Orbit-related sea level errors for TOPEX altimetry at seasonal to decadal timescales

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Abstract. Interannual to decadal sea level trends are indicators of climate variability and change. A major source of global and regional sea level data is satellite radar altimetry, which relies on precise knowledge of the satellite’s orbit. Here, we assess the error budget of the radial orbit component for the TOPEX/Poseidon mission for the period 1993 to 2004 from a set of different orbit solutions. The errors for seasonal, interannual (5-year), and decadal periods are estimated on global and regional scales based on radial orbit differences from three state-of-the-art orbit solutions provided by different research teams: the German Research Centre for Geosciences (GFZ), the Groupe de Recherche de Géodésie Spatiale (GRGS), and the Goddard Space Flight Center (GSFC). The global mean sea level error related to orbit uncertainties is of the order of 1 mm (8 % of the global mean sea level variability) with negligible contributions on the annual and decadal timescales. In contrast, the orbit-related error of the interannual trend is 0.1 mm yr\(^{-1}\) (27 % of the corresponding sea level variability) and might hamper the estimation of an acceleration of the global mean sea level rise. For regional scales, the gridded orbit-related error is up to 11 mm, and for about half the ocean the orbit error accounts for at least 10 % of the observed sea level variability. The seasonal orbit error amounts to 10 % of the observed seasonal sea level signal in the Southern Ocean. At interannual and decadal timescales, the orbit-related trend uncertainties reach regionally more than 1 mm yr\(^{-1}\). The interannual trend errors account for 10 % of the observed sea level signal in the tropical Atlantic and the south-eastern Pacific. For decadal scales, the orbit-related trend errors are prominent in a several regions including the South Atlantic, western North Atlantic, central Pacific, South Australian Basin, and the Mediterranean Sea. Based on a set of test orbits calculated at GFZ, the sources of the observed orbit-related errors are further investigated. The main contributors on all timescales are uncertainties in Earth’s time-variable gravity field models and on annual to interannual timescales discrepancies of the tracking station subnetworks, i.e. satellite laser ranging (SLR) and Doppler Orbitography and Radiopositioning Integrated by Satellite (DORIS).

1 Introduction

Sea level is an important indicator of climate variability and change. Based on tide gauge data using different techniques, the global mean sea level rise for the last century is estimated to be 1.2–1.9 mm yr\(^{-1}\) (Douglas, 1997; Church and White, 2011; Jevrejeva et al., 2008, 2014; Hay et al., 2015). Based on satellite altimetry data since 1993, the current rate of global mean sea level has been estimated to be more than 3 mm yr\(^{-1}\) (Cazenave et al., 2014; Ablain et al., 2016; Quartly et al., 2017). The main sources of the current rise are thermal expansion of the sea water and melting of glaciers and ice sheets. At interannual timescales, changes of terrestrial water storage imprint additionally on the global mean sea level (Llovell et al., 2011). Recent work (Watson et al., 2015; Fasullo et al., 2016) has focused on the detectability of accelerations in global mean sea level trends during the last decades. Regionally, sea level rates during the last 24 years show higher variability, they range from −1 to more than 10 mm yr\(^{-1}\). They are mainly linked to regional changes in the ocean’s density field, which might be induced by internal ocean variability, atmosphere–ocean interaction, or influx of
freshwater. Satellite altimeters are a unique source of global and regional sea level data and have been available continuously since the beginning of the 1990s. Precise orbits of altimetry satellites are a precondition for global and regional mean sea level investigations (Rudenko et al., 2012, 2014), and errors related to precise orbit determination (POD) are demonstrably one of the major error sources for global and regional sea level products (Ablain et al., 2015). A detailed description of the main factors contributing to the radial orbit errors is given by Fu and Haines (2013). The orbit errors have typically long wavelengths and may contain systematic contributions at seasonal to decadal timescales.

Couhert et al. (2015) investigated the main contributions to the radial orbit error budget for the Jason-1 and Jason-2 series based on Geophysical Data Records (GDR-D) at seasonal to decadal timescales for the second altimetry decade (2002–2013). According to their analysis, the orbit-related uncertainty of the global mean interannual and decadal trends is less than 0.1 mm yr$^{-1}$. As main factors for regional errors, they identified contributions from tracking data and reference frames (up to 8 mm) at seasonal timescales, contributions from tracking data (up to 3 mm yr$^{-1}$) and Earth’s time-variable gravity field (up to 2 mm yr$^{-1}$) at interannual timescales, and contributions from tracking data (up to 2 mm yr$^{-1}$) and Earth’s time-variable gravity field (up to 1.5 mm yr$^{-1}$) at decadal timescales. A corresponding assessment for the first altimetry decade (1992–2001) is still missing and is the rationale of this paper.

We assess the error budget of the radial orbit component for the TOPEX/Poseidon mission for the period 1993 to 2004 from a set of different orbit models. We have chosen TOPEX/Poseidon, since it is the reference altimetry mission used in the European Space Agency’s (ESA) Climate Change Initiative (CCI) Sea Level project over this time span (Ablain et al., 2016). We assess the radial orbit error budget at regional and global scales at seasonal, interannual, and decadal timescales by the analysis of three state-of-the-art orbit solutions derived and provided by different research teams from the German Research Centre for Geosciences (GFZ), the Groupe de Recherche de Géodésie Spatiale (GRGS), and the Goddard Space Flight Center (GSFC). Note that our assessment necessarily excludes contributions from errors common to these three orbits. However, since the three orbits were derived using various up-to-date models, the errors common to the three orbits should be rather low, which makes us confident that our error estimates represent most of the error. In our further analyses, we use test orbits calculated at GFZ to investigate the impact of uncertainties of the tracking station subnetworks, of the reference frame, and of the Earth’s time-variable gravity field models on the radial orbit component and hence the derived sea level.

A detailed description and assessment of the analysed orbits as well as specifications of the altimeter data processing are given in Sect. 2. Section 3.1 describes the methods implemented to assess the orbit errors for the different timescales and the corresponding results for global and regional scales. The estimates of the orbit-related error for global mean and regional sea level are given in Sect. 3.2 and 3.3, respectively. The specific orbit-related errors for ascending and descending passes are investigated in Sect. 3.4. In Sect. 3.5, we examine for which areas the orbit error reaches more than 10% of the corresponding sea level variability. The main findings are summarized and discussed in Sect. 4.

## 2 Orbit and altimetry data

### 2.1 Description of the analysed orbit solutions

Our aim is to assess the range and the characteristics of radial orbit errors on regional and global scales. Therefore, the differences between three independent state-of-the-art orbit solutions available for the TOPEX/Poseidon mission are analysed. All orbit solutions are derived in the International Terrestrial Reference Frame (ITRF) 2008 (Altamimi et al., 2011) and use satellite laser ranging (SLR) and Doppler Orbitography and Radiopositioning Integrated by Satellite (DORIS) tracking data but are based on different software and on distinct models. The actual multi-mission GFZ orbit solution VER11 (Rudenko et al., 2017) is used as a reference in this paper and is called REF hereafter. The GSFC std1504 orbit (Lemoine et al., 2010; Beckley et al., 2015) has been chosen by the ESA CCI Sea Level Phase 2 project and differs in many aspects from the GFZ orbit, regarding software as well as the suite of implemented models including another Earth’s gravity field model. As the third model, we have chosen the GRGS orbit solution (Soudarin et al., 2016), which is derived using models similar to those of the GFZ solution but employs another software package. The main models used for GFZ REF, GRGS, and GSFC std1504 orbits are described in Table 1. The main differences in these three orbit solutions are related to the choice of the Earth’s time-variable gravity (TVG) field models, ocean tide model, modelling of non-tidal atmospheric and oceanic gravity, and the treatment of geocentre variations in station displacements, as well as the constraints of the observation data (SLR/DORIS). While for the GRGS solution comparatively high weight is on the SLR data, for the GFZ solution there is higher weight on the DORIS data. Proper modelling of the Earth’s gravity field, in particular of its time-variable part, is crucial for the computation of precise orbits of altimetry satellites and has been shown to contribute to errors in regional sea level trends and seasonal signals (Rudenko et al., 2014; Esselborn et al., 2015). For the pre-GRACE (Gravity Recovery and Climate Experiment) period, the TVG field is poorly constrained. The weekly TVG solutions used for the GSFC orbit were derived up to degree and order 5 from the analysis of SLR and DORIS observations to 20 geodetic satellites starting from 1993 (Lemoine et al., 2016). The TVG part used for the GFZ REF (GRGS) orbits consists of the combina-
Table 1. The main models used for calculation of GFZ VER11, GSFC std1504, and GRGS orbits.

<table>
<thead>
<tr>
<th>Parameter</th>
<th>GFZ REF (VER11) orbit</th>
<th>GSFC std1504 orbit</th>
<th>GRGS orbit</th>
</tr>
</thead>
<tbody>
<tr>
<td>Polar motion and UT1</td>
<td>IERS EOP 08 C04 (IAU2000A) series with IERS diurnal and semi-diurnal variations</td>
<td>IERS bulletin A daily (consistent with ITRF2008), diurnal, and semi-diurnal variations</td>
<td>IERS EOP 08 C04</td>
</tr>
<tr>
<td>Station displacements due to annual geocentre variations</td>
<td>None</td>
<td>Ries (2013)</td>
<td>None</td>
</tr>
<tr>
<td>Non-tidal atmospheric loading effect on stations</td>
<td>Based on ECMWF ERA-Interim data</td>
<td>None</td>
<td>None</td>
</tr>
<tr>
<td>Ocean loading effect on stations</td>
<td>FES2004 (Lyard et al., 2006)</td>
<td>GOT4.10 (Ray, 2013)</td>
<td>FES2012 (Carrère et al., 2012)</td>
</tr>
<tr>
<td>Static Earth’s gravity field model</td>
<td>EIGEN-6S4 (Förste et al., 2016) degree/order (d/o 81–90</td>
<td>Updated harmonic piecewise fit weekly solutions (Lemoine et al., 2016) up to d/o 5</td>
<td>EIGEN-6S2 (Rudenko et al., 2014)</td>
</tr>
<tr>
<td>Earth’s time-variable gravity field model</td>
<td>EIGEN-6S4 degree 2: yearly value and drift term;</td>
<td>EIGEN-6S4 degree 2: yearly value and drift term;</td>
<td>EIGEN-6S2 degree 2: yearly value and drift term;</td>
</tr>
<tr>
<td></td>
<td>d/o 1–80: periodic (semi-) annual variations;</td>
<td>d/o 2–50: periodic (semi-) annual variations;</td>
<td>from 1 January 2003: yearly values and drift</td>
</tr>
<tr>
<td></td>
<td>from 15 August 2002: yearly values, drift terms and</td>
<td>from 1 January 2003: yearly values and drift terms for</td>
<td>terms for d/o 2–50</td>
</tr>
<tr>
<td></td>
<td>(semi-) annual variations for d/o 1–80</td>
<td>d/o 2–50</td>
<td></td>
</tr>
<tr>
<td>Ocean tide model</td>
<td>EOT11a (Savchenko and Bosch, 2012) up to d/o 80</td>
<td>GOT4.10 up to d/o 50</td>
<td>FES2012 up to d/o 50</td>
</tr>
<tr>
<td>Non-tidal atmospheric and oceanic gravity</td>
<td>GFZ AOD1B RL05 up to d/o 100 (Dobslaw et al., 2013),</td>
<td>ECMWF 6-hourly fields up to d/o 50</td>
<td>3-hourly ERA-Interim and TUGO R12 up to d/o 50</td>
</tr>
<tr>
<td></td>
<td>including ECMWF 6-hourly fields and OMCT</td>
<td></td>
<td></td>
</tr>
<tr>
<td>Atmospheric density model</td>
<td>MSIS-86 (Hedin, 1987)</td>
<td>MSIS-86</td>
<td>DTM 94, with the best available solar activity</td>
</tr>
<tr>
<td>Earth radiation and albedo</td>
<td>Knocke et al. (1988)</td>
<td>Knocke et al. (1988)</td>
<td>Albedo and IR pressure values interpolated from</td>
</tr>
<tr>
<td>Radiation pressure model</td>
<td>Tuned eight-panel (Cerri and Ferrage, 2016)</td>
<td>Tuned eight-panel</td>
<td>ECMWF 6hr grids</td>
</tr>
<tr>
<td>Tracking data</td>
<td>SLR, DORIS (Mendes and Pavlis, 2004)</td>
<td>SLR, DORIS (Mendes and Pavlis, 2004)</td>
<td>Thermo-optical coefficient from pre-launch box</td>
</tr>
<tr>
<td>DORIS tropospheric correction model</td>
<td>Vienna Mapping Functions 1 (Boehm and Schuh, 2004)</td>
<td>Vienna Mapping Functions 1</td>
<td>model</td>
</tr>
<tr>
<td>DORIS modelling</td>
<td>DORIS beacon frequency bias modelling</td>
<td>DORIS beacon phase centre</td>
<td>DORIS beacon phase centre</td>
</tr>
<tr>
<td>DORIS system time bias</td>
<td>Estimated once per arc</td>
<td>Estimated once per arc</td>
<td>None</td>
</tr>
<tr>
<td>Drag coefficients</td>
<td>Estimated once per arc</td>
<td>Estimated once per arc</td>
<td>Estimated every 12 h</td>
</tr>
<tr>
<td>Along- and cross-track empirical accelerations (once per revolution)</td>
<td>Estimated every 24 h</td>
<td>Estimated every 24 h</td>
<td>Estimated once per arc (3.5 days)</td>
</tr>
<tr>
<td>SLR antenna reference</td>
<td>LRA model (see note below)</td>
<td>LRA model (see note below)</td>
<td>X: 1.2429, Y: -0.0012, Z: 0.8783 in metres</td>
</tr>
<tr>
<td>DORIS antenna reference</td>
<td>Pre-launch</td>
<td>Pre-launch</td>
<td>Pre-launch</td>
</tr>
<tr>
<td>SLR/DORIS observation weight</td>
<td>3 cm/0.05 cm s⁻¹</td>
<td>10 cm/0.2 cm s⁻¹</td>
<td>1 cm/0.03 cm s⁻¹</td>
</tr>
</tbody>
</table>

tion of yearly coefficients, drift terms, and annual and semi-
annual variations for degrees and orders 1 to 80 (2 to 50)
derived from GRACE data and SLR measurements to the
Laser Geodynamic Satellite (LAGEOS)-1/2. The annual and
semi-annual coefficients used for the GFZ REF orbit are fit-
ted yearly starting from August 2002. For the pre-GRACE
period before August 2002 (January 2003), only the degree-
two terms exhibit yearly values and drift terms; however, the
annual and semi-annual variations, which were derived for
the GRACE period, are applied for degrees and orders 1–80
(2–50) (Rudenko et al., 2014, Förste et al., 2016).

The approach adopted for the estimation of the radial orbit
effects implies that errors common to all three orbits cannot be
detected. In particular, all three orbits rely on the ITRF2008
reference frame and basically the same set of tracking sta-
tions. To further estimate the orbit-related radial orbit er-
ror budget due to the most significant factors, we have de-

erived five test orbits based on the GFZ REF orbit. The er-
rors related to inconsistencies of the tracking data networks
are tested by using only one tracking network instead of
two. Since the GRGS orbit was derived without estimation
of the DORIS system time bias, we have studied the impact
of this bias on the radial orbit differences with special focus
on systematic differences between ascending and descending
passes. The effect of errors in the realization of the terres-
trial reference frame is tested by the implementation of the
most recent ITRF2014 version. The effects of uncertainties
in Earth’s TVG field models are tested by the implementa-
tion of the EIGEN-6S2 model which is the predecessor of
the EIGEN-6S4 model. For each case, the same background
models and estimated parameters were used as for the REF
orbit, except for those that represent the changes for the spe-
cific test case. The five test orbits and the differences with
respect to the GFZ REF orbit are

- SLR orbit: derived by using SLR tracking observations
  only;

- DORIS orbit: computed by using DORIS tracking ob-
  servations only;

- TBias orbit: calculated without estimation of the
  DORIS system time bias;

- ITRF14 orbit: calculated by using the information on
  station positions and velocities from ITRF2014 (Al-
  tamimi et al., 2016) instead of ITRF2008; and

- Geoid orbit: obtained by using the EIGEN-6S2
  (Rudenko et al., 2014), Earth’s gravity field model, in-
  stead of the EIGEN-6S4 model (Förste et al., 2016).
  Note that the Geoid orbit is based on the same gravity
  field model as the GRGS orbit.

2.2 TOPEX altimeter data

In order to assess the orbit accuracy at crossover points and
to relate the estimated errors to the total variability of the sea
level data, along-track TOPEX Sea Level v1.1 Essential Cli-
mate Variable (ECV) data (Ablain et al., 2015) released from
the ESA CCI Sea Level project have been included in the
analyses. The along-track data have been corrected for all
instrumental and geophysical effects by the state-of-the-art
models provided with the data. However, for some correc-
tions, updated models were applied. These include EOT11a
ocean tides and loading tides (Savcenko and Bosch, 2012),
solid Earth tides following the International Earth Rotation
and Reference Systems Service (IERS) 2003 Conventions,
and updated GPD+ wet tropospheric corrections (Fernan-
dez and Lazaro, 2016). The altimeter crossover differences
were calculated for each test orbit separately. For the calcula-
tion of sea level anomaly grids, the GSFC std1504 orbits
have been selected. The processing of the data, the crossover
point analyses, and collinear analyses, as well as the interpo-
lation to a regular grid, were performed using GFZ’s Altimet-
er Database and Processing System (ADS) Central (Schöne
et al., 2010).

2.3 Evaluation of the orbit solutions

In the following, the performance of the analysed orbits is
evaluated. For the GFZ orbit solutions, the consistency with
tracking data and at arc overlaps is assessed. Table 2 pro-
vides the main results of precise orbit determination of the
GFZ reference and test orbits, namely, the average values of
SLR and DORIS root mean square (rms) fits; radial, cross-
track, and along-track 2-day arc overlaps, illustrating the in-
ternal orbit consistency in these directions; and the number
of the arcs used to compute these values for the reference
and five test orbits. When using the same observation types
and weighting, smaller values of arc overlaps and observation
fits indicate improved orbit quality. Reduced radial arc over-
laps characterize reduced radial orbit error. SLR observations
were used at all 494 orbital arcs of five GFZ orbits, except for
the DORIS orbit for which no SLR observations were used
at all. Since DORIS data are available for TOPEX/Poseidon
only until 31 October 2004, these data were used at 459 or-
bital arcs preceding this date, except for the SLR orbit for
which no DORIS observations were used at all. All orbital
arcs for GFZ orbits are manoeuvre free. Thus, 2-day arc over-
laps were computed for 433 overlaps for the REF, TBias,
ITRF14, and Geoid orbits. In the case of the SLR and DORIS
orbits, a few gaps in the observations caused radial arc over-
lap larger than 0.5 m. Those arc overlaps have been excluded
from the statistics resulting in fewer arc overlaps shown for
these orbits in Table 2.

Figure 1 provides information on the SLR rms fit of the
reference and tests orbits, while Fig. 2 displays the radial
arc overlap of two consecutive 2-day orbit arcs. The four
orbits derived using SLR and DORIS observations provide comparable levels of average SLR rms fits (1.96–1.99 cm; Fig. 1). The smallest SLR rms fit (1.59 cm) but largest radial arc overlap (1.72 cm) are obtained for the SLR-only orbit, indicating a weak orbit quality over large geographical areas. The largest SLR rms fit is obtained for the TBias orbit. When no DORIS system time bias is estimated, inconsistencies between the timing of the observation system result in higher misfits. Among the five orbits derived using DORIS observations, a slightly increased average value of the DORIS rms fits (0.04795 cm s\(^{-1}\)) is obtained for the DORIS orbit derived using only DORIS observations (related to the weighting of observation types and the number of observations used) followed by the TBias orbit (0.04785 cm s\(^{-1}\)), while the other orbits derived using SLR and DORIS observations (REF, ITRF14, and Geoid) show comparable average values of the DORIS rms fits (0.04775–0.04778 cm s\(^{-1}\)). The smallest average value of the radial arc overlaps (Fig. 2) is obtained using the EIGEN-6S2 geopotential model (0.83 cm). The radial arc overlaps of the TOPEX/Poseidon orbit derived using only SLR data are 1.95 times larger than those of the orbit derived using only DORIS data. Using the reference frame ITRF2014 instead of ITRF2008 eliminates many outliers in the radial arc overlaps and therefore reduces the average value of the radial overlaps from 0.90 to 0.84 cm.

The DORIS system time bias is regularly estimated and applied during GFZ’s POD process to adjust the DORIS time system to the SLR time system. Zelensky et al. (2006) showed that there is a strong linear relationship between along-track orbit position and the DORIS time bias. The comparison of the fits and overlap values of the REF and the TBias orbits (Table 2) shows that the estimation of the DORIS time bias improves the orbit quality. The temporal behaviour of the DORIS system time bias derived for the TOPEX/Poseidon REF, ITRF14, and Geoid test orbits is in close agreement (Fig. A1) and resembles the estimation given by Lemoine et al. (2016). For the GFZ VER11 (REF) orbit, it indicates variations between −22.4 µs and +4.4 µs from 1992.73 to 1994.18, followed by a period of a linear trend of 35.11 µs yr\(^{-1}\) between 1994.18 and 1995.00 that ends with a jump from −28.65 to +1.98 µs around 1995.00. Then, the DORIS time bias shows two rather stable periods with a mean value of +3.70 µs with a standard deviation of 1.77 µs from 1995.0 to 1999.0 and a mean value of −1.32 µs with a standard deviation of 1.19 µs from 1999.0 to 2001.13, followed again by a period of a linear trend (−3.14 µs yr\(^{-1}\)) from 2001.13 to 2004.83. The mean value of the DORIS system time bias is 0.04 ± 0.36 µs for the DORIS test orbit, and it is equal to zero (not shown in the figure) for the TBias orbit.

For all orbit solutions, a crossover point analysis for the period April 1993 to September 2004 has been performed based on the altimeter data described in Sect. 2.2. Differences between the values of ascending and descending passes at crossover points are caused by oceanic variability and errors related to the measurements, the orbit, and the ap-
Table 3. Crossover point analysis: median and rms values of global mean height differences for maximum time lapses of 5 days for all orbit solutions during the period April 1993–September 2004. The highest and lowest values of each quantity are marked in bold.

<table>
<thead>
<tr>
<th>Crossover differences</th>
<th>REF</th>
<th>GSFC</th>
<th>GRGS</th>
<th>SLR</th>
<th>DORIS</th>
<th>TBias</th>
<th>ITRF14</th>
<th>Geoid</th>
</tr>
</thead>
<tbody>
<tr>
<td>Median (mm)</td>
<td>−3.1</td>
<td>−1.6</td>
<td>−3.0</td>
<td>−2.7</td>
<td>−4.7</td>
<td>−3.6</td>
<td>−2.8</td>
<td>−2.1</td>
</tr>
<tr>
<td>rms (mm)</td>
<td>49.8</td>
<td>49.5</td>
<td>51.3</td>
<td>51.2</td>
<td>50.7</td>
<td>49.8</td>
<td>49.8</td>
<td>49.7</td>
</tr>
</tbody>
</table>

Using ITRF2014 instead of ITRF2008 does not change the rms of crossover differences but improves their median values. The Geoid orbit solution exhibits clearly improved ascending/descending differences (−2.1 mm) as well as a slight reduction of the rms values. A further analysis of the temporal evolution of the ascending/descending differences reveals that these improvements take place in the pre-GRACE period before August 2002.

3 Estimation of the orbit-related sea level error

Sea level varies on typical temporal and spatial scales that are often connected to the driving processes. At the same time, orbit errors are not randomly distributed but exhibit also typical temporal and spatial patterns. Here, we apply statistical methods in order to assess the errors related to the orbit solutions for global and regional sea level at seasonal to decadal timescales.

3.1 Methods

In order to estimate the orbit-related errors in sea level height, the differences between the radial components of the GFZ REF orbit and the two independent orbit solutions (GSFC and GRGS) have been analysed. To assess the effect of uncertainties in the reference system, in the realization of the tracking station networks and in Earth’s time-variable gravity on the radial error budget, we have evaluated the differences of the radial orbit components between the GFZ’s REF and ITRF14, SLR, DORIS, and Geoid test orbits. Since the radial orbit components map directly to the derived sea level heights, we consider the differences presented here to represent estimates of the orbit-related sea level error. However, since the orbit error analysis is based on orbit differences, any error common to all three orbits will be lacking in our assessment.

The differences of the radial orbit components at the time of the altimetry measurement (1 Hz, ~6.7 km on the ground) are calculated and interpolated to a global 1° × 1° grid for every cycle (9.92 days). In general, we merge both ascending and descending passes in our calculations. In addition, we analyse ascending and descending passes for some orbit combinations separately. In order to study the global mean differences between the radial orbit components and their temporal evolution, global mean rms values per cycle are derived. They are calculated as the square root of the spatial
weighted mean of all squared radial orbit differences on the $1^\circ \times 1^\circ$ grid for the respective cycle.

Since we are not interested in the orbit error itself, but rather in the effect of radial orbit errors on global and regional sea level, we treat the radial orbit differences the same way as the sea level values from altimetry. For the estimation of global mean errors, the gridded radial orbit differences are averaged (with area weighting) over the ocean ($\pm 67^\circ$ latitude). Starting from these global mean orbit differences, global mean rms values relative to the temporal mean of the series are calculated as an estimate for the orbit-related error of the global mean sea level. Decadal trends, annual and semi-annual signals, and the corresponding formal errors are estimated by a least-square fit. The seasonal errors are derived from the amplitudes of the annual signal. As a measure of errors at interannual timescales, we calculate the rms of the 5-year running trend series of the radial orbit differences. Since the time series is only 11 years long, it is not possible to derive statistically sound estimates of the decadal trend. Here, the errors of decadal trends are assumed to correspond to the absolute values of the trends fitted to the series of the radial orbit differences. For the estimation of regional upper bound errors, at each grid point, rms values relative to the local temporal mean, annual cycle, rms of the 5-year running trend, and decadal trends are calculated in correspondence to the global analyses. From the $1^\circ \times 1^\circ$ grid, the maximum values over the ocean are extracted to estimate regional upper bound errors.

In order to relate the estimated errors to the total variability of the sea level data, TOPEX altimeter data have been included as well. The data and the processing are described in Sect. 2.2. From the gridded sea level anomalies, seasonal, interannual, and decadal trends were derived using the methods described above.

### 3.2 Global mean errors

In the following, we investigate the orbit-related global sea level error, differentiating between the total error and its annual, interannual, and decadal components. The TBias orbit differences are not included in these analyses but will be further investigated for the study of changes between ascending and descending passes (Sect. 3.4). The time series of the global mean rms of gridded radial orbit differences per cycle are shown in Fig. 3 for all orbit solutions relative to GFZ’s REF orbit. The largest differences occur between the REF and the GRGS orbits; the smallest changes occur for the ITRF14 test orbit. Most orbit differences are dominated by subseasonal variability; only for the Geoid and ITRF14 orbits the rms per cycle series are governed by seasonal and decadal periods. For the Geoid, GSFC, and GRGS orbit differences relative to the REF orbit, the rms series exhibit a seasonal cycle, which is an indication for seasonal orbit differences on regional scales. The rms of the REF minus Geoid orbit difference is decreased after August 2002, indicating that the main differences between the two orbits originate from the pre-GRACE period. In contrast, the differences between the REF and the ITRF14 orbits slightly increase from 2000 onwards.

From the time series of global mean orbit differences over the oceans, the rms, annual cycle, 5-year trend variability, and decadal trend differences are calculated and used as an estimate of the orbit-related error on different timescales. These orbit errors are summarized in Table 4 for all orbit models together with the corresponding values derived from altimetric sea level anomalies. The global mean rms of the radial orbit differences between the REF and GRGS (GSFC) orbits amounts to 1.2 (1.1) mm, which corresponds to 8% of the global mean sea level variability of 13.0 mm. The restriction to one tracking station subnetwork leads to large changes of the orbit; for the DORIS (SLR) orbit solution, the rms values of the radial differences with respect to the REF orbit amount to 1.8 mm (0.7 mm), which exceeds the size of the estimated total orbit errors. This highlights the importance of manifold, precise, and consistent tracking data for accurate global mean sea level estimates. The substitution of the Earth’s gravity field model (EIGEN-6S4 by EIGEN-6S2) and the ITRF realization (ITRF2008 by ITRF2014) accounts for 0.2 and 0.3 mm, respectively, of the mean rms orbit errors. A spectral analysis of the global mean radial differences (Fig. A2) exhibits peaks at ~60 days for all but the GRGS and TBias orbit differences and at ~90 and ~170 days for the SLR and DORIS orbit differences. An annual component can be observed for the GRGS and Geoid orbit differences. Since the annual amplitude is less than 1 mm only, it can be neglected and is not included in Table 4. The time series of the 5-year running trends of the global mean radial
orbit differences over the ocean are shown in Fig. 4 for the various orbit combinations. All curves range between $-0.3$ and $+0.2$ mm yr$^{-1}$, and show at least one zero crossing and imply interannual changes of the estimated decadal sea level trends. The corresponding curve of the 5-year running trends for the global mean sea level (not shown) ranges between 4 mm yr$^{-1}$ in the year 1997 and 2.6 mm yr$^{-1}$ in 2000. Before 1998, the GSFC and GFZ solutions are close to each other and both suggest smaller sea level trends for this period than the GRGS solution. After that, trends derived from GFZ orbits are weaker than the ones derived from GSFC orbits and stronger than the ones derived from GRGS. The maximum interannual trend variability of 0.1 mm yr$^{-1}$ occurs between the REF and GRGS orbits (Table 4) which amounts to almost 30% of the corresponding value derived for the global mean sea level curve (0.37 mm yr$^{-1}$). An error of this size might interfere with the estimation of global mean sea level acceleration. Hence, relative to the GFZ orbits, the use of the GSFC (GRGS) orbits would result in a slightly increased (decreased) acceleration of the global mean sea level curve during the TOPEX period. Since the exclusive use of the DORIS tracking station leads to interannual trend variability of 0.11 mm yr$^{-1}$, inconsistencies of the tracking stations’ subnetworks might explain large portions of the observed global mean interannual variability. The errors of the interannual trend variability are for all orbit combinations higher than for the decadal trends. The global mean decadal trends (calculated over the full mission time) are mostly significant than for the annual trend variability. The errors of the interannual trend variability of 0.1 mm yr$^{-1}$ in the year 1997 and 2.6 mm yr$^{-1}$ in 2000. Before 1998, the GSFC and GFZ solutions are close to each other and both suggest smaller sea level trends for this period than the GRGS solution. After that, trends derived from GFZ orbits are weaker than the ones derived from GSFC orbits and stronger than the ones derived from GRGS. The maximum interannual trend variability of 0.1 mm yr$^{-1}$ occurs between the REF and GRGS orbits (Table 4) which amounts to almost 30% of the corresponding value derived for the global mean sea level curve (0.37 mm yr$^{-1}$). An error of this size might interfere with the estimation of global mean sea level acceleration. Hence, relative to the GFZ orbits, the use of the GSFC (GRGS) orbits would result in a slightly increased (decreased) acceleration of the global mean sea level curve during the TOPEX period. Since the exclusive use of the DORIS tracking station leads to interannual trend variability of 0.11 mm yr$^{-1}$, inconsistencies of the tracking stations’ subnetworks might explain large portions of the observed global mean interannual variability. The errors of the interannual trend variability are for all orbit combinations higher than for the decadal trends. The global mean decadal trends (calculated over the full mission time) are mostly significant but can be further neglected, since they are well below the uncertainty of the corresponding global mean decadal sea level trend ($\pm 0.5$ mm yr$^{-1}$).

### Regional errors

The maximum regional errors derived from the analysis of the gridded orbit difference series over the oceans are summarized in Table 5. The TBias orbit differences are not included in these analyses but will be further investigated for the study of changes between ascending and descending passes (Sect. 3.4). Regionally, the maximum radial orbit differences on the $1^\circ \times 1^\circ$ grid between the REF and GRGS (GSFC) orbits amount to 10.7 (7.4) mm. The exclusive use of only one tracking station subnetwork leads to distinct changes with rms values of 9.3 (7.2) mm for the DORIS (SLR) subnetwork. This suggests that for the weighting factors applied with GFZ’s REF orbit especially inhomogeneity in the SLR station subnetwork has the potential to produce notable regional orbit errors.

Annual difference signals with respect to the REF orbit are most prominent for the GSFC and GRGS solutions, while they are negligible for the SLR, DORIS, and ITRF14 orbits. The corresponding patterns of the annual amplitudes for the differences of REF versus GSFC, GRGS, and Geoid orbits and of GRGS versus GSFC orbits are shown in Fig. 5. The observed patterns for the GSFC and GRGS orbit differences consist of a dipole with centres in the south-eastern Indian Ocean and the Caribbean. Since the two centres are phase shifted by half a year, the effect on the global mean differences is marginal. The pattern coincides with the patterns already shown to be related to the use of AOD1B products (Rudenko et al., 2016a) and different time-variable gravity fields for TOPEX/Poseidon POD (Esselborn et al., 2015). However, the annual differences between the REF and Geoid orbits can only explain part of the observed differences between the REF and GRGS orbits. In addition, the annual dif-
Table 5. Estimates of regional maximum orbit-related errors for the total and seasonal signals, and interannual and decadal trends. Values are derived from the radial orbit differences for REF minus SLR, REF minus DORIS, REF minus ITRF14, REF minus Geoid, REF minus GSFC, and REF minus GRGS for the period April 1993–June 2004. Details on the estimation method are given in Sect. 3.1.

<table>
<thead>
<tr>
<th>Regional maximum error</th>
<th>REF-SLR</th>
<th>REF-DORIS</th>
<th>REF-ITRF14</th>
<th>REF-Geoid</th>
<th>REF-GSFC</th>
<th>REF-GRGS</th>
</tr>
</thead>
<tbody>
<tr>
<td>rms (mm)</td>
<td>7.2</td>
<td>9.3</td>
<td>2.4</td>
<td>3.5</td>
<td>7.4</td>
<td>10.7</td>
</tr>
<tr>
<td>Annual amplitude (mm)</td>
<td>1.4</td>
<td>2.1</td>
<td>0.4</td>
<td>3.2</td>
<td>5.4</td>
<td>5.6</td>
</tr>
<tr>
<td>rms 5-year trend (mm yr(^{-1}))</td>
<td>0.5</td>
<td>0.6</td>
<td>0.2</td>
<td>0.4</td>
<td>1.2</td>
<td>0.9</td>
</tr>
<tr>
<td>(\Delta) decadal trend (mm yr(^{-1}))</td>
<td>0.2</td>
<td>0.4</td>
<td>0.2</td>
<td>0.4</td>
<td>1.0</td>
<td>0.7</td>
</tr>
</tbody>
</table>

Figure 5. Annual amplitude of the radial orbit differences for REF minus GSFC, REF minus GRGS, GRGS minus GSFC, and REF minus Geoid. The regions with formal errors larger than the fitted value are masked out (white). The maximum amplitude difference is given in Table 5.

Differences between GRGS and GSFC orbits are quite small and show no distinct pattern. Another plausible source of the relatively strong signal for the GSFC and GRGS orbit cases is the differences in the annual corrections for station coordinates by geocentre motion corrections and non-tidal atmospheric loading. A careful consideration of the relevant models used for the POD of these three orbits suggests that the observed differences originate in part from the non-tidal atmospheric loading effect on the stations which was applied for the GFZ but not the GRGS and GSFC orbits. There is evidence that the annual signal from the EIGEN-6S2 gravity field model is closer to the gravity field solution applied for the GSFC orbits than to the one from EIGEN-6S4 – at least in the pre-GRACE period (Fig. 5).

The patterns of the interannual variability of the regional trends are shown in Fig. 6 for all orbit differences. The trend errors reach up to 1.2 (0.9) mm yr\(^{-1}\) for the GSFC (GRGS) orbit differences (Table 5). The patterns of the trend variability from the GSFC and GRGS differences show coinciding maxima in the regions around South America and Australia. The differences for the Geoid orbit show similar features even though the absolute trend variability is smaller (up to 0.4 mm yr\(^{-1}\)). For the SLR and DORIS orbit differences, the patterns of interannual variability (Fig. 6) are patchy and oriented along individual tracks. For the ITRF14 solution, the trend variability is slightly increased at high latitudes (up to 0.2 mm yr\(^{-1}\)). The patterns of the interannual trend variability derived from the GFZ test orbits suggest that differences in the TVG modelling and contributions from the tracking systems are the most plausible sources of the observed regional differences of trend variability between REF, GSFC, and GRGS orbits.

The strongest regional changes in the decadal trend (Fig. 7 and Table 5) are observed for the differences between the REF and GSFC orbits (up to 1.0 mm yr\(^{-1}\)). For the GSFC orbit, high absolute decadal trend differences tend to coincide with maximum seasonal differences but not with maximum interannual variability. The differences between the REF and GRGS orbit trends reach 0.7 mm yr\(^{-1}\) at maximum, and the patterns of maximum annual amplitudes, interannual and decadal trend differences coincide. The differences between the REF and Geoid orbit trends resemble these patterns. However, the trend values are smaller (up to 0.4 mm yr\(^{-1}\)) and can explain only about half of the observed decadal trend differences. The decadal trends related to EIGEN-6S2 and EIGEN-6S4 differences during the
Figure 6. The rms of 5-year running trend differences of the radial orbit components for REF minus GSFC, REF minus GRGS, REF minus SLR, REF minus DORIS, REF minus ITRF14, and REF minus Geoid for the period April 1993–June 2004. The global mean rms of the differences over the ocean is given in Table 4.

Figure 7. Decadal trend differences of radial orbit components for REF minus GSFC, REF minus GRGS, REF minus SLR, REF minus DORIS, REF minus ITRF14, and REF minus Geoid for the period April 1993–June 2004. Regions with formal errors larger than the fitted value are masked out (white). The global mean trend difference over the ocean is given in Table 4.

TOPEX period presumably originate from the modelling of the TVG after August 2002, since before drift terms are only applied to degree-2 terms. The degree-2 terms, in turn, are defined by SLR data and show close agreement between the two TVG models for the pre-GRACE period. The ITRF14 orbit differences drift locally by a rate of up to 0.2 mm yr$^{-1}$ with positive values in the Southern Hemisphere and negative values in the Northern Hemisphere, indicating a drift in the
z component between the reference system realizations. The observed values are in good agreement with the combined change of scale and rate of the z component of the transformation between ITRF2008 and ITRF2014 (Altamimi et al., 2016). The regional decadal trends for the SLR and DORIS orbit differences are patchy and rather related to particular tracks without consistent long-wavelength behaviour. Higher trends of up to 0.4 mm yr\(^{-1}\) emerge for the DORIS orbit. The patterns of the decadal trend differences derived from the GFZ test orbits suggest that differences in the TVG modelling are the most plausible source of the observed regional decadal trend differences between REF, GSFC, and GRGS orbits.

3.4 Differences between ascending and descending passes

The crossover point analysis (Table 3) reveals considerable global mean differences between ascending and descending passes for most orbits. Fu and Haines (2013) showed that orbit errors might induce diverging drifts for sea level derived from ascending and descending passes. In the following, we study whether there are systematic changes to the results obtained so far when ascending and descending passes are investigated separately. Therefore, for a subset of orbit solutions, the same analyses were performed as before but for data sets derived from ascending and descending passes only. Since the DORIS orbit reveals the most pronounced median ascending/descending differences, we have chosen to study the REF minus DORIS and the REF minus TBias orbit differences further. During the POD of the GRGS orbit, the DORIS system time bias has not been estimated; therefore, we include the GRGS orbit in the analysis as well. However, in contrast to the previous analysis, we study the difference of Geoid minus GRGS instead of REF minus GRGS in order to exclude the effects of different time-variable gravity fields from the analysis.

The global mean radial orbit differences for ascending and descending passes are for all three cases in the range of ±12 mm (Fig. A3). The ascending and descending radial orbit differences are significantly anti-correlated. The correlation coefficient is almost −1 for the REF minus TBias case, almost −0.8 for the Geoid minus GRGS case, and still −0.5 for the REF minus DORIS case. The correlation is further increased for periods of more than 1 year. The REF minus TBias global mean time series resembles the DORIS system time bias applied for the REF orbit (Fig. A1). The global mean radial differences for the Geoid minus GRGS case reveal similar features as well. All three orbit differences exhibit diverging global mean radial differences for ascending and descending tracks after the year 2000. The interannual trend variability and decadal trends derived from the analysis of the global mean radial orbit difference series over the oceans are summarized in Table 6 for the merged, ascending, and descending passes. If the ascending and descending passes are analysed separately, the interannual trend variability is increased by at least 5 times for the corresponding orbit differences. Ascending passes exhibit higher variability than the descending. The differences for the global mean decadal trends between ascending and descending passes are a multiple of the values for the merged data and reach up to 0.6 mm yr\(^{-1}\), where both data sets are drifting in opposite directions.

The regional patterns of the decadal trend differences for ascending and descending passes are shown in Fig. 8. The DORIS orbit differences reveal a striking spread between the decadal trends of the ascending and descending passes. The trends are opposite for ascending and descending passes for most areas of the global ocean and reach regionally absolute values of up to 0.8 mm yr\(^{-1}\). Trends for the REF minus TBias orbit differences are very similar but smaller than the REF minus DORIS orbit ones. The corresponding analysis for the Geoid minus GRGS orbit differences shows again very similar features as for the DORIS differences. This indicates that discrepancies in the reference systems of the tracking stations (distribution of tracking stations, observation sampling, etc.) might give rise to long-wavelength orbit errors being anti-correlated for ascending and descending passes. Relevant contributions originate from uncertainties of the timing of the DORIS measurements. Increasing time biases are related to increasing along-track position errors and seem to be transferred to radial orbit errors. This mechanism is not fully understood, but a further analysis is beyond the scope of this paper. The uncertainties are especially pronounced in tropical and subtropical regions. On regional scales, the interannual and decadal trend errors derived from ascending/descending passes separately can be many times higher than the values derived from the merged data. Even though such effects tend to cancel, whenever both components are merged, they might still introduce considerable errors in regional studies that are based on along-track data, e.g. at calibration sites.

3.5 Regional orbit errors and sea level variability

Our analysis exhibits large-scale patterns of the orbit-related error. Errors for interannual to decadal sea level trends of more than 1 mm yr\(^{-1}\) might hamper the interpretation of the observed sea level variability from altimetry, at least apart from the large oceanic currents. In order to define regions where the orbit-related error should be considered when analysing sea level data from TOPEX, we have determined areas with orbit errors of at least 10% of the corresponding sea level value. Figure 9 shows the sea level variability, seasonal signal, and interannual and decadal trends derived from the ESA CCI TOPEX altimeter data for those regions where the orbit error amounts to at least 10% of the corresponding sea level value. Taking into account the total orbit-related error, about half the ocean is affected. This includes especially calm oceanic regions, whereas for energetic regions, like the Antarctic Circumpolar Current, tropical Pacific, and
Table 6. Differences of interannual trend variability and decadal trend for merged, ascending, and descending passes related to the orbit solution. Values are derived from the mean radial orbit differences over the oceans for Geoid minus GRGS, REF minus DORIS, and REF minus TBias for the period April 1993–June 2004. Values for ascending and descending passes are given in brackets.

<table>
<thead>
<tr>
<th></th>
<th>Global mean differences</th>
<th>Geoid-GRGS</th>
<th>REF-DORIS</th>
<th>REF-TBias</th>
</tr>
</thead>
<tbody>
<tr>
<td></td>
<td>rms 5-year trend (mm yr(^{-1}))</td>
<td>0.10 (0.62, 0.48)</td>
<td>0.11 (0.53, 0.38)</td>
<td>0.02 (0.55, 0.57)</td>
</tr>
<tr>
<td>Δ decadal trend (mm yr(^{-1}))</td>
<td>−0.02 (0.30, −0.34)</td>
<td>−0.06 (0.20, −0.27)</td>
<td>−0.01 (0.10, −0.13)</td>
<td></td>
</tr>
</tbody>
</table>

Figure 8. Decadal trend differences of radial orbit components for ascending (a) and descending (b) passes for Geoid minus GRGS, REF minus TBias, and REF minus DORIS for the period April 1993–June 2004. Regions with formal errors larger than the fitted value are masked out (white). The global mean trend difference over the ocean is given in Table 6.

4 Summary and conclusions

We have investigated the radial orbit error budget associated with three state-of-the-art orbit solutions from GFZ, GSFC, and GRGS over the first altimetry decade (1993–2004). It is crucial to know the accuracy of these early altimeter data in order to judge the reliability of long-term sea level trends and of estimates of the acceleration of global mean sea level rise. For this purpose, we have chosen the TOPEX/Poseidon mission, since it is the reference altimetry mission used in the ESA CCI Sea Level project over this time span. We estimate the orbit errors from the radial orbit differences which implies that errors common to all orbits cannot be detected. However, since the three orbits were derived using various up-to-date models, the errors common to the three orbits should be rather low, which makes us confident that our error estimates represent most of the error. A set of five test orbit solutions derived at GFZ is used to estimate the contributions of the most significant factors to the error budget. We have focused on the impact of uncertainties of the tracking station subnetworks (SLR and DORIS), of the DORIS system time bias, of the reference frame, and of the Earth’s time-variable gravity field models on the radial orbit component and hence the derived sea level. The estimates of the radial orbit errors at seasonal, interannual (5-year), and decadal timescales are

the western boundary currents of the Northern Hemisphere, the dynamic ocean signal is much larger than the orbit error. For the seasonal signal, mainly the Southern Ocean is concerned. Critical regions for the estimation of the interannual variability are the tropical and subtropical Atlantic and the south-eastern Pacific. For decadal scales, the orbit-related trend errors are prominent in a couple of regions including the South Atlantic, western North Atlantic, central Pacific, and south-eastern Indian Ocean, but also several marginal seas including the Mediterranean, Red Sea, Yellow Sea, and Sea of Japan.
Figure 9. The rms of sea level, annual amplitude, rms of interannual (5-year) running trend, and decadal trends from TOPEX altimeter data for the period February 1993–October 2005. Colour coded are sea level values for which the local orbit errors (estimated from GFZ minus GRGS) reach more than 10% of the local sea level values. All other regions are masked out (white).

given in Table 4 for the global mean sea level and in Table 5 for the regional sea level.

According to our study, the contribution of orbit uncertainties to the error of the global mean sea level during the TOPEX period are of the order of 1.2 mm, which corresponds to 8% of the variability of the global mean sea level (13 mm). The global mean annual (seasonal) component of the radial error is well below 1 mm and can be neglected. The orbit-related errors of the decadal trends are up to 0.08 mm yr\(^{-1}\) and should not induce any significant artificial global mean sea level trends. However, on timescales of 5 years, the trend variability may reach up to 0.1 mm yr\(^{-1}\), which amounts to almost 30% of the corresponding sea level variability (0.37 mm yr\(^{-1}\)), and could potentially hamper the detection of sea level acceleration from the altimeter data. The major contributions to this error (0.04–0.11 mm yr\(^{-1}\)) are, most probably, discrepancies of the station subnetworks (DORIS or SLR) used. The contributions of Earth’s time-variable gravity field model and the ITRF realization (ITRF2008 versus ITRF2014) to the global mean error are of only minor importance (0.03 mm yr\(^{-1}\)).

These values are in line with the mean upper bound orbit errors given by Couhert et al. (2015) derived for Jason-1 and Jason-2 orbits for the second altimetry decade (2002–2012).

For regional scales, the maximum rms of the gridded radial orbit error is 11 mm. This error is indicative of the orbit-related sea level error on the 1°×1° grid and hence notably less than the actual radial orbit error reached for TOPEX/Poseidon (e.g. Marshall et al., 1995). The 11 mm orbit-related sea level error includes a large fraction of sub-seasonal variability which is not the subject of this study. The regional upper bound error of the seasonal signal is 6 mm, of the interannual trend variability 1.2 mm yr\(^{-1}\), and of the decadal trend 1 mm yr\(^{-1}\). Errors for interannual to decadal sea level trends of more than 1 mm yr\(^{-1}\) might hamper the interpretation of the observed sea level variability from altimetry. For about half of the ocean outside the energetic regions (e.g. Antarctic Circumpolar Current, tropical Pacific, the Gulf Stream system, and Kuroshio system), the orbit-related errors reach at least 10% of the observed sea level variability. For the seasonal signal, mainly the Southern Ocean is concerned. Critical regions for the estimation of the interannual variability are the tropical and subtropical Atlantic and the south-eastern Pacific. For decadal scales, the orbit-related trend errors are prominent in a couple of regions including the South Atlantic, western North Atlantic, central Pacific, and south-eastern Indian Ocean, but also several marginal seas including the Mediterranean, Red Sea, Yellow Sea, and Sea of Japan.

When using ascending and descending passes separately, the interannual and decadal trend errors can reach multiples of the values derived from the merged data. This is the case for global mean values as well as for regional values. The corresponding large-scale pattern is coherent for low and medium latitudes and is strongly anti-correlated for ascending and descending passes. Even though such effects tend to cancel, whenever both components are merged, they might
still introduce considerable errors in regional studies, that are based on along-track data, e.g. at calibration sites.

Orbit errors related to discrepancies between the tracking station subnetworks (distribution of tracking stations, observation sampling, etc.) are studied based on GFZ’s SLR, DORIS, and TBias orbit solutions. Using SLR and DORIS observations for TOPEX POD together reduces (improves) the rms of the altimetry single-satellite crossover differences considerably (2–3 %), though the DORIS observations seem to aggravate the median differences between ascending and descending passes. The proper estimation of the DORIS system time bias has proven to be a critical factor for the minimization of this effect. The most significant changes are observed for the DORIS orbit solution, suggesting that uncertainties of the SLR station subnetwork should have the most prominent effects on the orbit accuracy – at least for GFZ’s orbit solutions. This fact is, most probably, related to the weighting factors applied to the observations within the GFZ orbit determination process. Using the latest reference frame (ITRF2014) instead of the predecessor (ITRF2008) slightly improves the accuracy of the TOPEX/Poseidon orbit solution. The contribution of the uncertainties in the ITRF realization to the regional upper bound error is only marginal. Errors induced by uncertainties of the Earth’s time-variable gravity field model are studied on the base of GFZ’s Geoid orbit solution. The orbit evaluations show that the Geoid orbit performs slightly better than the REF orbit in the pre-GRACE period due to differences in the periodic annual and semi-annual variations applied to the TVG field models. Uncertainties of the gravity field model give rise to orbit errors at all analysed periods. We estimate regional upper bound errors of ~ 3 mm for the seasonal signal and of 0.4 mm yr⁻¹ for the interannual trend variability and the decadal trend. This accounts for about 60 % of the seasonal, about 30 % of the interannual, and about 40 % of the decadal orbit error which are related to differences between EIGEN-6S2 and EIGEN-6S4.

The regional upper bound radial orbit errors obtained from our study are by factors of 2 to 5 smaller than the ones reported by Couhert et al. (2015) for the period 2002 to 2012. This might partly reflect recent improvements of the stability of reference frames which results in smaller changes from ITRF2008 to ITRF2014 than previously from ITRF2005 to ITRF2008. However, the accuracy of the Earth’s time-variable gravity model and the tracking observations for the 1990s should be inferior to more recent periods. The error related to the uncertainties of the tracking station subnetworks might be underrated in our study since all analysed orbits rely on basically the same set of tracking observations. The effect of uncertainties of the TVG field might be underestimated as well, since both EIGEN-6S4 and EIGEN-6S2 model the TVG field in the pre-GRACE period by periodic annual and semi-annual variations derived from GRACE plus annual values and drift terms for degree-2 terms derived from SLR measurements. In contrast, the TVG field

Data availability. The GFZ VER11 orbits (Rudenko et al., 2016b) can be accessed at GFZ Data Services via http://pmd.gfz-potsdam.de/portal/. The GSFC std1504 orbits are available at ftp://cddis.gsfc.nasa.gov/pub/misc/test/JasonOrbits/gsfc/tp-orbits/. The GRGS orbits are available at ftp://cddis.gsfc.nasa.gov/pub/doris/products/orbits/grg/. The five GFZ test orbits (SLR, DORIS, TBias, ITRF14, and Geoid) of TOPEX/Poseidon can be obtained upon request to the authors. ESA’s along-track TOPEX Sea Level ECV data can be obtained via http://www.esa-sealevel-cci.org/products.
Appendix A

Figure A1. DORIS system time bias of TOPEX/Poseidon REF, DORIS, ITRF14, and Geoid orbits.

Figure A2. Power spectra of the global mean radial orbit differences over the oceans for REF minus GSFC, REF minus GRGS, GRGS minus GSFC, REF minus SLR, REF minus DORIS, REF minus ITRF14, REF minus Geoid, and REF minus TBias. Vertical dashed lines mark periods of 59, 85, and 170 days, and 1 and 5 years.
Figure A3. Global mean radial orbit differences over the oceans per cycle and 1-year box-car filtered for Geoid minus GRGS, REF minus DORIS, and REF minus TBias separately for ascending (blue, cyan) and descending (yellow, red) tracks. The cross-correlation coefficient between the ascending and descending passes for the original and the filtered series is given in the lower part of each graph.
References


