the original calculation, [Monaghan et al. \(1986\)](#) assumed an extremely
 774 high ^{10}Be concentration in dust (5×10^8 at/g) compared to recent observations (2.1×10^8 at/g on average)
 775 ([Shen et al., 2010](#)). Hence, we recalculated the recycled $^{10}\text{Be}\%$ (shown here) using a dust ^{10}Be
 776 concentration of 2.1×10^8 at/g.

777 3. All the ^{10}Be data are normalized to the AMS standard KNSTD07 ([Nishiizumi et al., 2007](#)), but the
 778 correction for temporal variability in ^{10}Be production is not done here. We refer to [Table SD1](#) in
 779 [Supplementary Dataset](#) for production-normalized fluxes.

780

781

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803

Table SD1 Compiled ¹⁰Be depositional flux data from precipitation collections and the correspon

Station ¹	Country/region	Lat. ²	Lon.	Observation time span	
		°N	°E	Start time	End time
<i>New Zealand (35-50°S)</i>					
Dunedin (1997)	New Zealand	-45,867	170,500	01.10.1997	1998/9/30
Gracefield (1996)	New Zealand	-41,250	174,917	01.10.1996	30.09.1997
Gracefield (1997)	New Zealand	-41,250	174,917	01.10.1997	30.09.1998
Leigh	New Zealand	-36,283	174,800	01.10.1997	30.09.1998
<i>India, etc (10-30°N)</i>					
Trinidadian	Trinidad	10,361	-61,250	1987	1988
Bangalore(1980)	India	12,967	77,583	01.07.1980	15.06.1981
Ahmedabad (1980)	India	23,067	72,633	15.06.1980	04.07.1981
Ahmedabad (1979)	India	23,067	72,633	22.06.1979	14.06.1980
Miami	USA	25,757	-80,197	01.08.1980	01.08.1981
Agra (1980)	India	27,167	78,033	25.06.1980	03.07.1981
Agra (1979)	India	27,167	78,033	04.07.1979	22.06.1980
Delhi (1980)	India	28,583	77,200	20.06.1980	10.07.1981
Delhi (1979)	India	28,583	77,200	14.06.1979	19.06.1980
<i>Kikai Island (~28°N)</i>					
Site.2 (2000)	Kikai island	28,319	129,985	05.06.2000	24.06.2001
Site.2 (2001)	Kikai island	28,319	129,985	24.06.2001	29.06.2002
Site 1 (2000)	Kikai island	28,324	129,928	05.06.2000	24.06.2001
Site 1 (2001)	Kikai island	28,324	129,928	25.06.2001	29.06.2002
<i>USA/Switzerland (30-50°N)</i>					
College Station	USA	30,628	-96,335	17.10.1980	09.11.1981
Berkeley	USA	37,871	-122,273	10.12.1980	15.12.1981
Salt Lake City	USA	40,760	-111,892	04.12.1980	04.12.1981
New Haven	USA	41,304	-72,929	24.12.1980	07.01.1982
Argonne	USA	41,706	-87,982	15.09.1980	15.09.1981
Jungfrauoch (1998)	Switzerland	46,320	7,590	01.01.1998	31.12.1998
Jungfrauoch (1999)	Switzerland	46,320	7,590	01.01.1999	31.12.1999
Jungfrauoch (2000)	Switzerland	46,320	7,590	01.01.2000	31.12.2000
Jungfrauoch (2001)	Switzerland	46,320	7,590	01.01.2001	31.12.2001
Jungfrauoch (2002)	Switzerland	46,320	7,590	01.01.2002	31.12.2002
Jungfrauoch (2003)	Switzerland	46,320	7,590	01.01.2003	31.12.2003
Jungfrauoch (2004)	Switzerland	46,320	7,590	01.01.2004	31.12.2004
Tacoma	USA	47,248	-122,444	09.08.1980	03.08.1981
Dubendorf (1998)	Switzerland	47,250	8,270	01.01.1998	31.12.1998
Dubendorf (1999)	Switzerland	47,250	8,270	01.01.1999	31.12.1999
Dubendorf (2000)	Switzerland	47,250	8,270	01.01.2000	31.12.2000
Dubendorf (2001)	Switzerland	47,250	8,270	01.01.2001	31.12.2001
Dubendorf (2002)	Switzerland	47,250	8,270	01.01.2002	31.12.2002
Dubendorf (2003)	Switzerland	47,250	8,270	01.01.2003	31.12.2003
Dubendorf (2004)	Switzerland	47,250	8,270	01.01.2004	31.12.2004
<i>Below are stations with observation period <9 months</i>					
Dunedin(1997)	New Zealand	-45,867	170,500	01.03.1997	30.09.1997
Hokitika	New Zealand	-42,717	170,967	01.08.1998	30.11.1998
Kaitoke	New Zealand	-39,959	175,089	01.01.1997	31.05.1997
Trivandrium(1980)	India	8,483	76,950	01.04.1980	15.08.1980

Trivandrium(1979)	India	8,483	76,950	09.07.1979	14.11.1979
Bangalore(1979)	India	12,967	77,583	25.06.1979	19.10.1979
Bombay	India	18,900	72,817	20.06.1979	21.11.1979
Khandala	India	18,900	73,917	23.06.1979	11.08.1979
Cuttack	India	20,800	85,933	21.06.1979	04.11.1979

- Note: 1. Stations are sorted and grouped based on latitude. Observations of multiple years at the s
2. When coordinates are not provided in the original study (e.g. USA & Trinidad), we estimate L;
3. Observation period is estimated based on the starting time and end time of sample collection. S
4. The ^{10}Be concentrations used for flux calculation have been normalized to the same AMS stan
5. The average solar modulation factor (ϕ) during sample collection period is sourced from www.
6. All flux data are normalized to a modern ϕ of 501.76 MeV, using the modelled relationship bet
7. ϕ -normalized $[\text{}^{10}\text{Be}]_{\text{rain}}$ is used in Fig. 6.
8. The ϕ -normalized flux was further normalized to the average Holocene ^{10}Be production rate, i.
9. The fluxes are calculated based on Eq. (5) in the main text and then normalized to the average
10. GCM-derived fluxes for the average Holocene climate conditions are from Heikkilä and von]

ding modelled flux

Observation period ³	Precipitation rate	Depositional flux ⁴	1 sigma	AMS standard-normalized [¹⁰ Be] _{rain}	1 sigma
day	cm/yr	at/cm ² /yr	at/cm ² /yr	at/cm ³	at/cm ³
365	68	1,81E+06	9,13E+03	2,67E+04	1,35E+02
365	117	2,57E+06	9,13E+03	2,20E+04	7,79E+01
365	104	2,44E+06	9,13E+03	2,35E+04	8,80E+01
365	126	2,40E+06	1,83E+04	1,91E+04	1,45E+02
730	184	3,60E+05	0,00E+00	1,95E+03	0,00E+00
350	78	4,56E+05	3,65E+04	5,85E+03	4,68E+02
385	73	2,83E+05	2,74E+04	3,87E+03	3,75E+02
359	64	1,22E+06	6,39E+04	1,91E+04	9,98E+02
366	118	1,60E+06	3,62E+05	1,35E+04	3,06E+03
374	64	1,12E+06	2,74E+04	1,75E+04	4,28E+02
355	30	2,32E+06	9,12E+04	7,72E+04	3,04E+03
386	69	2,11E+06	6,39E+04	3,05E+04	9,26E+02
372	40	2,35E+06	7,30E+04	5,88E+04	1,82E+03
385	253	3,20E+06	6,94E+04	1,26E+04	2,74E+02
371	202	2,70E+06	5,46E+04	1,34E+04	2,71E+02
385	195	2,29E+06	9,49E+04	1,17E+04	4,86E+02
370	170	1,79E+06	3,98E+04	1,06E+04	2,34E+02
389	115	2,17E+06	2,53E+05	1,89E+04	2,21E+03
371	61	1,25E+06	3,26E+05	2,03E+04	5,29E+03
366	58	1,47E+06	2,80E+05	2,53E+04	4,81E+03
380	105	2,82E+06	4,79E+05	2,68E+04	4,55E+03
366	91	2,23E+06	8,32E+05	2,44E+04	9,10E+03
365	120	2,20E+06	1,73E+05	1,83E+04	1,44E+03
365	122	2,05E+06	1,73E+05	1,68E+04	1,42E+03
305	129	2,49E+06	1,73E+05	1,92E+04	1,33E+03
335	141	2,33E+06	1,73E+05	1,65E+04	1,22E+03
365	139	2,23E+06	1,73E+05	1,61E+04	1,24E+03
365	112	1,87E+06	1,73E+05	1,67E+04	1,54E+03
365	122	2,30E+06	1,73E+05	1,89E+04	1,42E+03
360	100	3,58E+06	3,16E+05	3,56E+04	3,15E+03
335	81	1,84E+06	2,01E+05	2,27E+04	2,49E+03
305	106	2,65E+06	2,01E+05	2,49E+04	1,89E+03
365	89	2,65E+06	2,01E+05	2,97E+04	2,26E+03
365	138	2,90E+06	2,01E+05	2,10E+04	1,46E+03
365	113	2,05E+06	2,01E+05	1,82E+04	1,78E+03
365	67	1,85E+06	2,01E+05	2,76E+04	3,01E+03
335	101	2,22E+06	2,01E+05	2,20E+04	1,99E+03
214	63	1,57E+06	9,13E+03	2,50E+04	1,45E+02
122	295	5,24E+06	2,74E+04	1,78E+04	9,28E+01
151	104	3,57E+06	7,30E+04	3,43E+04	7,02E+02
137	157	1,23E+06	2,74E+04	7,85E+03	1,74E+02

129	175	1,25E+06	1,00E+05	7,14E+03	5,74E+02
117	124	7,85E+05	3,65E+04	6,33E+03	2,94E+02
155	184	1,16E+06	5,47E+04	6,30E+03	2,98E+02
50	355	2,49E+06	1,00E+05	7,02E+03	2,83E+02
137	106	6,66E+05	2,74E+04	6,26E+03	2,57E+02

same station are separated and the starting time (year) is added behind the station name

at. & Lon. from Google Earth.

Samples with observation period < 9 months are not used for statistical analysis in this study.

Standard KNSTD07.

[.faa.gov/data_research/research/med_humanfacs/aeromedical/radiobiology/heliocentric/](https://www.faa.gov/data_research/research/med_humanfacs/aeromedical/radiobiology/heliocentric/)

between ϕ and global-averaged production rate (Masarik and Beer, 2009) as calculated for modern condition

e. multiplied by a factor of ~1.23 according to Steinhilber et al. (2012).

Holocene ^{10}Be production rates.

Blanckenburg (2015). The dataset is rescaled to an average Holocene production rate.

Average ϕ^5	ϕ -normalized flux ⁶	1 sigma	ϕ -normalized [¹⁰ Be] _{rain} ⁷	1 sigma
MeV	at/cm ² /yr	at/cm ² /yr	at/cm ³	at/cm ³
429	1,74E+06	8,77E+03	2,56E+04	1,30E+02
369	2,39E+06	8,46E+03	2,04E+04	7,22E+01
429	2,34E+06	8,77E+03	2,26E+04	8,45E+01
429	2,31E+06	1,75E+04	1,84E+04	1,40E+02
488	3,57E+05	\	1,94E+03	\
907	5,48E+05	4,38E+04	7,02E+03	5,62E+02
901	3,39E+05	3,28E+04	4,64E+03	4,49E+02
719	1,36E+06	7,10E+04	2,12E+04	1,11E+03
911	1,92E+06	4,35E+05	1,63E+04	3,68E+03
907	1,35E+06	3,29E+04	2,11E+04	5,14E+02
719	2,58E+06	1,01E+05	8,59E+04	3,38E+03
907	2,53E+06	7,67E+04	3,67E+04	1,11E+03
721	2,62E+06	8,13E+04	6,55E+04	2,03E+03
904	3,83E+06	8,32E+04	1,51E+04	3,29E+02
830	3,14E+06	6,36E+04	1,56E+04	3,16E+02
904	2,75E+06	1,14E+05	1,41E+04	5,82E+02
830	2,09E+06	4,64E+04	1,23E+04	2,73E+02
924	2,62E+06	3,06E+05	2,28E+04	2,67E+03
913	1,50E+06	3,92E+05	2,44E+04	6,37E+03
921	1,78E+06	3,38E+05	3,05E+04	5,81E+03
909	3,39E+06	5,76E+05	3,22E+04	5,46E+03
909	2,68E+06	1,00E+06	2,93E+04	1,09E+04
446	2,13E+06	1,67E+05	1,78E+04	1,40E+03
578	2,14E+06	1,80E+05	1,75E+04	1,47E+03
909	2,99E+06	2,07E+05	2,31E+04	1,60E+03
819	2,70E+06	2,00E+05	1,91E+04	1,42E+03
868	2,64E+06	2,04E+05	1,90E+04	1,47E+03
903	2,24E+06	2,07E+05	2,00E+04	1,85E+03
637	2,46E+06	1,85E+05	2,02E+04	1,52E+03
911	4,31E+06	3,81E+05	4,29E+04	3,79E+03
446	1,78E+06	1,95E+05	2,20E+04	2,41E+03
578	2,75E+06	2,10E+05	2,59E+04	1,97E+03
909	3,18E+06	2,42E+05	3,57E+04	2,71E+03
819	3,36E+06	2,34E+05	2,43E+04	1,69E+03
868	2,43E+06	2,38E+05	2,15E+04	2,11E+03
903	2,22E+06	2,41E+05	3,31E+04	3,60E+03
642	2,39E+06	2,16E+05	2,36E+04	2,14E+03
360	1,45E+06	8,42E+03	2,30E+04	1,34E+02
459	5,12E+06	2,68E+04	1,74E+04	9,07E+01
360	3,29E+06	6,73E+04	3,16E+04	6,47E+02
774	1,40E+06	3,12E+04	8,94E+03	1,99E+02

781	1,43E+06	1,15E+05	8,16E+03	6,55E+02
781	8,96E+05	4,17E+04	7,23E+03	3,36E+02
768	1,32E+06	6,22E+04	7,15E+03	3,38E+02
742	2,80E+06	1,13E+05	7,88E+03	3,18E+02
781	7,61E+05	3,13E+04	7,15E+03	2,94E+02

is (Heikkilä et al., 2013b).

Production-normalized flux⁸	1 sigma	Precipitation data fitting-based ¹⁰Be flux⁹	GCM-derived flux (average Holocene)¹⁰	1 sigma
at/cm ² /yr	at/cm ² /yr	at/cm ² /yr	at/cm ² /yr	at/cm ² /yr
2,14E+06	1,08E+04	1,86E+06	1,22E+06	1,70E+05
2,94E+06	1,04E+04	3,11E+06	1,37E+06	1,80E+05
2,88E+06	1,08E+04	2,75E+06	1,37E+06	1,80E+05
2,84E+06	2,16E+04	3,00E+06	1,49E+06	1,80E+05
4,40E+05 \		1,61E+06	8,35E+05	3,24E+05
6,75E+05	5,40E+04	6,93E+05	1,57E+06	3,14E+05
4,17E+05	4,04E+04	8,36E+05	6,55E+05	3,22E+04
1,67E+06	8,75E+04	7,33E+05	6,55E+05	3,22E+04
2,37E+06	5,36E+05	1,57E+06	1,69E+06	3,14E+03
1,66E+06	4,05E+04	9,33E+05	7,36E+05	3,44E+05
3,17E+06	1,25E+05	4,37E+05	7,36E+05	3,44E+05
3,12E+06	9,44E+04	1,10E+06	7,14E+05	5,04E+05
3,23E+06	1,00E+05	6,40E+05	7,14E+05	5,04E+05
4,72E+06	1,02E+05	3,98E+06	1,58E+06	3,45E+05
3,87E+06	7,84E+04	3,17E+06	1,58E+06	3,45E+05
3,38E+06	1,40E+05	3,07E+06	1,58E+06	3,45E+05
2,57E+06	5,71E+04	2,67E+06	1,58E+06	3,45E+05
3,23E+06	3,77E+05	2,09E+06	1,71E+06	2,12E+05
1,85E+06	4,82E+05	1,53E+06	1,37E+06	1,99E+05
2,19E+06	4,17E+05	1,53E+06	1,59E+06	9,20E+05
4,17E+06	7,09E+05	2,80E+06	1,65E+06	2,30E+05
3,30E+06	1,23E+06	2,44E+06	2,20E+06	4,70E+05
2,62E+06	2,06E+05	3,31E+06	1,60E+06	8,00E+04
2,63E+06	2,21E+05	3,36E+06	1,60E+06	8,00E+04
3,68E+06	2,55E+05	3,57E+06	1,60E+06	8,00E+04
3,33E+06	2,47E+05	3,89E+06	1,60E+06	8,00E+04
3,25E+06	2,51E+05	3,83E+06	1,60E+06	8,00E+04
2,76E+06	2,55E+05	3,09E+06	1,60E+06	8,00E+04
3,03E+06	2,28E+05	3,36E+06	1,60E+06	8,00E+04
5,30E+06	4,69E+05	2,78E+06	1,45E+06	2,03E+05
2,19E+06	2,40E+05	2,24E+06	1,60E+06	7,88E+04
3,39E+06	2,58E+05	2,95E+06	1,60E+06	7,88E+04
3,91E+06	2,98E+05	2,47E+06	1,60E+06	7,88E+04
4,14E+06	2,88E+05	3,83E+06	1,60E+06	7,88E+04
2,99E+06	2,93E+05	3,13E+06	1,60E+06	7,88E+04
2,73E+06	2,97E+05	1,86E+06	1,60E+06	7,88E+04
2,94E+06	2,66E+05	2,80E+06	1,60E+06	7,88E+04
1,78E+06	1,04E+04	1,73E+06	1,21E+06	1,75E+05
6,30E+06	3,30E+04	7,95E+06	1,26E+06	2,25E+05
4,05E+06	8,29E+04	2,71E+06	1,37E+06	1,81E+05
1,73E+06	3,84E+04	1,36E+06	2,10E+06	4,73E+05

1,76E+06	1,41E+05	1,52E+06	2,10E+06	4,73E+05
1,10E+06	5,13E+04	1,10E+06	1,57E+06	3,12E+05
1,62E+06	7,66E+04	1,80E+06	1,22E+06	8,40E+04
3,45E+06	1,39E+05	3,47E+06	1,22E+06	8,40E+04
9,37E+05	3,85E+04	1,11E+06	1,11E+06	1,24E+05

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