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1 **A 3D shear-wave velocity model of the upper mantle beneath China and surrounding**  
2 **areas**

3

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15

16 **Abstract**

17

18 We present a three-dimensional model of shear wave velocity for the upper mantle of China

19 and the surrounding region by analyzing 50338 vertical component multi-mode Rayleigh

20 wave seismograms, recorded at 144 permanent and more than 300 temporary broadband

21 stations in and around China. The procedure involves combination of 1-D path average

22 models obtained by modeling each Rayleigh waveform up to the 4th higher mode in a

23 tomographic inversion scheme. The dense station network and the use of multi-mode analysis

24 help to achieve a lateral resolution of a few hundred kilometers down to 400 km depth. The  
25 seismic lithosphere, as it is defined by the crust and the high velocity mantle lid, is to the first  
26 order thin in east China and thick in the west, with a high velocity lid extending down to  
27 about 200 km depth beneath much of the Tibet-Pamir plateau. Beneath India, the thickness of  
28 the seismic lithosphere gradually increases from ~100 km in south India to more than 150 km  
29 in north India, where it underthrusts the Tibetan plateau to approximately the Jinsha River  
30 Suture. High velocity lid extending down to 100-150 km depth are also observed in the Tarim  
31 basin, Sichuan basin and Ordos block. In the eastern part of the North China craton the  
32 seismic lithosphere is probably close to or thinner than 70 km. Adjacent to these areas, the  
33 high velocity lid in the eastern Yangtze craton and South China fold system extend down to  
34 70-80 km depth. A large-scale subhorizontal high velocity body is observed at depths of 150-  
35 350 km beneath the entire East China cratonic areas. This high velocity body might be the  
36 remnant of a delamination process which resulted in the decratonization of the North China  
37 and the Yangtze cratons.

38

39 **Key words:** Rayleigh wave, surface wave, tomography, fundamental mode, higher modes,  
40 China.

41 **1. Introduction**

42

43 China is an assembly of ancient continental fragments separated by fold belts, which were  
44 accreted from the late Proterozoic to the Cenozoic (e.g., Huang, et al., 1980). Its present  
45 tectonics has been profoundly shaped by the Indo-Asian continental collision in the southwest  
46 and the subduction of the NW Pacific plate in the east with resistance by the Siberian shield in  
47 the north (Fig. 1). China has three major Precambrian cratons: the North China craton (NCC,  
48 also called Sino-Korean craton), the Yangtze craton (YC, also called South China craton) and  
49 the Tarim block. The interactions among different blocks have formed the tectonic features  
50 seen today and caused many intraplate earthquakes (Ma, 1987; Ma, et al., 1984; Liu et al.,  
51 2007; Yin and Harrison, 2000). The convergence of Indian and Eurasian plates, which started  
52 50 million years ago, has created the world's largest plateau and is pushing the crust and  
53 mantle lithosphere out of its way to the east (e.g., Royden et al., 2008). The NW Pacific and  
54 Philippine Sea plate subduction zone produced substantial heterogeneity in the mantle beneath  
55 east China, as well as widespread uplift, volcanism and extension. North China and Mongolia  
56 comprise the major part of the Central Asian Orogenic Belt (CAOB), which was accreted  
57 from the Neoproterozoic to the Mesozoic due to the resistance of the Siberian shield (Windley  
58 et al., 2007). All of these events have left their imprint on the upper mantle structure.  
59 Unraveling the tectonic history and understanding the tectonic processes require a better  
60 knowledge of the China lithosphere.

61

62 High-viscosity lithospheric plates moving over a lower-viscosity asthenosphere is a basic  
63 element of plate tectonics. The terms lithosphere and asthenosphere were originally defined



64 with reference to rheology, with the lithosphere essentially behaving as elastic solid, and the  
65 asthenosphere deforming as a viscous fluid (Barrell, 1914). Later on, additional terms like  
66 thermal, chemical or seismic lithosphere have been introduced (Anderson, 1995; Eaton et al.,  
67 2009), with the seismic lithosphere being defined as the high velocity lid overlying a low  
68 velocity asthenosphere. The bottom of the high velocity lid, and therefore the seismic  
69 lithosphere, can be defined as the point of inflection overlying the low-velocity layer in the  
70 upper mantle velocity-depth relationship (Eaton et al., 2009). For the present work we adopt  
71 this definition.

72

73 Regional body wave tomography is sensitive to lateral variations but has poor vertical  
74 resolution in the shallow mantle, due to smearing along near-vertical propagation paths. The  
75 resolution normally begins at a depth roughly equal to the average inter-station distance. In  
76 contrast, the dispersion of surface waves provides a good vertical resolution (in general few  
77 tens of kilometers) of S-wave velocity. However, sampling continental upper mantle requires  
78 long period surface waves with typical path lengths greater than 2000 km (Sieminski et al.,  
79 2004; Li et al., 2008; Priestley and Tilmann, 2009), which limits horizontal resolution to at  
80 least a few hundred kilometers. In regional surface wave tomography, surface waves are often  
81 analyzed at periods shorter than 150 s, where the fundamental mode is mostly sensitive to the  
82 top 200 km. This sensitivity can be extended to a depth of more than 400 km by including  
83 higher modes (Debayle, 1999; Lebedev and Nolet, 2003; Priestley et al., 2006; Feng et al.,  
84 2010).

85

86 China is a very suitable place for surface wave study, as there are not only a lot of earthquakes

87 in plate boundary zones around China, but also many intraplate earthquakes within China.  
88 Fundamental mode surface wave studies of China have reached a resolution of several  
89 hundred kilometers showing features correlated with the large geological units (Romanowicz,  
90 1982; Griot et al., 1998; Ritzwoller and Levshin, 1998; Curtis et al., 1998; Huang et al., 2003;  
91 Friederich, 2003; Feng and An, 2010). These studies generally agree that the lithosphere  
92 reaches a thickness of more than 200 km in western China and thins to less than 100 km in  
93 eastern China. However, there are differences at a more regional scale. For example, Griot et  
94 al. (1998) and Huang et al. (2003) observed a thick lithosphere beneath the Tibetan plateau,  
95 while others reported on a thin mantle lid (Romanowicz, 1982) or a missing lithosphere  
96 (Friederich, 2003) beneath central and northern Tibet. The resolution can be improved by  
97 including surface wave overtones and by increasing the station density. Lebedev and Nolet  
98 (2003), Priestley et al. (2006) and Feng et al. (2010) have shown that the upper mantle  
99 structure of eastern Asia can be better constrained by fitting multi-mode surface waveforms,  
100 although they have only used few stations in China for which waveform data were available.  
101 Here we follow the approach used in Priestley et al. (2006) and apply the multi-mode surface  
102 wave tomography to 47 evenly spaced permanent broadband stations in China, for which the  
103 instrument response is well known. In order to increase the number of stations available we  
104 also used all publicly available temporary experiment data in China. The increased number of  
105 stations makes the inter-station spacing less than 300 km in east China (Fig. 2). The use of  
106 Rayleigh waves analyzed at periods longer than 50s for path lengths greater than a few  
107 thousands of kilometers provides a lateral resolution of several hundred kilometers extending  
108 to a depth of 400 km.  
109

110 In the present work we have performed both isotropic and anisotropic inversions. The  
111 isotropic components in both cases are very similar. We discuss here the isotropic component  
112 of the anisotropic inversion because it is less subject to biases. There is a few percent of  
113 azimuthal anisotropy in the uppermost 200 km. Interpreting the anisotropic pattern requires  
114 intensive resolution tests and will be done in another paper.

115

## 116 **2. Data and Methods**

117

118 Our data comprise Rayleigh waves in vertical component seismograms. We utilize the  
119 fundamental mode and overtones up to rank 4. Fig. 3 shows the sensitivity kernels of the  
120 Rayleigh-waves of different modes and at different periods. While the fundamental-mode  
121 Rayleigh waves theoretically exist for all frequencies, the higher modes are limited to higher  
122 frequencies. It can be seen that the sensitivity of the fundamental mode is generally limited to  
123 the upper mantle, whereas the higher modes provide additional sensitivity below the mantle  
124 transition zone. Note that Fig. 3 is based on a theoretical calculation using a PREM model  
125 modified to remove the low velocity layer (Priestley et al., 2006). The depth sensitivity of the  
126 observed data can be reduced by diverse factors such as missing frequency content, lack of  
127 higher mode excitation, mode conversion due to sharp lateral velocity contrasts and  
128 regularization of the inversion.

129

130 Waveform data from more than 400 stations (Fig. 2), which have been operational at some  
131 time between 1999 and 2007 have been requested from different agencies. We requested  
132 waveform data of 47 broadband stations from the Chinese Earthquake Network Center, many

133 of which have not previously been used for this kind of study. In addition, data from more  
134 than 300 temporary stations in China and nearly 100 stations around China were requested  
135 from the IRIS and GEOFON data centers. The selected distribution of stations helped in  
136 achieving a good path coverage and azimuthal distribution (Fig. 2).

137

138 The technique used for constructing a 3D Sv model proceeds in two distinct stages. It was  
139 previously employed in a number of regional scale surface wave tomography studies (e.g.  
140 Debayle, 1999; Debayle and Kennett, 2000; Heintz et al., 2005; Pilidou et al., 2004; Priestley  
141 et al., 2006).

142

### 143 **2.1. Waveform inversion for a path-average 1D model**

144

145 In the first stage we model each waveform by a 1D shear wave velocity model representing an  
146 average seismic structure from a source to a receiver. We use the automated version (Debayle,  
147 1999) of the waveform inversion approach of Cara and L ev eque (1987). An original aspect of  
148 this approach is the introduction of a set of secondary observables, built up from the wave-  
149 forms as the primary data of the inversion. Compared to the strongly nonlinear problem of in-  
150 verting for waveforms directly, these secondary observables have only a mild nonlinear de-  
151 pendence on the model parameters. This property minimizes the dependence on the starting  
152 model and reduces the number of iterations needed to find a 1D depth-dependent model,  
153 which predicts waveforms compatible with the observed surface wave seismogram. In detail,  
154 the observed vertical component data are cross-correlated with pure-mode synthetics com-  
155 puted for a reference model for the fundamental and four higher modes. The resulting cross-

156 correlograms are filtered at different frequencies using Gaussian filters, and then their envel-  
157 ope function is calculated. A set of secondary observables is then selected by sampling values  
158 taken at different time lags on the envelope on each of the actual cross-correlograms: one  
159 value at each significant local maximum of the envelope function and two values on the  
160 flanks of those maxima. The automated selection of these secondary observables is discussed  
161 in details by Debayle (1999) and Debayle and Ricard (2012). The waveform inversion  
162 matches these secondary observables with synthetic secondary observables. Once the envel-  
163 ope-based secondary observables are fitted, the inversion proceeds by fitting the phase of the  
164 cross-correlogram at the maximum of the envelope function. The cross-correlation with pure-  
165 mode synthetics obviates the need for explicit mode separation of the observed data and the  
166 automated selection of secondary observables makes the inversion scheme robust in the case  
167 of a mode being insufficiently excited or strongly overlapping with another mode.

168

169 This approach requires two basic assumptions that the observed seismograms can be represen-  
170 ted in terms of multi-mode surface waves that propagate independently and that they do so  
171 along great circle paths. The necessary condition for the validity of the first assumption is that  
172 the medium should be varying smoothly (Woodhouse and Jun, 1974) and for the second that  
173 the lateral velocity variations should not be too large. Kennett (1995) finds that the validity of  
174 the path-averaged approximation holds good at period greater than 30 s and for surface waves  
175 propagating beneath continents at regional distances (with typical path lengths of  $\sim 4000$   
176 km). For the case where surface waves cross major structural boundaries (continent-ocean  
177 transition) he suggests to use the path-average approximation only at periods greater than 50  
178 s. In addition, results from other studies (Trampert and Woodhouse, 1995; Yoshizawa and

179 Kennet, 2002) show that for the fundamental and first higher modes, off-great circle propaga-  
180 tion can be neglected at periods greater than 40 s and at epicenter-station distances smaller  
181 than 10000 km. For these reasons, and considering the large variation in crustal thickness,  
182 with extremely thick crust in Tibet and a thinner oceanic crust at the margins of the study re-  
183 gion, we choose to process surface waves in the period range 50-160 s. Removing the shortest  
184 periods ( $< 50$  s) reduces the effect of strong lateral variations in the shallow part. By limiting  
185 our analysis to long period surface waves, we can work safely with the great circle approximi-  
186 ation and expect that propagation or site effects can be neglected.

187

188 The waveform fitting procedure is automatic for each seismogram. The period range used  
189 here is 50-160 s for the fundamental and up to four higher Rayleigh wave modes, depending  
190 on their signal-to-noise ratio (SNR). At each period, the SNR is deemed adequate if the ratio  
191 between the maximum amplitudes of the envelopes of the signal and noise is greater than 3.  
192 The signal is evaluated in five bandwidths centered at periods of 50, 70, 90, 120, and 160 s.  
193 The inversion is performed for the upper mantle structure with the crustal structure being  
194 fixed. We used path average crustal models from the 3D global 3SMAC crustal model (Nataf  
195 and Ricard, 1995) and the smoothed PREM model for the mantle and compute the stress  
196 displacement functions necessary to build the initial synthetic seismograms using the code  
197 from Takeuchi and Saito (1972). Source parameters are taken from the Global CMT catalog.  
198 The inversion is considered successful if the final synthetics matches well the observed  
199 seismogram and if the inversion converges towards a unique and stable velocity model  
200 (Debayle, 1999). Debayle (1999) uses a chi-square misfit parameter to ensure that the  
201 secondary observables are well fitted. In addition, to make sure that the final model provides a

202 good fit to the actual seismogram, he computes  $E_{res/act}$ , the energy of the residual signal over  
203 the energy of the actual signal, and  $Re$ , the energy reduction of the residual signal between the  
204 initial and last iteration. These energies are summed over the signal for group velocities  
205 between 3.5 and 6 km/s. The inversion is considered successful if the misfit parameter is  
206 smaller than 3 and if  $E_{res/act} < 0.3$  or  $Re > 90\%$ . The output of this inversion scheme is a 1-D  
207 path average shear wave velocity model along each great circle path. In this study we  
208 obtained 50338 1D path averaged models. Fig. 4 shows the distribution of path lengths. For  
209 this data set we achieved more than 100 paths crossing each  $2^\circ$  by  $2^\circ$  for the entire study area,  
210 and more than 500 paths almost everywhere in China (Fig. 2).

211

212 One half of the rays have path length shorter than 6000 km (Fig. 4). We also included paths  
213 longer than 6000 km in the study to increase ray coverage, although longer paths involve  
214 larger Fresnel zone and are more susceptible to off great-circle deviations and multi-pathing.  
215 We tested this effect by repeating the entire analysis by either using only shorter paths (<6000  
216 km) or using all paths and found that the bias due to longer paths is not recognizable (see  
217 Supplementary Figs. 1 and 2). Comparison of Figs. 8 and S2 show that the effect of longer  
218 paths is small and does not affect the interpretation of our tomographic model. Our preferred  
219 model is the one that include longer paths, as we believe that the benefit of the additional ray  
220 coverage is greater than the bias due to off-great circle propagation.

221

222 Because of the imposition of the *a priori* crustal model, anomalies above the Moho are not  
223 constrained by the surface wave data but simply reflect the *a priori* model. Anomalies  
224 immediately below the Moho are in principle resolved, but will suffer from significant

225 artifacts if there are discrepancies between the actual and assumed crustal structures. This  
226 trade-off has been evaluated by Debayle and Kennett (2000), Pilidou et al. (2005) and  
227 Priestley et al. (2006), and is likely to be small at depths greater than 100 km.

228

229 In Fig. 5 we show examples of the waveform inversion for four selected paths. The phase of  
230 the surface waveforms has been matched very well for paths 1, 2, and 4. For path 3 the delay  
231 in the fundamental wave mode is still slightly underpredicted because the regularisation  
232 discourages extreme changes to the reference model, and an error bar of 5 % is used on the  
233 phase of the cross-correlogram function (Debayle, 1999), allowing a slight misfit between the  
234 observed and synthetic waveform. In fact, path 3 represents the worst case of the selected  
235 waveforms that is just at threshold of the acceptance. The amplitudes are not quite as well  
236 matched, as these depend more strongly on three-dimensional structure, and therefore cannot  
237 be matched perfectly by a 1D pure path inversion. For path 4 only changes in the upper 200  
238 km of the model are required to fit the waveform data, while deeper portion of the models  
239 were modified by the inversion for the other paths.

240

## 241 **2.2. Tomographic inversion**

242

243 In the second stage we combine the 1-D velocity models in a tomographic inversion using a  
244 scheme developed by Debayle and Sambridge (2004) for massive datasets. This scheme is an  
245 extension of the continuous regionalization algorithm of Montagner (1986) using the  
246 inversion approach of Tarantola and Valette (1982). This algorithm produces both the  
247 isotropic component of 3D Sv-wave speed heterogeneity and the azimuthal anisotropy. A



248 laterally smooth model is obtained by imposing correlation between neighboring points with  
 249 the use of a Gaussian covariance function. Indeed, the Tarantola and Valette (1982) approach  
 250 can be seen as a way of finding the model that gives the best fit to the data while keeping it as  
 251 “close” as possible to the *a priori* information. The smoothness of the inverted model in  
 252 poorly sampled regions is therefore mostly constrained by the width of the Gaussian  
 253 covariance function, while in regions with higher ray density the need for a satisfactory data  
 254 fit allows a rougher model. The Gaussian covariance function between two points  $r$  and  $r'$  is :  
 255 
$$C_{m0}(r, r') = \sigma_r \sigma_{r'} \exp(-\Delta_{r,r'}^2 / 2L_{corr}^2) ,$$
  
 256 where  $\Delta_{r,r'}$  is the distance between  $r$  and  $r'$ ,  $\sigma_r$  is the standard deviation in point  $r$  and  $L_{corr}$  is the  
 257 correlation length (Montagner, 1986).  $L_{corr}$  controls the horizontal smoothness of the model  
 258 and  $\sigma_r$  controls the amplitude of the model perturbation at a geographical point  $r$ . The Earth  
 259 model is discretized in  $1^\circ$  by  $1^\circ$  cells, which is much smaller than the surface wave  
 260 wavelength or Fresnel zone at our period of interest. Although the inversion problem is  
 261 strongly underdetermined (the number of independent pieces of information contained in the  
 262 data is less than the number of model parameter), the inversion problem is stabilized by the  
 263 use of appropriate regularization. A reasonable way to choose  $L_{corr}$  is to make sure that the  
 264 surface of width  $2L_{corr}$  centered around each of the ray paths ensures a good coverage of the  
 265 study area. With our dataset, this condition is fulfilled, even when  $L_{corr}$  corresponds to the  
 266 shortest wavelength ( $\sim 200$  km for 50 s Rayleigh waves). After several trials, we set  $L_{corr}$  to  
 267 250 km. This value allows us to exploit the information contained in our shortest period  
 268 Rayleigh waves (Sieminski et al., 2004), while keeping a relatively smooth model. Following  
 269 Montagner (1986) and Debayle and Sambridge (2004), we argue that the choice of  $L_{corr}$ , which  
 270 is based on the data wavelength, is more physically based than choosing damping parameters

271 in a classical inversion scheme. We show in Fig. S10 the effect of changing  $L_{\text{corr}}$  to 400 km.  
272 Although we obtain a smoother model, the large scales structures that are discussed in this  
273 paper are preserved. The *a priori* standard deviation is set to  $\sigma = 0.05$  km/s according to  
274 expected values at regional (Nishimura and Forsyth, 1989; Debayle and L  v  que, 1997;  
275 Debayle et al., 2001) and global scales (Debayle and Ricard, 2012). Following Montagner  
276 (1986) and Debayle and Sambridge (2004), we argue that the choice of  $L_{\text{corr}}$ , and of the *a*  
277 *priori* standard deviation, which are based on the data wavelength and on previous observed  
278 shear wave heterogeneities, are more physically based than choosing damping parameters in a  
279 classical inversion scheme.

280

281 In the next sections we will discuss the reliability of the model using different tests before  
282 presenting and discussing the preferred model.

283

### 284 **3. Resolution tests and reliability of the model**

285

286 In stage one of calculating 1D path-average model, artifacts can arise from errors in the  
287 assumed crustal model and source parameters. Synthetic tests in prior work suggest that the  
288 effects of crustal corrections with different crustal models, e.g., the 3SMAC or the CRUST2.0  
289 (<http://igppweb.ucsd.edu/~gabi/rem.html>), are indistinguishable at depths larger than 100 km  
290 and are minor at shallower depths (Debayle and Kennett, 2000; Pilidou et al., 2004; Priestley  
291 et al., 2008b). In addition, a good (redundant) path density and azimuthal coverage like what  
292 we have for this work is a basic requirement for reducing the influence of errors in source  
293 parameters. Cara and L  v  que (1987) already demonstrated the weakness of the dependence

294 on the reference model, so that we can safely start the inversion from a unique upper mantle  
295 model (a smooth version of PREM) with a crustal part adapted to each path.  
296

297 For the second stage of determining the 3D model using a continuous regionalization scheme,  
298 we computed the *a posteriori* error on the tomographic model, using the formalism of  
299 Tarantola and Valette (1982) adapted to the continuous regionalization by Montagner (1986).  
300 Maps and cross sections in the 3D distribution of the *a posteriori* error are shown in Fig. 6. It  
301 is well known that the *a posteriori* error, which is obtained from the square root of the  
302 diagonal terms of the *a posteriori* covariance matrix, is a useful tool in order to quantify  
303 model resolution. The *a posteriori* covariance matrix  $C_m$  is related to the *a priori* covariance  
304 matrix  $C_{mo}$  by  $C_m = (I-R)C_{mo}$ , where  $I$  is the identity matrix and  $R$  is the resolution matrix.  
305 Therefore, the error estimates depend on the *a priori* covariance matrix. We note that it is also  
306 true for  $R$ , and any resolution test will depend on the choices made for the correlation length  
307 and *a priori* standard deviation. Providing that our *a priori* choices are reasonable, the *a*  
308 *posteriori* errors are a very useful tool to guide the interpretation of the seismic model.

309 Regions where the *a posteriori* error is close to the *a priori* error (i.e.  $\sigma = 0.05$  km/s) can be  
310 regarded as the regions of poor resolution ( $R$  close to 0). Due to the dense path coverage and  
311 the presence of higher modes, the *a posteriori* error value is less than 0.035 km/s for the entire  
312 study region. This means that S-wave perturbations larger than  $\pm 0.035$  km/s ( $\sim 0.8\%$   
313 considering an average velocity of 4.4 km/s) are significant at the 68% confidence level and  
314 that perturbations greater than  $\pm 0.07$  km/s ( $\sim 1.6\%$ ) are significant at the 95% confidence  
315 level. We indicate the 0.8% and 1.6% contour levels on the cross-sections (Figs. 9 and S5) as  
316 visual guide to areas of significant heterogeneity at 68% and 95% confidence level,

317 respectively. As the *a posteriori* error is actually smaller than 0.035 km/s in many parts of the  
318 study region (Fig. 6), these contours represent a conservative estimate.

319

320 For the second stage we checked the dependence of the results on the *a priori* and starting  
321 models (note that the *a priori* and starting models are identical) with a simple analytical test  
322 (flat model resolution test). We built two input synthetic models by adding a uniform shear  
323 wave perturbation of 5% and 15% to the *a priori* model and performed 3D tomographic inver-  
324 sions in order to recover these two synthetic models (Fig. S9). From the output, it is evident  
325 that our *a priori* choices ( $L_{\text{corr}}=250$  km;  $\sigma = 0.05$  km/s) allow us to retrieve the flat model  
326 uniformly for the area of interest. In supplementary Fig. S9, we color in white a very narrow  
327 interval (approx. 1%) around the target value of the model. The smearing around the edges of  
328 the map can be interpreted as the effect of the width of the Gaussian correlation length.

329

330 The image resolution is not so easy to quantify, but the checkerboard tests can provide a  
331 qualitative measure of our ability to resolve a particular input model. We conducted a number  
332 of checkerboard tests to examine the ability of the selected data set to recover velocity  
333 anomalies of different size. Fig. 7 shows the test with seismic anomalies extending over  
334 500x500 km in the middle of the map horizontally and 100 km vertically. Alternating high  
335 and low velocity anomalies with magnitudes of  $\pm 6\%$  are spread over the entire volume,  
336 separated by  $\sim 500$  km wide zero percent anomalies. We calculated synthetic Rayleigh wave  
337 seismograms for the same ray paths, source parameters, frequency contents as in the observed  
338 data and carried out the same inversion procedure. At shallow depth ( $< 200$  km) the input  
339 model can be almost completely recovered. At depths of 200-400 km nearly half of the

340 magnitude of the anomalies can be recovered. The synthetic test shown in Fig. 7 gives an  
341 intuitive representation of our ability to recover a particular model from our ray coverage and  
342 *a priori* choices at the scale length of the individual blocks. However, as shown by L ev eque et  
343 al. (1993) such a test does not demonstrate that other synthetic models with larger size  
344 structure will be necessarily be, or even equally well, retrieved. For this reason, we performed  
345 a further synthetic test with checker dimension of 1000x1000 km horizontally and 100 km  
346 vertically (Fig. S8). The geometries of these larger scale anomalies are also well retrieved. We  
347 therefore assume that seismic anomalies larger than 500 km in horizontal and 50-100 km in  
348 vertical direction are reasonably well resolved by our data in the uppermost 400 km.

349

350 The study area is very heterogeneous with dramatic variations in crustal thickness (Li et al.,  
351 2006; Zhang et al., 2011). Beneath the orogenic belts of Tibet, Tien Shan and Pamir the  
352 maximum crustal thickness reaches more than 80 km, measured by different seismic means  
353 (e.g., Kind et al., 2002; Li et al., 2006; Zhang et al., 2011). At shallow depth (less than 100  
354 km) the 3D inversion may be affected by errors in our *a priori* knowledge of the crust. The  
355 strongest biases are expected in regions with the thickest crust. To test the effect of the thick  
356 crust we have removed the paths that cross a rectangular area in western China where the  
357 crustal thickness is mostly over 60 km (mainly Tibet-Pamir-Tien Shan orogenic belts, see Fig.  
358 S3) and repeat the tomographic inversion. This results in 51% of all paths being removed,  
359 leading to a dramatic loss of coverage west of  $\sim 105^\circ$ , and of course a total loss of information  
360 in the excluded area. We display maps between 50 and 150 km depth in Fig. S4 and two  
361 vertical sections EE" and FF" in Fig. S5 which can be compared with Fig. 8 and 9. The  
362 eastern part of the maps at 100 and 150 km depths are very similar to the final model (Fig. 8)

363 inverted using the entire dataset. Comparison of Figs. 9 and S5 demonstrates that the  
364 anomalies resolved at the 68% and 95% confidence level have a very similar shape indicating  
365 that Sv velocity perturbations are robust in this part of the model, not significantly biased by  
366 paths crossing regions with thick crust and robust against a removal of about half of our  
367 dataset. The large scale pattern of sections EE" and FF" (Fig. 9) is also preserved after the  
368 inversion with the reduced data set. In particular, high velocity anomalies are visible beneath  
369 the Ordos block and Sichuan basin (from sub-crustal depth to >200 km), the Songliao basin  
370 (at >100km depth) and for the South China fold system (thin lithosphere at <75 km depth).  
371 Previous tests based on the same approach to waveform fitting (e.g. Priestley et al., 2008b)  
372 suggest that even beneath Tibet, the influence of a fixed crust is likely to be small at depths  
373 larger than 100-125 km.

374

#### 375 **4. Observations**

##### 376 **4.1. Horizontal Sections**

377

378 Fig. 8 shows 6 horizontal sections of the isotropic Sv velocity perturbations of the 3D  
379 inversion at depths from 100 to 300 km.

380

381 The slice at 100 km depth shows high velocities in India, Tarim basin, Sichuan basin, Ordos  
382 block and Songliao basin and wide-spread low velocities elsewhere including Tibet, the  
383 CAOB and in the oceanic areas. In the northwest Pacific subduction zones the low velocity  
384 anomalies following plate boundaries should represent the mantle wedges. In the depth range  
385 of 100-200 km we can recognize the downgoing slab by high velocities. The west China and

386 CAOBS are characterized by low velocity anomalies at shallow depths. At 100 km depth, the  
387 low velocity anomaly in Tibet is sharply bounded by the Indian plate to the south, by the  
388 Tarim basin to the north and by the Sichuan basin and Ordos block to the east, which show up  
389 as high velocity anomalies typical of continental lithospheric mantle. The low velocities in the  
390 central part likely arise due to a combination of a too low crustal thickness in the *a priori*  
391 model and possibly low velocities in the shallowest upper mantle in northern Tibet.

392

393 S wave anomalies at 125-200 km depth largely reflect the variation in the thickness of the  
394 seismic lithosphere. High velocity anomalies indicate the presence of mantle lithosphere. Low  
395 velocities are mainly associated with the asthenosphere. In general west China including  
396 Tibet, Tien Shan-Pamir, Sichuan basin, and Ordos block is characterized by thick lithosphere  
397 while the lithosphere beneath east China is thin. The high velocity anomaly in the mantle  
398 lithosphere beneath much of Tibet and the Pamirs extends to a depth of 200 km. At 125 km a  
399 low velocity zone can be observed in northeastern Tibet, focused along the western Kunlun  
400 Fault (KF) and Jinsha River Suture (JRS) south of Tarim and Qaidam basins. At the same  
401 location reduced velocities (slow compared to their immediate surroundings, not relative to  
402 the reference model) can be discerned to a depth of 175 km. North central India has high  
403 velocities down to a depth of 150 km, while high velocities beneath the rest of India can only  
404 be seen to shallower depth (<125 km). The Songliao basin in NE China also has high  
405 velocities.

406

407 The amplitude of S wave anomalies reduces significantly at depths below 200 km. However,  
408 the resolution tests also indicate that the magnitude of recovered anomalies is reduced by

409 about 50%, such that the real change in the magnitude of anomalies is hard to quantify. In  
410 marked contrast to the structure at 150-200 km depth, low velocity asthenosphere dominates  
411 Tibet and the surrounding orogenic regions (e.g. Tien Shan). These low velocities appear to be  
412 surrounded by a (partially broken) ring of higher than average velocities. At 300 km depth  
413 high velocity is widespread in east China in a large area west of the Pacific subduction zone.  
414 The Pacific subduction zone including Taiwan and Japan are characterized by high velocity  
415 anomalies.

416

#### 417 **4.2. Vertical Sections**

418

419 We created 6 vertical sections crossing major tectonic units (Fig. 9). Sections AA", BB" and  
420 CC" are approximately in east-west direction, while sections DD", EE" and FF" are in north-  
421 south direction. Locations of the sections are indicated on the 100 km depth horizontal map in  
422 Fig. 8. For each section we plotted S velocity perturbation as well as the absolute velocity.  
423 Seismic anomalies depend on the reference model, and features such as low velocity zones  
424 (i.e. negative velocity gradients with depth) are more readily identified on absolute velocity  
425 profiles but smaller anomalies can be more easily identified in the perturbation image, such  
426 that both types of representation are complementary.

427

428 Section AA" extends from southernmost Pamir and central Tibet through the Sichuan basin  
429 and South China fold system (SCFS) to the Philippine Sea. A pronounced high velocity  
430 anomaly is observed beneath entire Tibet down to a depth of 200 km and is interpreted as the  
431 mantle lithosphere. The shallow low-velocity anomaly beneath Tibet is sharply bounded to the



432 east by the mantle lithosphere of the Sichuan basin, seen as high velocity body down to a  
433 depth of 175 km. Beneath the Pamir and Hindu Kush high velocity exists down to a depth of  
434 250-300 km. All these anomalies are significant at the 95% confidence level and can be  
435 clearly observed in relative and absolute velocity images. From the continental region east of  
436 the Sichuan basin to the oceanic area of the Philippine Sea the upper mantle is characterized  
437 by low velocity, indicating a thin lithosphere (<100 km). The high velocity beneath Taiwan  
438 extends from 100 km depth to at least 300 km. Our long period surface waves cannot resolve  
439 a thin slab next to a dominant low velocity mantle wedge. At larger depths low velocity  
440 anomalies are weaker, the high-velocity signature of the slab dominates and is picked up by  
441 our data. However, it is smeared horizontally by our long period dataset.

442

443 Section BB" passes through two cratonic regions (Tarim basin and NCC) and extends to the  
444 Tien Shan to the west and to the Philippine Sea to the east. The Tien Shan and the Qilian  
445 (QFS) orogenic belt are known to have thick crust whereas the Tarim basin has a normal  
446 continental crustal thickness (Li et al., 2006; Zhang et al., 2011). High velocity mantle  
447 lithosphere can be clearly seen beneath the Tarim basin in the velocity perturbations as well as  
448 in the absolute velocity. In the velocity perturbation image, the high velocity body appears to  
449 extend to the west beneath the Tien Shan and to the east beneath the QFS. High velocity  
450 mantle lithosphere exists beneath the Ordos block, which constitutes the western part of the  
451 NCC. The high velocity lid beneath the eastern NCC is too thin to be observed in our model,  
452 and is underlain by low mantle velocities between 75 and 175 km depth. Farther east the  
453 subducted oceanic slab of the Philippine Sea and the Pacific plates are associated with high  
454 velocity anomalies, in good agreement with the slab seismicity.

455

456 In section CC" the low velocity beneath the CAO B centered at the Hangay Dome can be seen  
457 as a low velocity anomaly reaching a depth of 100 km at the southern tip of Baikal lake.

458 Songliao basin in northeast China appears to have a deep lithospheric root as seen in the  
459 velocity perturbation, but has no distinct sub-lithospheric low velocity zone. The high velocity  
460 signature of the Pacific subducted slab is observed in the mantle beneath Japan and NE China.

461

462 The India-Eurasia collision zone can best be examined on section DD", which cuts through  
463 the Indian plate, central Tibet, the Hangay Dome and southern Siberia. The most significant  
464 feature of our mantle cross-section is the northerly dipping high velocity body, suggesting the  
465 Indian mantle lithosphere underthrusts much of Tibet until the JRS. Beneath the JRS and KF,  
466 south of the Qaidam basin, the high velocity mantle layer suddenly jumps to a shallower  
467 depth by about 50 km, which we interpret as the start of the Eurasian mantle lithosphere. At  
468 either edge of the thick mantle lithosphere beneath the plateau subvertical high velocity  
469 bodies are clearly visible at depths between 300 and 400 km, corresponding to the ring of  
470 faster velocities seen in map view (Fig. 8, 300 km depth slice).

471

472 Section EE" links the Sichuan basin and the Ordos block, which form the western parts of  
473 two cratons, the NCC and YC. The lithosphere of both cratons is significantly thicker (>150  
474 km) in the west, while it is much thinner (<70-80 km) in the east (compared to section FF"  
475 also see E-W sections AA" and BB"). In the section the Sichuan basin and Ordos block have a  
476 different velocity signature. The Sichuan basin is a more pronounced high velocity body, both  
477 in perturbation and in absolute velocities, extending to a depth of ~175 km. The Ordos block

478 extends to a depth of ~150 km and is less pronounced in the section of absolute velocity.

479

480 Section FF" is located in east China passing through the SCFS, the eastern parts of the YC

481 and the NCC and Songliao basin. The mantle lithosphere beneath the NCC is too thin to be

482 resolved. At very shallow depth (70-80 km) beneath the SCFS and YC high velocities are

483 visible, which we interpret as mantle lithosphere. Beneath Songliao basin high velocities

484 reach a depth of 300 km and spread to the north and south directions in the depth range of

485 150-400 km beneath the YC and NCC.

486

## 487 **5. Discussions**

488

### 489 **5.1. Segmentation of lithospheric blocks over China**

490

491 China consists of Precambrian cratons separated by Phanerozoic fold belts. However, the

492 thickness of the lithosphere does not follow the geographic locations of these tectonic units. It

493 is known from numerous studies (Huang et al., 2003; Lebedev and Nolet, 2003; Priestley et

494 al., 2006; Feng and An, 2010; Feng et al., 2010; Obrebski et al., 2012) that the lithosphere is

495 thin in east China and thick in the west, roughly divided by the North-South Gravity

496 Lineament (NSGL). The NSGL is a major gravity gradient, 100 km wide, which marks the

497 border between west and east China with distinct topographic, tectonic and seismic properties

498 and therefore has been recognized for a long time to be important in the evolution of eastern

499 Asia (Xu, 2007). Across the NSGL, the Bouguer gravity anomaly increases rapidly from -100

500 mGal in the west to -40 mGal in the east. Our result (Figs. 8-10) confirms earlier

501 observations but with more detailed information. The depth slice at 100 km in Fig. 8 clearly  
502 highlights the mantle lithospheric roots of the cratonic blocks, indicated by high velocities.  
503 These include the Tarim basin, west NCC (Ordos block) and west YC (Sichuan basin). The  
504 lithospheric roots extend to ~150 km depth beneath Ordos block and to ~175 km depth  
505 beneath Tarim basin and Sichuan basin. These cratonic blocks form the north and east borders  
506 of the Indian-Asian collision zone and have acted as rigid blocks resisting the plate motion  
507 and guiding lithospheric deformation around them during the collision (see Clark and Royden,  
508 2000; Royden et al., 2008). The northward moving Indian plate has a thickness of 100-175  
509 km with its thickest part in north central India adjacent to Tibet. The seismic lithosphere  
510 beneath much of the Pamir-Tibetan plateaus has doubled its thickness during the Indo-Asian  
511 collision with a maximum thickness over 200 km beneath the Tibetan plateau.

512

513 In the eastern portion of the NCC and YC, the lithosphere is too thin to be well observed by  
514 large-scale surface wave studies (Priestley et al., 2006; Lebedev and Nolet, 2003; Huang et  
515 al., 2009; Obrebski et al., 2012). We observe a weak high velocity signature which may mark  
516 the bottom of the high velocity mantle lithosphere beneath the east YC, but are still missing  
517 that of the NCC. Fig. 9 (cross section FF") suggests that the lithosphere of the east YC is 70-  
518 80 km thick, while the lithosphere of the east NCC would be thinner than ~70 km and  
519 therefore not resolved by our data. A thin lithosphere beneath the cratonic areas in east China  
520 is supported by receiver function studies. The base of the lithosphere beneath the east NCC  
521 was estimated by receiver functions as shallow as ~60 km, whereas it is ~10 km deeper in the  
522 east YC (Sodoudi et al., 2006; Chen et al., 2008; Chen, 2009). However, although our test  
523 (Figs. S3-S5) suggests that our data pick up high velocities of a shallow mantle lithosphere in

524 east China, it is important to keep in mind that this high velocity signature is only constrained  
525 locally at the 68% confidence level, and that shallow structure down to about 100 km depth  
526 may trade-off with crustal structure, an effect which is not accounted for in our tests.

527

## 528 **5.2 Sub-lithospheric structure**

529

530 The subducted Pacific slab has been clearly imaged by body wave tomography (e.g., Huang  
531 and Zhao, 2006; Li et al., 2008). In some places, we pick up the signature of a high velocity  
532 oceanic slab with surface waves, but the lateral resolution is not as good as that of body  
533 waves. In the northwest Pacific subduction zones the subducted oceanic lithosphere can be  
534 followed to a depth of 200-300 km (Figs. 8 and 9). The wide-spread low velocity anomalies  
535 following plate boundaries at shallower depths (<100 km) are likely to represent mantle  
536 wedges. In the depth range of 100-300 km we can recognize the downgoing slab by high  
537 velocities. At larger depths (greater than 300 km) the resolution is insufficient to clearly image  
538 the slab. A pronounced high velocity body is resolved at the 95% confidence level beneath  
539 Taiwan at a depth range of 150-300 km. High velocity anomalies have also been observed by  
540 body wave tomography (Huang et al., 2010) and surface wave tomography (Sibuet et al.,  
541 2004) and were interpreted as evidence for a subducted Eurasian slab beneath Taiwan. Ai et  
542 al. (2007) observed a thickening in the mantle transition zone beneath Taiwan, which is  
543 compatible with the high velocity Ryukyu slab penetrating the mantle transition zone.

544

545 No high velocity mantle lithosphere is recognized along the CAO that extends from the  
546 Altai Mountains to the east to the Pacific Ocean. Widespread low velocity anomalies exist

547 below the crust to a depth of 300 km (Fig. 8 and Fig. 9 sections C-E). Kustowski et al. (2008)  
548 also observed a low velocity upper mantle below the CAO. The most prominent low  
549 velocity anomaly is located beneath the Hangay Dome and is visible down to a depth of 150  
550 km, as also observed by Priestley et al. (2006).

551

552 In the Songliao basin, northeast China, there is no significant low velocity zone in the sub-  
553 lithospheric mantle. Instead, a weak high velocity anomaly is constrained at the 68%  
554 confidence level from about 100 km depth to a depth of 300 km and connects in the depth  
555 range of 150-350 km to a large-scale subhorizontal high velocity body that spreads from  
556 below the Songliao basin ~500 km to the north and more than 2000 km to the south  
557 underlying the entire YC and NCC (Fig. 8 and Fig. 9, sections CC" and FF"). In the next  
558 section we interpret this large-scale high velocity body, which is constrained with a  
559 confidence level greater than 68%, as delaminated mantle lithosphere underlying the entire  
560 cratonic area beneath east China.

561

### 562 **5.3 Destruction of the lithosphere beneath the east China cratonic areas**

563

564 The eastern portion of China mainland comprises different tectonic units, including two major  
565 cratons (NCC and YC), Songliao Basin and Xing'an Ranges north of NCC and the South  
566 China Fold System (SCFS) south of YC (see Fig. 1). A common feature of the entire region is  
567 that the lithosphere is thin. The NS trending NSGL marks a sharp transition of the thickness of  
568 the lithosphere from east to west. We observed the base of the high velocity mantle lid  
569 beneath the YC and SCFS at ~80 km depth. Low velocities beneath the NCC indicate a

570 thinner lithosphere there. Beneath the Songliao Basin, although we did not observe a  
571 significant low velocity zone in the sub-lithospheric mantle, Zheng et al. (2011) and Sun et al.  
572 (2010) showed with ambient noise tomography that the lithosphere is ~70 km thick in the  
573 area. We have seen that trade-offs with crustal structure may bias our estimation of the shear  
574 wave velocity in the uppermost 100 km, so that resolving a seismic lithosphere thinner than  
575 70 km with our approach is difficult.

576

577 The NCC and YC are two ancient cratons, which have formed and consolidated during  
578 Paleoproterozoic time and collided during Triassic time (Yang et al, 2010; Zhu et al., 2012).  
579 Diamond-bearing kimberlites erupted at ~470 Ma provide evidence of a thick (>200 km)  
580 lithosphere at that time. However, the Cenozoic basalts sampled a thin lithosphere of 80-120  
581 km thickness (e.g., Menzies and Xu, 1998; Griffin et al., et al., 1998; Kusky et al., 2007; Yang  
582 et al., 2010; Zhu et al., 2012). Therefore, it is commonly agreed that the thick lithosphere  
583 beneath the ancient cratons of the NCC and YC has been destroyed in the Mesozoic and the  
584 depleted cratonic lithospheric root has been removed. However, the extent and mechanism of  
585 the destruction of the lithospheric keel is an open debate (see Zhu et al., 2012, for a review).  
586 The region is characterized by high heat flow and extensive seismicity (Ma et al., 1984;  
587 Wesnousky et al., 1984). The extension experienced since the Mesozoic, together with  
588 Cenozoic volcanism in this area (Menzies & Xu, 1998; Yang et al., 2010), caused by  
589 delamination or thermal erosion of the thick lithospheric root (Kusky et al., 2007), may be  
590 responsible for the lithospheric thinning. Different tectonic events proposed to be responsible  
591 for the decratonization of the NCC and YC include the Inda-Eurasia collision, mantle plume  
592 activity, the collision of these two cratons and the west Pacific plate subduction. Recent

593 studies tend to agree that the latter is the major trigger for the decratonization (reviewed by  
594 Zhu et al., 2012).

595

596 We observed widespread high velocities, constrained with a confidence level greater than  
597 68%, at 150 to 350 km depth underlying the whole area of Songliao Basin and the eastern  
598 portion of the North China and Yangtze Cratons (Fig. 9, CC" and FF"). Priestley et al. (2006)  
599 reported on observation of a high-velocity feature beneath Songliao Basin. Obrebski et al.  
600 (2012) also observed a fast anomaly at 200 km depth beneath the NCC and interpreted it as a  
601 possible delaminated lithospheric root of the NCC. Our result confirms their observation but  
602 we show that the high velocity anomalies are distributed over a much larger area, 100-200 km  
603 in thickness and ~3000 km in length. Up to three local maxima (underneath the YC, NCC and  
604 Songliao basin) can be discerned in the wider high velocity anomaly, and it is possible that  
605 smearing makes more localized anomalies under these three regions appear connected. The  
606 high velocity zone in our mantle model is parallel to the Pacific subduction zone, underlies  
607 the entire region of the eastern China cratons, and is bordered to the west by the NSGL. It is  
608 still unclear how and when the NSGL formed, or whether it is even related to the lithospheric  
609 deformation. However, the NSGL marks the western border of the sub-lithospheric high  
610 velocity zone along its ~3000 km length in our model.

611

612 Various studies have shown that the eastern portion of the North China Craton does not possess  
613 a lithospheric root similar to those commonly seen beneath almost all other cratons (Menzies  
614 & Xu, 1998; Kusky et al., 2007; Zhu et al., 2012). The timing of root loss is not well  
615 constrained but seems to be ca. 140-120 Ma (Kusky et al., 2007). Two main scenarios



616 proposed in previous studies are density foundering or delamination (Gao et al., 2004; Kusky  
617 et al., 2007; Windley et al., 2010; Xu et al., 2013) and thermal-chemical erosion (Menzies and  
618 Xu, 1998). Recycled continental crust has been found in the NCC, suggesting that the lower  
619 part of cratonic mantle lithosphere has delaminated since the Mesozoic (Gao et al., 2004; Xu  
620 et al., 2013). Delamination is a sudden process that seems to agree with the timing of the  
621 lithospheric root loss of the NCC, however, a simultaneous response of the entire cratonic  
622 areas to this catastrophic loss is not yet clear (Kusky et al., 2007). Among different  
623 mechanisms, the decratonization of the North China Cratons is very likely related to the  
624 Mesozoic Pacific subduction beneath eastern China (Xu, 2007; Zhu et al., 2012). This  
625 hypothesis is supported by our observation that the large-scale high velocity body in sub-  
626 lithospheric mantle is parallel to the NSGL and the Pacific subduction zone. This high  
627 velocity anomaly may represent a remnant lithospheric root, which initially formed the lower  
628 part of cratonic mantle lithosphere and has delaminated since the Mesozoic. If this is the case,  
629 it is not clear how buoyant, depleted lower lithosphere can flounder nor is it clear how the  
630 delaminated material can persist in the upper mantle for such a long time duration. It is  
631 surprising that lithospheric material from a delamination event in the Mesozoic still remains  
632 in the upper mantle. Maybe the underlying subducted oceanic slab prevents its further  
633 descent. Subducted oceanic slab has been found to stagnate in the mantle transition zone  
634 below east China (e.g., Huang and Zhao, 2006; Li et al., 2008), providing a possible barrier  
635 for the delaminated lithospheric block from sinking deeper into the mantle. In fact, parts of  
636 the delaminated body might have contacted and joined the subducted oceanic slab. Although  
637 our observation favours the delamination model, the fundamental debates between  
638 delamination and erosion still remain.

639

#### 640 **5.4 The India-Asia collision zones**

641

642 We observe a thick high velocity mantle lid underlying much of the Tibetan plateau (Figs. 8,  
643 9). High velocity underthrust Indian plate has also been observed by global body wave  
644 tomography (e.g., Li et al., 2008), but with this technique seen to be limited to the southern  
645 Tibetan plateau (see also Kind and Yuan, 2010). Regional body wave tomography may reveal  
646 more lateral variation in the mantle lithosphere (e.g., Liang et al., 2012), but is less sensitive  
647 to a subhorizontal lithosphere with a large lateral extent. Most surface wave studies observed  
648 high mantle velocities over much of the plateau (e.g., Huang et al., 2003; Lebedev and Nolet,  
649 2003; Priestley et al., 2006; 2008a). This discrepancy can be largely explained by different  
650 resolutions of the body wave and surface wave studies (Nunn et al., 2014). However, the  
651 surface wave studies of Friederich (2003) as well as Feng and An (2010) and Feng et al.  
652 (2010) infer a low velocity zone in the upper 200 km beneath north Tibet. Agius and Lebedev  
653 (2013) infer low velocities to ~150 km depth in northern Tibet from two-station dispersion  
654 measurements, with high velocities at larger depths, an observation which can partially  
655 reconcile with both set of apparently contrasting surface wave observations. Observations of  
656 Pn (McNamara et al., 1997; Hearn et al., 2004; Liang and Song, 2006) and Sn propagation  
657 (Barazangi and Ni, 1982; McNamara et al., 1995; Barron and Priestley, 2009) point to low  
658 velocities and high attenuation immediately below the Moho north of the BNS. This might  
659 correspond to the localized low velocity zone we observe along the JRS and KF south of  
660 Tarim and Qaidam basins (see map at 125 km depth in Fig. 8) and also to the shallow  
661 northward dip of the top of the high velocity layer visible in cross-section D-D" (Fig. 9).

662 Using surface waves recorded by temporary experiments within Tibet, Ceylan et al. (2012)  
663 observed high velocity, probably Indian plate underthrusting Tibet up to about 34°N,  
664 coincident with the JRS in Eastern Tibet. Supporting evidence for a less attenuative and  
665 probably higher velocity mantle at depths of more than ~130 km comes from the frequency  
666 dependence of Sn attenuation in Northern Tibet (Barron and Priestley, 2009). Ceylan et al.  
667 (2012) additionally revealed a very localised deep-seated low velocity zone below and just  
668 north of the Kunlun Mountains, extending in depth to the limit of their resolution (~220 km).  
669 The body wave tomography images of Liang et al. (2012) also show this low-velocity feature  
670 whose N-S width is only around 200 km and is therefore not reliably imaged by our surface  
671 wave tomography. However, we did not observe a large-scale low velocity mantle zone in  
672 north Tibet that could represent a major upwelling of asthenosphere. Instead, the localized  
673 mantle low velocity zone probably marks the northern border of either underthrusting India or  
674 the overthickened Tibetan lithosphere, separating the Tibetan plateau from the Tarim basin  
675 and Tien Shan. As the properties of Indian and Tibetan mantle lithosphere might be quite  
676 similar, it is not possible to distinguish between these alternatives using seismological data.  
677

678 In Fig. 10 we compare the surface wave models with those obtained by P- and S-receiver  
679 functions (Zhao et al., 2010). Along all the sections we marked the positions of the Moho and  
680 LAB observed by Zhao et al. (2010) on the absolute velocity profiles derived by the surface  
681 wave inversion (Fig. 10). Below the crust the high velocity mantle lithosphere often finds  
682 agreement with the line drawings of the receiver function LAB. This is the case for the Indian  
683 LAB, represented by the white dashed lines in the southern part of the sections. The Asian  
684 LAB, which is observed from S-receiver functions and is indicated by the dashed lines in the

685 central and northern part of the sections, is only matching the base of the high velocity body  
686 on the east line (R01), but is not seen by the surface waves on the central (R02) and west lines  
687 (R03). The lower lateral resolution might prevent surface waves from observing the fine  
688 structure seen by receiver functions, which are more sensitive to the interior structure of  
689 incipient fragmentation of the Indian slab below Tibet (see Liang et al., 2012, Ceylan et al.,  
690 2012). It is also possible that receiver functions do not see the base of the lithosphere, but a  
691 mid-lithospheric structure, as reported for the North American craton (Abt et al., 2010; Kumar  
692 et al., 2012).

693

## 694 **6. Conclusion**

695

696 We derived 3-D upper mantle absolute shear wave velocities by modeling fundamental and  
697 higher mode waveforms of surface waves. We extended the multi-mode surface tomography  
698 of East Asia of Priestley et al. (2006) by adding more permanent stations within China and  
699 constrained the study area to China and its close vicinity. The reduced inter-station distances  
700 enabled us to reduce the lateral smoothing by using a smaller Gaussian correlation length  
701 during the regionalization approach, thus increasing lateral resolution. We created a 3D S  
702 velocity model over China with a good resolution from the top of the upper mantle to a depth  
703 of ~400 km. Compared to earlier studies, velocity anomalies are better and more sharply  
704 defined in the model. Similar to Priestley et al. (2006) the velocity perturbation decreases  
705 from ~10% at shallow depths to ~2% at depths of 300-400 km. Although synthetic recovery  
706 tests indicate that magnitude of anomalies below 200 km is not fully recovered, the reduction  
707 in percentage anomaly is too large to be explained by the reduced resolution alone. At 100-

708 200 km depth the model is sensitive to the lateral variation of the thickness of the mantle  
709 lithosphere. The seismic lithosphere is generally thinner in East China and thick in West  
710 China. It reaches a thickness of 200 km beneath the Pamir-Tibetan plateaus. Also observed as  
711 relatively thick seismic lithosphere (>100km) are the Indian plate, Sichuan basin, Ordos block  
712 and Tarim basin. The lithosphere in the eastern part of the Yangtze craton is as thin as 70-80  
713 km, whereas the lithosphere of the North China craton is too thin to be resolved. Beneath  
714 these two cratons, an extensive high velocity body at depths of 150-350 km is observed and  
715 resolved with a confidence level greater than 68%. A possible interpretation is to associate this  
716 body with the remnant ancient lithospheric material from a large scale delamination event that  
717 was proposed to be the cause of the decratonization of Eastern China. This suggestion needs  
718 to be probed in future studies, though.

719

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728

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937

938 **Figure captions:**

939

940 Figure 1: Topography map of China and adjacent regions with major tectonic units. Red and  
941 blue lines define borders of major tectonic units. Black dashed line denotes the North-South  
942 Gravity Lineament (Xu, 2007). Abbreviation are: SGFS, Songpan-Ganzi fold system; QFS,  
943 Qinling fold system; QDFS, Qing-Dabie fold system; SB, Sichuan basin; OB, Ordos block;  
944 KF, Kunlun fault; JRS, Jinsha-River suture; BNS, Bangong-Nujiang suture, YZS, Yarlung-  
945 Zangbo suture.

946

947 Figure 2: Map of stations and ray coverage. Triangles denote the seismic stations used in the  
948 study. The path density is defined by number of paths crossing a grid of  $2^{\circ} \times 2^{\circ}$ .

949

950 Figure 3: Rayleigh wave sensitivities as a function of depth at different periods for the  
951 fundamental and the first four higher modes.

952

953 Figure 4: Histogram of path lengths in the final dataset.

954

955 Figure 5: Example of waveform inversion for 4 paths. Locations of the paths are indicated in  
956 the upper panel. The middle panels show the initial (red) and final models (green). Waveforms  
957 are shown in the lower panels. Black curves are observed data, bandpass filtered within 50-  
958 160 s. Red and green curves are synthetic waveforms for the initial and final models,  
959 respectively.

960



961 Figure 6: *A posteriori* error map showing (a) horizontal sections for depths of 100, 125, 150,  
962 175, 200 and 300 km and (b) vertical sections along the same profiles as shown for the model.  
963 The *a posteriori* error values are less than  $0.035 \text{ km s}^{-1}$  for most of the region (compare the *a*  
964 *priori* value of  $0.05 \text{ km s}^{-1}$ ).

965

966 Figure 7: Horizontal and vertical slices of the checker board resolution test. Alternating high  
967 and low velocity perturbations with a size of  $500 \times 500 \text{ km}$  in horizontal and  $100 \text{ km}$  in depth  
968 and a magnitude of 6% are separated by zero-anomaly background in the input model.

969

970 Figure 8: Horizontal sections of the Sv-velocity perturbations at depths of 100, 125, 150, 175,  
971 200 and 300 km. The percentage anomalies for depths of 100-175 km and for depths of 200-  
972 300 km are denoted by different scales. Where the crust is thicker than about 60 km in the  
973 3SMAC model, the perturbations in the 100 km slice partly reflect the starting model.

974 Locations of the 6 profiles shown in Fig. 9 are indicated on the map at 100 km depth. Gray  
975 lines mark borders of major tectonic units from Fig. 1. Green dashed line denotes the North-  
976 South Gravity Lineament.

977

978 Figure 9: Cross-sections of Sv velocity perturbation and absolute velocity along 3 EW lines  
979 A-C and 3 NS lines D-F. Locations of the profiles are indicated in Fig. 8. Color scales for  
980 relative and absolute velocities are indicated at the bottom. Surface topography is plotted on  
981 top of each profile with major tectonic units indicated. Black dots denote the relocated  
982 earthquakes from the EHB catalog (Engdahl et al., 1998) within 100 km either side of the  
983 profile. Contour lines surround the areas whose anomalies (i.e. deviations from the reference

984 model) are significant at least at the 68% (dashed lines) and 95% (solid lines) confidence  
985 levels. Abbreviations: QFS, Qilian fold system; CAO, Central Asia Orogenic Belt; SCFS,  
986 South China fold system; YC, Yangtze craton; NCC, North China craton.

987

988 Figure 10: Comparison of the upper mantle model with three receiver function profiles in  
989 Tibet (Zhao et al., 2010). Along each profile, the upper panel shows the topography and the  
990 lower panel shows the upper mantle absolute velocities along the profile. Locations of the  
991 Moho and the LAB, derived by Zhao et al. (2010) from P and S-receiver functions,  
992 respectively are indicated in the velocity profiles.

993

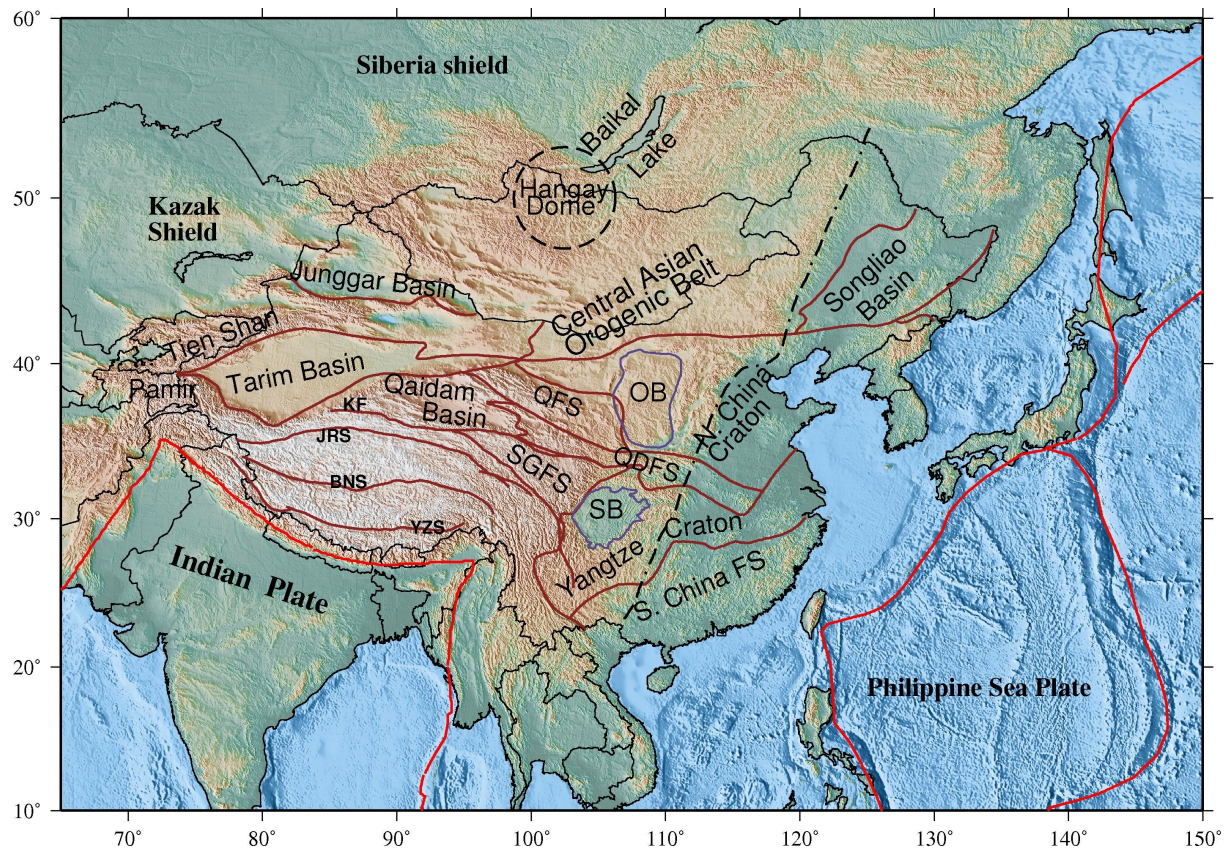


Fig. 1

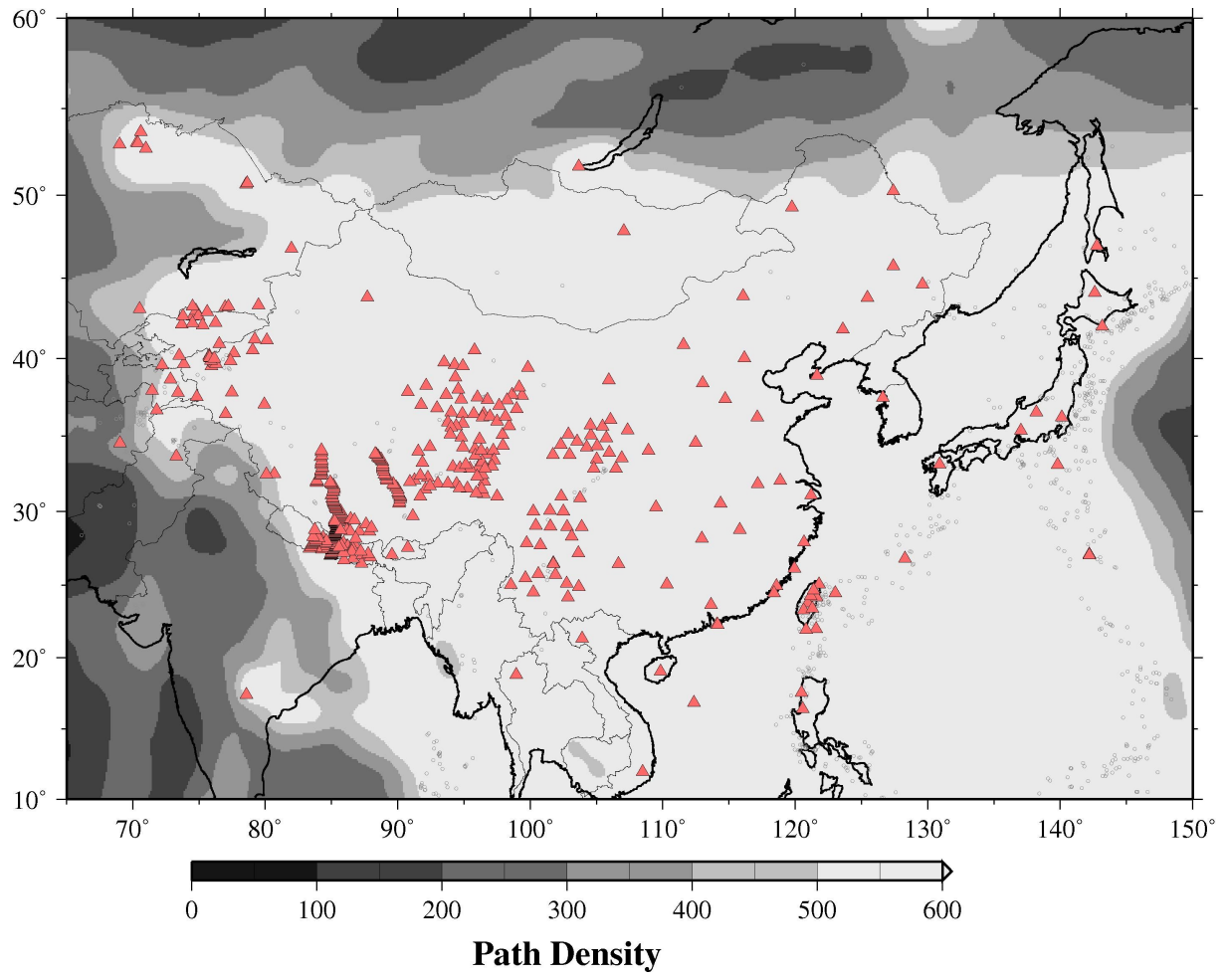


Fig. 2

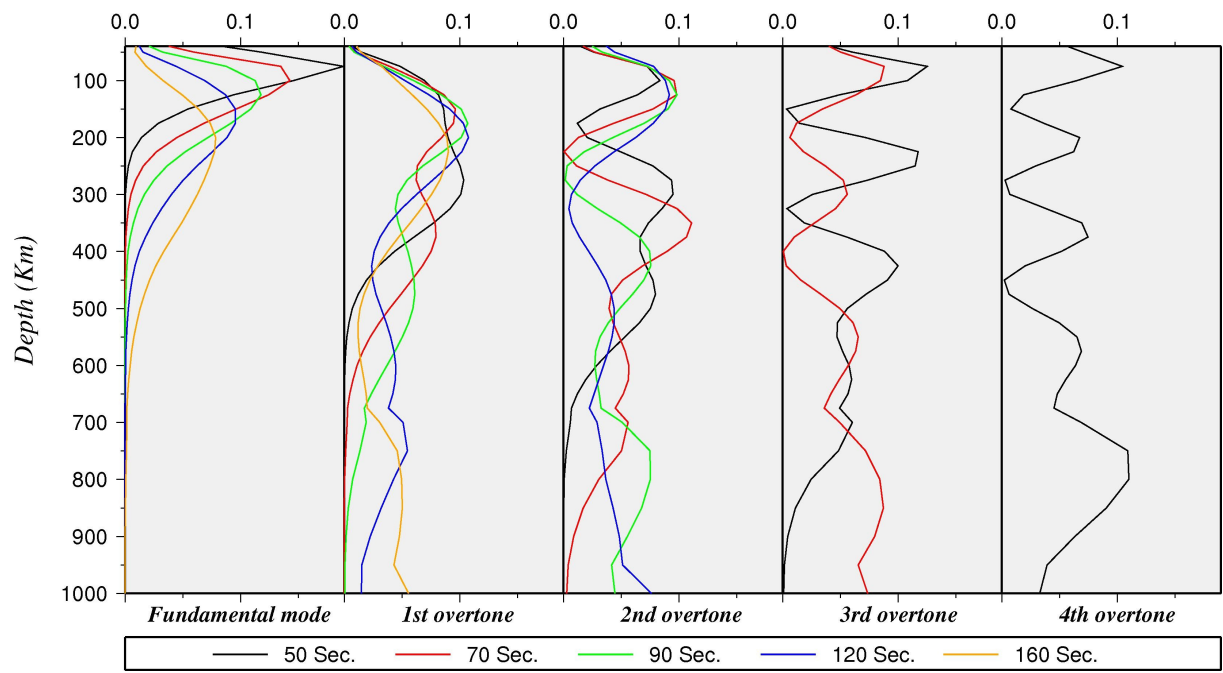


Fig. 3

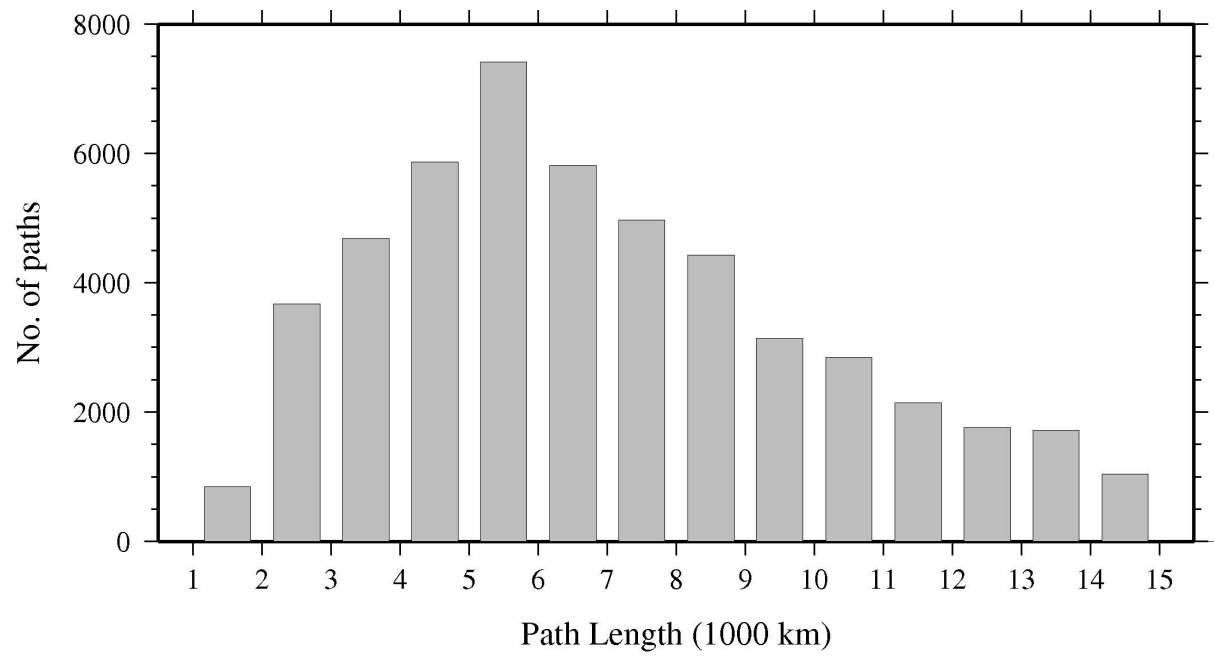


Fig. 4

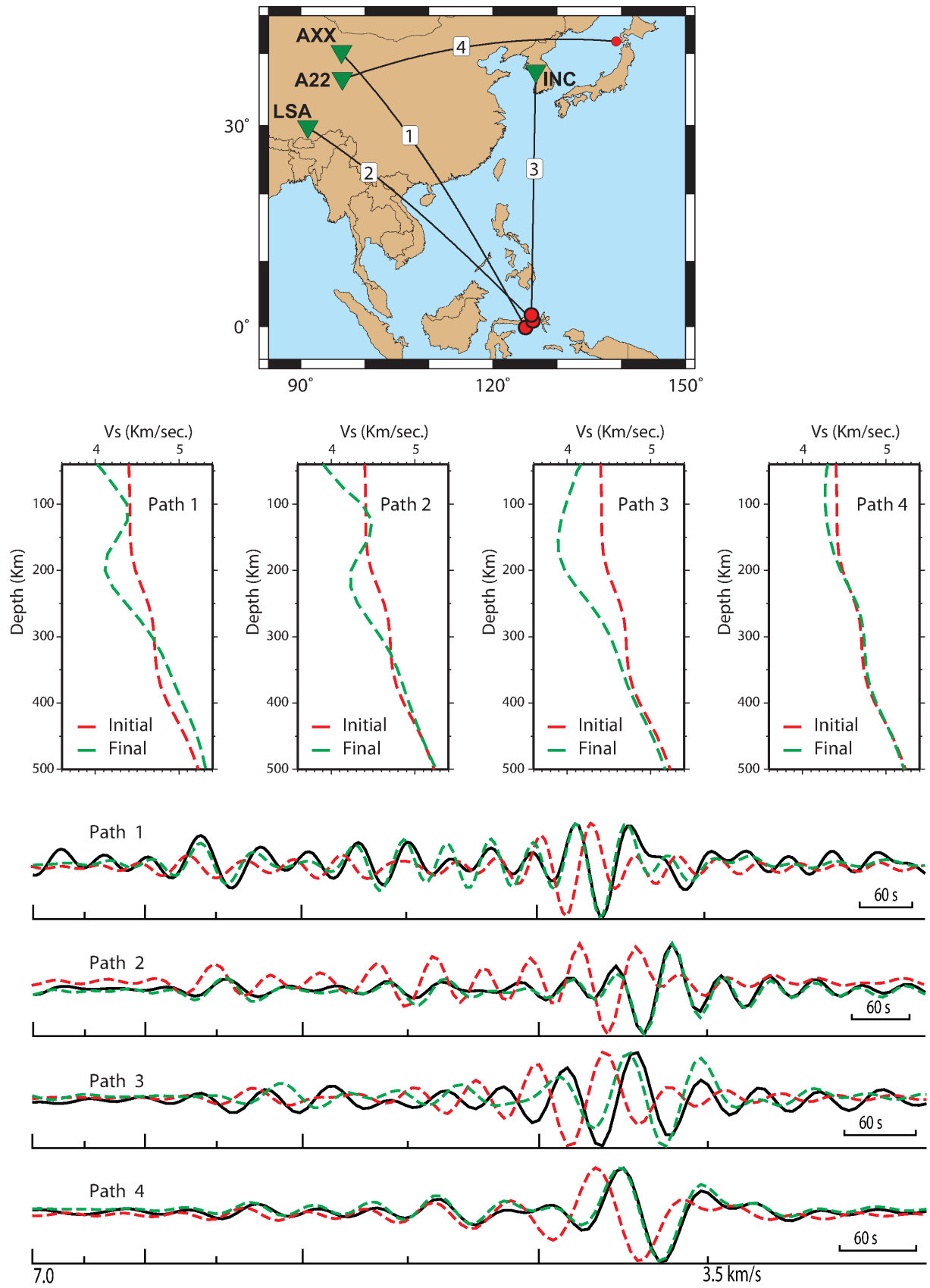


Fig. 5



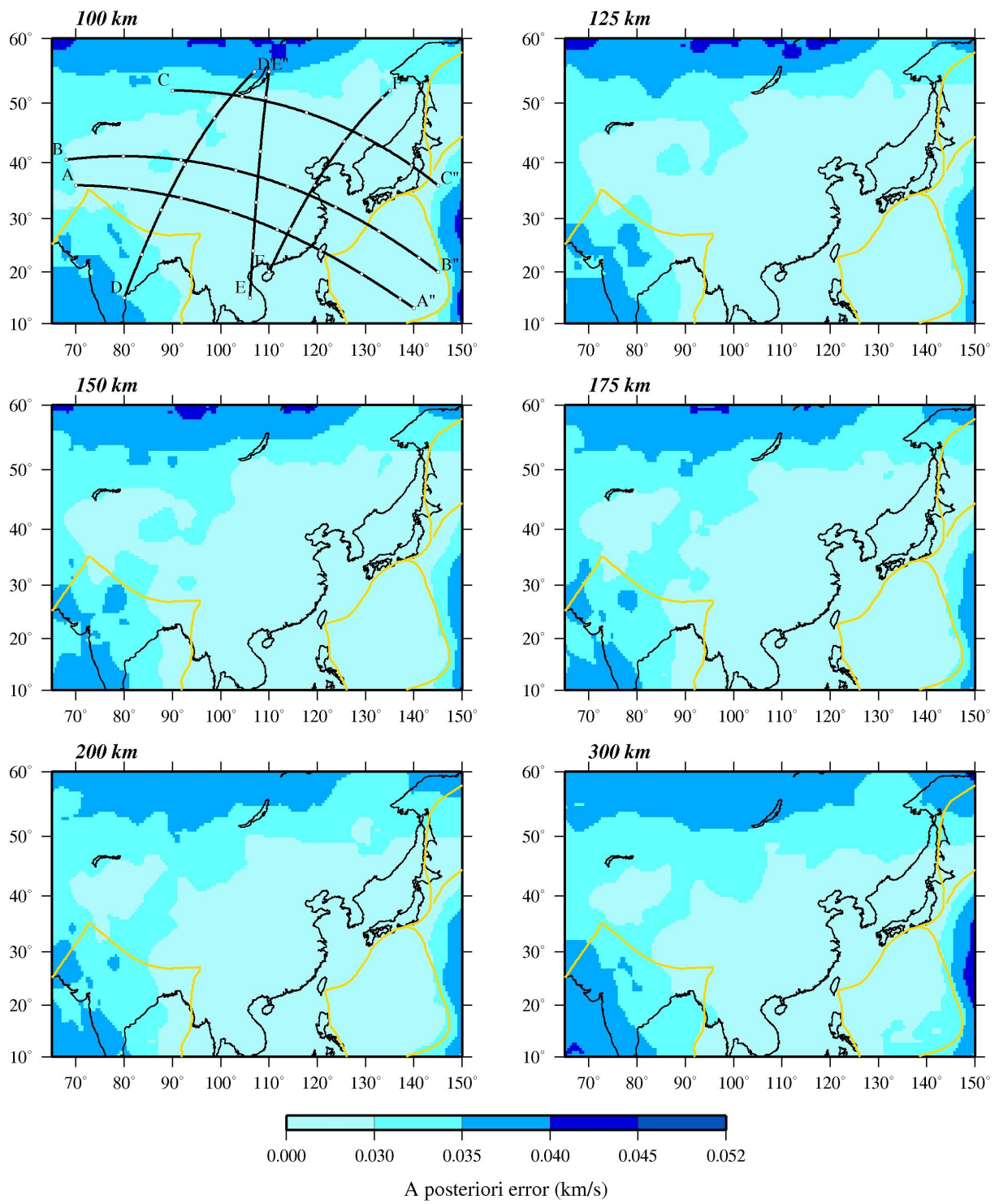


Fig. 6a



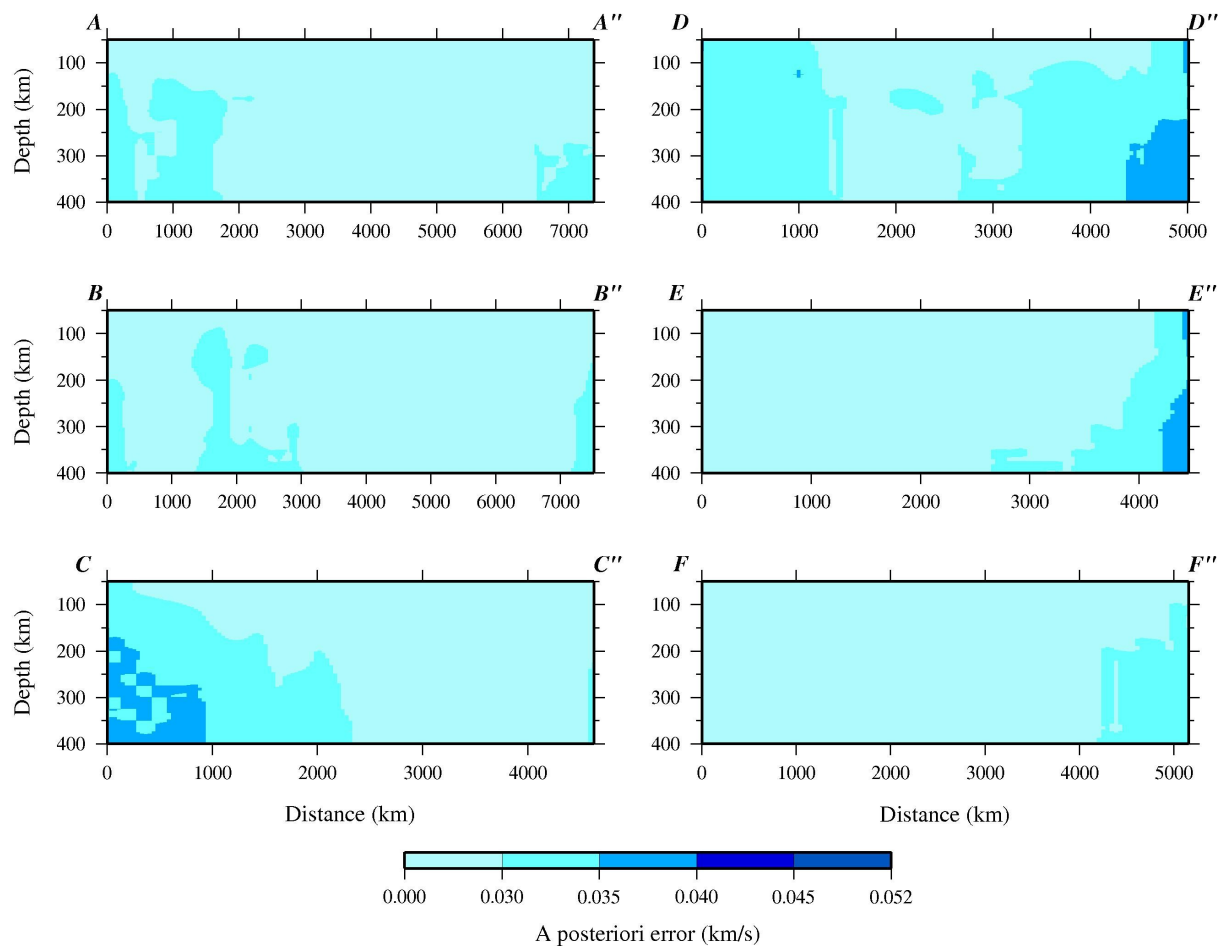


Fig 6b

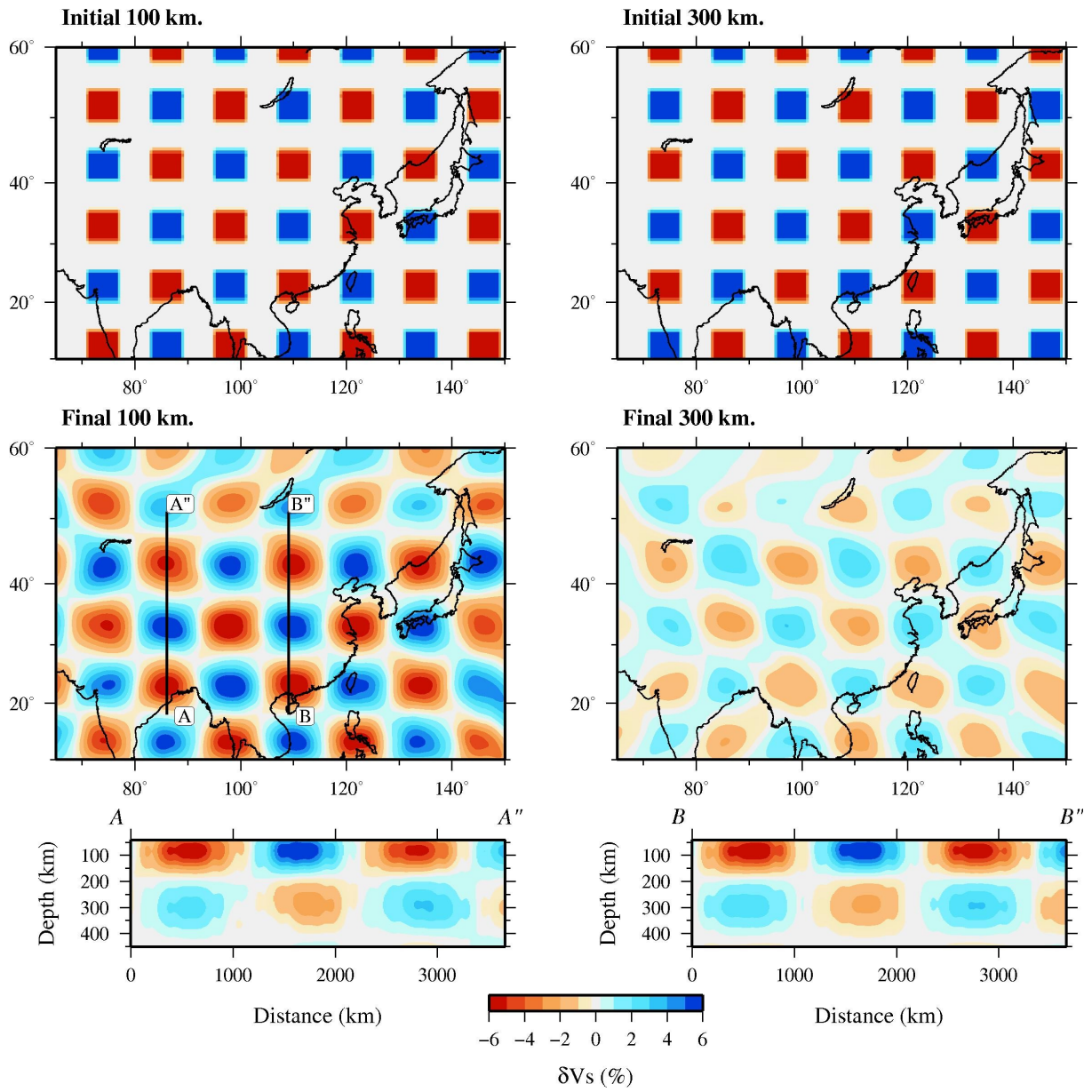


Fig. 7

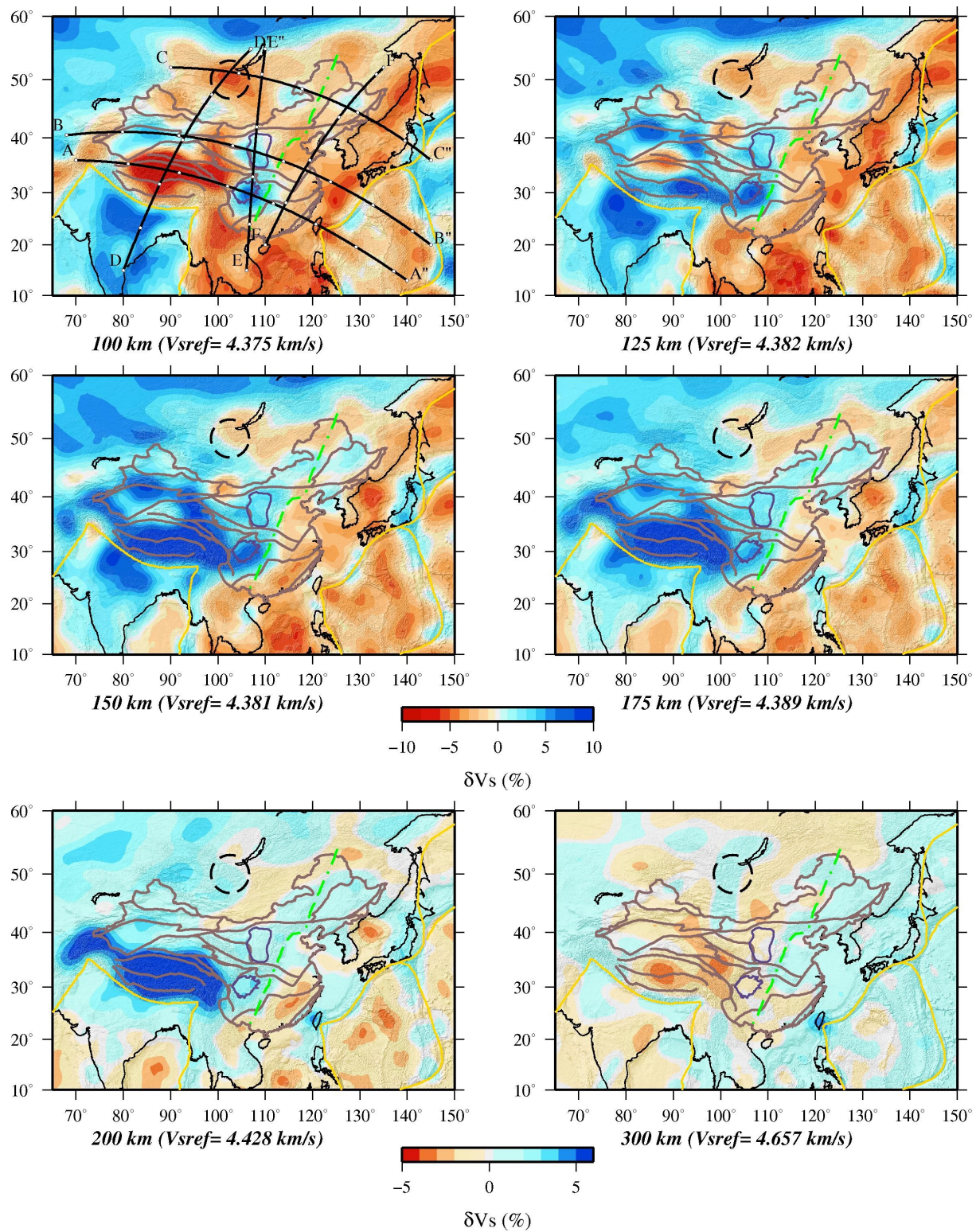


Fig. 8



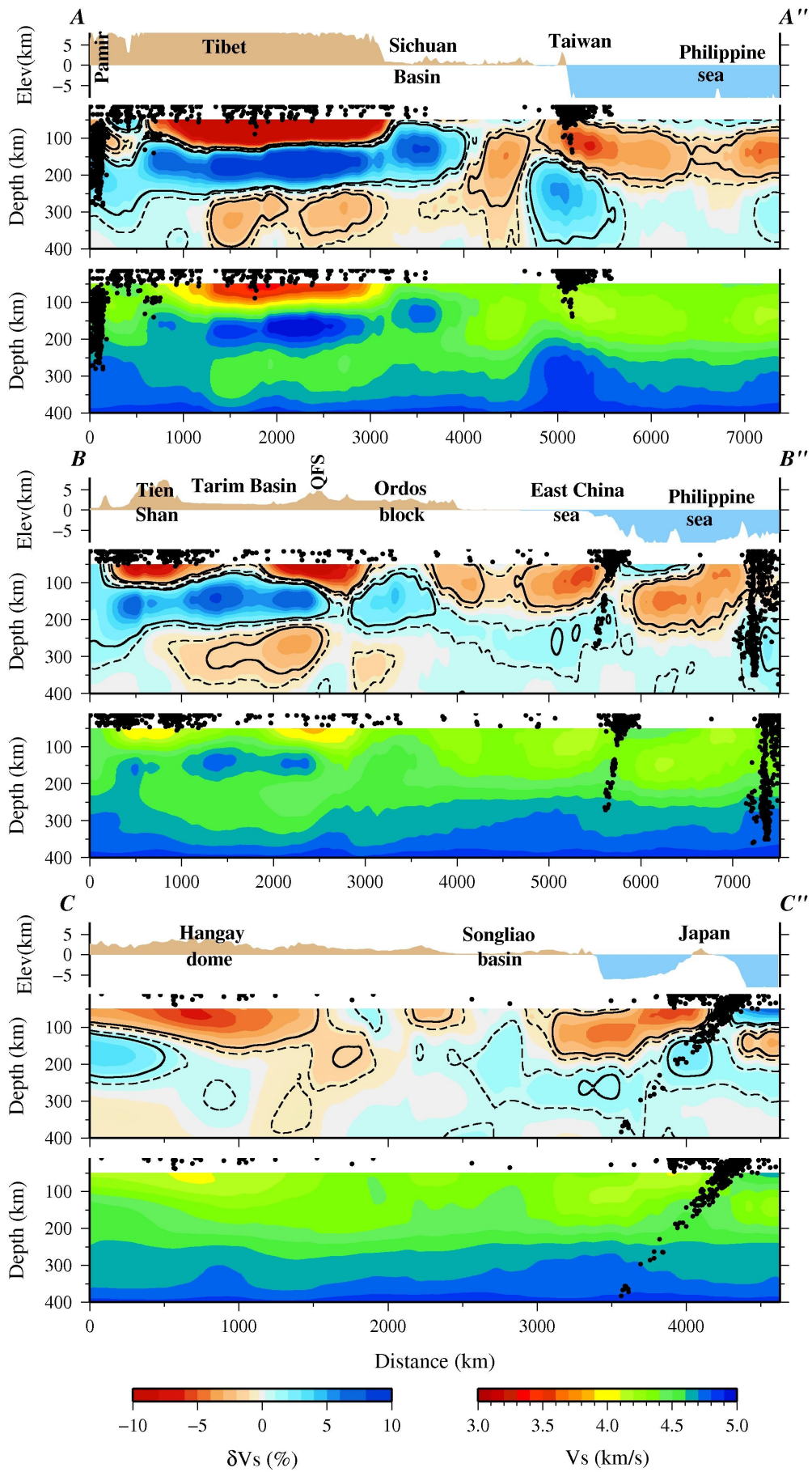


Fig. 9

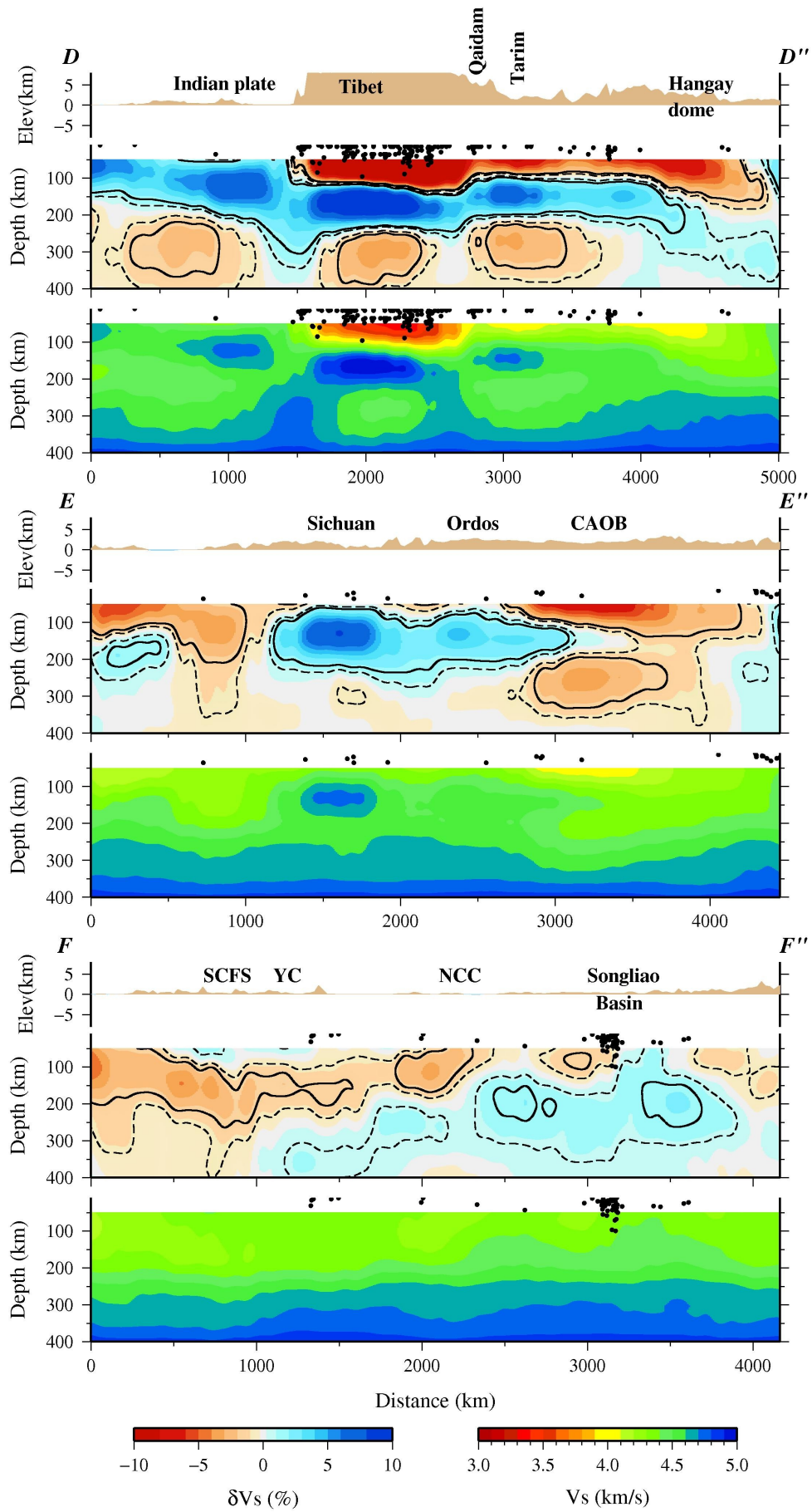


Fig. 9, continued

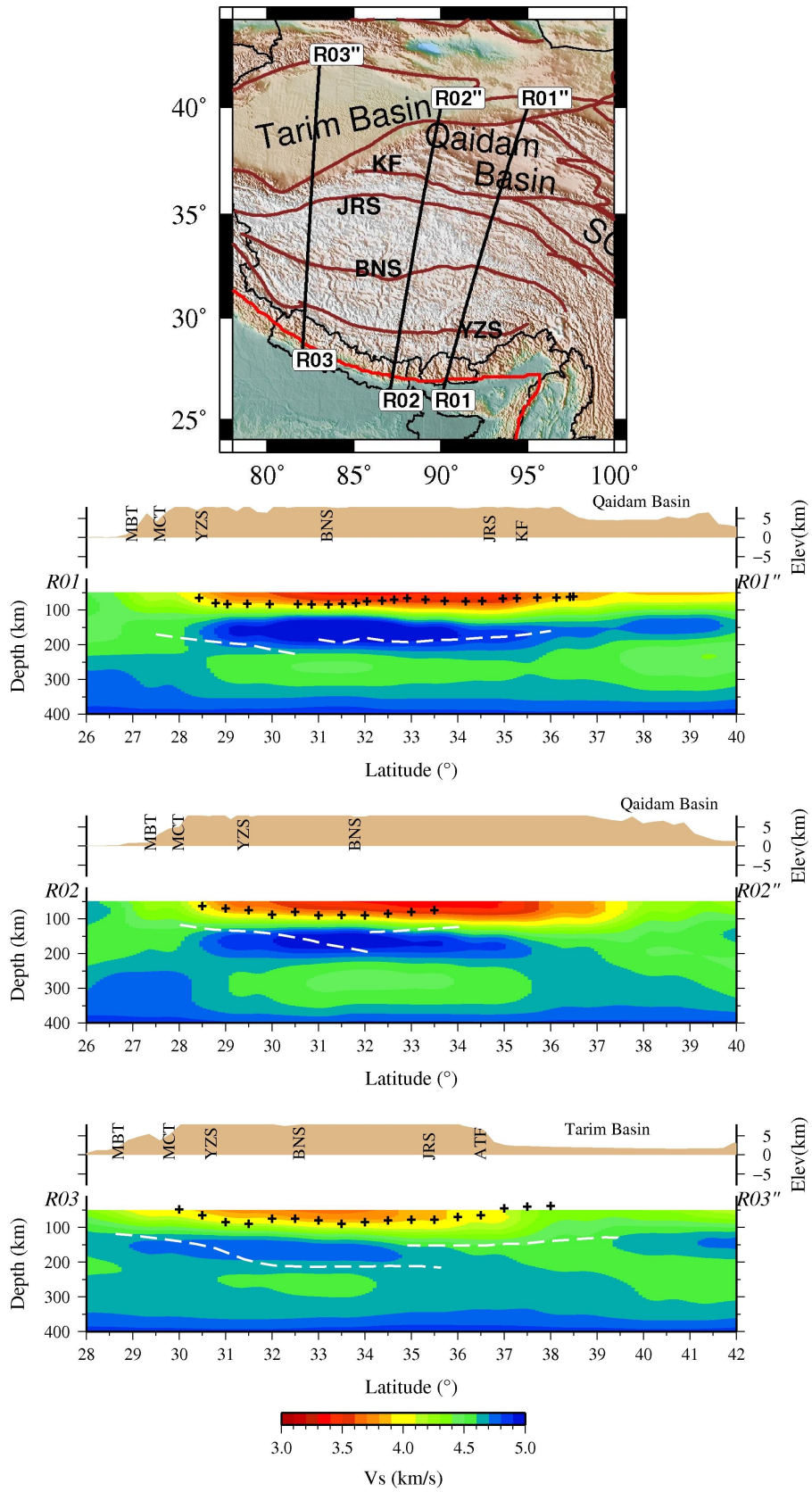


Fig. 10